



Contribution of WFIRST to Astronomy in the Next 30 years

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AIAA session “Astronomy from Space 30 Years in the Future”

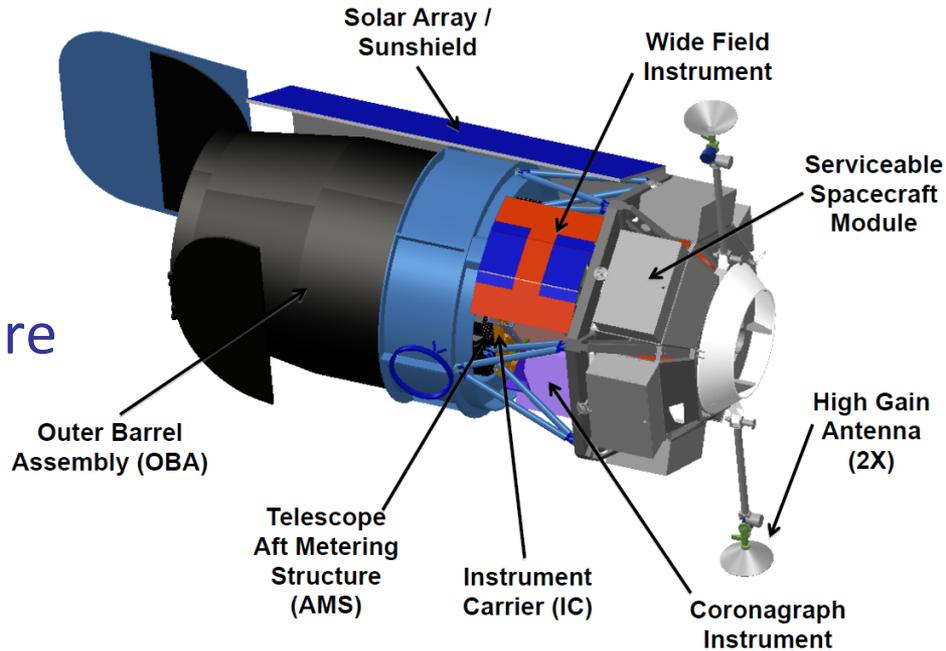
San Diego, CA

6 August 2014

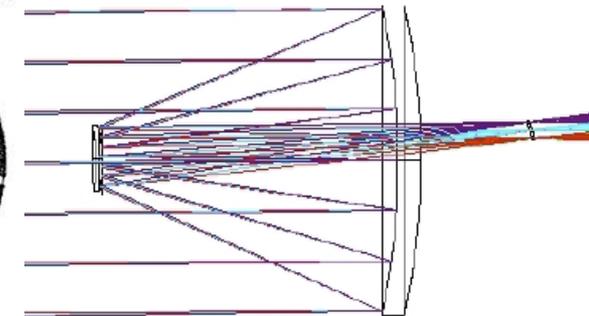
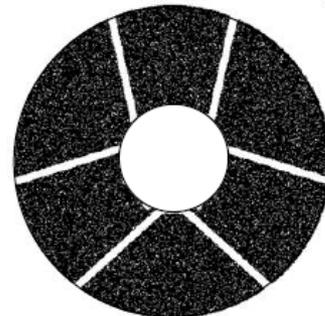
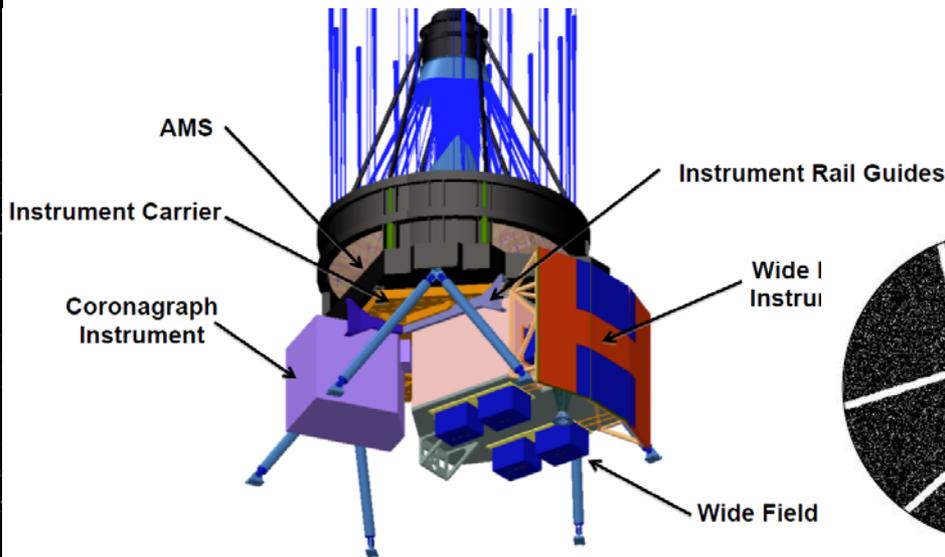
- WFIRST: top-ranked large space project (Decadal 2010)
 - Original vision was for a “1.5-m telescope, wide FOV, NIR imaging, low-resolution spectroscopy”
 - Current vision is for a 2.4-m telescope, wide FOV, visible-NIR imaging, low-resolution spectroscopy, & coronagraph
 - 5-year mission at GEO
- Coronagraph technology: top-ranked medium space project
 - Larger telescope permits adding coronagraph instrument
 - Science plus tech demo
 - 6-year mission at GEO
- New start FY-17
- Launch early 2020s

Ref.: “New Worlds, New Horizons in Astronomy and Astrophysics”, pp. 16-20, NRC, NAS, 2010

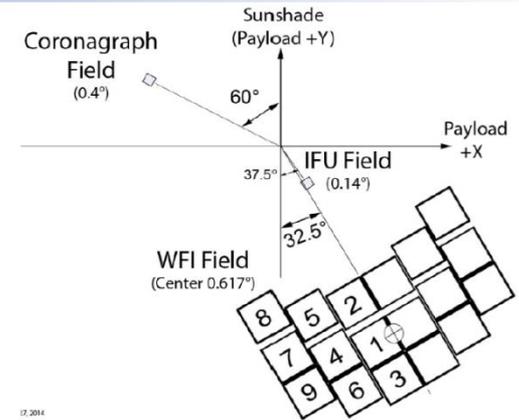
- 2.4-m donated telescope already exists
- 3-mirror anastigmat (TMA)
- On-axis secondary
- ULE lightweighted optics
- 270 K operating temperature
- GEO orbit, 24-hour period



WFIRST-AFTA SDT Interim Report



- 18 H4RG HgCdTe detectors
- FOV is 0.28 deg^2 (100x JWST, 200x HST)
- 0.7 – 2.0 micron bandpass
- 6-filter imaging
- grism and IFS spectra

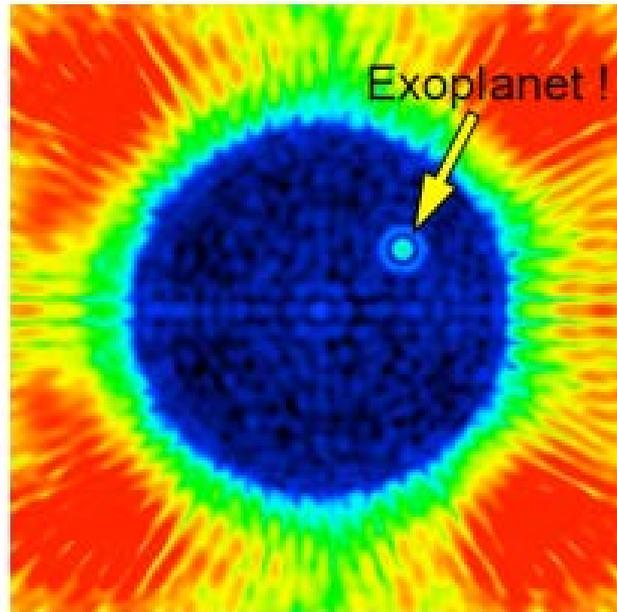


WFIRST-AFTA Deep Field

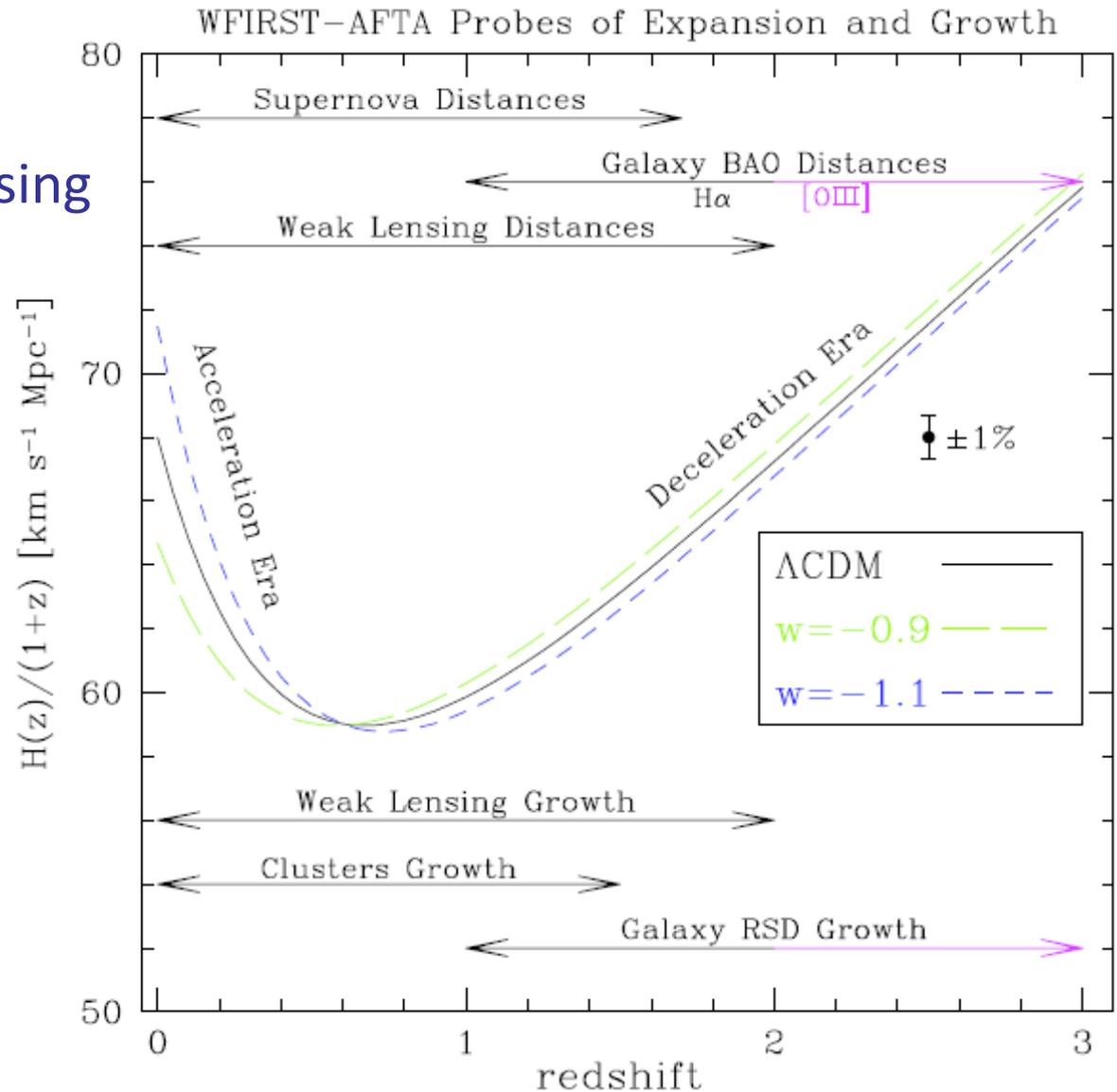
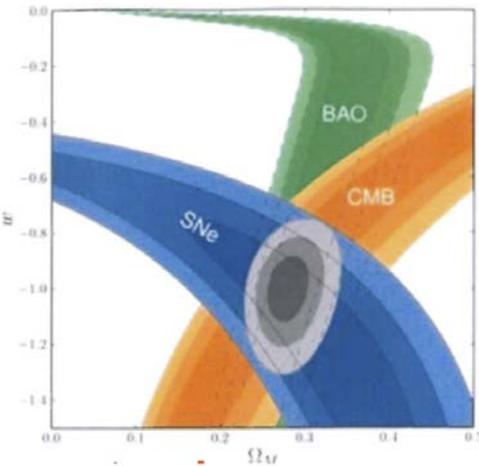



Hubble Deep
Field

- Coronagraph
 - electron-multiplying ccd detectors
 - 0.4 – 1.0 micron bandpass
 - 6-filter imaging, 10% bands
 - 4-filter IFS spectra, resolution 70
 - 10^{-9} contrast (after post-processing)
 - 100 milliarcsec inner working angle at 400 nm



- Baryon acoustic oscillations
- Gravitational lensing
- Supernovae





Top-level questions of the field:

1. Is cosmic acceleration caused by a new energy component or by the breakdown of General Relativity (GR) on cosmological scales?
2. If the cause is a new energy component, is its energy density constant in space and time, or has it evolved over the history of the universe?

WFIRST-AFTA addresses these questions using multiple methods to measure the history of cosmic expansion and structure growth, tightly constraining the properties of dark energy, the consistency of GR, and the curvature of space.

Supernova Survey: Distance measurements, $z = 0 - 1.7$.

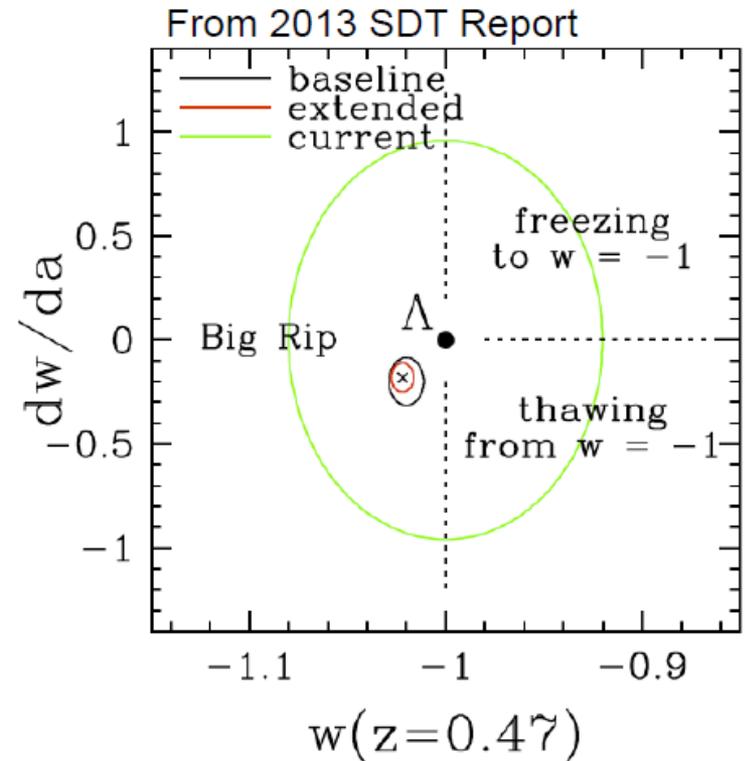
Weak Lensing Survey: Growth of structure from cosmic shear, galaxy-galaxy lensing, abundance of massive clusters.

Galaxy Redshift Survey: Distance and expansion rate from baryon acoustic oscillations, growth of structure from redshift-space distortions.

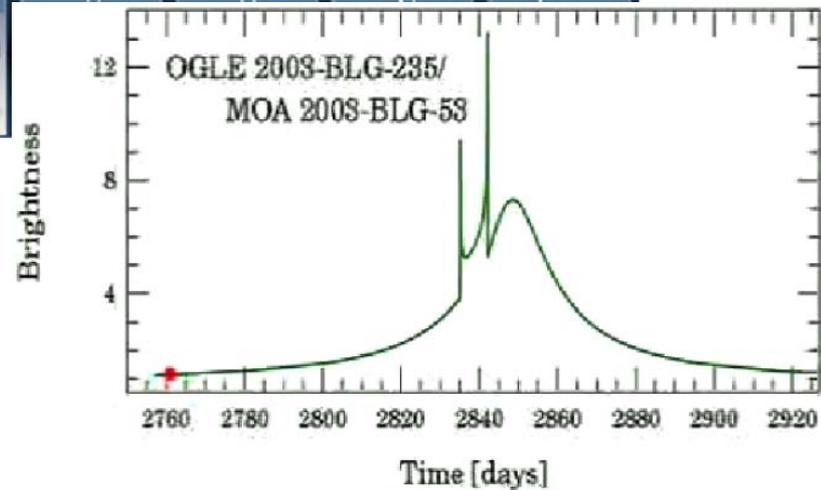
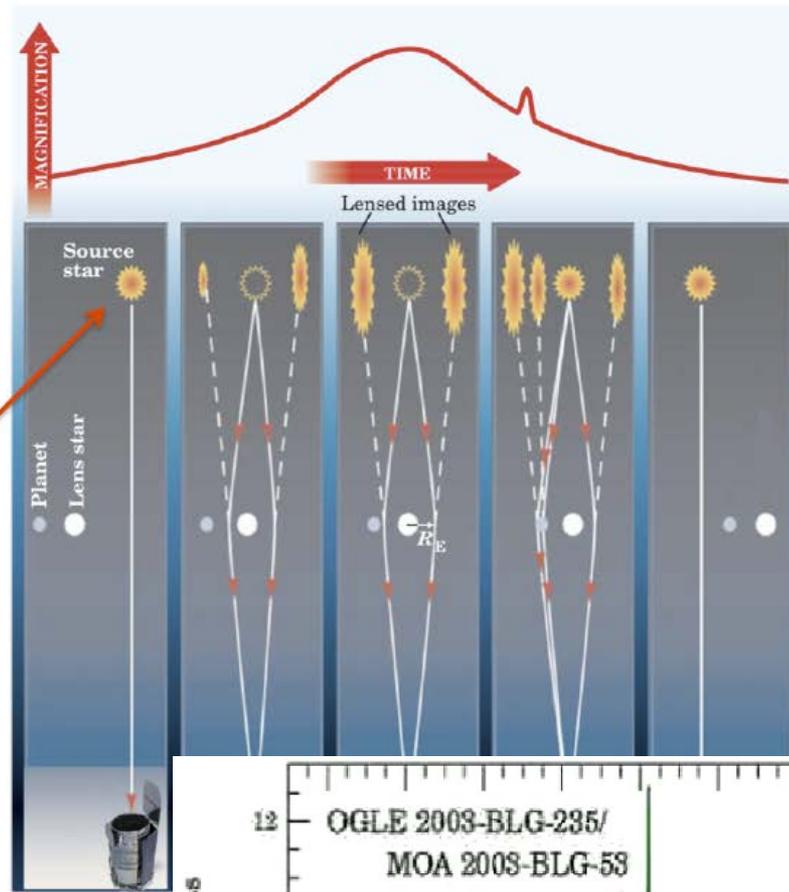
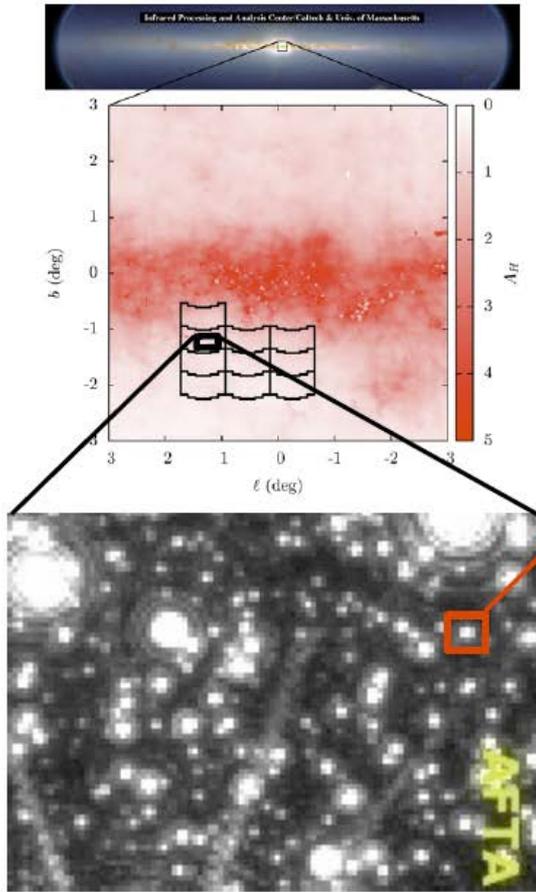
Emphasis on cross-checks and control of systematics at every level.

WFIRST-AFTA will improve cosmological measurements by 1-2 orders of magnitude over current data, with greater redshift leverage, control of systematics, and cross-checks of methods.

Potential to reveal surprises below the sensitivity of current data, confirm them internally with cross-checks, and investigate their physics by combining expansion and growth probes.

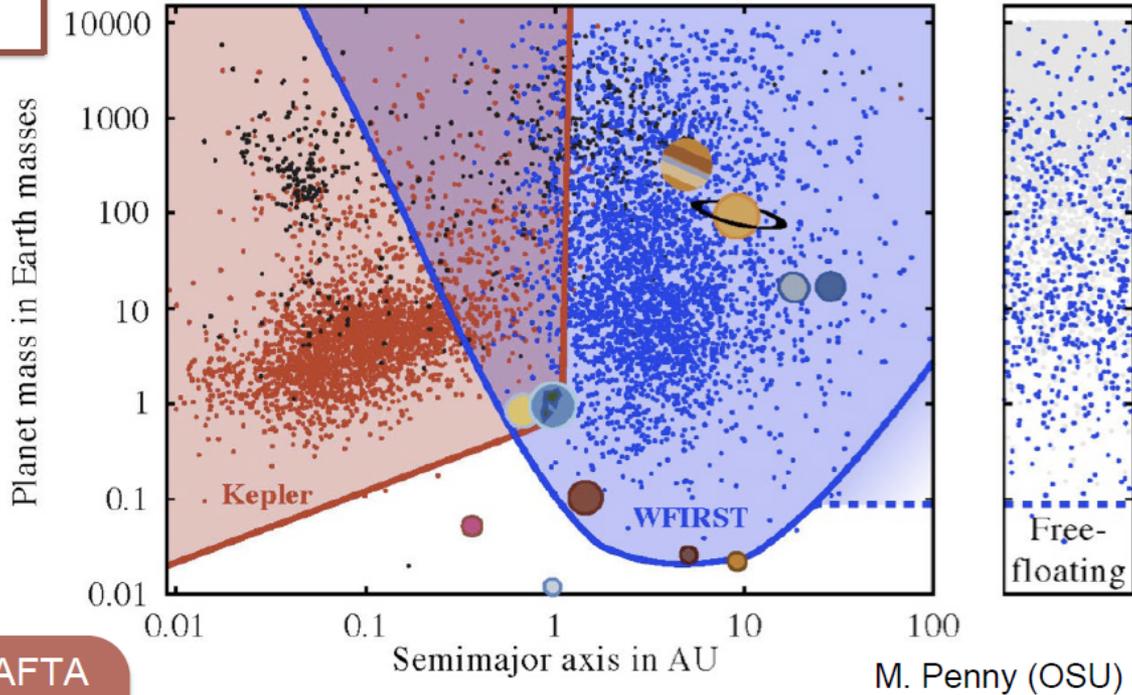


Forecast dark energy constraints from baseline & extended programs, compared to current knowledge. Distinct regions of plane represent fundamentally different physics.





Combined with space-based transit surveys, WFIRST-AFTA completes the statistical census of planetary systems in the Galaxy.



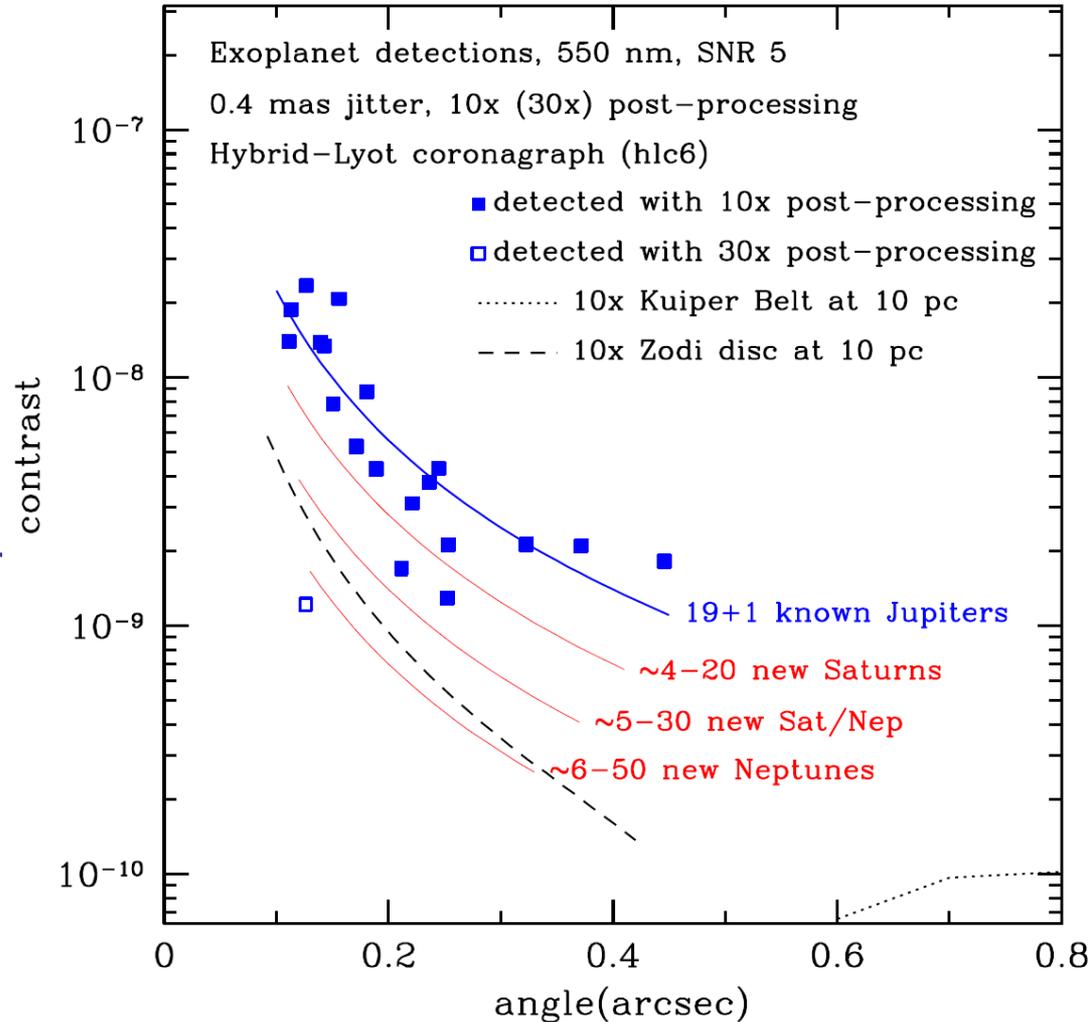
WFIRST-AFTA perfectly complements Kepler, TESS, and PLATO.

- ~3000 planet detections.
- 300 with Earth mass and below.
- Hundreds of free-floating planets.

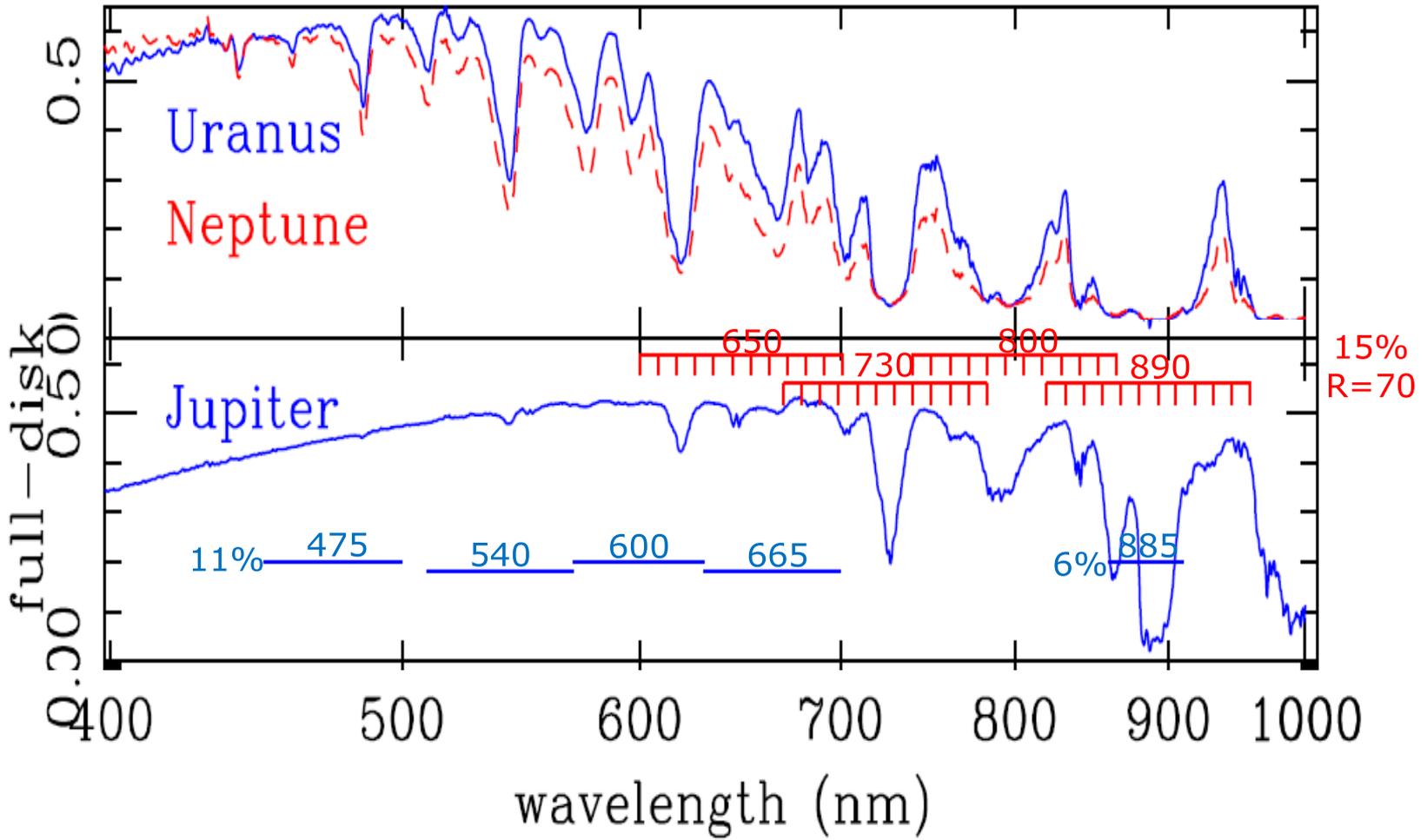
WFIRST-AFTA is more capable than the IDRM design

- 1.6 times larger planet yields
- Factor of two better sensitivity to Earth-mass planets.
- Improved ability to measure masses and distances to the microlensing host stars.

- Detect and obtain spectra of a dozen known RV planets
- Search for new planets, from Jupiters to Neptunes
- Detect and image exozodi and Kuiper-type debris disks

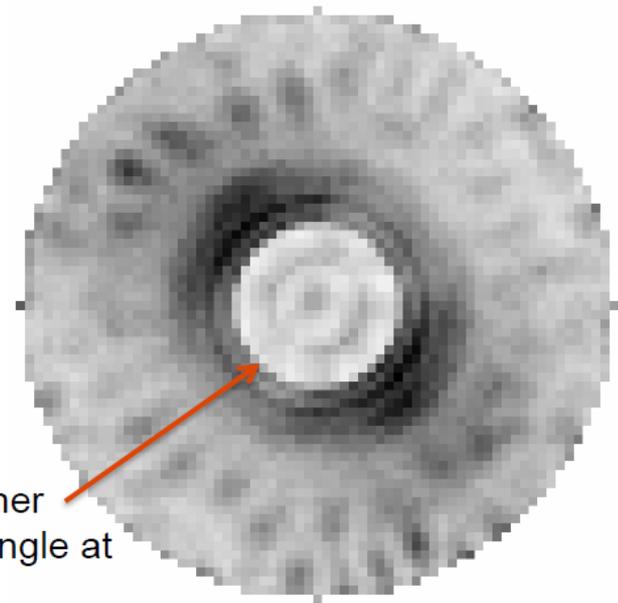


yrv145-hlc6c.pdf



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- Circumstellar disks reveal the locations of planets and trace the history of collisions
- Few disks below 100x the solar system dust level (100 zodi) have been detected, and none have resolved images
- WFIRST-AFTA will detect disks down to 10 zodi around nearby stars;
 - Important for planetary systems and for future Earth imaging missions
- WFIRST-AFTA + LBT-I are needed: LBT-I gives total dust amounts; WFIRST-AFTA then gives reflectance; both together give debris properties

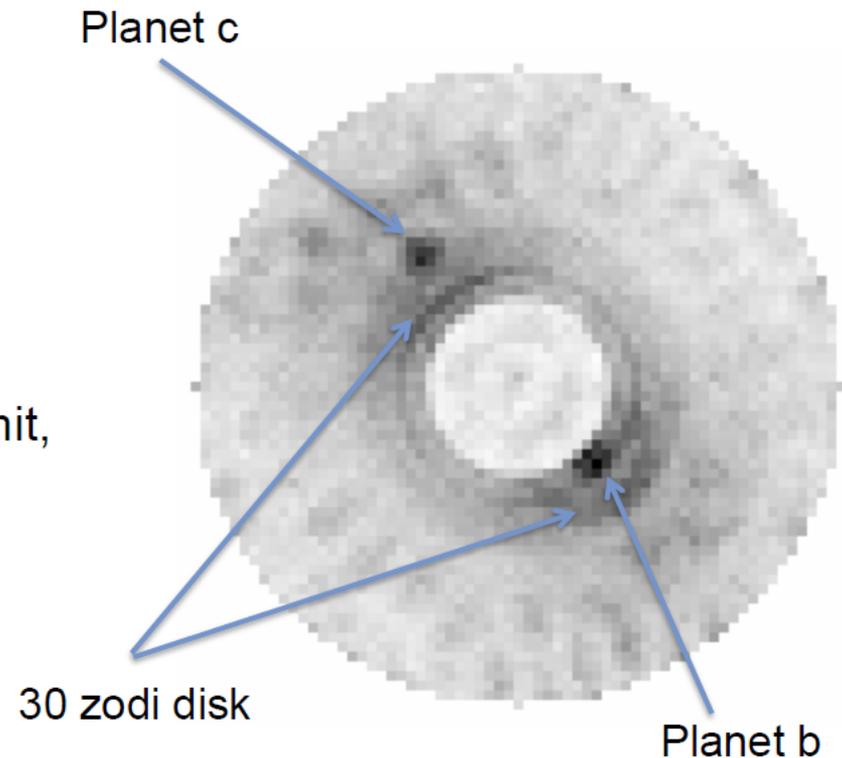


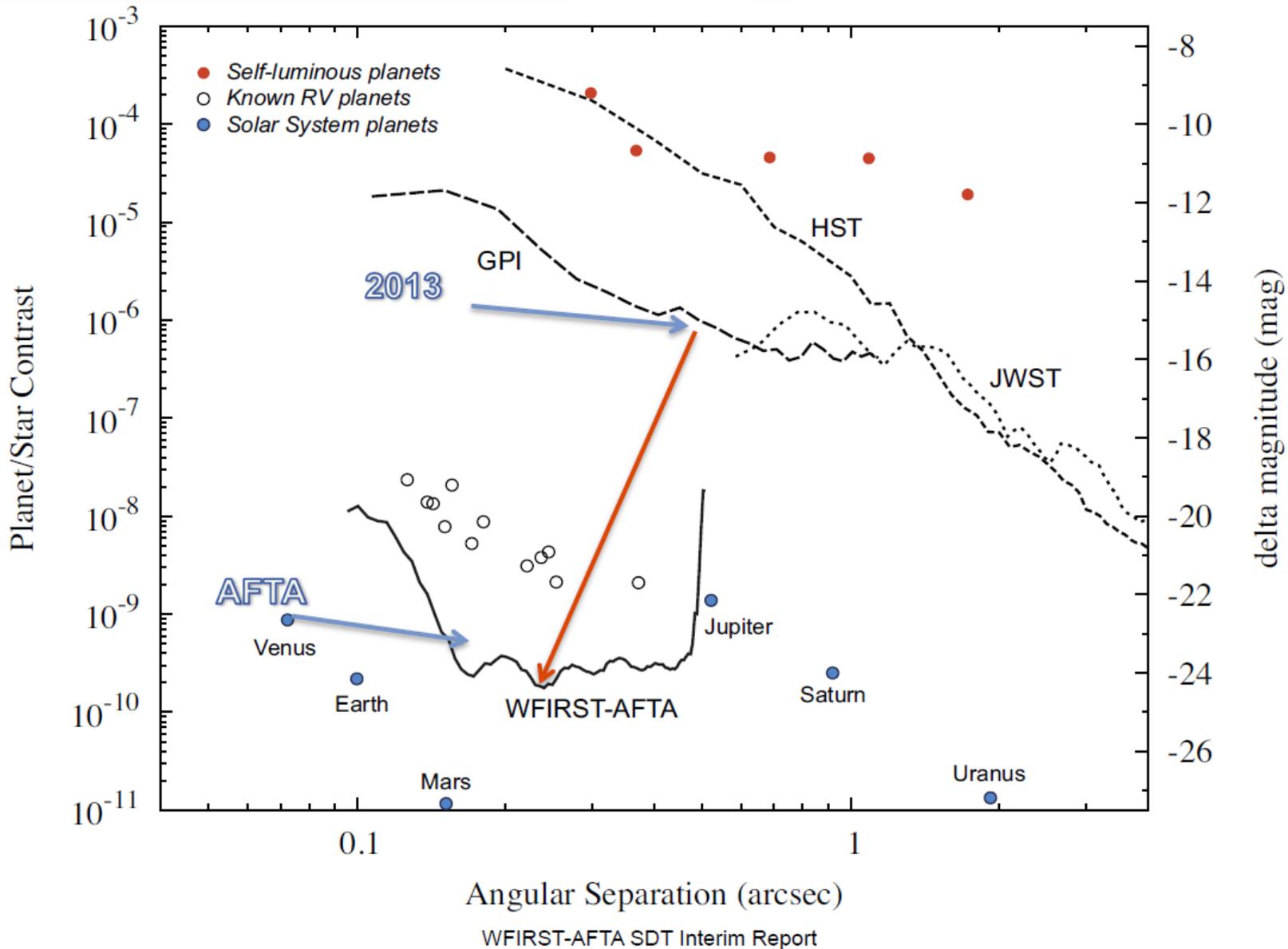
1.1 AU inner working angle at $d = 8$ pc

Simulation of 20 zodi disk WFIRST-AFTA image (24 h at 8 pc)

- The WFIRST-AFTA coronagraph will give us the first reflected light visible images of the planetary systems of nearby stars
- 47 UMa System with known RV planets (~Jupiter masses)
- G1V star at 14 pc
- Planet b has SMA = 2.1 AU, planet c has SMA = 3.6 AU
- Assume 30 zodi dust (628 zodi measured 3 sigma upper limit, Millan-Gabet et al. 2011)
- Assume incl 60 d, PA 45 d, pl. albedo 0.4, pl. orbit -90 d & 70 d

Simulation of a 10 hour exposure with HL coronagraph (0.4 mas jitter / 10 x speckle suppression, 550 nm 10% BW)







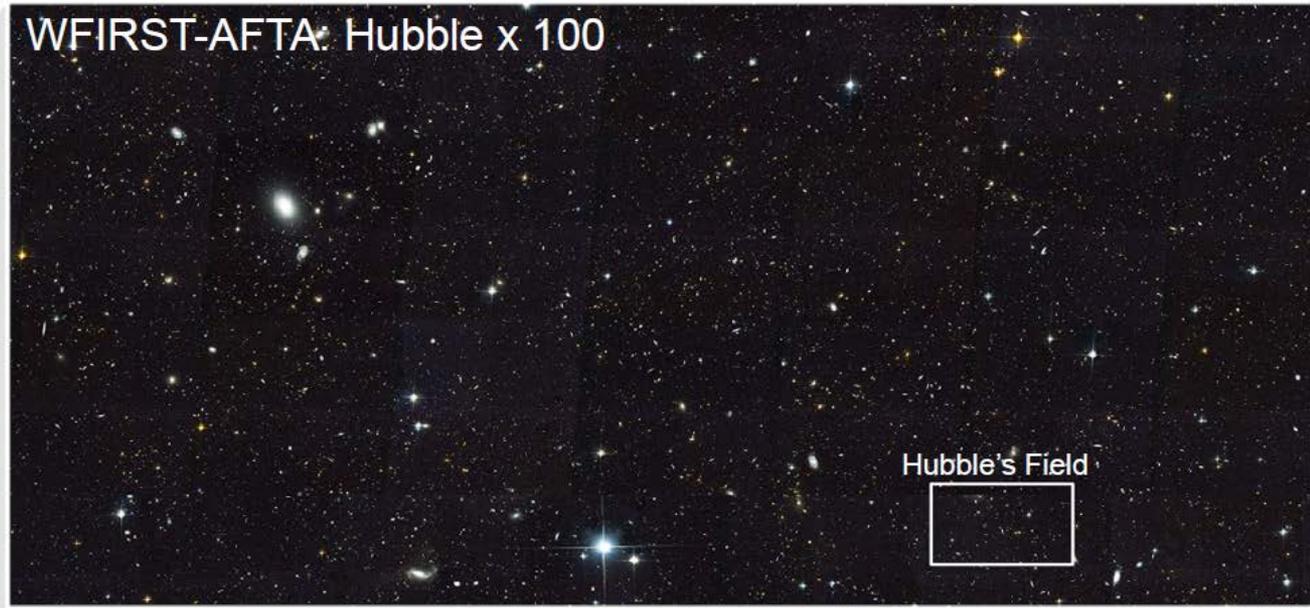
General Observer science (25%): example of capability



The Hubble Ultra Deep Field
seeing the Universe, 10,000 galaxies at a time



WFIRST-AFTA: Hubble x 100



A WFIRST-AFTA Deep Field
A New Window on the Universe - **1,000,000** galaxies at a time

04/30/2014

WFIRST-AFTA SDT Interim Report



AFTA WFIRST
Wide-Field Infrared Survey Telescope

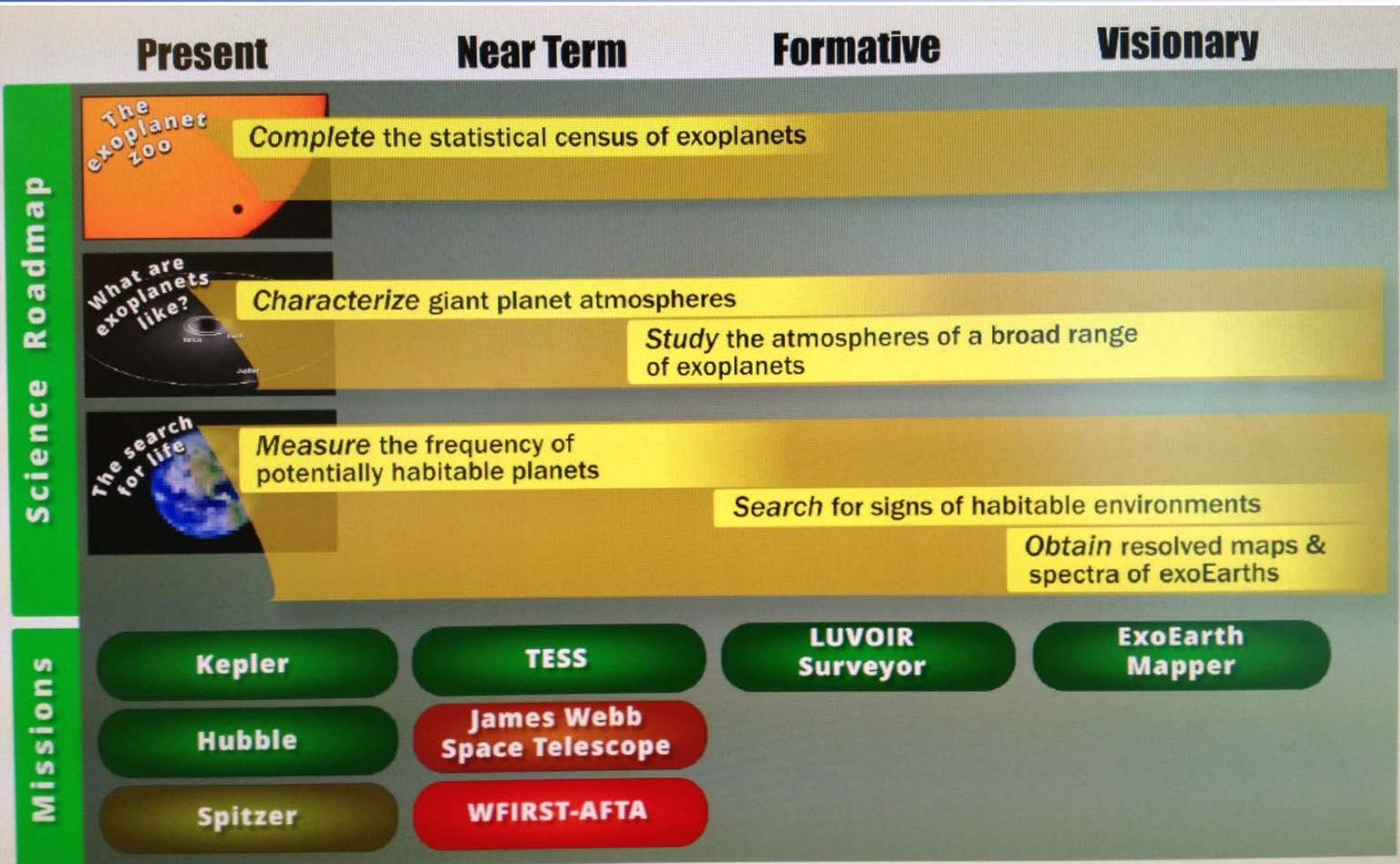


Figure 2.16 Schematic of the Exoplanets Roadmap, with science themes along the top and a possible mission sequence across the bottom. Credit: F. Reddy (NASA GSFC)



Science Goal 2 of "30-Year Plan": How did we get here?



AFTA WFIRST
Wide-Field Infrared Survey Telescope

	Present	Near Term	Formative	Visionary
Science Roadmap	 <p>Discover nearby planetary nurseries</p> <p>Measure disk structure & location of water</p>			
	 <p>Map the entire Milky Way</p> <p>Uncover the archaeology of all nearby galaxies</p>			
	 <p>Find the first black holes</p> <p>Characterize early black holes & their feedback</p> <p>Image accretion disks of black holes</p>			
	 <p>Image the first black holes</p> <p>Characterize the first starlight spectroscopically</p> <p>Map the epoch of reionization</p>			
Missions	Hubble	LSST	Gravitational Wave Surveyor	Gravitational Wave Mapper
	Spitzer	Extremely Large Telescopes	X-ray Surveyor	X-ray Mapper
	Herschel	James Webb Space Telescope	LUVOIR Surveyor	ExoEarth Mapper
	ALMA	WFIRST-AFTA	Far Infrared Surveyor	Cosmic Dawn Mapper



Goal 3 of "30-Year Plan": How does our Universe work? **JPL**

AFTA WFIRST
Wide-Field Infrared Survey Telescope

	Present	Near Term	Formative	Visionary
Science Roadmap	 <p>Measure dark energy & history of cosmic growth</p> <p>Probe the epoch of inflation</p> <p>Completely characterize the CMB</p>			<p>Map structure at reionization</p> <p>Measure cosmic expansion history with standard sirens</p>
	 <p>Constrain neutron star equation of state</p> <p>Understand black-hole-powered engines</p>		<p>Map black holes using gravitational waves</p> <p>Measure black hole masses & spins</p>	<p>Image the shadows of black hole event horizons</p>
	 <p>Image sources detected by aLIGO</p>		<p>Chart supermassive black hole mergers</p> <p>Search for electroweak-era gravitational waves</p>	<p>Hear the Big Bang</p>
Missions	Fermi	NICER	Gravitational Wave Surveyor	Gravitational Wave Mapper
	Chandra & XMM-Newton	LSST	X-ray Surveyor	Black Hole Mapper
	Hubble	James Webb Space Telescope	Far Infrared Surveyor	Cosmic Dawn Mapper
	WMAP & Planck	Gaia	CMB Polarization Surveyor	
		WFIRST-AFTA		

- WFIRST science in the 2020s is foundational for astrophysics science in the 2030s, in all 3 big-theme areas:
 1. exoplanet statistics, detections, and spectra lead to the search for habitable planets & maps of exoEarths
 2. maps of the Milky way lead to characterizing early black holes, the first starlight, & the epoch of reionization
 3. measuring dark energy, cosmic growth, and dark matter leads to measuring black holes with gravitational waves, black hole mergers, & echoes of the Big Bang



Thank you!

AFTA WFIRST
Wide-Field Infrared Survey Telescope