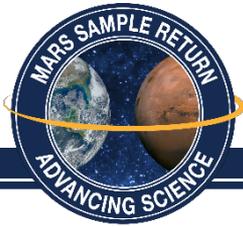


# Why Mars Sample Return is of Increasingly Compelling Interest to the International Mars Exploration Community

**David Beaty**<sup>1</sup>, Elliot Sefton-Nash, Sanjay Vijendran, Michael Meyer, Brandi Carrier<sup>1</sup>, Monica Grady, Harry McSween, Charles Edwards<sup>1</sup>, Thiag Kumaraswamy, and Yael Asher

<sup>1</sup>Jet Propulsion Laboratory, California Institute of Technology



# 1. Martian Meteorites

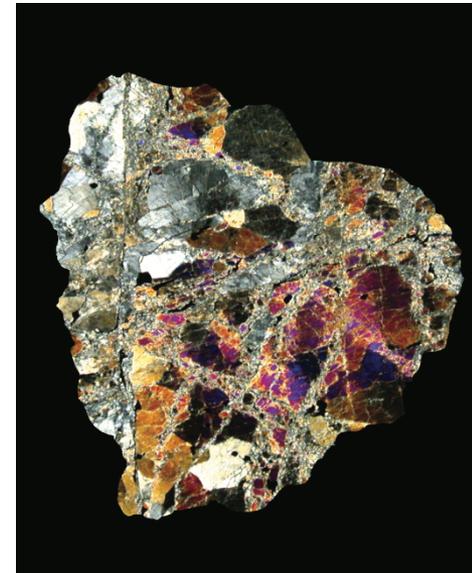
- The number of martian meteorites has now grown to over 100!
- Mars has joined the Moon and Vesta as the best characterized solar system bodies (other than Earth), in part because we have samples of them to study in the laboratory, for example:
  - Mars has a chondritic bulk composition and it is volatile depleted.
  - Mars has a unique oxygen isotopic fingerprint
  - Martian crust consists mostly of basaltic lavas and their fractionation products

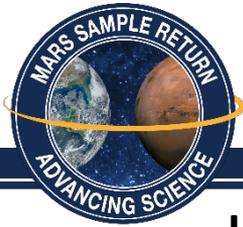
***Valuable, but we have learned their limits.***

*From McSween et al. (2018)*

*Pre-Decisional - For planning and discussion purposes only*

*ALH 84001*





# Meteorite Studies: Limitations

## Limited Sampling of Rock Types

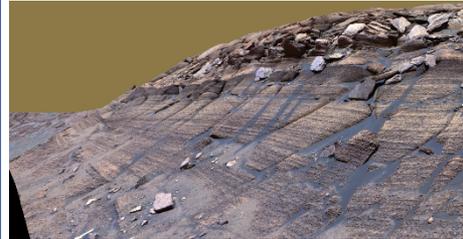
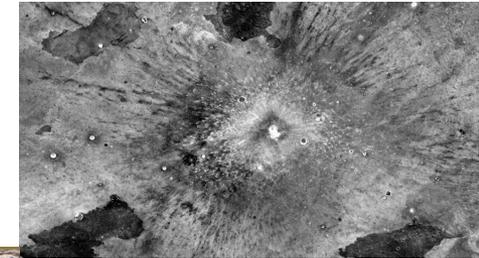
- Fluvial, lacustrine, evaporative, aeolian, granular, hydrothermal, or pyroclastic materials have NOT been sampled as meteorites
- Some of these are the most likely to contain organic matter and/or biomarkers



*NWA 7034  
regolith breccia-the lone  
sedimentary meteorite*

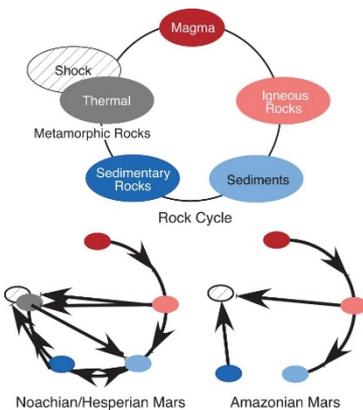
## Lack of Geological Context

- No source craters have been confidently identified



- Cannot be used to underpin planetary chronology

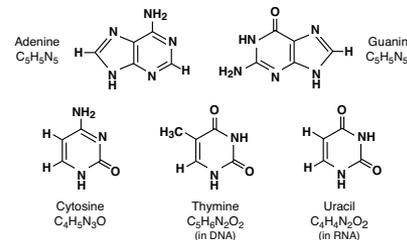
## Biased Sampling of Rock Ages



- Igneous rocks from the first half of martian history are virtually unsampled
- Igneous rocks represent the starting point for all other rock types in the martian rock cycle and are key to differentiation and the original crustal composition.

## Alteration via ejection process

- Some scientific investigations depend on analyzing samples that have not been strongly heated (volatiles, organics, some mineralogy)

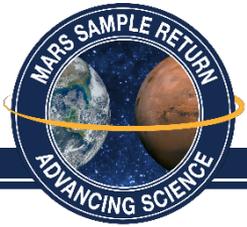


Amino Acids



Sulfates with varying H<sub>2</sub>O

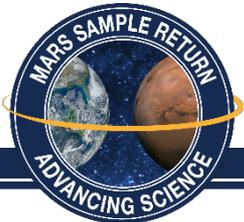
*From McSween et al. (2018)*



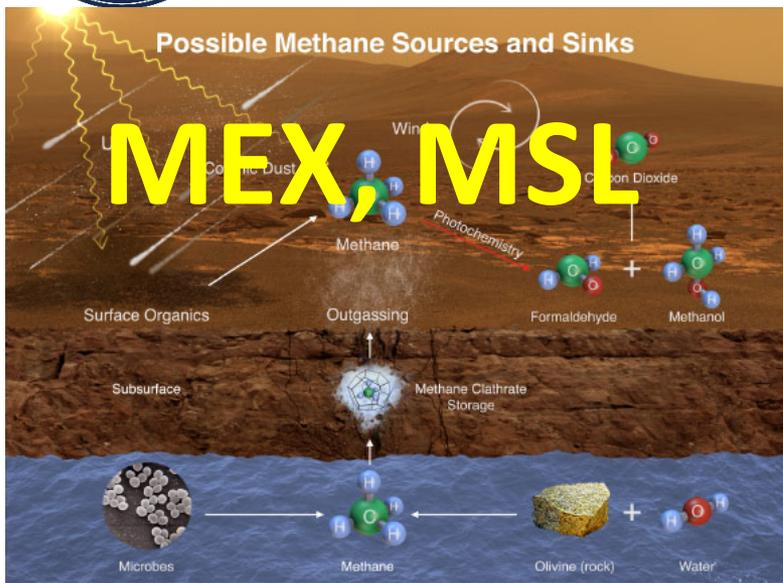
## 2. Mission Results

- Fabulous **new discoveries** have been obtained by the in-situ & orbital) exploration program. These results have emphasized the additional value that would come from analyzing Mars samples using the **greater spatial resolution** and great **diversity of instrumentation** available in terrestrial laboratories.

***Results have raised important **new** hypotheses that cannot be **fully** tested without samples.***



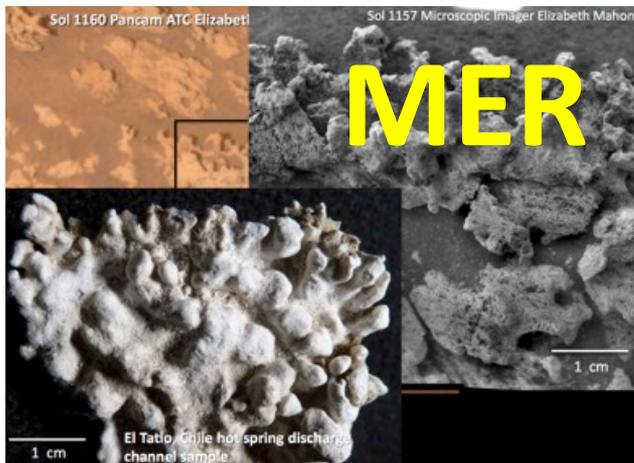
# Mission Results – Habitable Environments



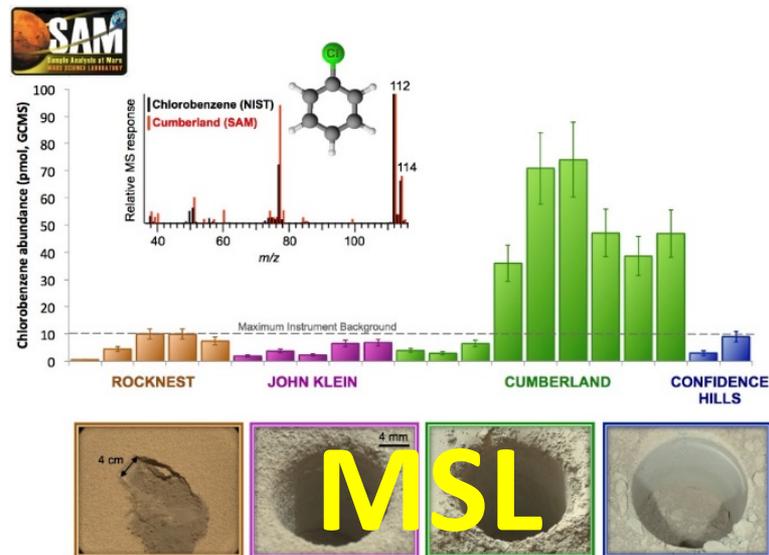
A habitable fluvio-lacustrine environment at Gale Crater

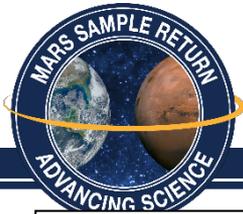
“Mars is episodically producing methane from an additional unknown source”

Organics in the near sub-surface



Hydrothermal systems & silica deposits

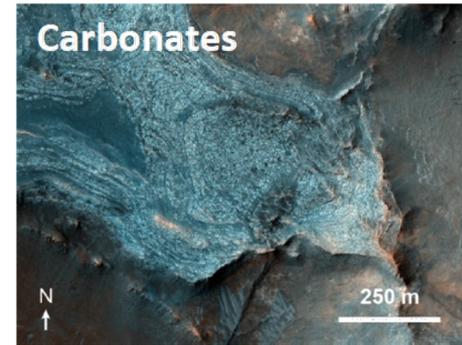




# Mission Results – Minerals Detected

	Class	Group/mineral/phase	Formula	
Primary	Framework silicates	Olivines	$(\text{Mg, Fe})_2\text{SiO}_4$	
		Orthopyroxenes	$((\text{Mg, Fe})_{0.95+x}, \text{Ca}_{0.05-x})\text{Si}_2\text{O}_6$	
		Clinopyroxenes	$(\text{Ca, Mg, Fe})\text{Si}_2\text{O}_6$	
		Plagioclase feldspars	$(\text{Ca, Na})(\text{Al, Si})\text{AlSi}_3\text{O}_8$	
		Alkali feldspars	$(\text{K, Na})\text{AlSi}_3\text{O}_8$	
	Sulfides	Pyrrhotite <sup>b</sup>	$\text{Fe}_{1-x}\text{S}$	
		Pyrite/marcasite <sup>c</sup>	$\text{FeS}_2$	
Oxides	Magnetite <sup>d</sup>	$\text{Fe}_{3-x}\text{Ti}_x\text{O}_4$		
	Ilmenite <sup>d</sup>	$\text{FeTiO}_3$		
Secondary	Oxides	Hematite	$\text{Fe}_2\text{O}_3$	
		Goethite <sup>d</sup>	$\text{FeO}(\text{OH})$	
		Akaganeite <sup>b</sup>	$\text{Fe}(\text{O, OH, Cl})$	
	Phyllosilicates (clay minerals)	Fe/Mg smectites (e.g., nontronite, saponite)	$(\text{Ca, Na})_{0.3-0.5}(\text{Fe, Mg, Al})_{2-3}(\text{Al, Si})_4\text{O}_{10}(\text{OH})_2 \cdot n\text{H}_2\text{O}$	
		Al smectites (e.g., montmorillonite, beidellite)	$(\text{Na, Ca})_{0.3-0.5}(\text{Al, Mg})_2(\text{Al, Si})_4\text{O}_{10}(\text{OH})_2 \cdot n\text{H}_2\text{O}$	
		Kaolin group minerals (e.g., kaolinite, halloysite)	$\text{Al}_2\text{Si}_2\text{O}_5(\text{OH})_4$	
		Chlorite	$(\text{Mg, Fe}^{2+})_5\text{Al}(\text{Si}_3\text{Al})\text{O}_{10}(\text{OH})_8$	
		Serpentine <sup>e</sup>	$(\text{Mg, Fe})_3\text{Si}_2\text{O}_5(\text{OH})_4$	
		High-charge Al/K phyllosilicates (e.g., muscovite, illite)	$(\text{K, H}_3\text{O})(\text{Al, Mg, Fe})_2\text{Al}_x\text{Si}_{4-x}\text{O}_{10}(\text{OH})_2$	
	Other hydrated silicates	Prehnite	$\text{Ca}_2\text{Al}(\text{AlSi}_3\text{O}_{10})(\text{OH})_2$	
		Analcime	$\text{NaAlSi}_2\text{O}_6 \cdot \text{H}_2\text{O}$	
		Opaline silica ( $n > 0$ ), quartz ( $n = 0$ )	$\text{SiO}_2 \cdot n\text{H}_2\text{O}$	
	Carbonates	Mg/Ca/Fe carbonates	$(\text{Mg, Fe, Ca})\text{CO}_3$	
	Sulfates	Kieserite ( $\text{MgSO}_4 \cdot \text{H}_2\text{O}$ ); szomolnokite ( $\text{FeSO}_4 \cdot \text{H}_2\text{O}$ ); Fe(II)-, Fe(III)-, and Mg-polyhydrated sulfates	$(\text{Fe, Mg})\text{SO}_4 \cdot n\text{H}_2\text{O}$	
		Gypsum ( $n = 2$ ), bassanite ( $n = 0.5$ ), anhydrite <sup>b</sup> ( $n = 0$ )	$\text{CaSO}_4 \cdot n\text{H}_2\text{O}$	
		Alunite	$\text{KAl}_3(\text{SO}_4)_2(\text{OH})_6$	
		Jarosite	$\text{KFe}_3(\text{OH})_6(\text{SO}_4)_2$	
		Not a named mineral	$\text{Fe}^{3+}\text{SO}_4(\text{OH})$	
		Chlorides	Chlorides	e.g., $\text{NaCl}$ , $\text{MgCl}_2$
		Perchlorates	Perchlorates <sup>f</sup>	e.g., $(\text{Mg, Ca})(\text{ClO}_4)_2$

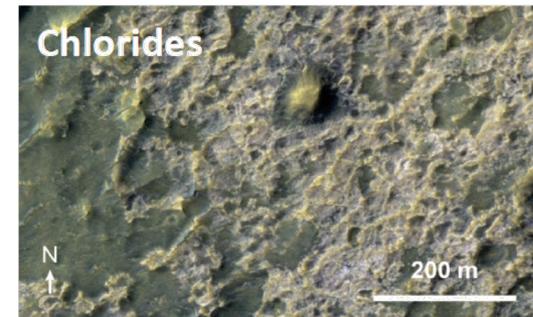
[Ehlmann et al., Annu. Rev. earth Planet. Sci. 2014]



MRO/HIRISE



MRO/HIRISE

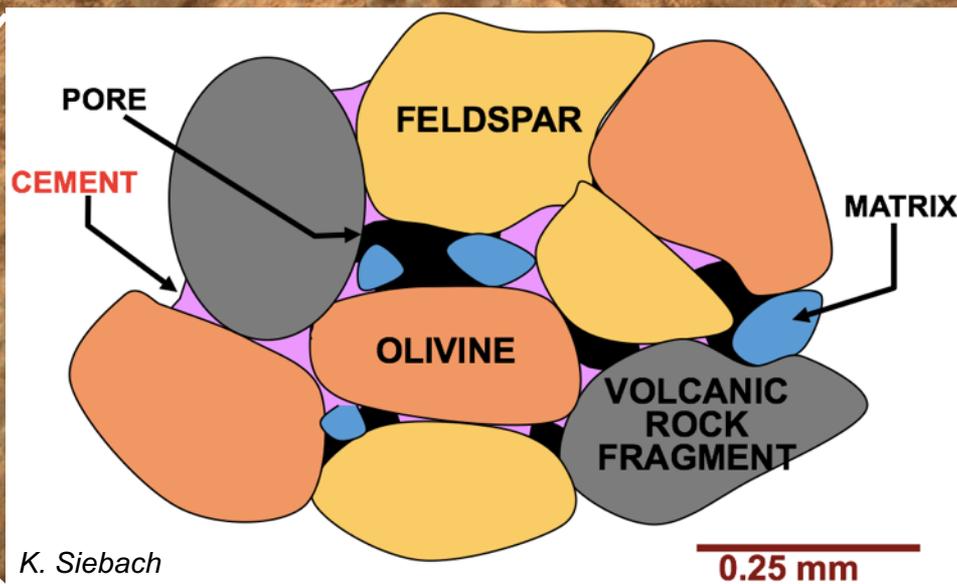


MRO/HIRISE

**How do the minerals relate to each other?**

# We need to reduce our spatial focus

## Modern “grain-by-grain” sediment analysis



### Volcanic Frags (Provenance)

- source lithologies
- nature / history of mantle sources
- source ages

### Feldspar (Provenance)

- source lithologies
- age / age history of source

### Olivine/Pyrox. (Provenance)

- source lithologies
- T/P history of sources

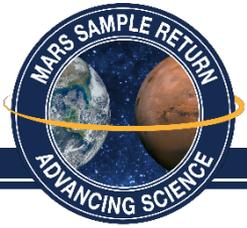
### Matrix (Provenance & Diagenesis)

- source lithologies
- fluid chemistry / fluid interactions
- clay mineral origins and diagenesis

### Cements (Diagenesis)

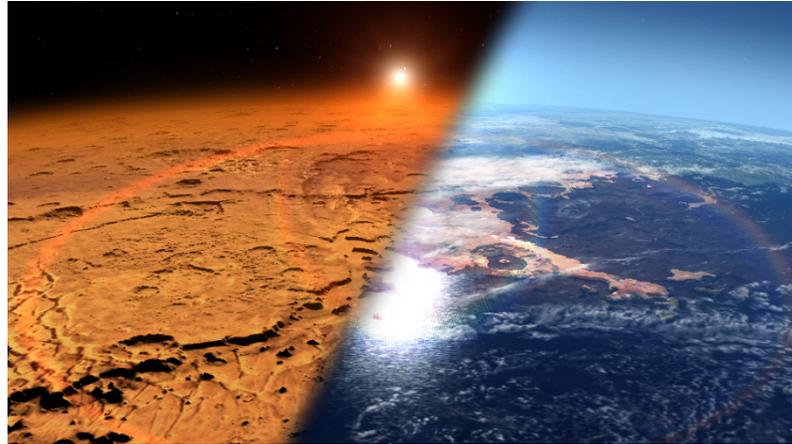
- cement “stratigraphy”
- fluid chemistry (e.g., pH, Eh)
- fluid sources / fluid histories

*From McLennan et al. (2018)*



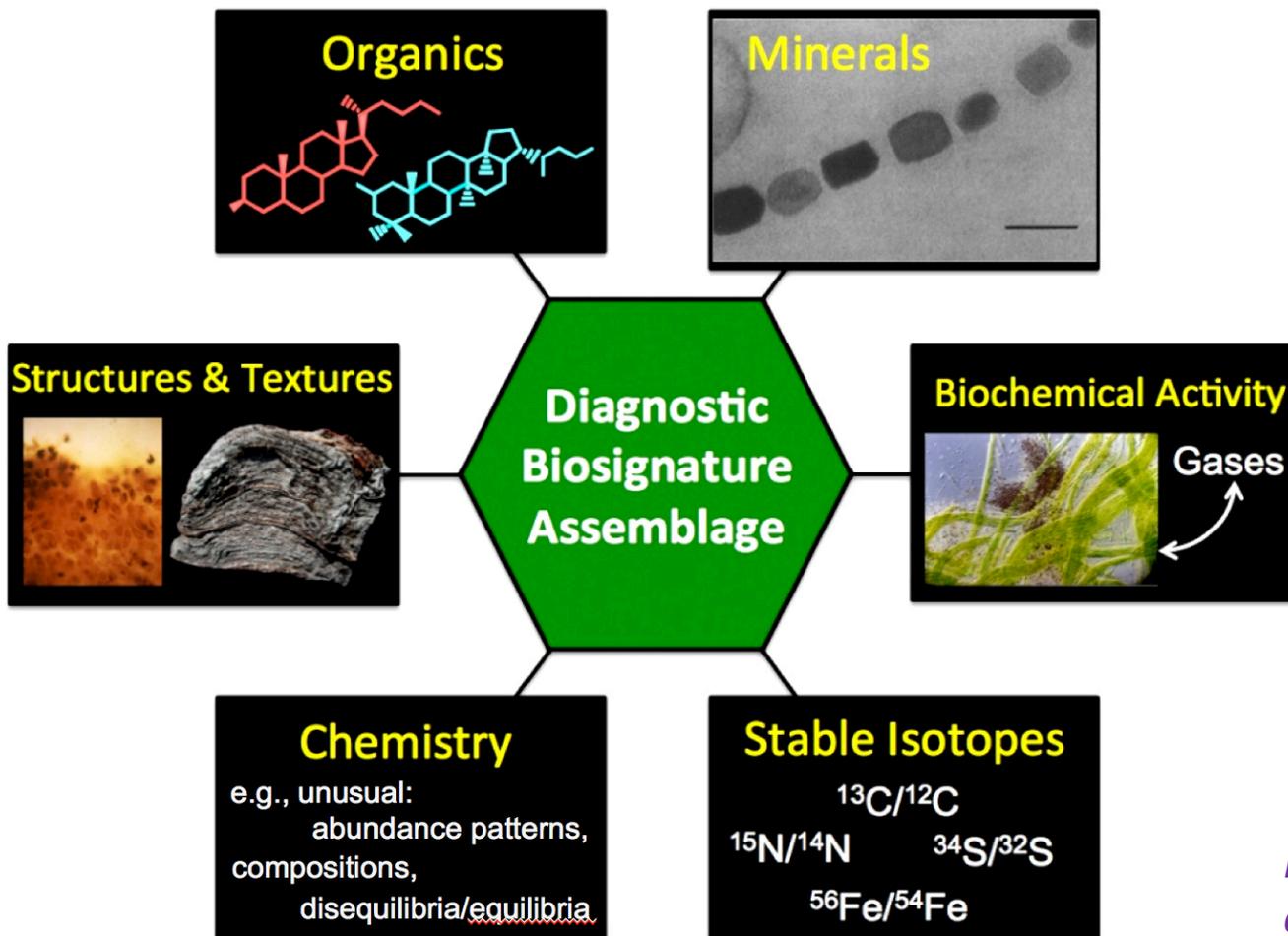
## 3. Astrobiology

- Although the **search for extra-terrestrial life** has long been the key driver of Mars exploration, we now have a much better understanding of the **potential for preservation** in the geologic record, the **evolution of Mars as a habitable planet**, and (by means of the study of **biochemistry on Earth**) the details of biological processes.



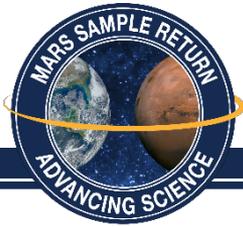
***Samples allow for far more effective search strategies.***

# Biosignature Categories



*From Des Marais et al. (2018)*

**Coordinated analyses of several categories of potential biosignatures can greatly strengthen our interpretations of their origins and significance**



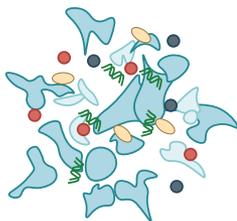
# Why RSS is Important to Astrobiology

## Samples require special handling

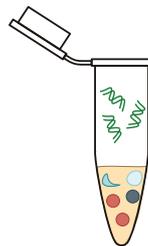
- Soil and mineral samples are difficult, requiring optimization
- Processing is a multi-step process requiring human attention
- Low biomass samples require extra precision



Prepare samples



Mechanical, chemical and enzymatic cell lysis



Purify DNA



Prepare for Sequencing

**Example: DNA extraction and sequencing**

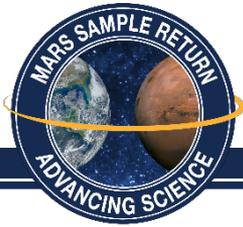
## Investigation pathway is discovery-dependent

### **We don't know:**

- What form life might take
- Whether it would be alive upon return to Earth
- Its ability to grow and metabolize on Earth
- Whether it is related to life on Earth

**The answers to these questions would dictate downstream experiments**

*From Mackelprang et al. (2018)*



## 4. Human Exploration

We have a greatly improved understanding of the ways in which the **risks, performance** and **cost** of putting humans on Mars can be improved by acquiring advance information—especially sample-related information.

### *Dust on Mars*



*20-km high dust devil, viewed from orbit (MRO)*

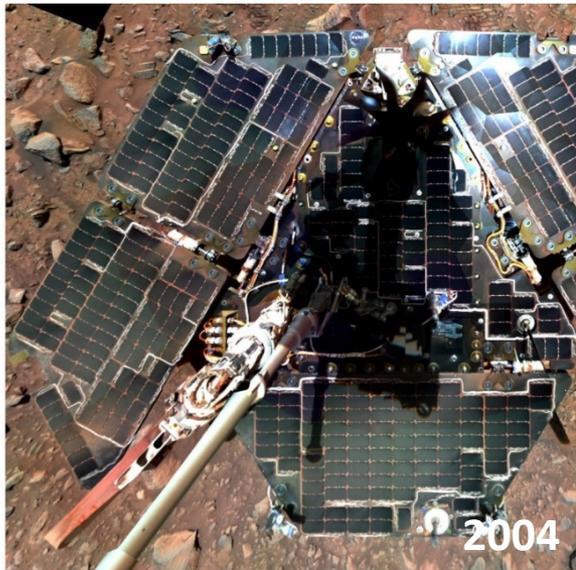
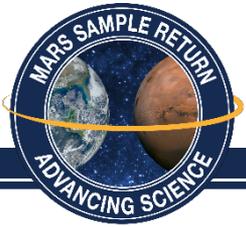


Photo Credit: NASA/JPL



*From Harrington et al. (2018)*

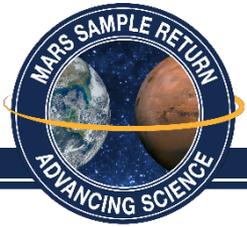


# Biological Hazards – Crew and PP

- Apollo demonstrated pervasiveness of surface dust
  - Longer duration martian missions will exacerbate
- What is the threat that a martian organism (or community of organisms) poses to mission (e.g. pathogenicity to the crew, ability to compete for or spoil critical resources, AND threat of degrading hardware/systems performance)?
- **Evaluate martian biohazards related to the Planetary Protection of Earth**



*From Harrington et al. (2018)*

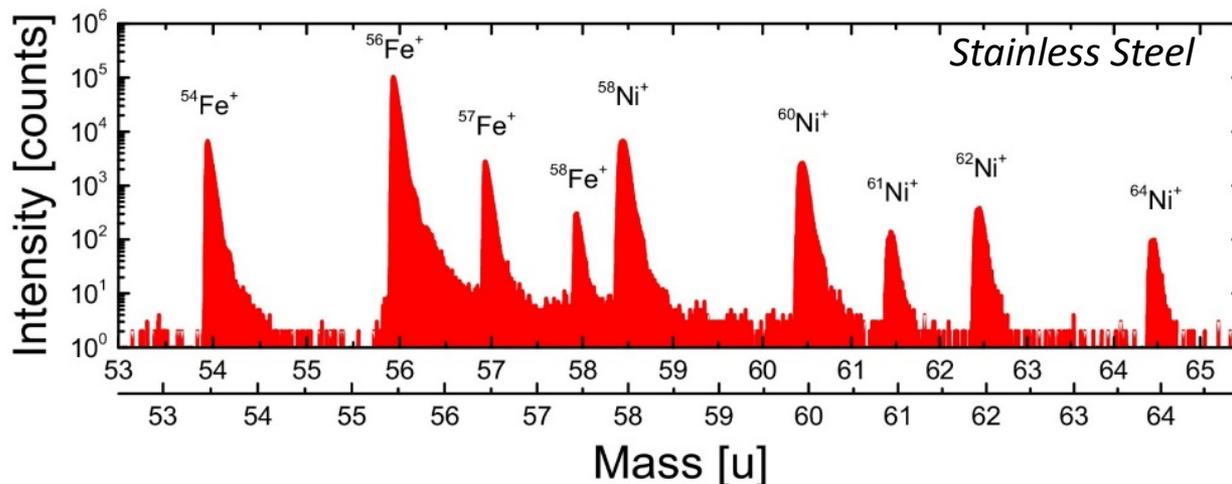


## 5. Instruments/Sample Prep

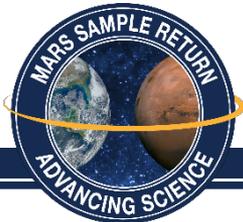
- We are seeing unprecedented improvements in our ability to prepare and analyze very small samples. Highly visible recent examples are the work on Hayabusa asteroid samples (JAXA) and the Stardust comet samples (NASA).

### EXAMPLE

CHILI, the Chicago Instrument for Laser Ionization, a new resonance ionization mass spectrometer developed for isotopic analysis at **high spatial resolution** and **high sensitivity** of **small samples**

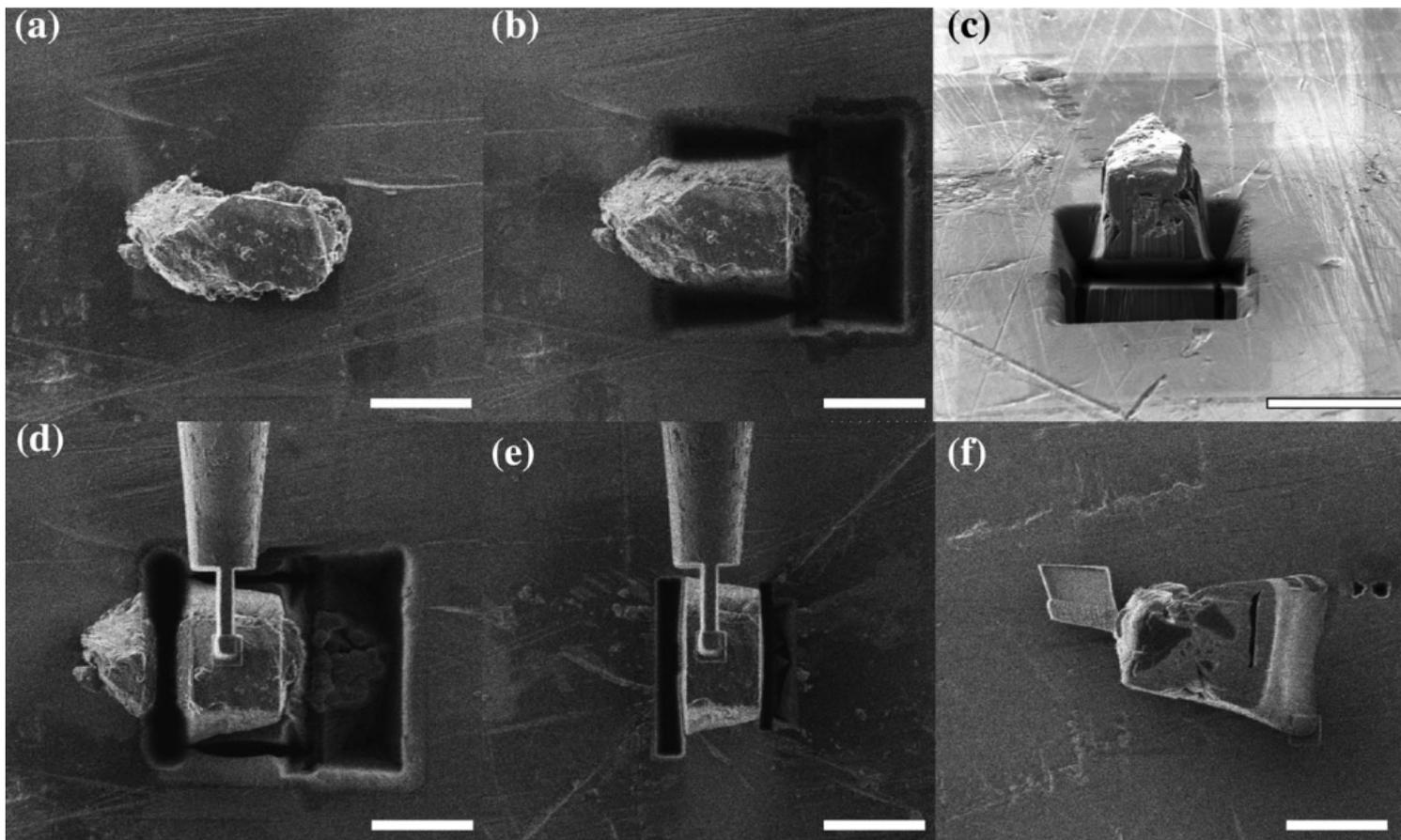


From Stephan et al. (2016) *Int. Jour. Mass Spec.*, v.407, p.1-15.

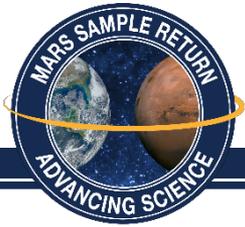


## 5. Instruments/Sample Prep

Investigation of cutting methods for small samples of Hayabusa and future sample return missions. Hayabusa grain RA-QD02-0265. All scale bars are 10 microns.

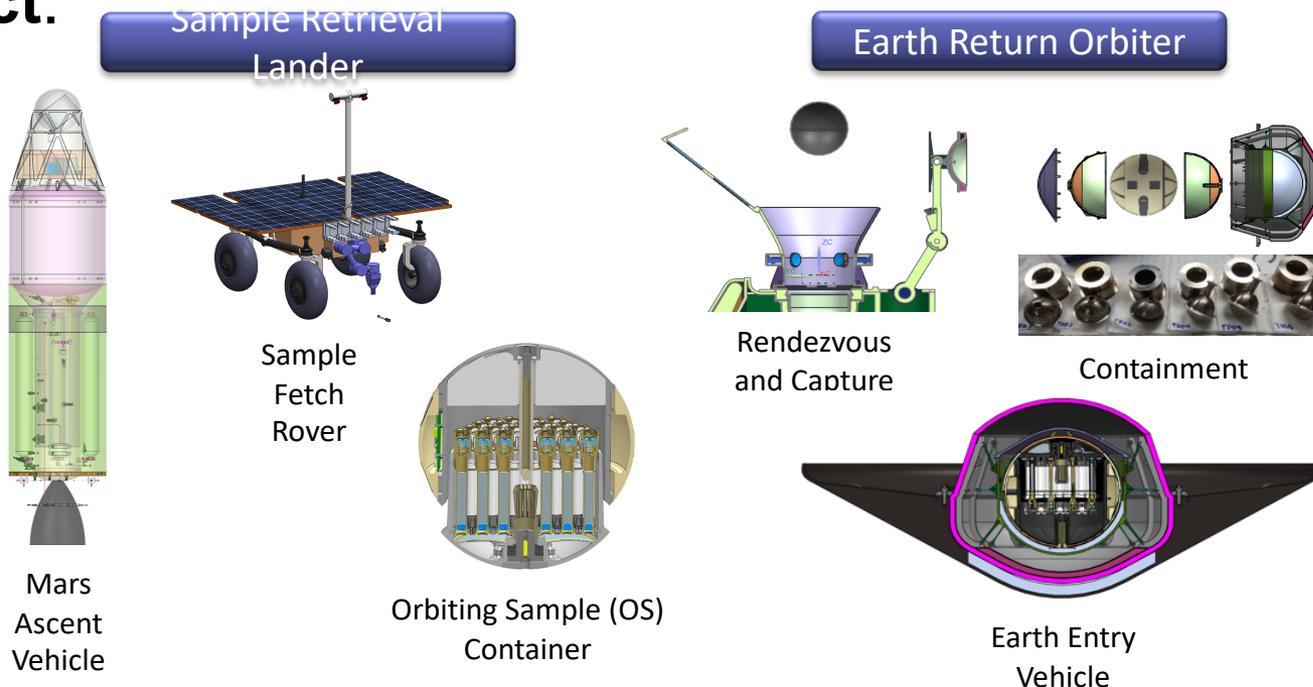


*From Uesugi et al. (2014) MAPS 49, p. 1186-1201*



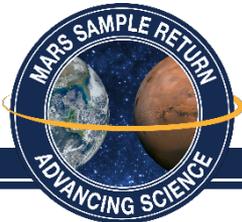
# 6. Engineering

Over the past decade there have been substantial **improvements in the capability** of the world's space agencies to acquire and preserve samples (most notably, the M-2020 sampling system), the development of the **Mars Ascent Vehicle**, and critical progress in **breaking the chain of contact**.



*Future mission concepts-From Edwards et al. (2018)*

*Pre-Decisional - For planning and discussion purposes only*



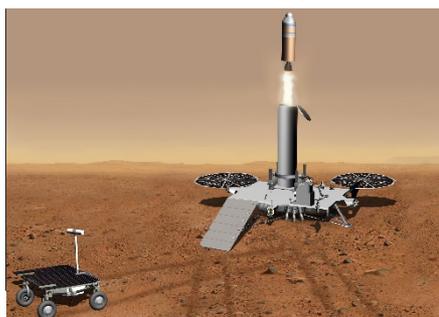
# Mars Sample Return

## SCIENCE



- Civilization-scale science
- Samples: the gift that keeps on giving
- Definitive scientific results
- Only way to advance critical sectors of planetary science & astrobiology

## ENGINEERING



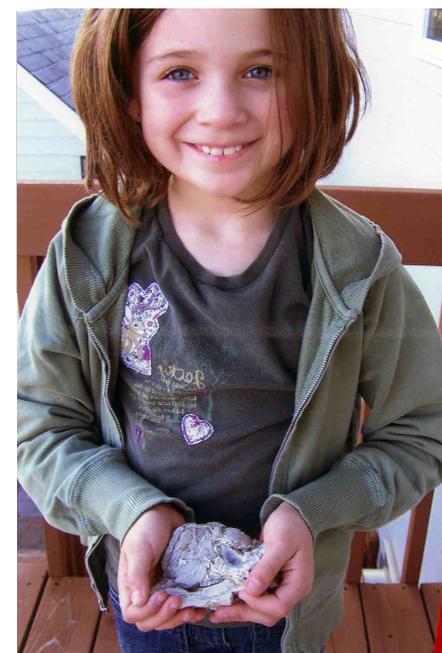
- Unique technical challenges drive unprecedented innovation
- Advances will benefit future robotic and human missions.
- Crucible for engineering as a discipline.

## PREPARATION



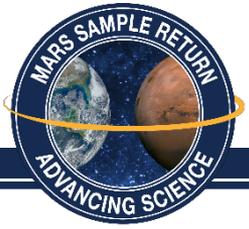
- Prepare for humans to Mars
- Inform planetary protection policy evolution to enable future missions

## INSPIRATION



- Inspire and train the next generation
- Magnet for international cooperation

***We have the opportunity and motivation to carry out MSR on an international basis***



# Back-Ups