1. INTRODUCTION

The NASA/JPL airborne precipitation radar APR-2 (cross-track scanning, dual frequency - 14 and 35 GHz, Doppler and dual polarization, see Sadowy et al. (2003) for detailed description of the instrument) was operated on the NASA P-3 aircraft during the Wakasa Bay (Japan) experiment. The experiment conducted jointly by the U.S. AMSR-E and Japanese AMSR teams in January/February 2003, was designed to (1) validate both the AMSR and AMSR-E shallow rainfall and snowfall retrieval capabilities, (2) extend the database of rainfall properties needed to implement a comprehensive physical validation scheme, and (3) extend our understanding of rainfall structures through the use of new remote sensing technology. On 12 flights, more than 30 hours worth of precipitation systems were observed, including rain and snow events, both over ocean and over land. On 8 of these flights, APR-2 observed stratiform rain and the associated melting layer, whose impact on passive remote sensing of precipitation is one of the specific issues addressed by this experiment. In this study the signatures of the melting layer on APR-2 measurements from four flights are presented and discussed.

2. OBSERVATIONS AND DISCUSSION

APR-2 operates at 13.4 GHz (Ku band) and 35.6 GHz (Ka band). Copolarized power (HH), crosspolarized power (HV) and mean Doppler velocity (v) data are collected at both frequencies.

Figure 1 shows the typical multiparametric data collected (at nadir) from a stratiform event. Panels a and b show the equivalent reflectivity expressed in dBZ as calculated from HH. Calibration of Ku band was verified examining the measured backscatter cross section at 10° incidence angle under clear conditions. The average normalized cross section σ^2_HH over all flight days was of 7.5 dB which is in good agreement with previous studies. Calibration of Ka band was verified comparing the dual-frequency backscattered power of very light rain in non-attenuated conditions (i.e., where backscattering is in Rayleigh regime at both frequencies and therefore equal reflectivity is expected). In this example the typical peak brightness signature at Ku band is evident between 1.5 and 2 km altitude. On the other hand, the brightness signature at Ka band is significantly weaker (most often a step rather than a peak) because of the weaker sensitivity of Ka band to the large particles that cause the peak at lower frequencies. Also, the effect of attenuation at Ka band is evident in the area of strongest precipitation. Panel c shows the vertical velocity measured from the Doppler return of the Ku band signal. The sharp increase in vertical velocity (from approximately 2 m s−1 to 7 m s−1) is due to the increased fall speed of the liquid hydrometeors with respect to the frozen ones and it takes place immediately below the brightness signature. Disruption of the brightness caused by light convection is visible in the first 2 minutes of observation. Panels d and e show the Linear Depolarization Ratio at the two frequencies. The sensitivity thresholds for LDR measurements are -30 dB for Ku band and -17 dB for Ka band. Significant returns are observed only in correspondence of the melting layer because of the presence of wet-ice particles. Finally, panel f shows the flight track, the green and red circles correspond to the beginning and end of the sampling period, respectively.

Table 1 shows the mean and standard deviation of 5 melting layer parameters from 10 segments of APR-2 nadir observations collected on January 19th, 21st, 23rd, and 27th, 2003. In generating the statistics of Table 1, only data with a measured LDR_{Ka} peak were used (for a total of 4233 sample profiles). Statistics calculated from LDR_{Ka} and v were further limited to data where those parameters were available. The maximum altitude of the brightband was observed at almost 3 km on January 24th, the minimum was observed at 0.5 km (right above the sea surface signature) on January 21st. The negative correlation (~0.5) between brightband altitude (Z_{BB}) and maximum brightband reflectivity (Z_{Ka}) observed in previous studies (e.g., Durden et al. 1997) and attributed to the cooling effect of the melting process was confirmed within each segment but the one in January 21st, where the brightband rose gradually from the sea level up to around 1.5 km because of the changing mesoscale conditions.

<table>
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<th>PARAMETER</th>
<th>Jan 19</th>
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<th>Jan 27</th>
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<td>27.5</td>
<td>27.5</td>
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<tr>
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<td>-13.2</td>
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<td>-13.0</td>
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</table>

Table 1 Melting layer parameters grouped by day (mean/standard dev.). Z_{Ka} is the maximum observed reflectivities (dBZ), LDR_{Ka} are the maximum observed Linear Depolarization Ratios (dB) and h_{Ka} is the altitude at which the Z_{Ka} at Ku band was observed (km).
Overall, the conclusions drawn in Durden et al. (1997) from TOGA COARE were confirmed by the Wakasa Bay Experiment. In particular it was observed:
- the positive correlation between $Z_{\text{ref}}$ Ku and brightband thickness (~0.7);
- the negative altitude offset of the peak in LDR with respect to the peak in $Z_{\text{ref}}$ Ku (observed mean offset = -70 m, negatively correlated with $Z_{\text{ref}}$ Ku).

Furthermore, it was observed a strong (~0.8) correlation of the altitude and thickness of the LDR signature with the area of vertical acceleration of the particles (see e.g., sample profiles in Figure 2). This seems to confirm the interpretation offered in that study for the possible cause of the negative altitude offset. On the other hand, the altitude of Ka band LDR peak was in general between that of the reflectivity peak and the Ku band LDR peak, and it showed a weaker correlation with $Z_{\text{ref}}$.

This might be due to the fact that smaller particles reach their terminal velocity higher than the larger particles.

- ice reflectivity above the bright band is less than rain reflectivity below; the difference becomes more negative with increasing rain reflectivity.

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REFERENCES
