

# Absolute Charge Exchange Cross Sections for O<sup>5+</sup>, O<sup>6+</sup> and O<sup>7+</sup> Collisions with CO and CO<sub>2</sub>

R. Mawhorter<sup>1</sup>, N. Djurić<sup>2</sup>, J. MacAskill<sup>2</sup>, S. J. Smith<sup>2</sup>, A. Chutjian<sup>2</sup>, and I. D. Williams<sup>3</sup>



<sup>1</sup>Pomona College, Claremont, CA and JPL/Caltech, Pasadena, CA

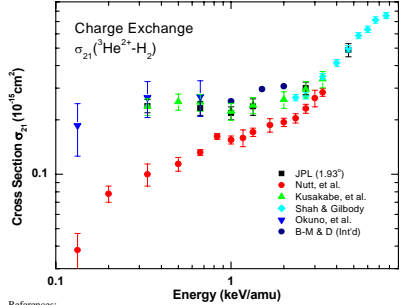
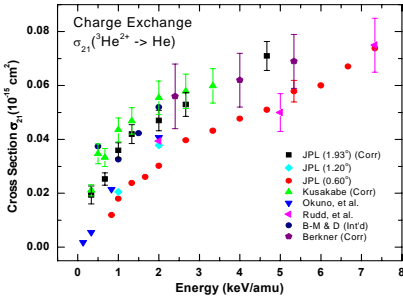
<sup>2</sup>Jet Propulsion Laboratory/Caltech, Pasadena, CA

<sup>3</sup>Queens' University Belfast, UK



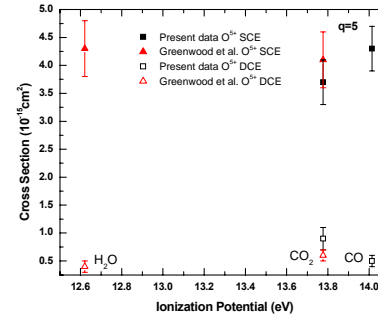
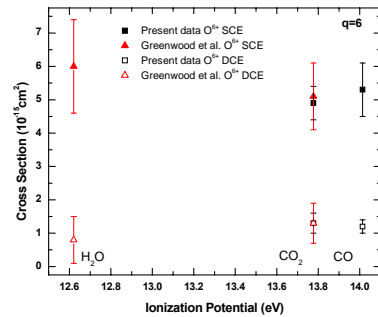
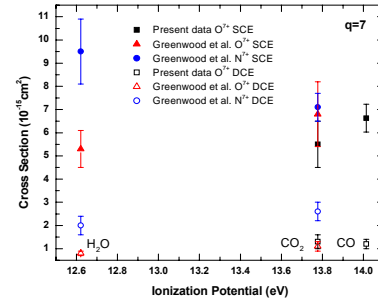
Motivated by ongoing EUV and X-ray studies of comets, we have continued our experimental investigations of absolute charge exchange cross sections for highly-charged ions present in the solar wind incident on cometary gases. These are the first measurements on the JPL charge exchange beam-line using a larger gas cell exit aperture. Data for O<sup>5+</sup> and O<sup>7+</sup> on CO<sub>2</sub> agree with earlier measurements [1], and are included in these new results for O<sup>5+</sup>, O<sup>6+</sup> and O<sup>7+</sup> on CO and CO<sub>2</sub>. The ion beam accelerating potential was 7 kV, which yields ion velocities consistent with the fast component of the solar wind. Agreement with earlier, smaller exit aperture measurements is also significant in demonstrating an independence from angular collection issues for these fast, heavy ions and targets. This was verified by studying collection angle-cross section effects for slow <sup>3</sup>He<sup>2+</sup> ions on He and H<sub>2</sub>.

## Acceptance Angle Studies

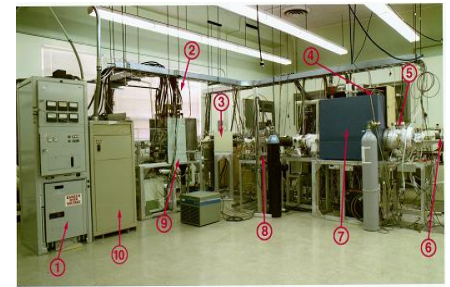


References:  
 JPL: unpublished data; Angles indicate exit aperture (half) acceptance angle  
 T. Kusakabe et al., J. Phys. Soc. Japan **59**, 1218 (1990).  
 K. Okuno et al., J. Phys. B **25**, L105 (1992).  
 W. L. Nutt et al., J. Phys. B **11**, 1457, (1978).  
 D. Bordeneuve-Montesquieu and R. Dagnac, J. Phys. B **25**, 2573 (1992), J. Phys. B **27**, 543 (1994).  
 M. E. Rudd et al., PRA **32**, 2128 (1985).  
 K. H. Berkner et al., Phys. Rev., **166**, 44 (1968).  
 M. B. Shah and H. B. Gilbody, Phys. B **7**, 256 (1974).

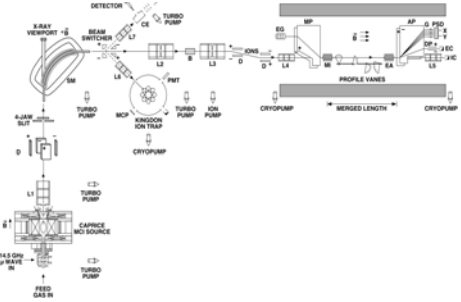
## JPL Experimental Results for O<sup>5+</sup>, O<sup>6+</sup> and O<sup>7+</sup>



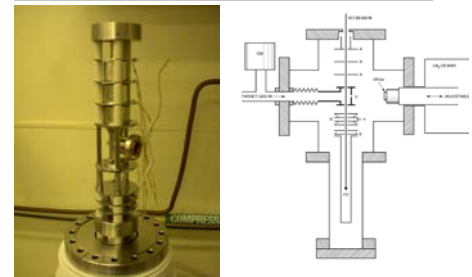
## The JPL Highly-Charged Ion Facility



- ① high-power microwave amplifier
- ② ECR current and cooling lines
- ③ selection magnet
- ④ solenoidal magnet
- ⑤ feedthrough collar
- ⑥ vanes drive
- ⑦ merged-beams chamber (in magnet bore)
- ⑧ f-value chamber
- ⑨ ECR
- ⑩ ECR magnet supply



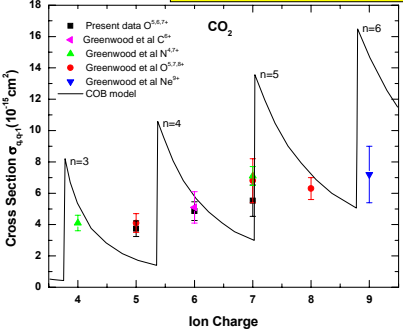
## Photograph and Schematic of CE Gas Cell



## Gas Flow Through Cell



## JPL Experimental Results, $\sigma$ ( $10^{-15}$ ) $\text{cm}^2$ , $E_i = qx7 \text{ keV}$

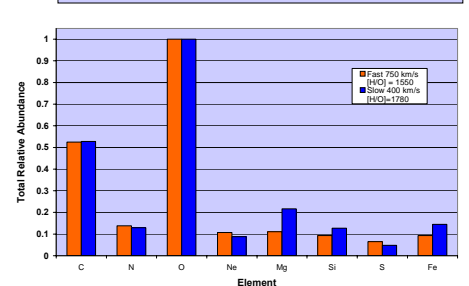


Gas/ion	C <sup>3+</sup>	C <sup>4+</sup>	N <sup>4+</sup>	N <sup>5+</sup>	O <sup>5+</sup>	O <sup>6+</sup>	O <sup>7+</sup>	O <sup>8+</sup>
H <sub>2</sub> O	$\sigma_{q=1}$	1.5±0.1	6.0±1.4	2.9±0.2	9.5±1.4	4.3±0.5	5.3±0.8	6.3±0.7
	$\sigma_{q=2}$	0.8±0.1	0.8±0.7	0.7±0.1	2.0±0.4	0.4±0.1	0.8±0.1	1.5±0.5
	$\sigma_{q=3}$				<0.3	0.3±0.2	<0.3	<0.3
CO <sub>2</sub>	$\sigma_{q=1}$	1.1±0.1	5.1±1.0	4.1±0.6	7.1±0.6	4.1±0.5	6.8±1.4	6.3±0.7
	$\sigma_{q=2}$	1.2±1.1	1.3±0.6	1.1±0.1	2.6±0.4	0.6±0.1	1.1±0.2	1.5±0.5
	$\sigma_{q=3}$				<0.3	0.6±0.1	0.3±0.1	<0.5
Present CO <sub>2</sub>	$\sigma_{q=1}$				3.7±0.4	4.9±0.5	5.5±1.0	
	$\sigma_{q=2}$				0.9±0.2	1.3±0.3	1.3±0.3	
	$\sigma_{q=3}$					0.4±0.1	0.4±0.2	
Present CO	$\sigma_{q=1}$				4.3±0.4	5.3±0.8	6.6±0.6	
	$\sigma_{q=2}$				0.5±0.1	1.2±0.2	1.2±0.2	
	$\sigma_{q=3}$					0.2±0.1	0.1±0.1	

## Comparisons of Theoretical Results<sup>(2)</sup> with JPL Experiments<sup>(1)</sup> for N<sup>7+</sup>, $\sigma$ ( $10^{-15}$ ) $\text{cm}^2$

	H <sub>2</sub> O	CO <sub>2</sub>	CO
JPL <sup>(1)</sup> experiment E=3 keV/amu	9.5 ± 1.4	5.3 ± 0.8	---
CTMC (Classical Trajectory Monte Carlo) E=2keV/amu	9	8	8
LZ (Landau-Zener) E=2keV/amu	2.1	2.4	2.1
COB (Classical Over-the-Barrier) E=2 keV/amu	12	3.0	3.1

## Solar Wind Abundances Schwadron & Cravens, ApJ 544,558(2000).



1. J. B. Greenwood et al., PRA **63**, 062707 (2001); J. B. Greenwood, et al., ApJ, **529**, 605 (2000); J. B. Greenwood, et al., Phys. Scr. **T110**, 358 (2004).

2. Hasan et al. Ap.J. **560**, L201, (2001).