

AN: **A32A-04**

TI: Remote measurements of tropospheric trace gases by solar absorption spectrometry

AU: G. C. Toon

EM: toon@mark4sun. jpl.nasa .gov

AU: B. Sen

EM: sen@mark4sun. jpl.nasa.gov

AU: J.-F. Blavier

AF: Jet Propulsion Laboratory, 4800 Oak Grove Drive, Pasadena, CA 91109, USA

EM: blavier@mark4sun. jpl .nasa.gov

AB: Recent balloon flights of the JPL MkIV Interferometer have obtained sunset spectra down to very low tangent altitudes (4 km) . Measured from an altitude of 39 km, these spectra are virtually identical to those that would be measured from space, and therefore can be used to assess the feasibility of spaceborne tropospheric remote sensing by the solar occultation technique. Each MkIV spectrum covers the entire 650 to 5650 cm⁻¹ spectral range simultaneously at 0.02 cm⁻¹ resolution. This wide spectral coverage and high resolution, in conjunction with the high signal-to-noise ratio inherent to the solar absorption technique, allows over 30 different gases to be measured simultaneously **in** the same airmass. These gases include several of tropospheric interest such as C₂H₆, C₂H₂, HCOOH, H₂O₂, H₂CO, CO, H₂O, **HDO, HCN, HN04, HNO3, NO2, NO** and O₃. This paper will discuss some of the factors which currently limit the accuracy of these retrievals and the possibility of future improvement.

SC: A

DE: 0365

DE: 0394

MN: 1997 Fall Meeting