



**JPL**

**2003 IEEE Aerospace Conference  
Big Sky, Montana**

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**SAFIR**

**The Single Aperture Far Infrared  
Observatory**

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March 13, 2003



# Genesis of SAFIR



**JPL**

pronounced "sapphire"!

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Huge science need and opportunity coupled with feasibility!

- SAFIR was recommended in the Decade Report for technology and concept development that would lead to future infrared missions.
- SAFIR was mentioned prominently in current Structure and Evolution of the Universe and Origins Theme Roadmaps.
- Recognized that large aperture, low temperature far infrared telescope is now achievable, especially with technology advances from JWST, SIRTf, and Herschel.
- Recognized SAFIR as a scientific successor to SIRTf and Herschel, and as a powerful scientific partner to TPF, JWST, and ALMA.

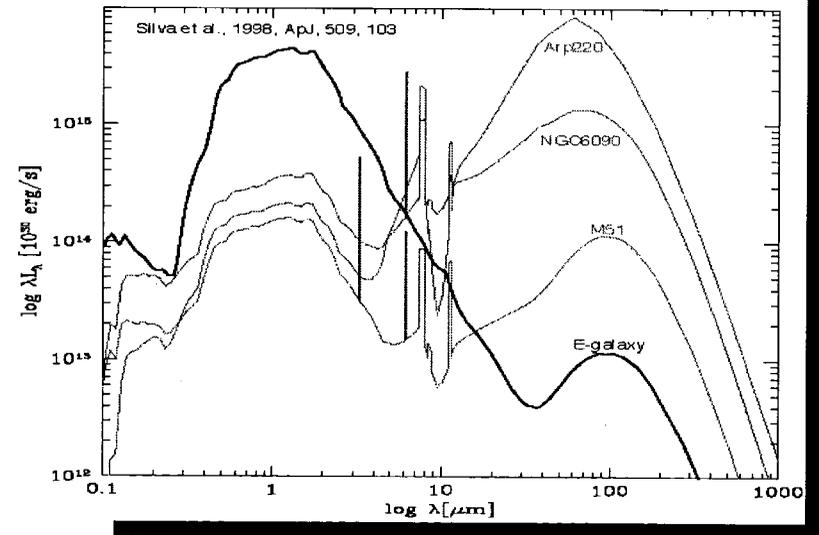
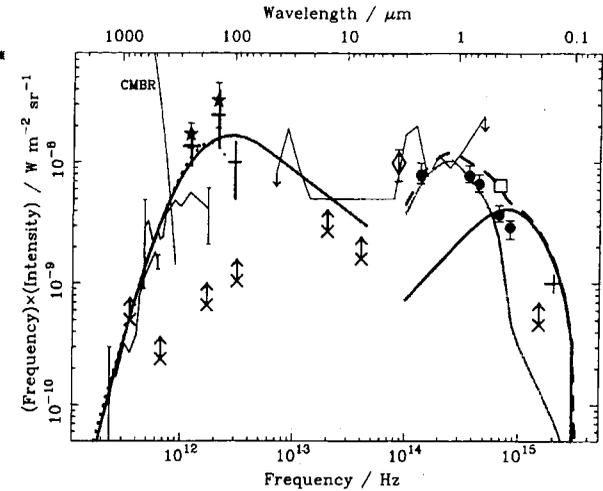


# The stage on which SAFIR plays ...

- Half the luminosity in the Universe is in far-IR!  
The young universe is redshifted there.
- Of the far-IR background, <1/3 is accounted for by discrete galaxies.
- Star formation -- near and far, now and long ago is an IR problem.
- The youngest primordial gas clouds will be visible only in the far-IR.
- Dust is nearly everywhere

*JWST will detect the first galaxies --  
SAFIR will understand why they hide!*

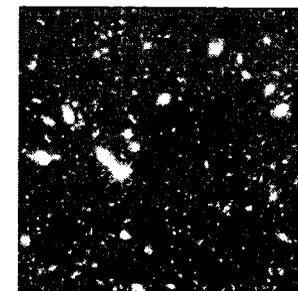
***Era of JWST and ALMA.  
SIRTF, SPICA, Herschel are done.***



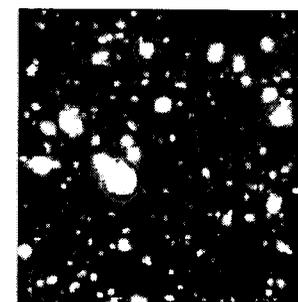


# SAFIR Key Science Drivers (*pre-SIRTF!*)

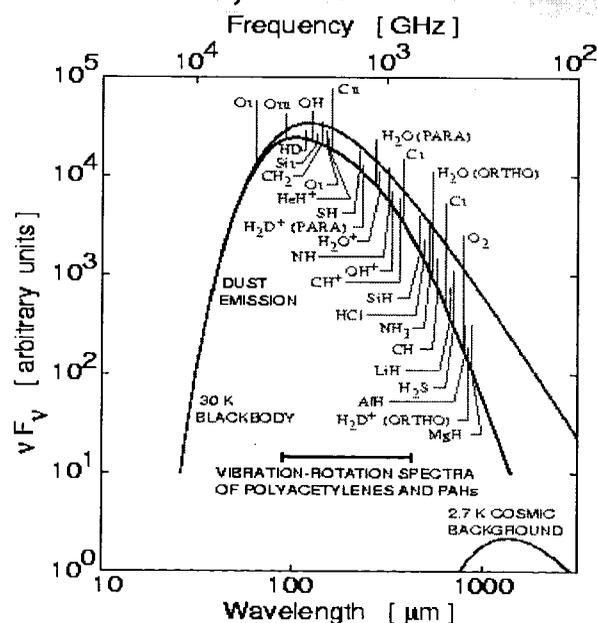
- Resolve the FIR background -- trace star formation to  $z > 5$  in an unbiased way, measuring redshifts directly.
- Understand how primordial material forms stars. Proto-bulges and -disk formation in pristine gas.  $H_2$  @  $z=20$ ?
- Understand role of active galactic nuclei in galaxy formation, and relevance to ULIRGS. Unification?



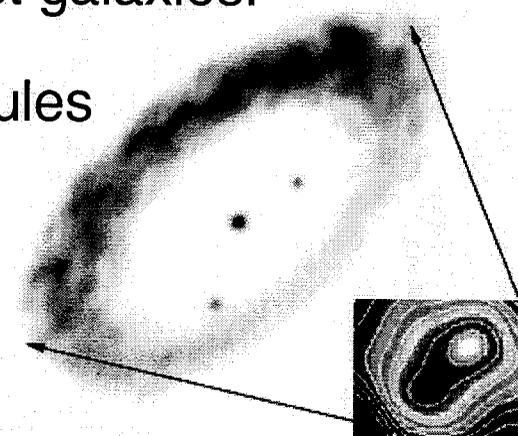
HDF



HDF at 1"



- Bridge gap between local high mass star formation and starburst galaxies.
- Track pre-biotic molecules from cores to planets.
- Identify voids in debris disks around stars.



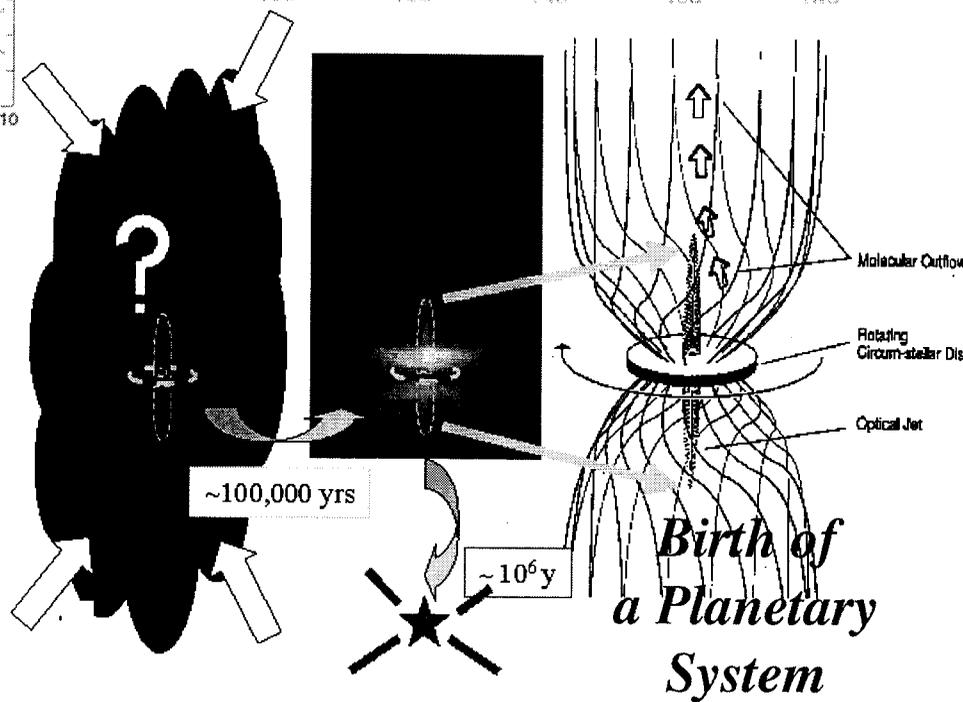
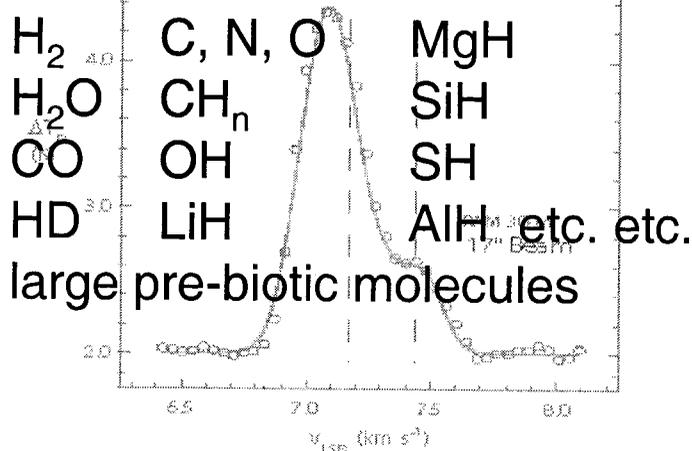
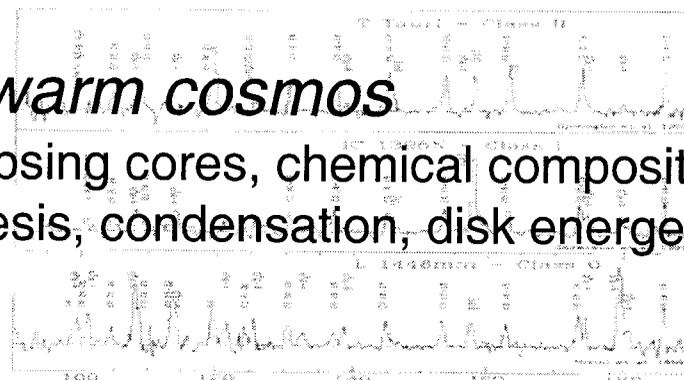
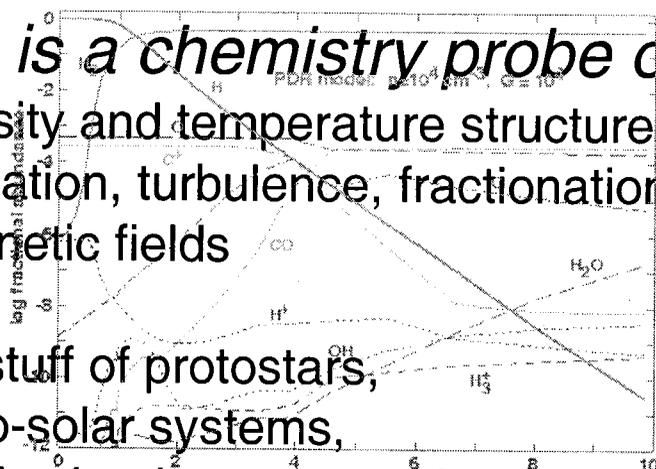


# SAFIR: Molecules to stars to planets to life?

## SAFIR is a chemistry probe of the warm cosmos

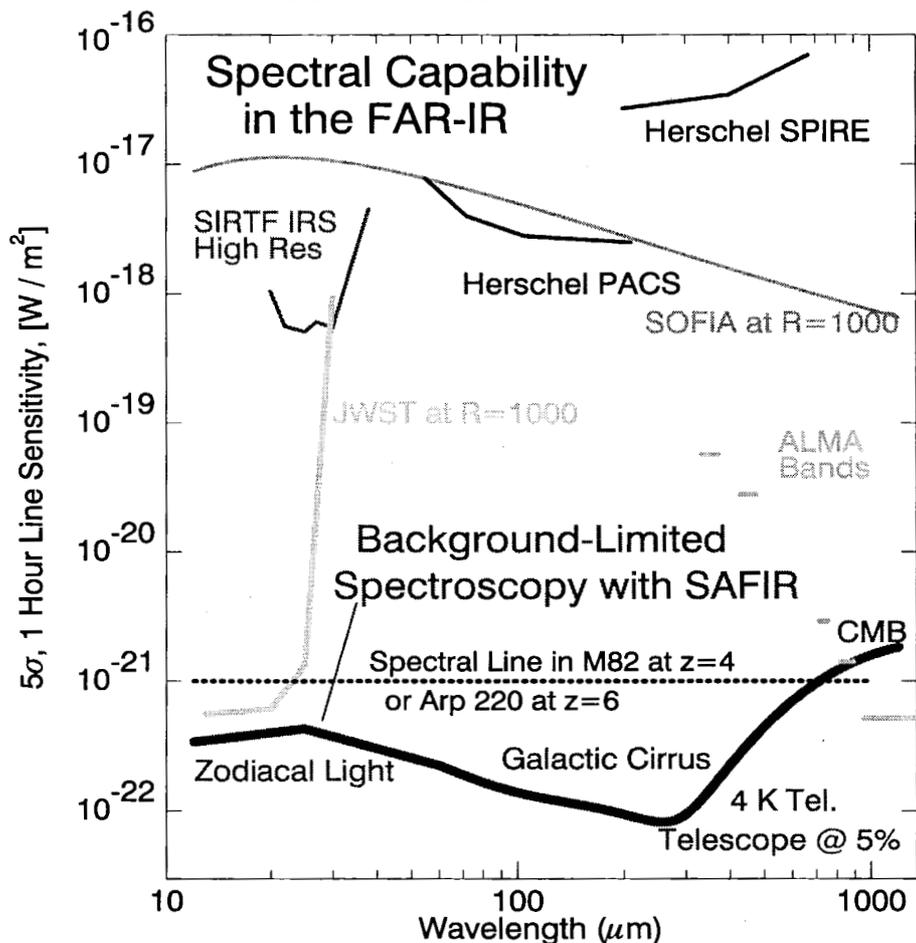
Density and temperature structure of collapsing cores, chemical composition, ionization, turbulence, fractionation, synthesis, condensation, disk energetics, magnetic fields

the stuff of protostars, proto-solar systems, debris clouds, comets, planets and the raw material of life





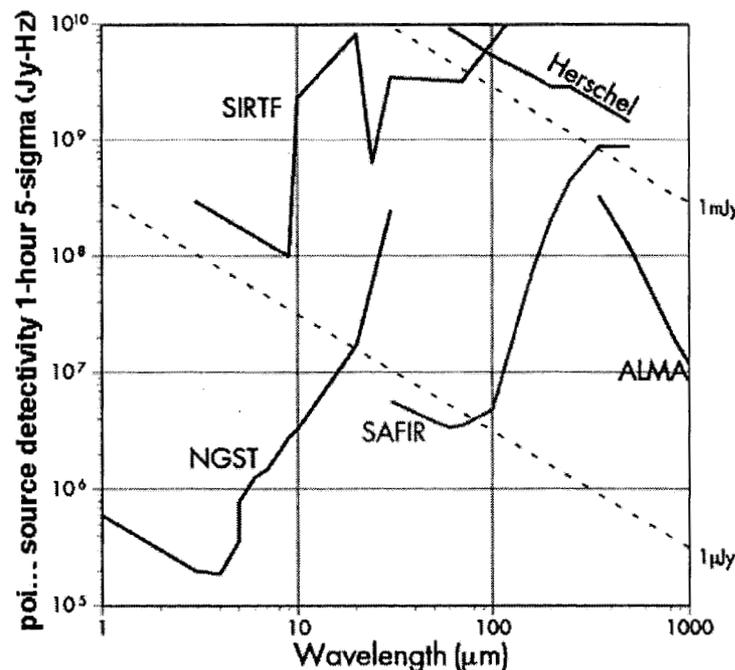
# SAFIR capabilities in comparison



*no confusion limits for spectroscopy!*

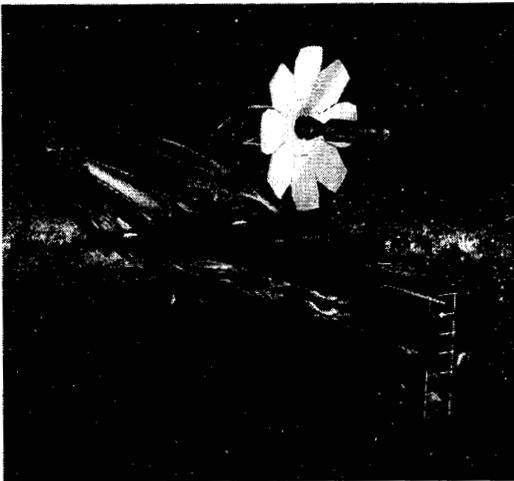
SAFIR will offer orders of magnitude improvement in

- spectroscopic sensitivity
- point source detectivity

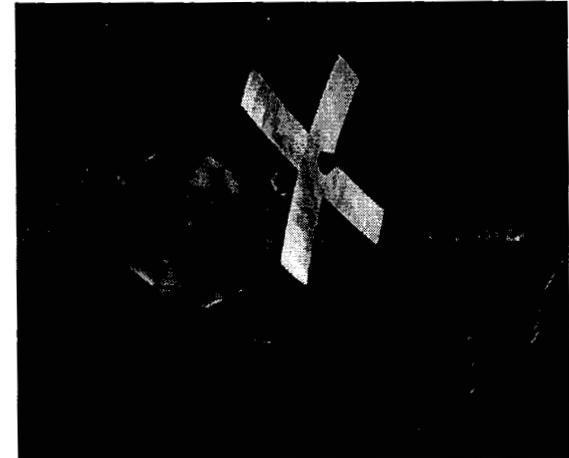




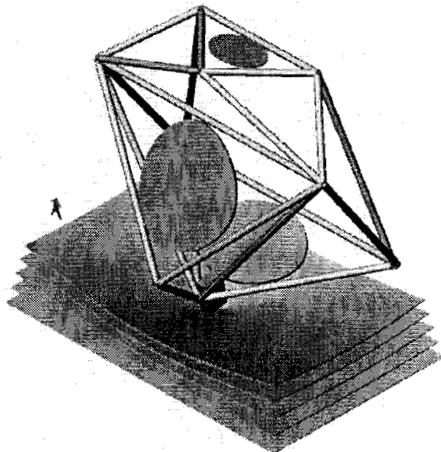
# Flavors of SAFIR



- JWST-like  
*max system validation*



- sparse aperture  
*maximize baselines  
deployment simplicity*



- “DART” w/ membrane mirrors  
*large aperture/weight ratio*

*commonality in technology needs*

→ deployment, active surface control  
→ large format, low noise detectors

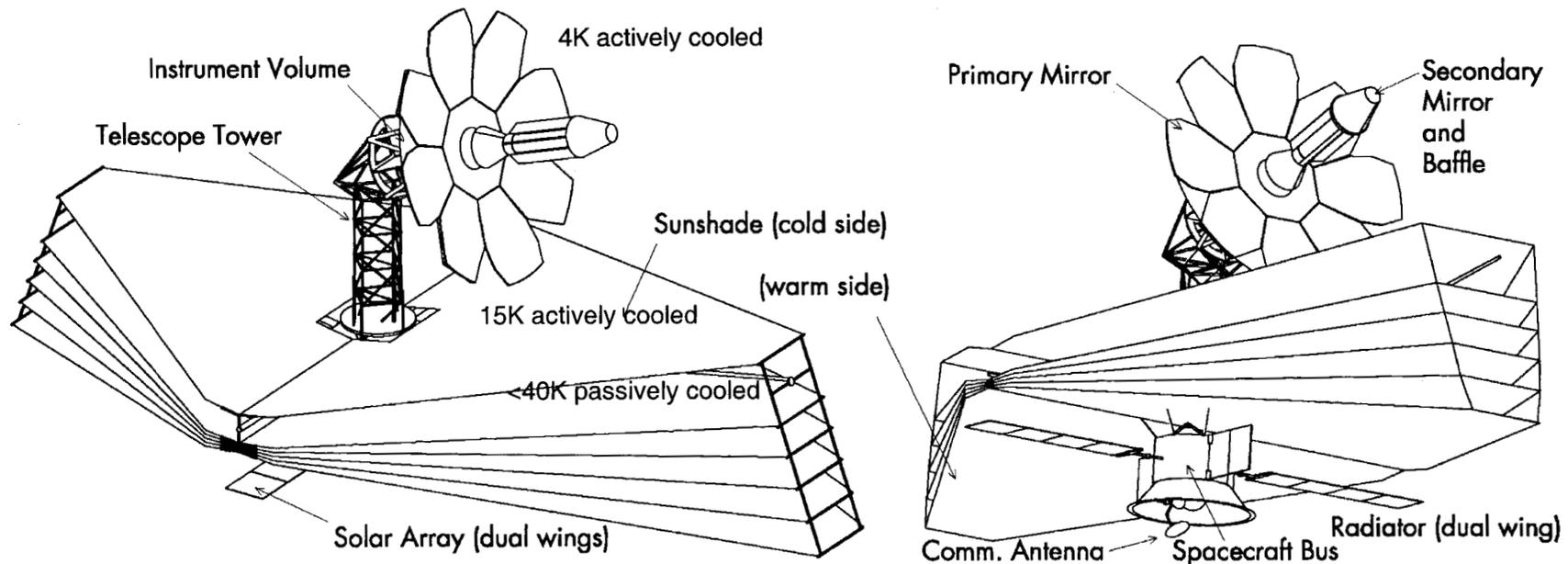
→ cryocoolers, thermal management  
→ large, lightweight optical structures



# A Thermal Strawman Design for SAFIR

(cooling is the biggest challenge... maybe we can do better?)

- <40K “JWST plus” sunshade
  - 15K actively cooled shield blocks sunshade;
  - 4K actively cooled telescope under shield;
  - 50 mK actively cooled focal plane;
- |           |         |
|-----------|---------|
| 1W lift   | } ~200W |
| 85mW lift |         |
| 10μW lift |         |



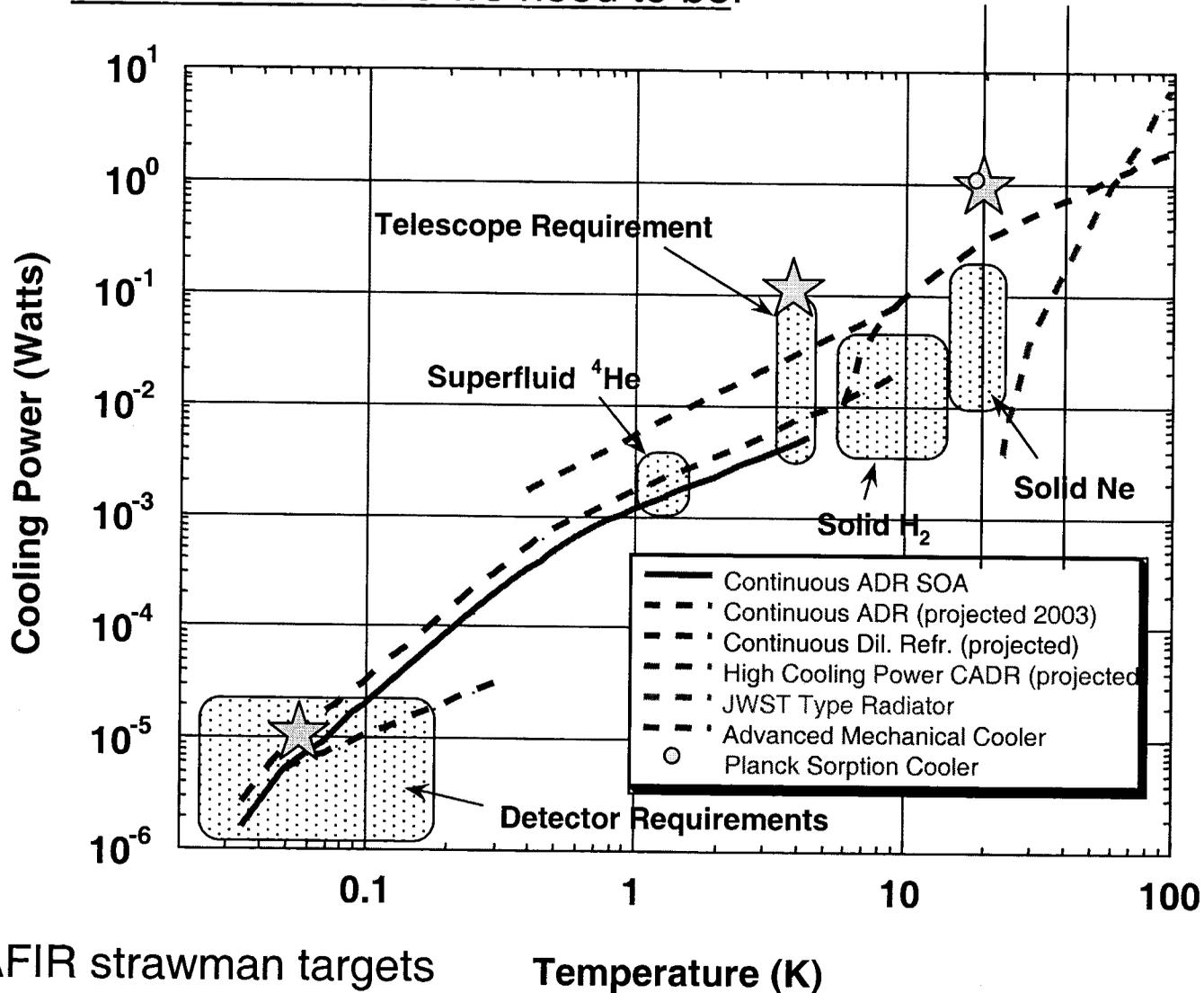
*SOA suggests that thermal requirements are achievable!*



# SAFIR Cryogenic Technology

we're not far from where we need to be!

cooled JWST-like shield sunshade

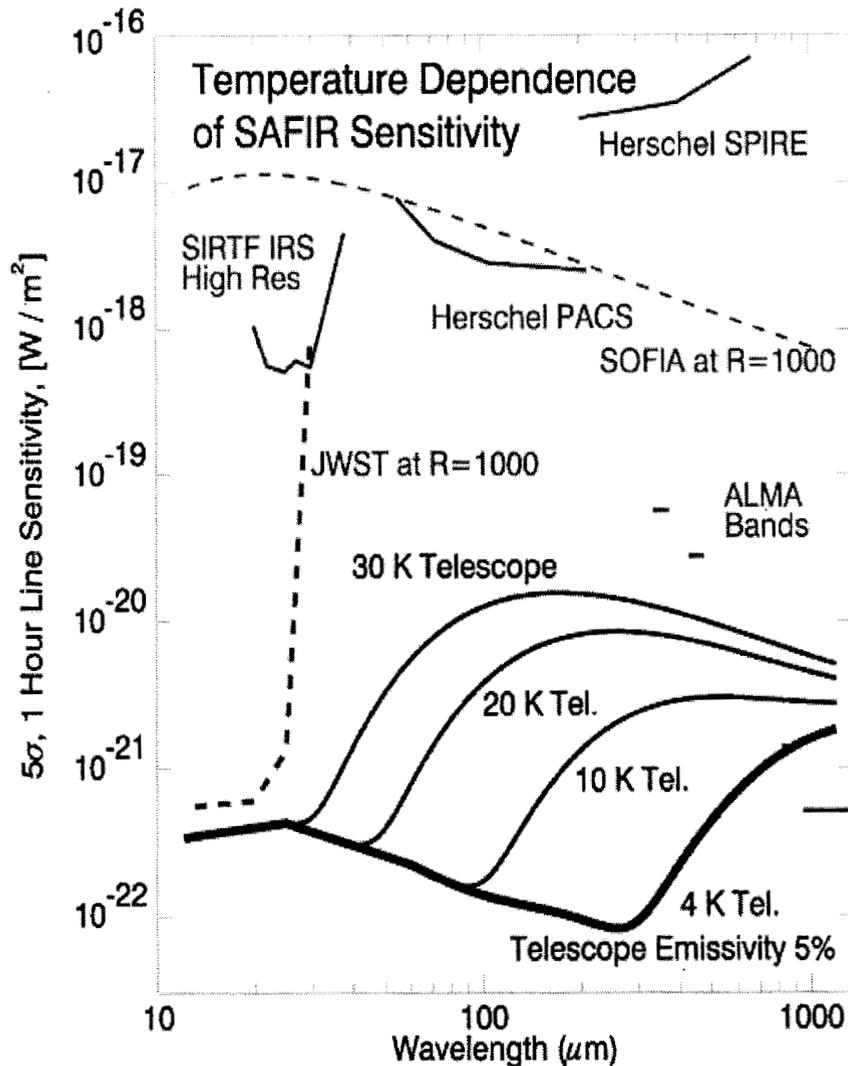


☆ SAFIR strawman targets

Temperature (K)



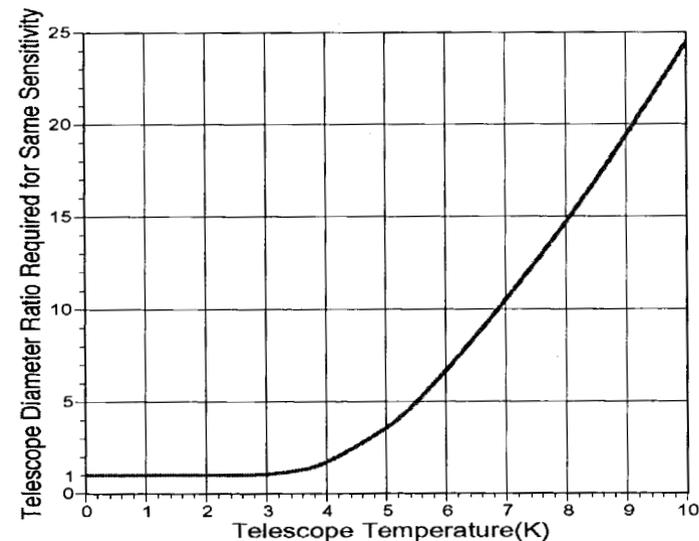
# But why 4K for SAFIR?



*Because it makes a big difference!*

A 4K scope is background-limited (zodi @  $<200\mu m$ , CMB @  $>200\mu m$ )

At these wavelengths, point source sensitivity is more dependent on temperature than on aperture!





# SAFIR *Observatory* Critical Technologies



incremental steps ...

- cryogenic, deployable large apertures
  - actuators, latches, mirror substrates  
(zero-G proof-of-concept highly desirable)
- optimized sun shield technology
  - material properties, refine designs  
(LEO or L2 proof-of-concept highly desirable)
- thermal transport technology
  - gas flow, capillary technology  
(zero-G proof-of-concept highly desirable)
- cryocooler technology
  - extension of ACTDP at 4-20K
  - augment existing ADR capabilities at 50mK-4K

(JWST heritage)

ST9 validation candidates

(JWST ConX heritage)

Code R technology investments



# SAFIR *Focal Plane* Critical Technologies

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- new spectrometer architectures (scaled-up versions of IR spectrometers are huge)
- focal plane cooling technologies for <100mK
- large-format ( $10^3$ - $10^4$  pixel) broadband arrays
  - semiconducting and superconducting (TES) bolometer arrays
  - Ge, Si BiB photoconductor arrays
  - SQPCs
- quantum noise-limited heterodyne spectrometers



# Summary

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- SAFIR will enable very compelling Origins *and* SEU science
- SAFIR is technologically challenging but within our grasp