

ORIGIN OF THE SOLAR WIND

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The origin of the solar wind, coronal magnetic field topology, and turbulence in the solar wind, are subjects of fundamental interest in research on the Sun and solar wind. Major changes in their understanding have taken place as a consequence of advances in radio occultation measurements of the solar corona. These unique measurements have also unified a wide range of diversified remote sensing and *in situ* measurements of the solar wind. The purpose of this talk is to summarize the new views of the Sun's atmosphere, and describe how they emerged solely from observations, i.e., without resorting to theoretical models.