

SpaceOps 2008

# Optical Communications from Planetary Distances

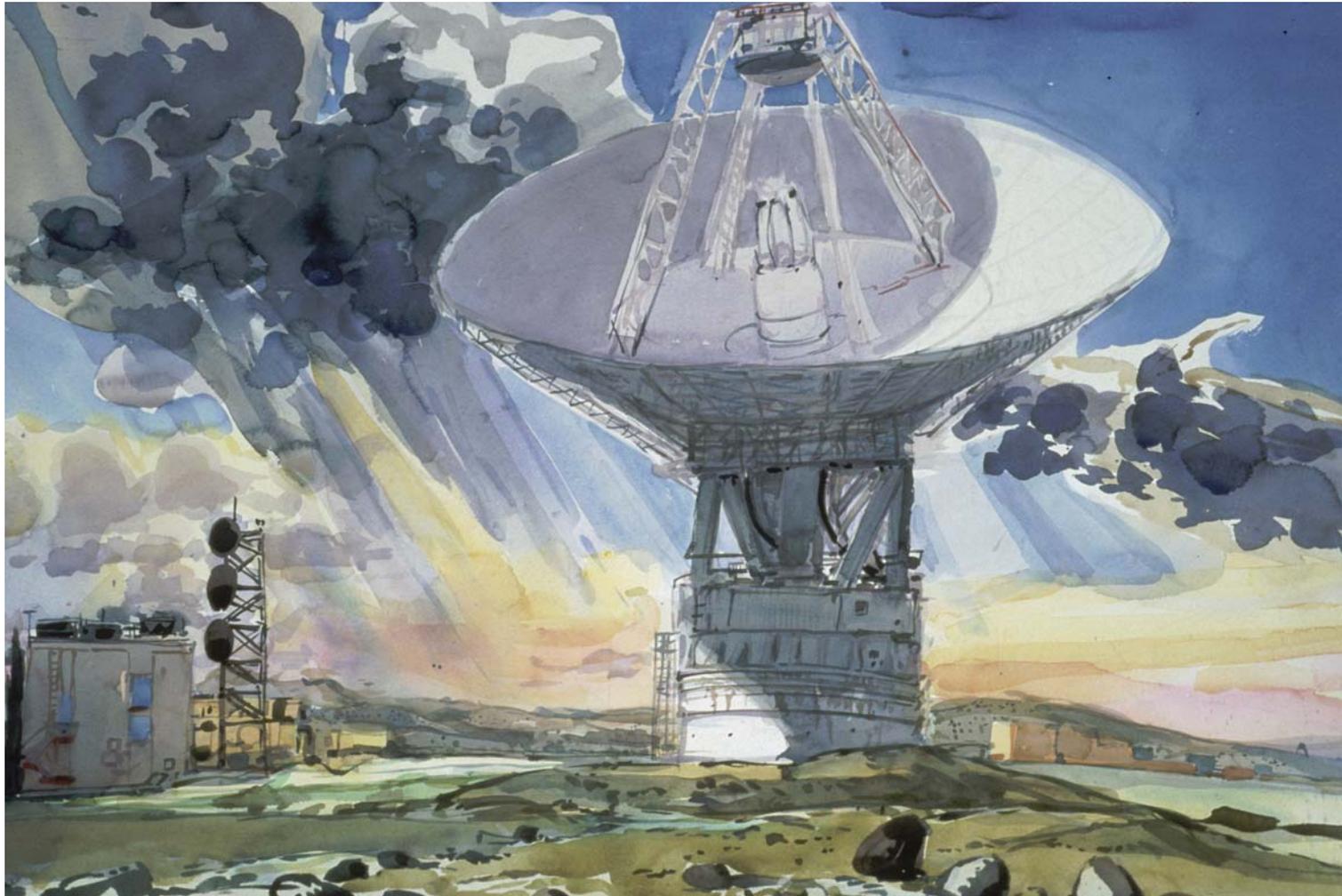
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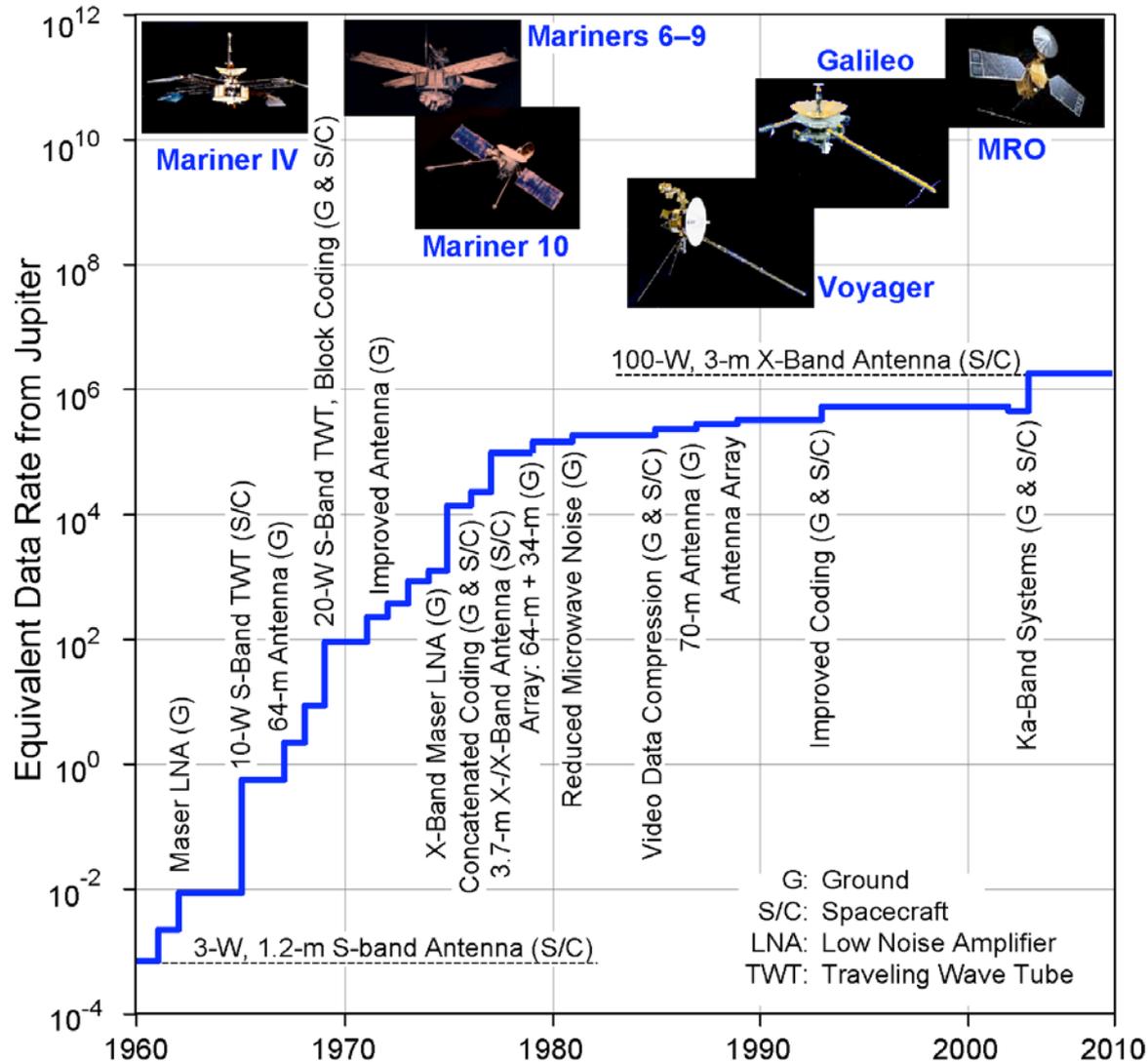


# Deep Space Network 70-m Antenna at Goldstone, California

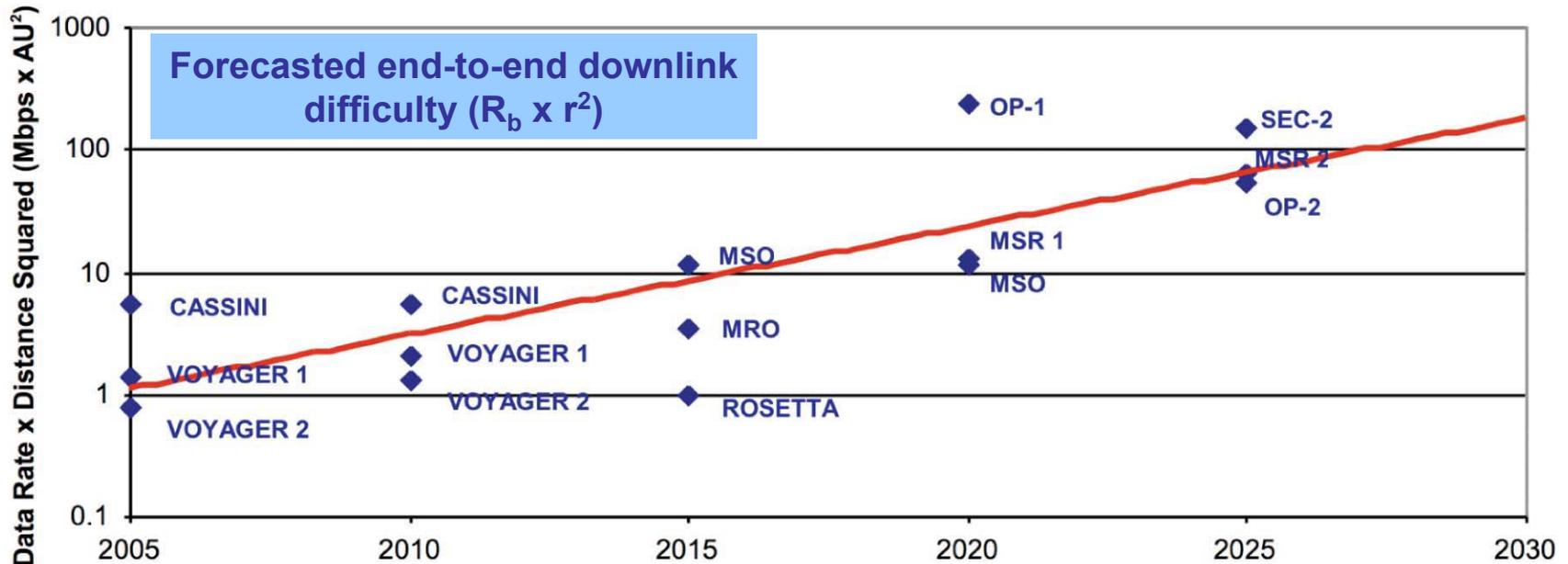
Artwork by Henk Pander



# History of Deep Space Telemetry



# Future of Deep Space Telemetry



- Maximum end-to-end downlink difficulty is expected to increase roughly two orders of magnitude by 2025
- Although radio frequency (RF) waves may be able to meet future demands, in practice, optical links may provide a better option for meeting such ominous goals.
- It will be difficult, if not impossible, to communicate at near gigabits per second from the deep space using RF means.

# Background

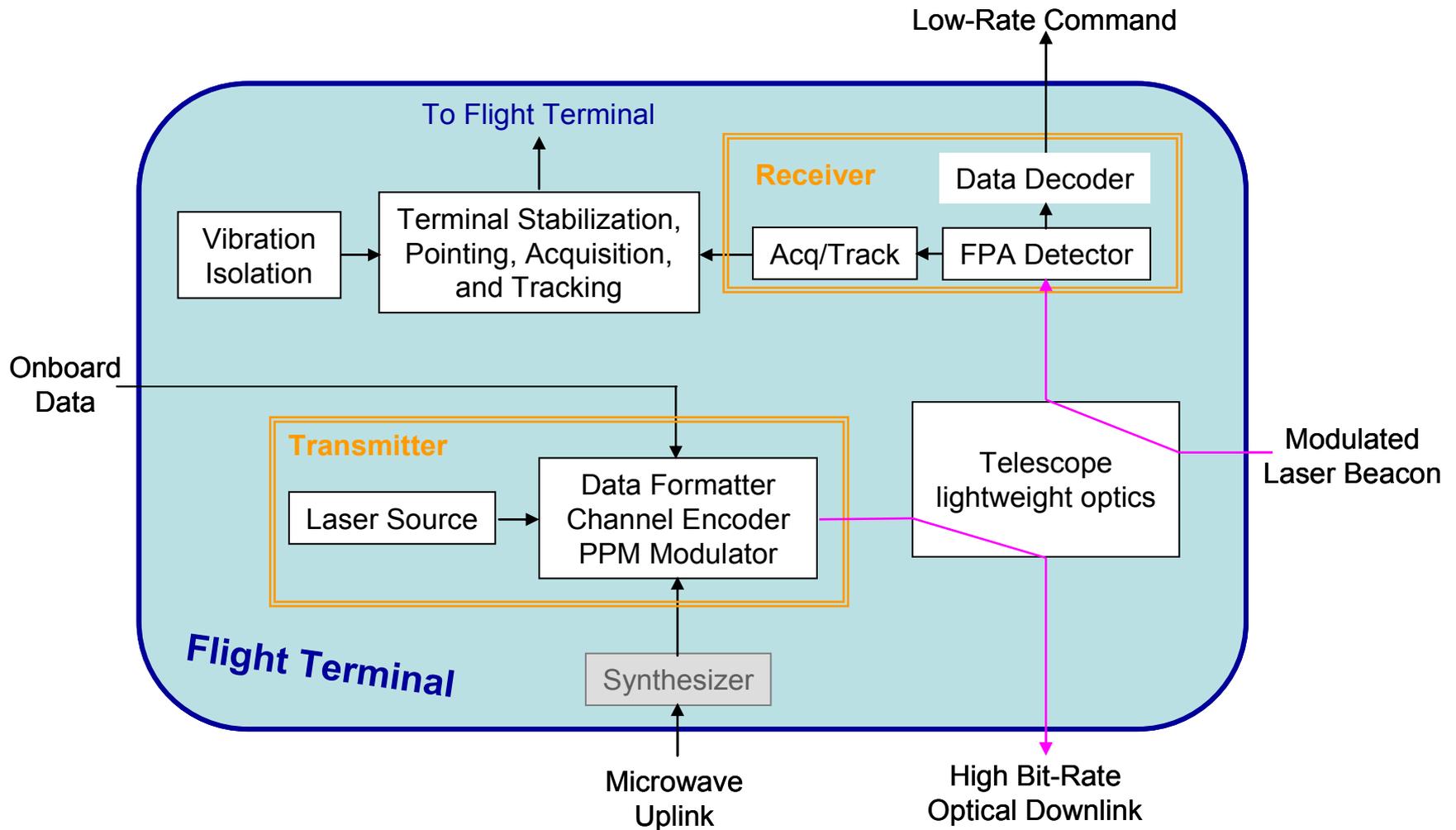
- An important challenge in developing technology for space applications is the identification and utilization of a flight opportunity to test the technology
  - no specific mission has been identified as the first user
- It is important to demonstrate the feasibility of deep space communications, albeit, at a low data rate.
- Feasibility can be proven by closing the link with adequate margin and good pointing and tracking performance
- The first generation of operational systems will use ground-based Earth terminals

# *Deep Space laser Communications*

## Challenges & Technical Solutions

- Optical Channel
  - Photon starved link due to vast planetary distances
  - High levels of background light during day time operation,
  - Reception in the presence of the atmosphere (attenuation, scintillation, signal loss, etc.)
- Technology solutions
  - Narrow laser beams focus energy for long distance applications
    - Space terminal vibration isolation and precise pointing
  - Pulse position modulation (PPM) at the transmit end
  - Photon counting at the receive end
  - Ground terminal static and dynamic corrections to combat telescope deformations and channel induced anomalies
  - Channel error correction via encoding

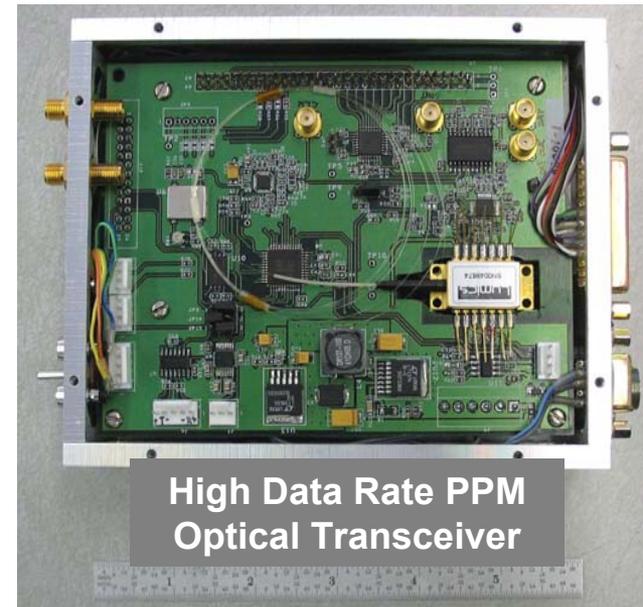
# Optical Flight Terminal Conceptual block Diagram



## *Optical Flight Terminal*

# High Data Rate PPM Transmitter

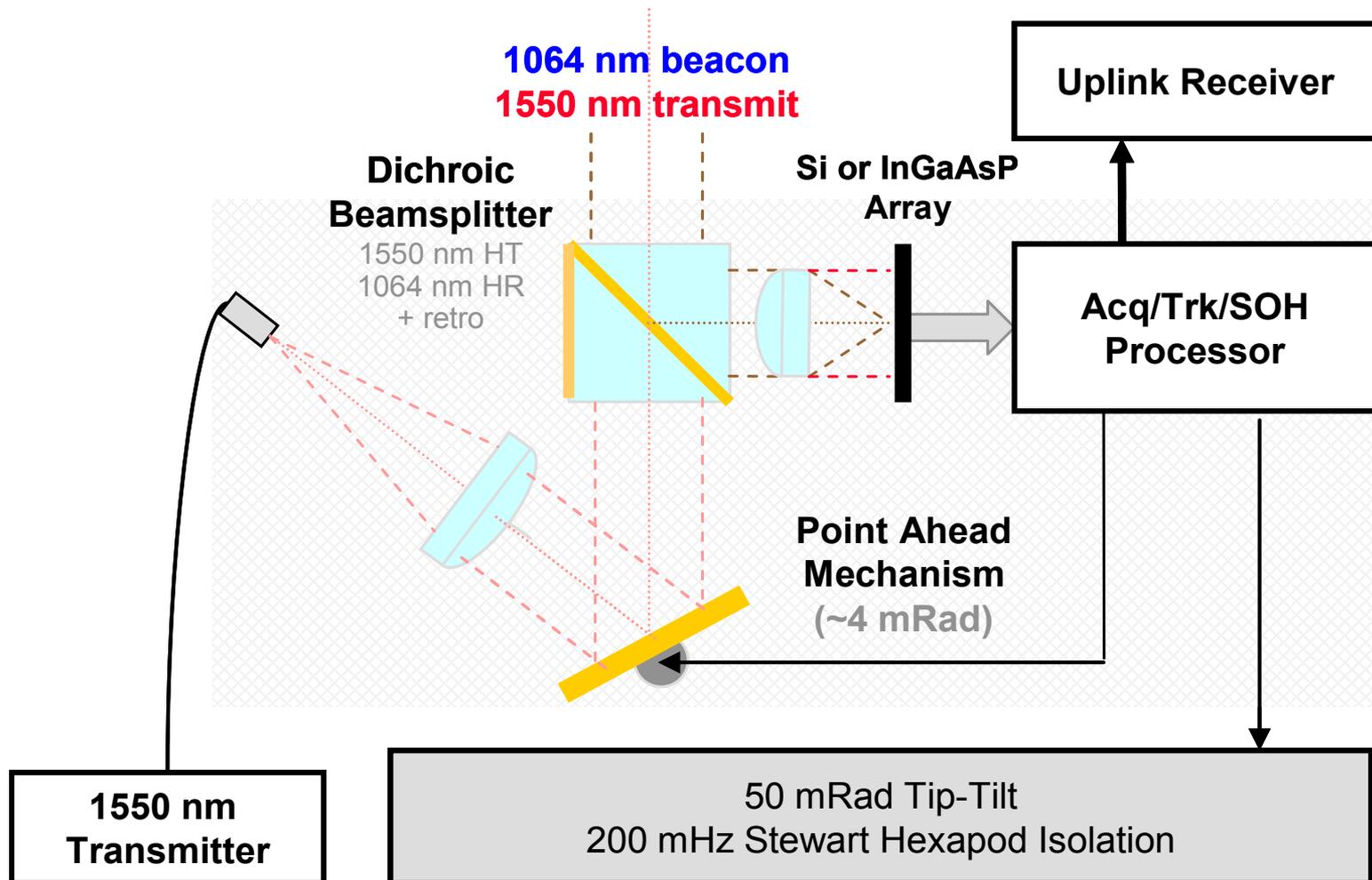
- High order PPM requires an efficient laser amplifier with a high peak to average power ratio
  - Yb fiber lasers operating near 1060 nm or Er/Yb fiber lasers operating near 1550 nm
- JPL is now implementing data rates in the 50 to 1200 Megabits per second range using sub-nanosecond PPM slot widths
- JPL has developed a series of FPGA encoders and modulators that operate a slot rates down to 100 picoseconds with real time data inputs to 1.2 gigabits per second.



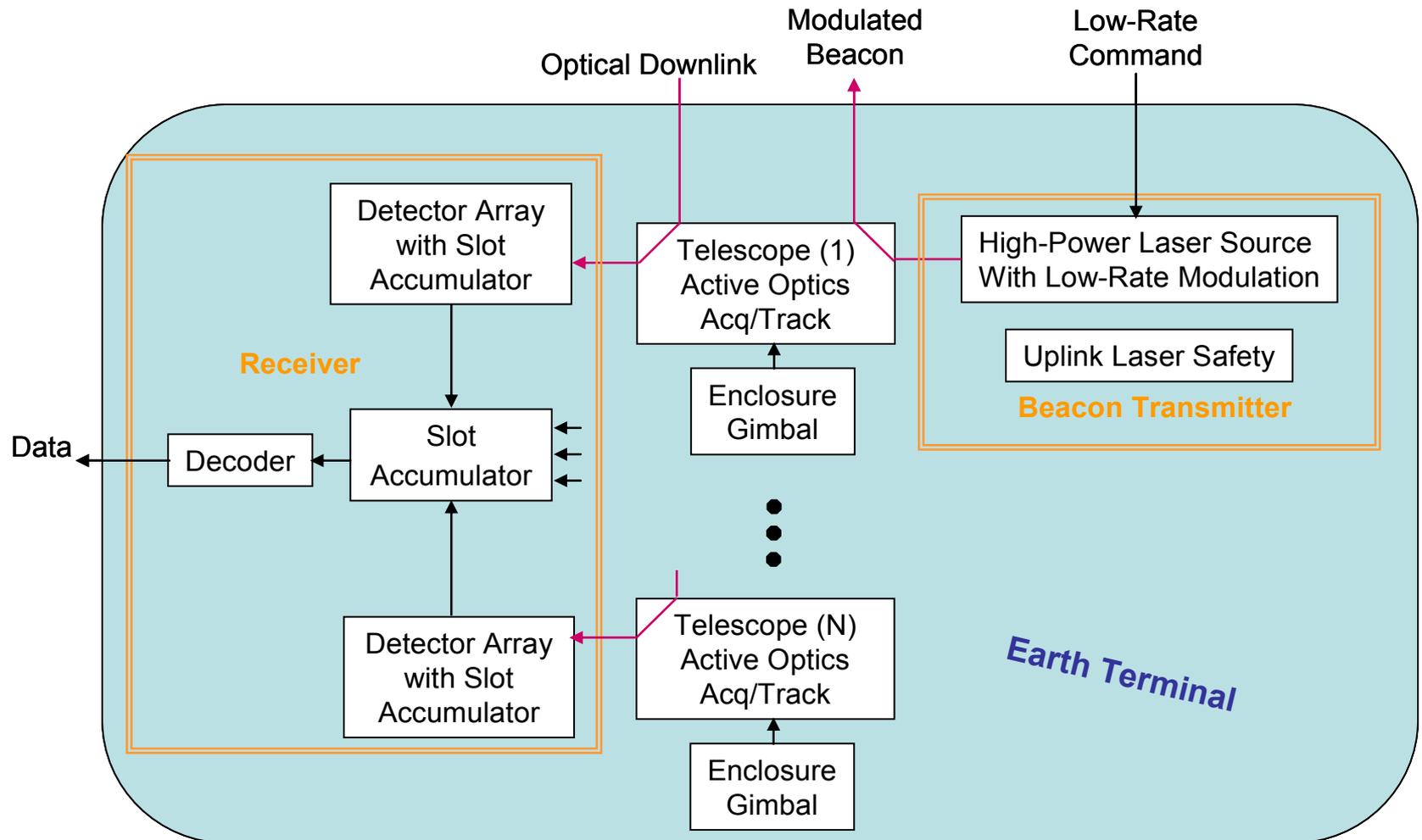
# *Optical Flight Terminal* **Flight Receiver**

- Flight photon counting detector qualification
  - InGaAs(P) single photon sensitive Negative Avalanche Feedback and Geiger mode detectors for space operation
- Development and demonstration of Doppler rate estimation algorithm in hardware based upon using GHz rate PPM slot clock

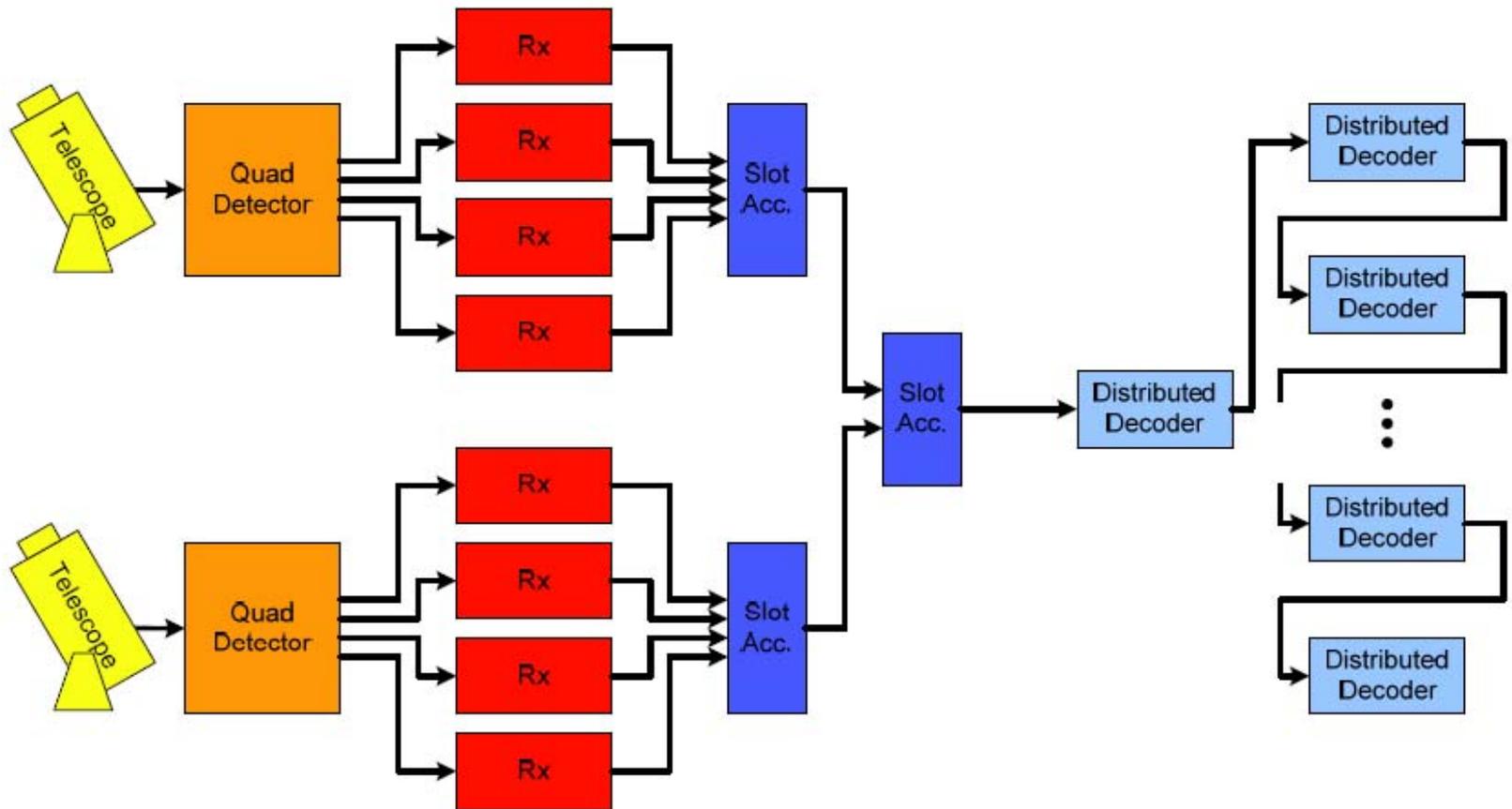
# Acquisition/Tracking/Pointing system based on Disturbance Free Platform concept



# Optical Ground Terminal Conceptual Block diagram



# Downlink Array Receiver Architecture

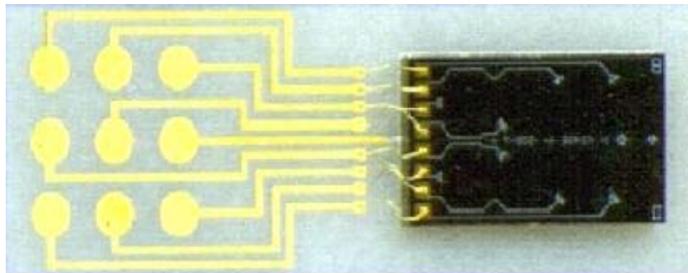


# High Efficiency Photon Counting Detector

- High single photon detection efficiency (SPDE) reduces spacecraft burden (less laser transmit power)
- Low output pulse jitter is required for reception of PPM signals at high data rates.
- Dark count rate requirements are less stringent than other photon counting applications due to the high background fluxes encountered during daytime link operations
- Detector count rate limitations (linearity) can be mitigated through the use of array architectures
- JPL has heavily invested in a state-of-the-art characterization facility for photon counting detectors
- The most viable technology presently commercially available is the InGaAs(P) intensified photodiode (IPD)
  - Was validated to photon count in the near-infrared with high efficiency (> 40% SPDE) and operate with minimal cooling (220K)
  - IPD is a vacuum tube device that suffers from photocathode degradation with a lifetime on the order of a few thousand hours.

# Detector, cont.

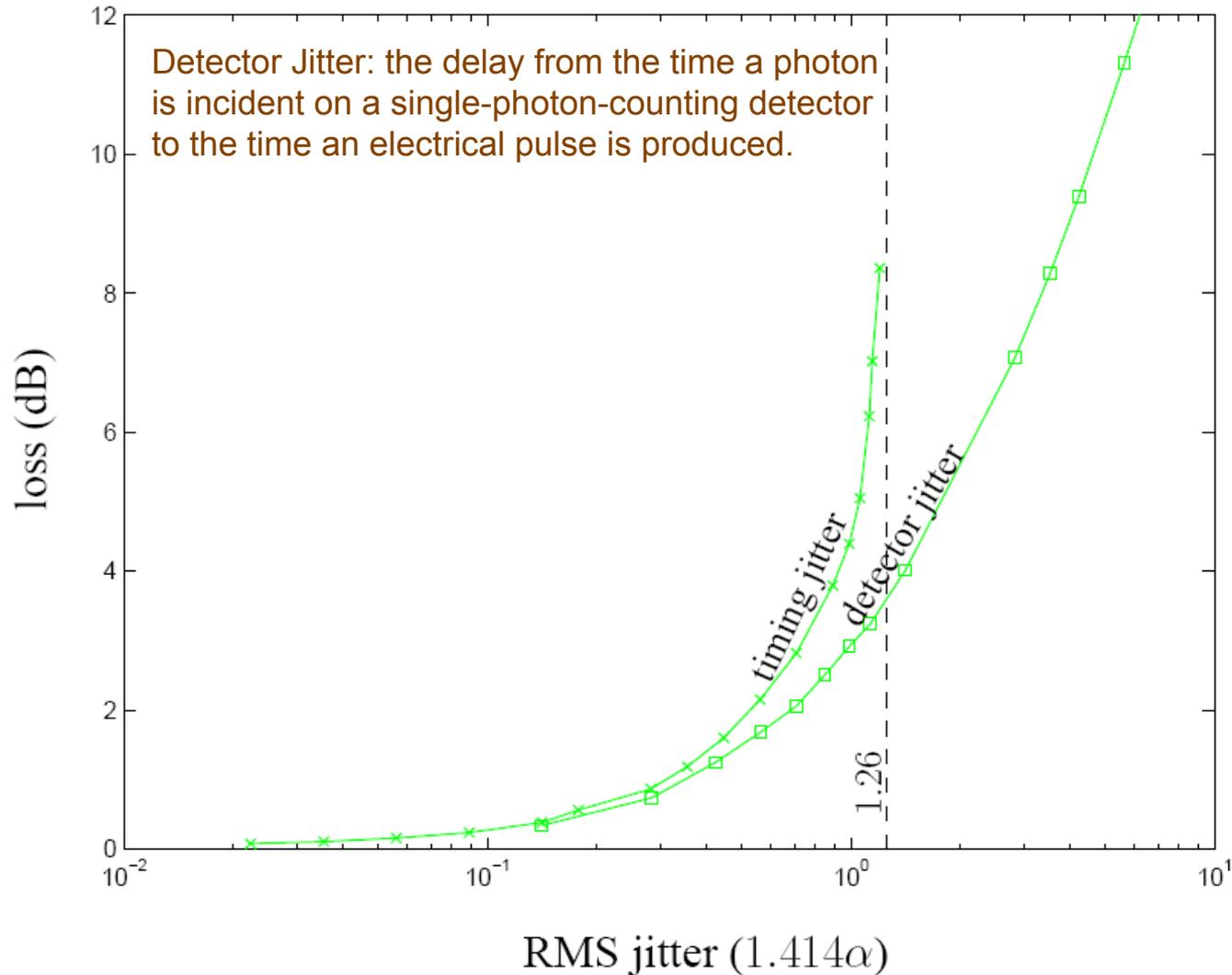
- Most promising has been demonstration Negative Avalanche Feedback (NAF) avalanche photodiodes with InGaAs absorbers for non-Geiger mode photon counting in the near infrared
- This technology has been previously demonstrated in silicon for visible wavelengths
- The NAF operational mode eliminates the requirement for the reset circuitry and read-out electronics that are required for semiconductor photon counters using Geiger mode operation.
- JPL has validated high gains (greater than  $10^5$ ) with excess noise factors near 1.0 in an InGaAs/InP material system and confirmed single photon response in sub-Geiger mode operation.



**Single Photon Detectors**

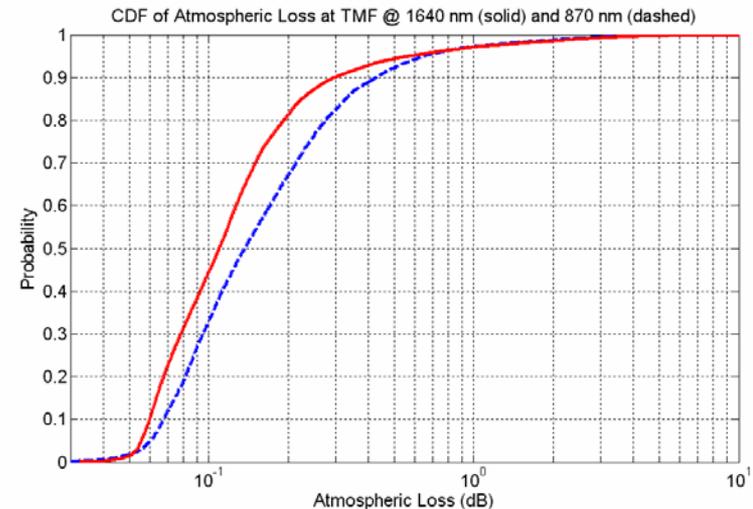
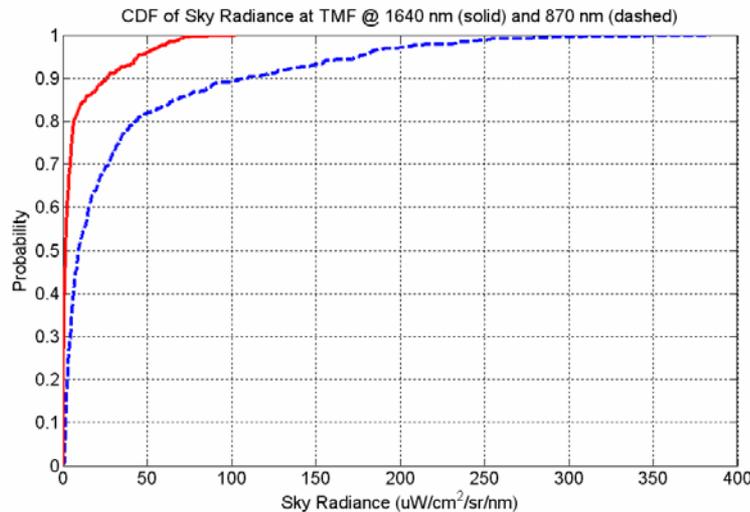
# Characterizing Receiver Anomalies

## Detector Jitter and Receiver Slot Synchronization Error



# Optical Channel

## Presence of Sunlight and Atmospheric Effects



- Atmospheric Visibility Monitoring instruments at Table Mountain
  - AERONET sun photometer (870, 1020, 1610  $\mu$ ) for sky radiance and attenuation monitoring
  - Telescope with MASS/DIMM for atmospheric turbulence monitoring
  - Infrared camera for day/night cloud coverage monitoring

# Summary

- Future deep space missions will demand telemetry services that can better served with optical links as opposed to radio links
- JPL is developing the first generation flight optical products for use by missions scheduled for launch in 2012 or later
  - Compact flight terminal
    - Flight terminal pointing and tracking
    - Space qualification
  - Large equivalent aperture ground receiver
    - Arrayed telescopes and arrayed detectors
  - Cryogenic station for detector characterization
  - End-to-end link characterization

**Cryogenic Probing Optical Detector Station Test Dewar**

