



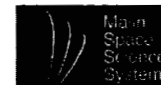
Inertial Measurements for Aerobraking Navigation (IMAN)

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Navigation Strategy for A/B will be similar to MGS and ODY

24/7 OD operations to reconstruct every drag pass, determine density

Daily strategy analysis to support ABM planning/decision

OD Process

Marshal inputs - ICs from last solution, attitude profile, atmos inputs

Collect tracking and SMFs after drag pass

Reconstruct previous orbit(s) drag pass densities using atmos model

Predict forward using MarsGRAM model with weekly wave determination

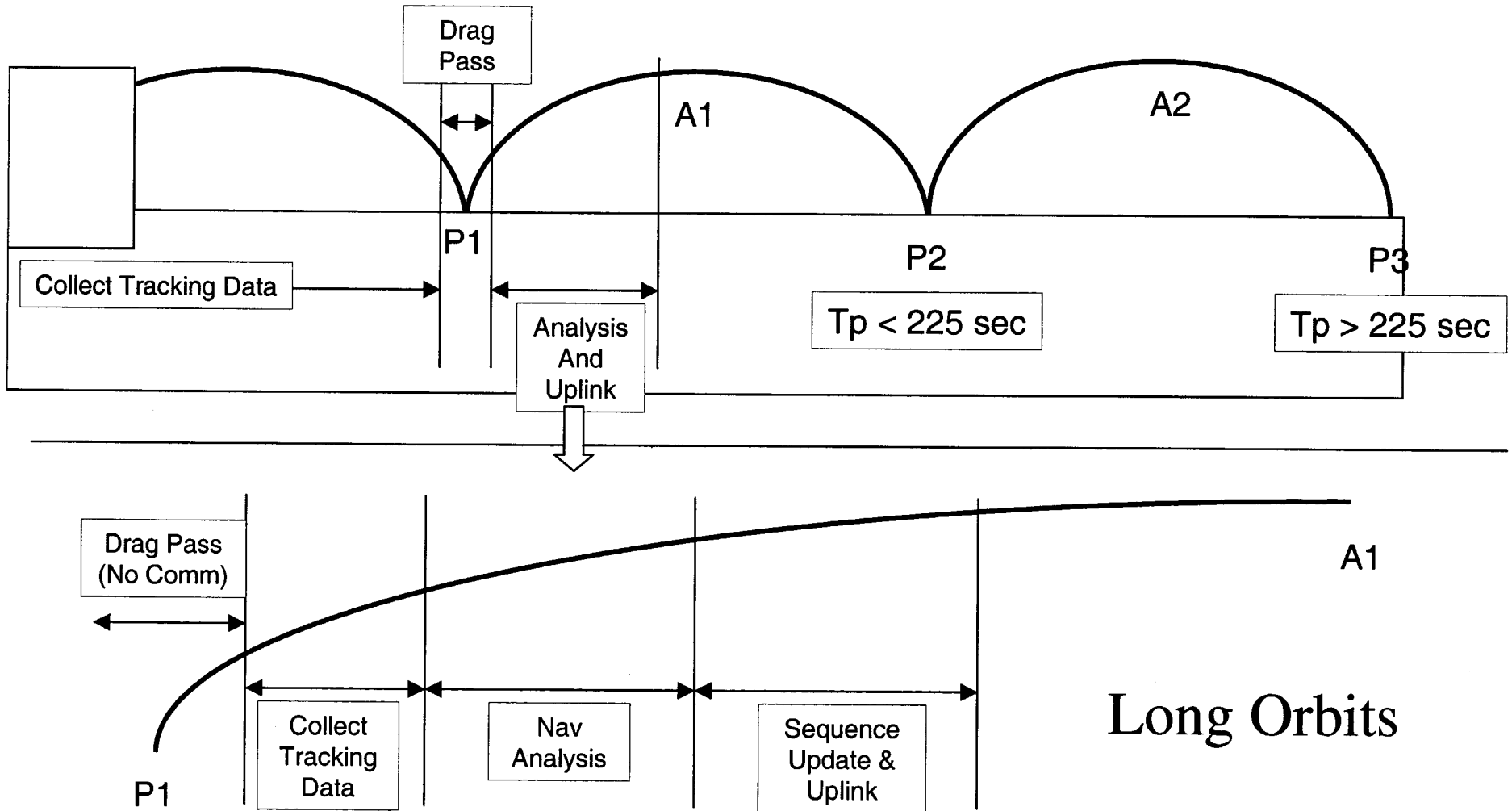
Deliver SPK and OPTG file

Daily Strategic Process

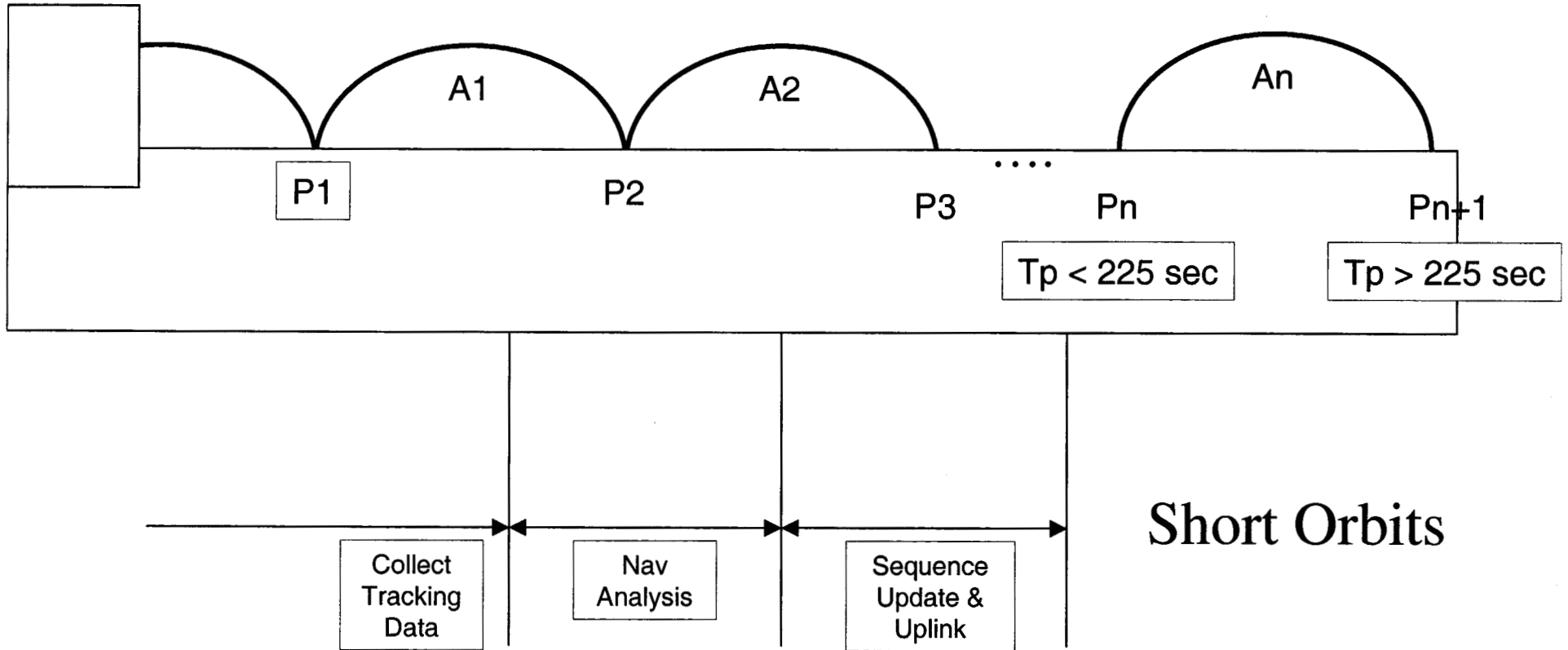
Runout deterministic OD solution, Compare to Baseline

Implement corridor control strategy, NAV recommends ABM decision

Aerobraking Navigation Process



Aerobraking Navigation Process





Aerobraking Nav Prediction Accuracy



- Requirement
 - Must predict Periapsis Time to within 225 sec
- Capability
 - Timing requirement uncertainty dominated by assumption on future drag pass atmospheric uncertainty
- Atmospheric Variability
 - Total Orbit-to-Orbit Atmospheric variability: 105% ($3\text{-}\sigma$)
- Periapsis timing prediction
 - To first order, the expected change in orbit period per drag pass will indicate how well future periapses can be predicted
 - This simplifying assumption is supported by OD covariance analysis

- Example
 - Total expected Period change for a given drag pass is 1000 seconds
 - Atmosphere could change density by 105%
 - Resulting Period change could be off by 105% = 1050 sec
 - If orbit Period is different by 1050 seconds, then the time of the next periapsis will be different by 1050 seconds
 - This fails to meet the 225 sec requirement
- Large Period Orbits
 - Period change per rev is large
 - Therefore can never predict more than 1 periapsis ahead within the 225 sec requirement with any confidence
- Small Period Orbits
 - Period change per rev is small (for example 30 seconds)
 - Therefore can predict several periapses in the future to within the 225 second requirement
 - Example: 105% uncertainty (31.5 sec) will allow ~ a 7 rev predict

Aerobraking Caveats

- The spacecraft must slew into the aerobraking orientation prior to each drag pass.
 - There is a loss of radiometric tracking precisely when the spacecraft “flies” through the most dynamically changing and unknown portion of its trajectory.
 - This leads to a significant increase in the spacecraft’s post-pass state (position and velocity) uncertainty.

- When reconstructing the drag pass, it is assumed that the spacecraft’s total change in velocity due to the atmospheric effects is purely due to drag.
 - In practice, the aerobraking orbits do not tend to fit the radiometric data unless residual noise is modeled as artificial dynamic acceleration events.
 - It is likely that the residual noise is due to the lack of modeling of aerodynamic lift and side-force.

- The aerobraking orbit reconstruction process is very time consuming (i.e. lasting several hours for each orbit) and workforce intensive (9 navigators for the Mars Odyssey aerobraking operations).

- All of the spacecraft events occur on a ground-generated timeline (i.e. a sequence of commands). At times, up to 3 sequences must be generated and successfully uplinked to the spacecraft every 24 hours.
 - The personnel required to perform this task constitutes an additional operational cost.

- Spacecraft events take place at times relative to the predicted time of periapse. Any error in this prediction could lead to:
 - aerobraking corridor control maneuver errors and thus inefficient propellant usage.
 - aerobraking drag pass attitude configuration slewing at off-nominal times, capable of inducing inadvertent compensative thruster firings, and thus another source of inefficient propellant usage (inadvertent safe-mode entry triggering is another possible outcome).

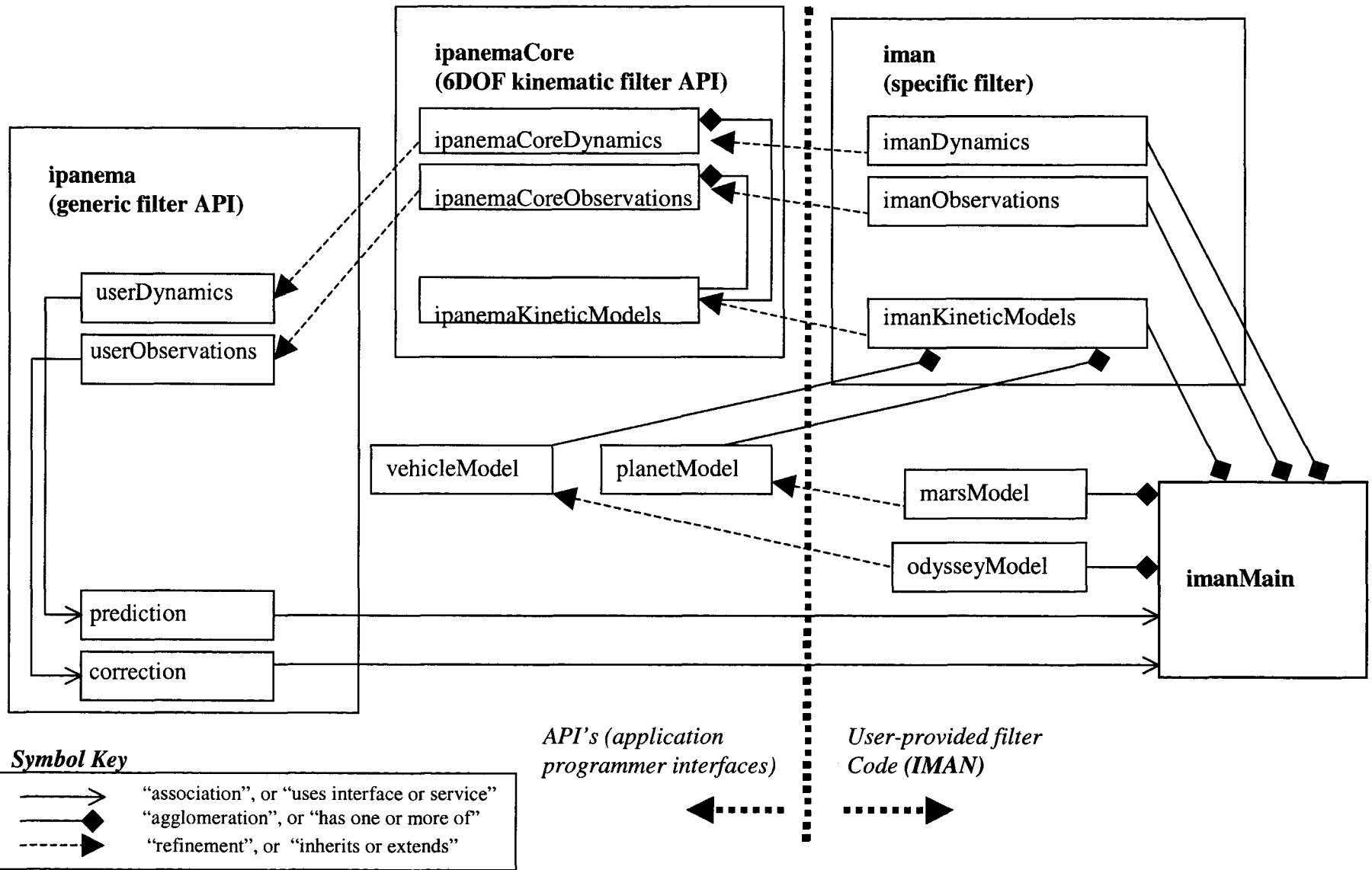
- Questions:
 - Is there a way to avoid the traditional tracking data gap?
 - Is there a way to improve our atmospheric models?
 - Is there a way to be less sensitive to tracking problems after a drag pass?
 - Is there a way to increase our knowledge of the spacecraft state?

- Answers:
 - There may be if the state is being measured through the drag pass.
 - Perhaps, if we can estimate for atmospheric model parameters of interest during a drag pass.
 - Yes, if the drag pass is measured as it occurs instead of after.
 - Yes, if the drag pass data type used is “strong” enough.
 - Data Noise and Observation/State Partial.

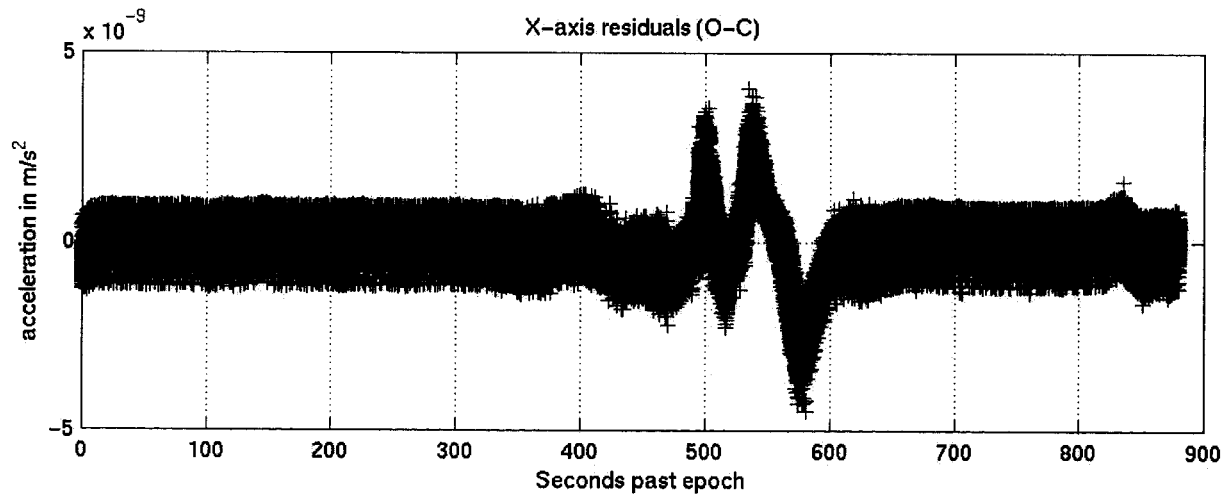
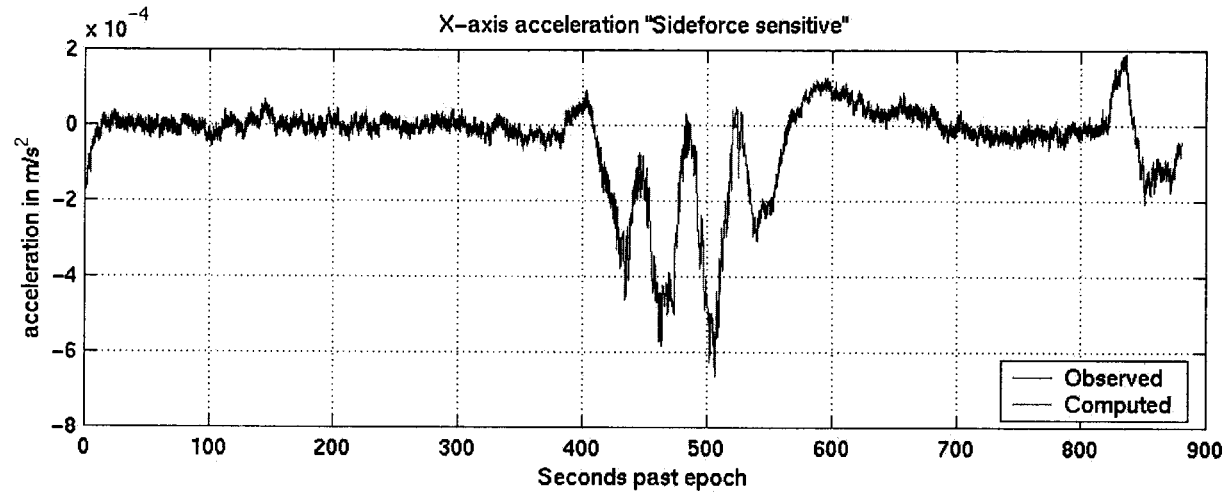
- 6-DOF drag pass reconstruction via IMU data processing could provide the answers.

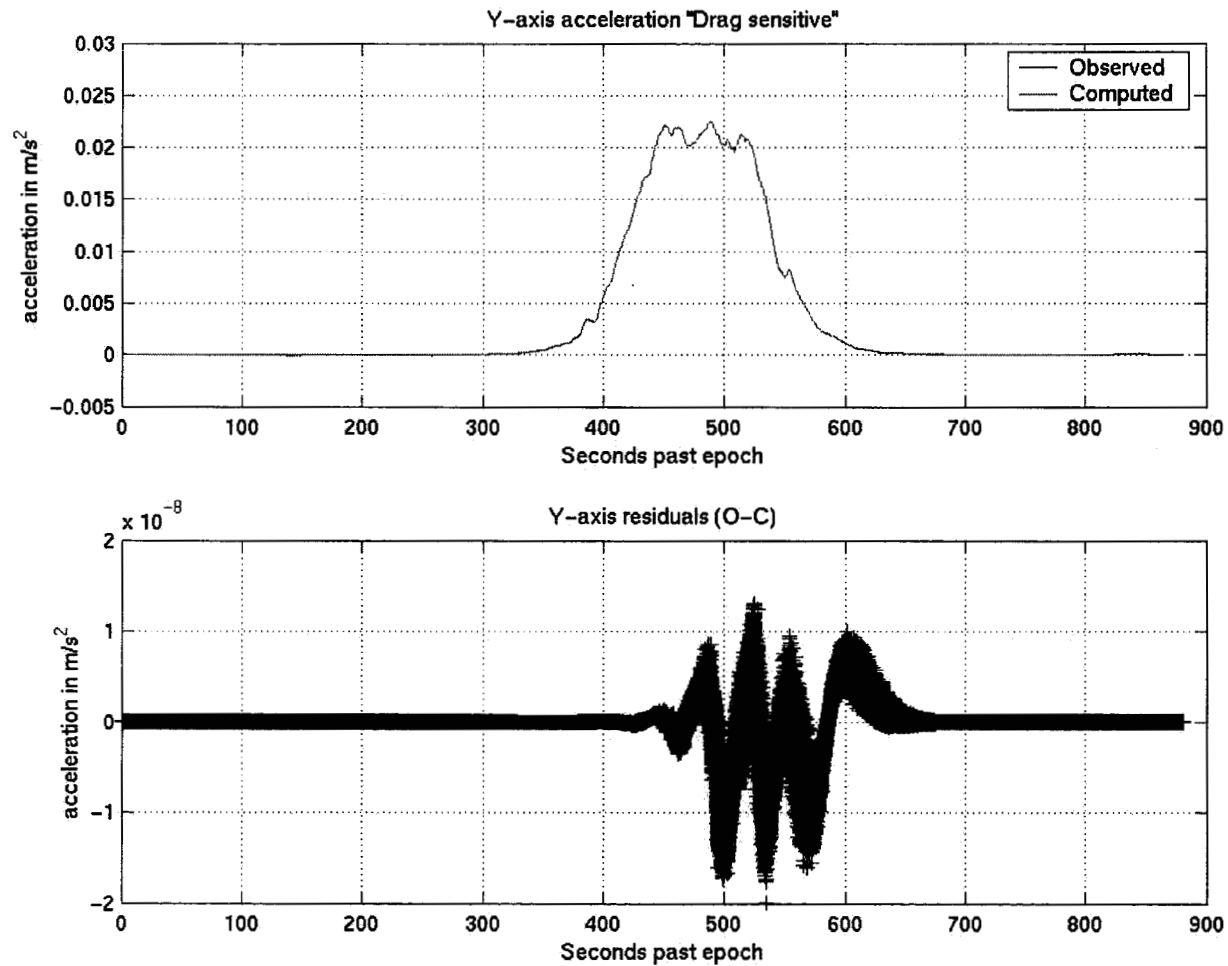
- IPANEMA (Java)
 - 6-DOF Kinematic State Filter with DMC
 - 19 Kinematic States
 - $X = [r, v, a, q, \omega, \alpha]$
 - Capable of ingesting user defined force, torque, spacecraft, and planetary models and state parameters

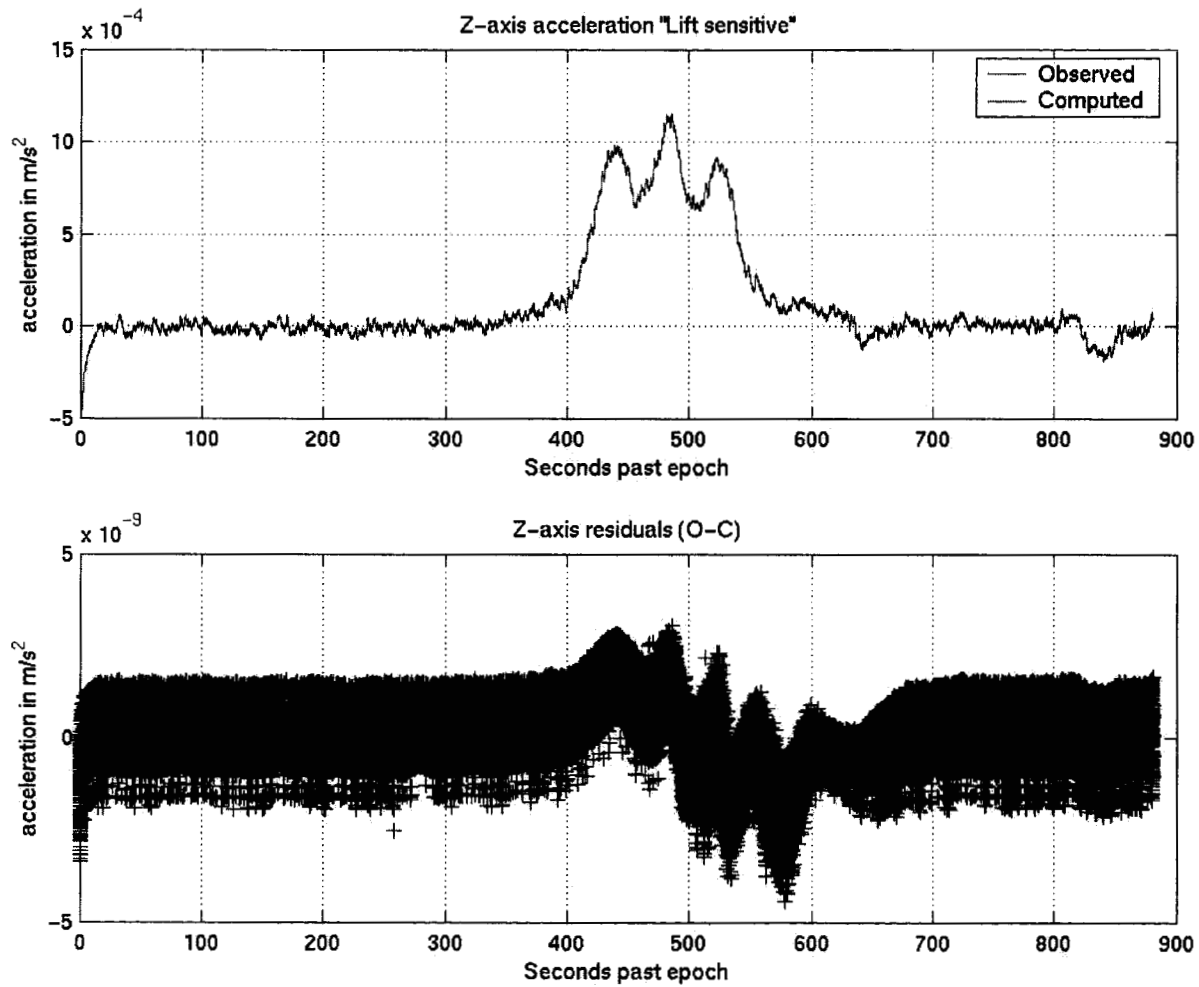
- IMAN (Java/Python)
 - Aerobraking-tailored force, spacecraft, and planetary atmosphere model with additional state parameters

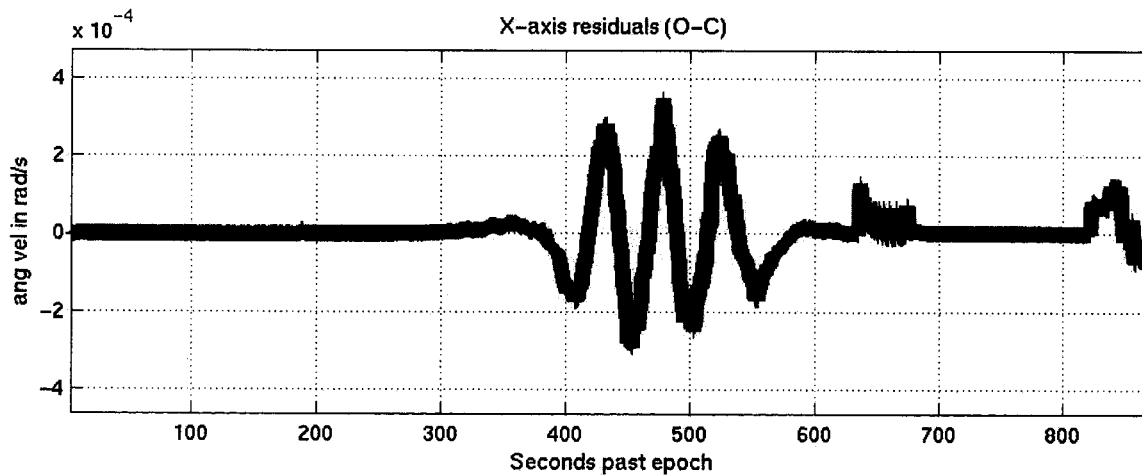
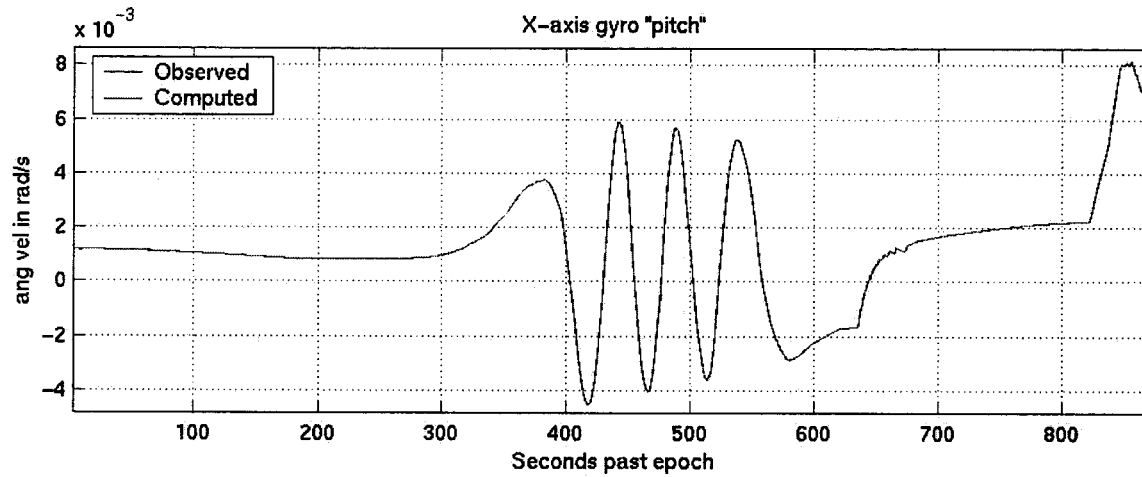


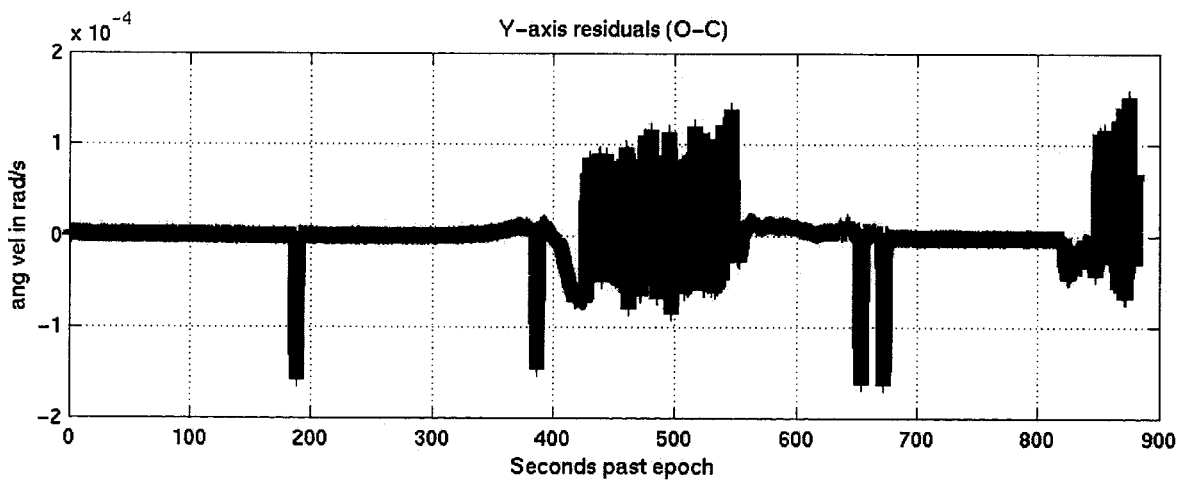
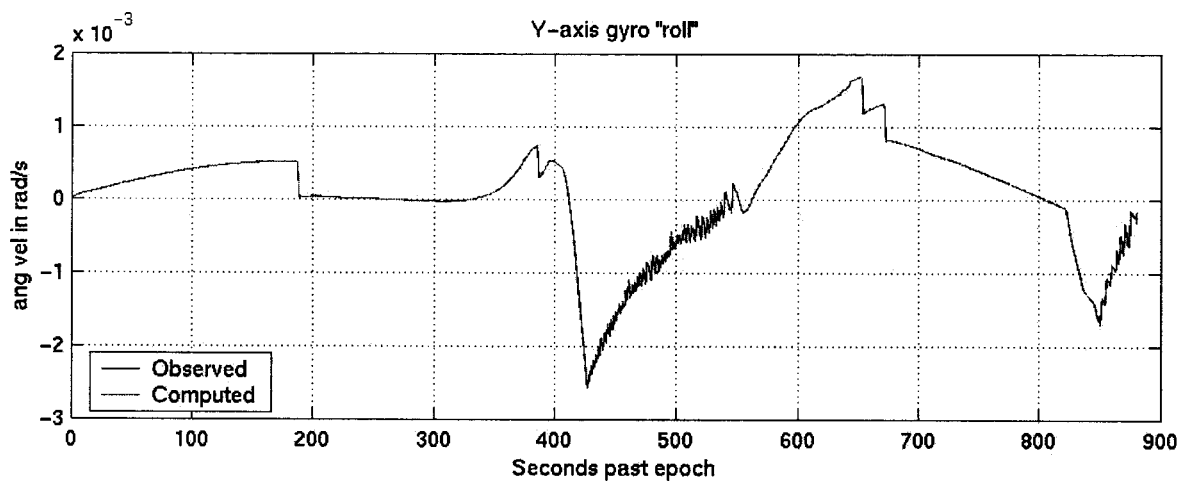
- Obtain pre drag-pass state and covariance based on best pre drag-pass radiometric fit.
- Generate predicted atmospheric model parameters for upcoming drag-pass.
 - Could be generated based on MarsGRAM.
- Obtain IMU data from telemetry
 - May have to be pre-processed.
- Initialize IMAN with pre drag-pass state and covariance and obtain post drag-pass state and covariance estimate based on 6-DOF fit.

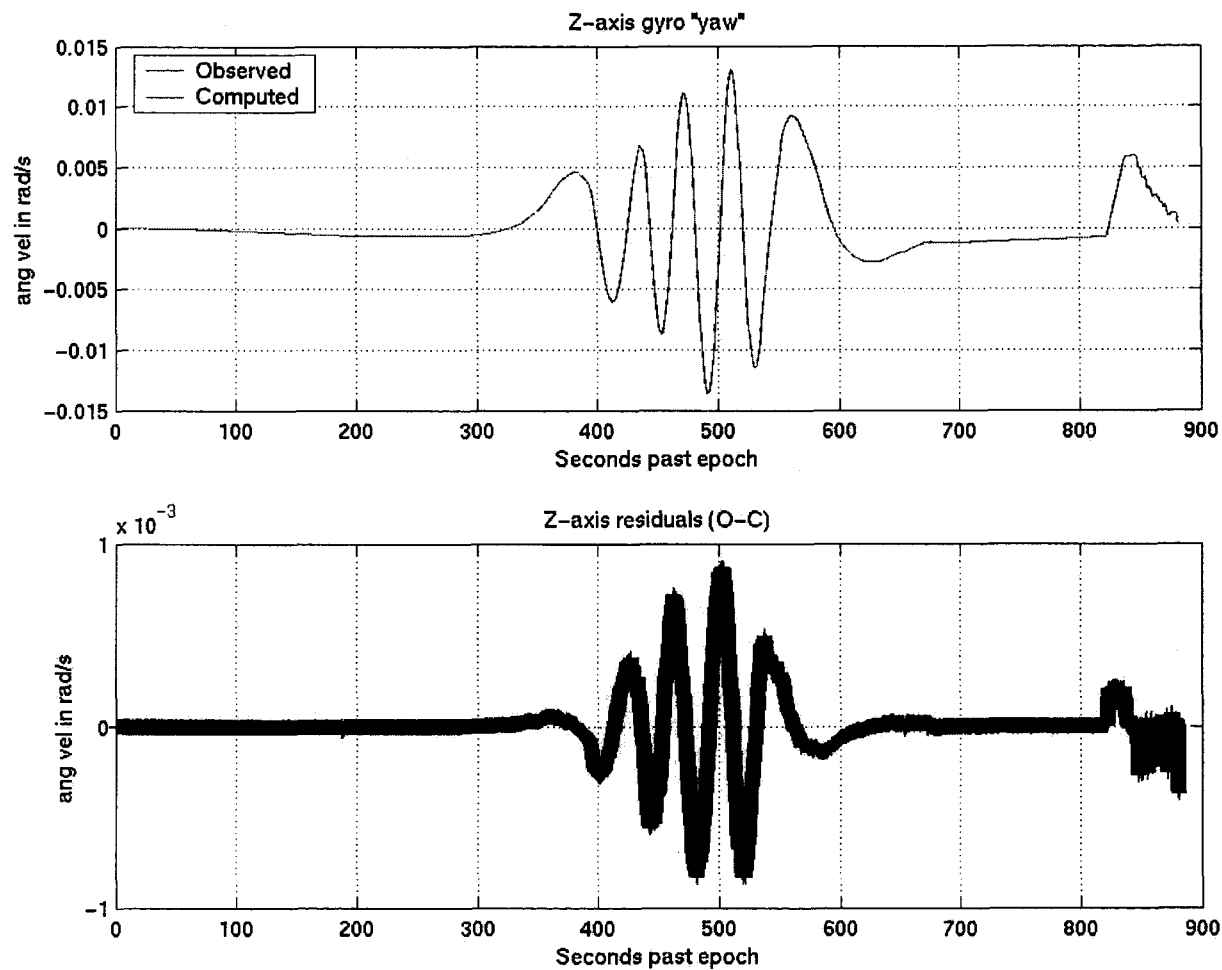


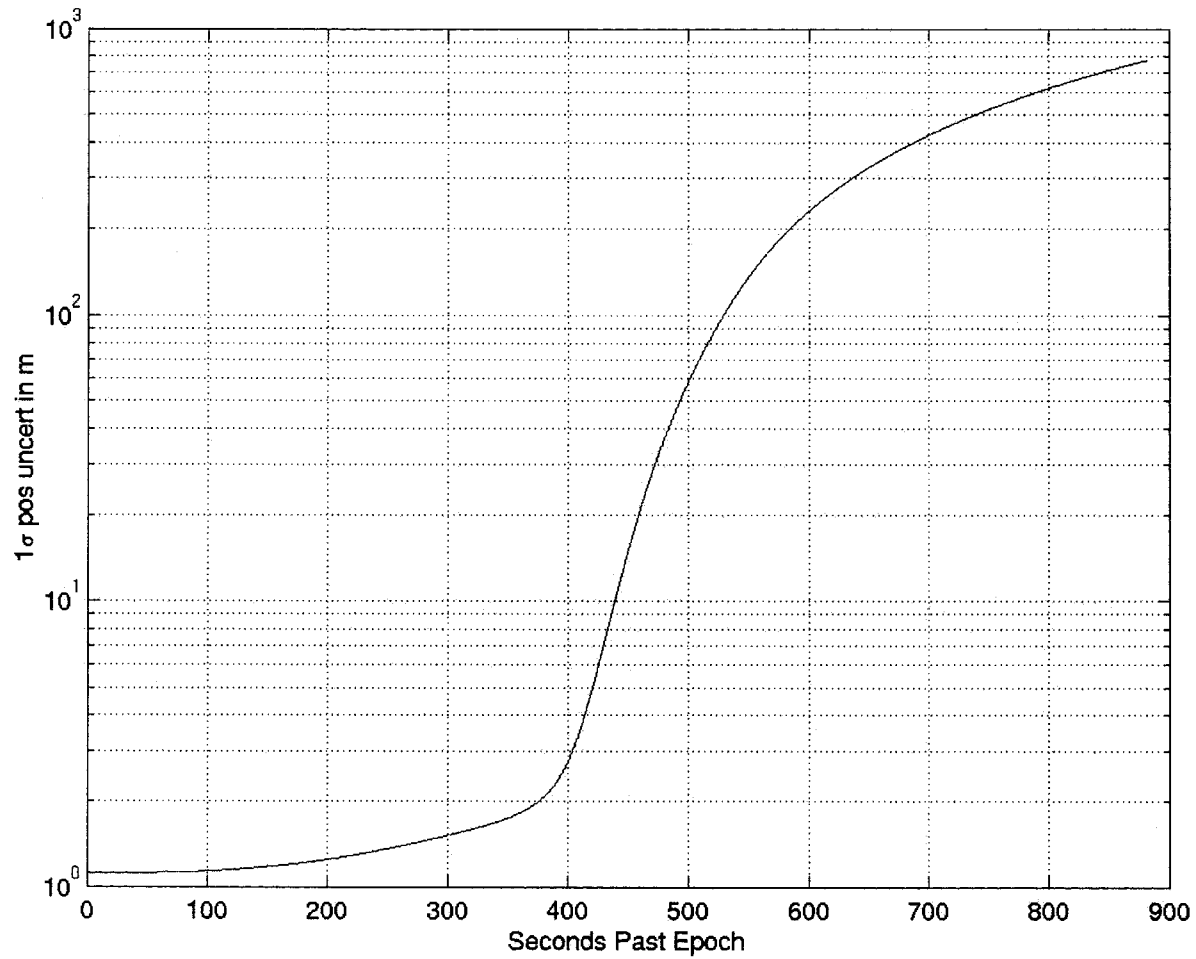












- The results thus far show:
 - It is possible to fit an aerobraking trajectory in a 6-DOF sense, by using IMU data as a navigation measurement.
 - A pure kinematic state covariance shows realistic uncertainties in the state given some nominal atmospheric model.

- Things to be done (next steps):
 - add in a higher fidelity gravity model (perhaps up to a 4x4 field).
 - add in observation model states, namely estimate for things like base density and scale height.
 - perform more filter tuning.
 - compare results against navigation team solutions.
 - fit representative aerobraking orbits of various orbital periods (e.g. 16 hour, 6 hour, 3 hour).
 - determine sensitivity to *a priori* covariance
 - try various measurement compression rates (i.e. 50 Hz, 20 Hz, 10 Hz, 1 Hz)