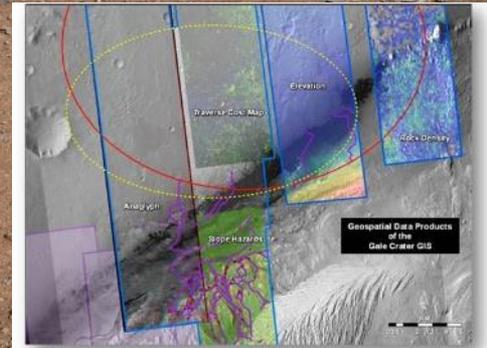
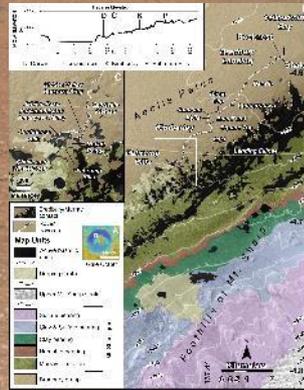
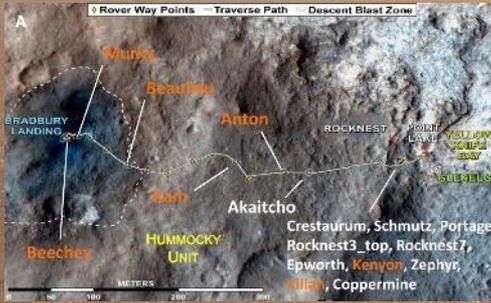


Mapping Curiosity's Traverse from Bradbury Landing to Mt. Sharp

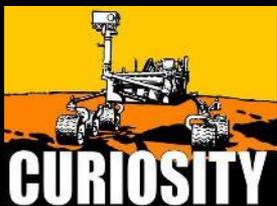


 Jet Propulsion Laboratory
California Institute of Technology

NASA/JPL-Caltech/MSSS

F. Calef III¹, L. Edgar², H. Gengl¹, T. Parker¹, S. Rowland³, J. Schroeder¹, and K. Stack¹

¹Jet Propulsion Laboratory – California Institute of Technology, ²USGS, ³Univ of Hawaii

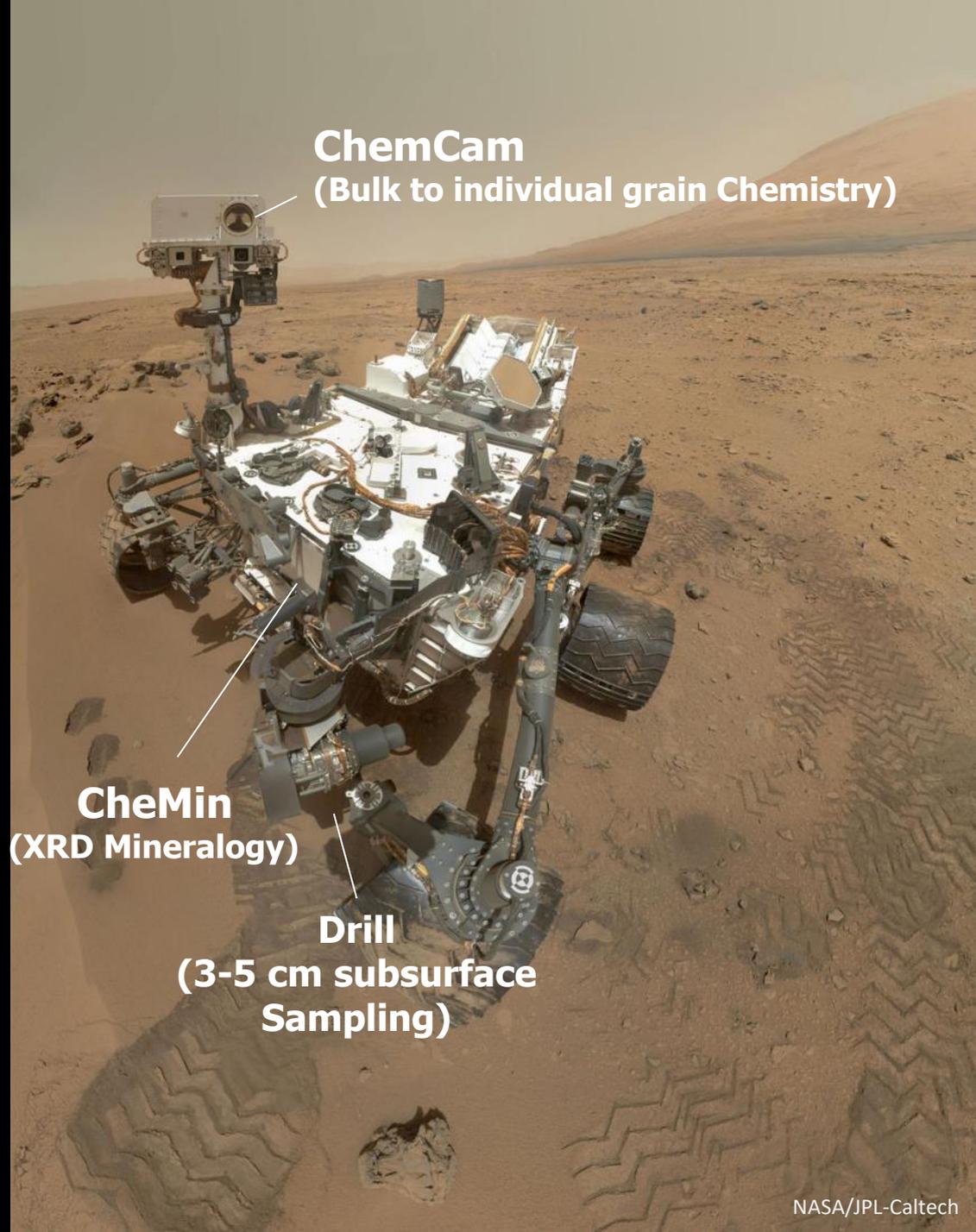


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Curiosity's primary scientific goal is to explore and quantitatively assess a local region on Mars' surface as a potential habitat for life, past or present

- Biological potential
- Geology and geochemistry
- Water, weather, and climate
- Radiation levels and hazards

**Curiosity self-portrait at Rocknest
Assembled from 55 MAHLI images
Shows four scoop trenches
and wheel scuff**



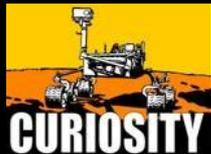
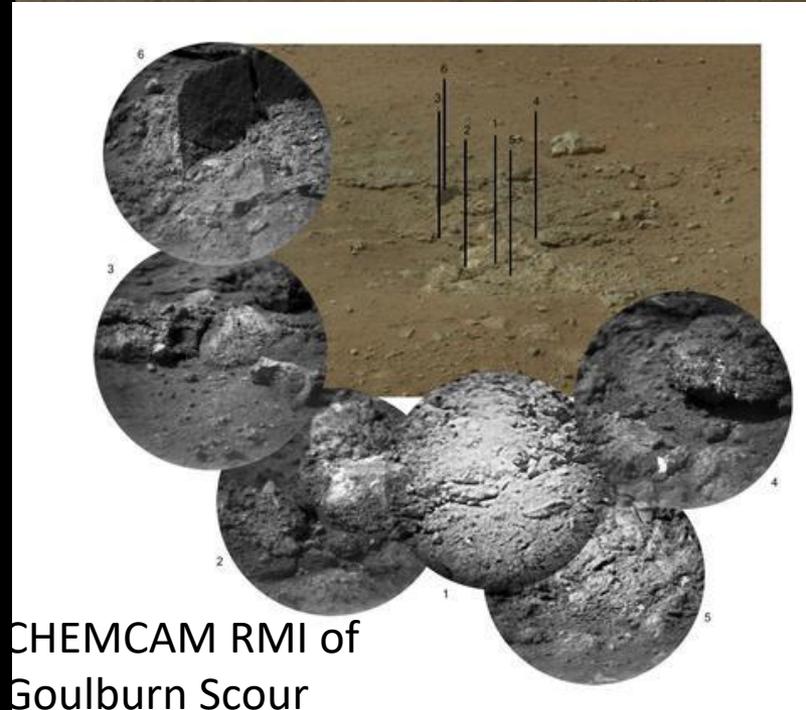
ChemCam
(Bulk to individual grain Chemistry)

CheMin
(XRD Mineralogy)

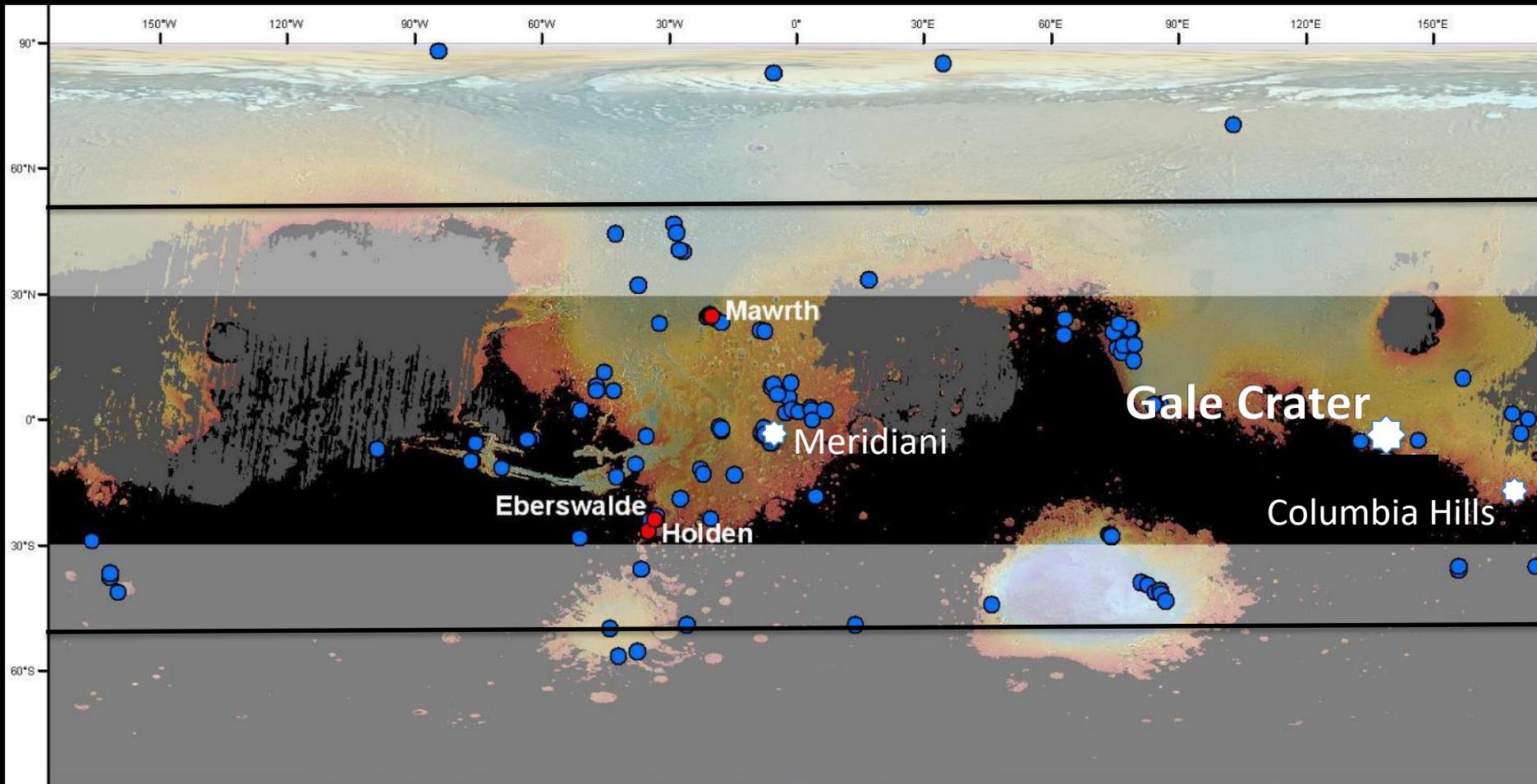
Drill
(3-5 cm subsurface
Sampling)

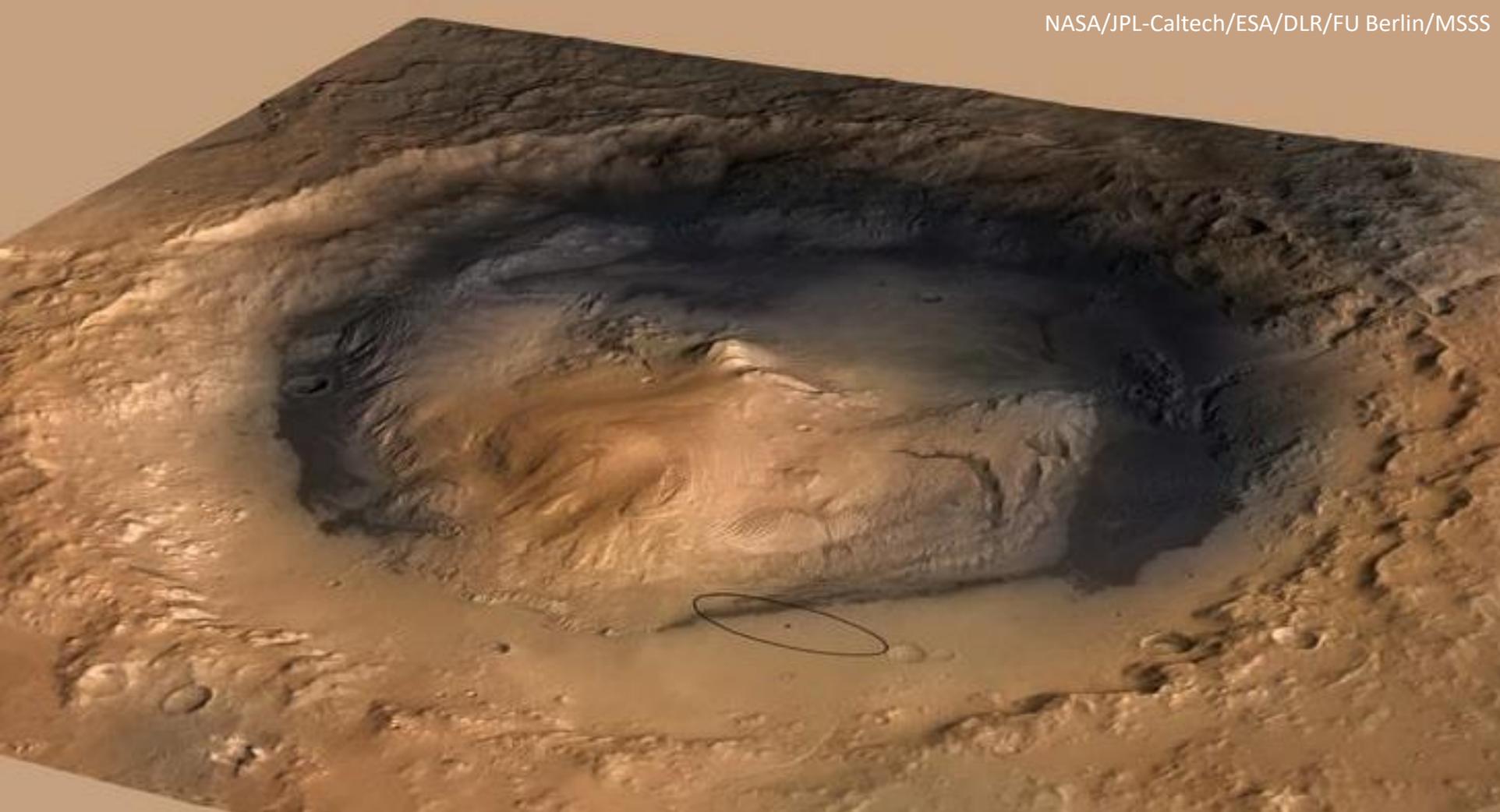
Mapping for MSL

1. Datasets are becoming increasingly large
 - I. >100 GBs of mapping data.
 - II. >10000s science instrument readings
2. Missions last “forever”
 1. Mars Pathfinder 30 days -> 85 days
 2. Mars Exploration Rover (Opportunity) 90 days -> 10 years and STILL GOING.
 1. MSL: 2 Earth years -> WHO KNOWS!
3. Rover Mission Needs
 - Where are we now?
 - Why are we here?
 - Where are we going?
 - Why are we going there?

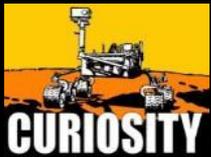


Where we've gone on Mars

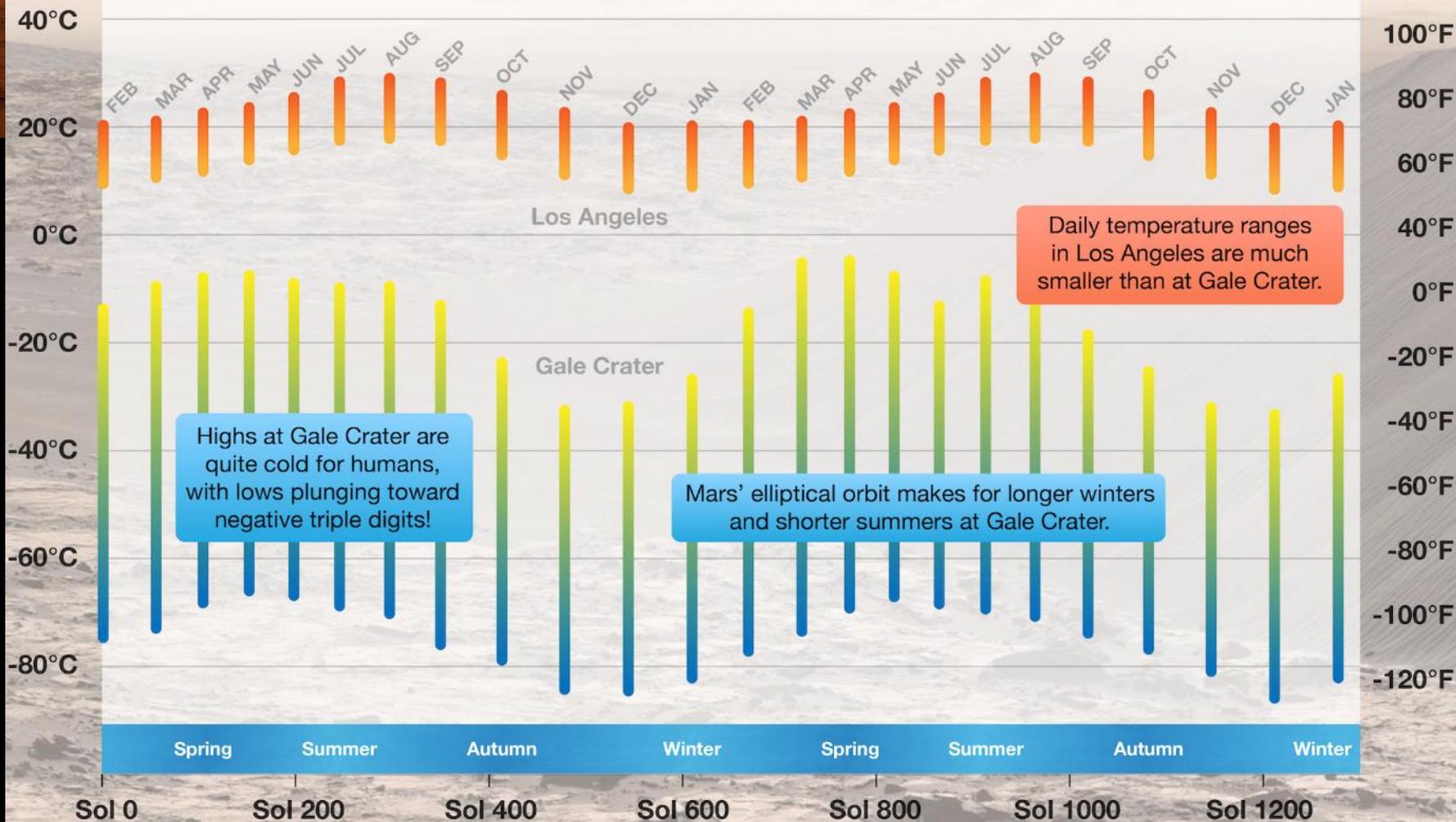




155-km Gale Crater contains a 5-km high mound of stratified rock. Strata in the lower section of the mound are composed of clays and sulfates, while the upper mound is dry, suggesting transition from 'wet' Mars to 'dry' Mars (Late Noachian to Early Hesperian?).



Seasonal Temperature Ranges at Gale Crater (with temperatures in Los Angeles at equivalent seasonal points)



NASA/JPL-Caltech/CAB(CSIC-INTA)

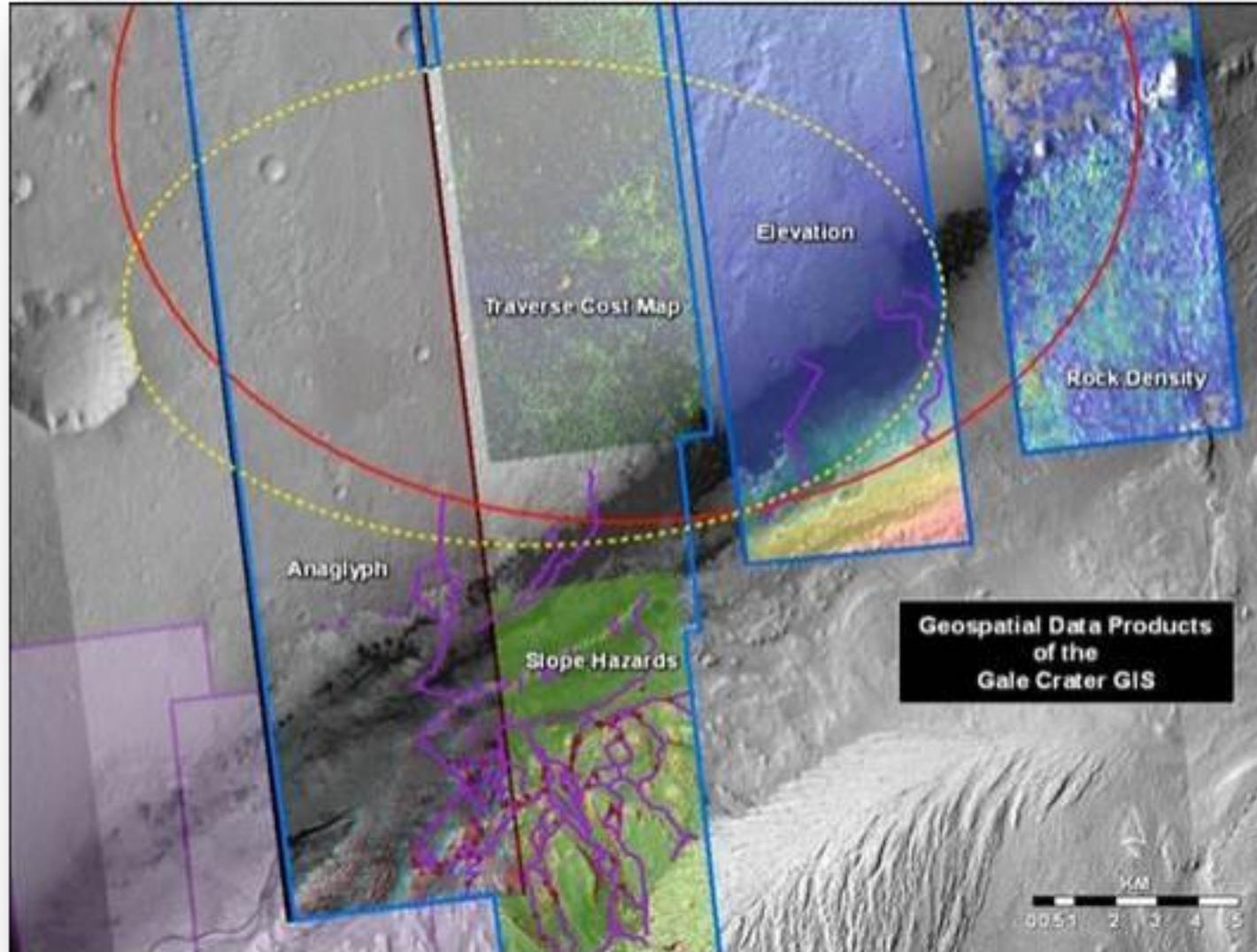


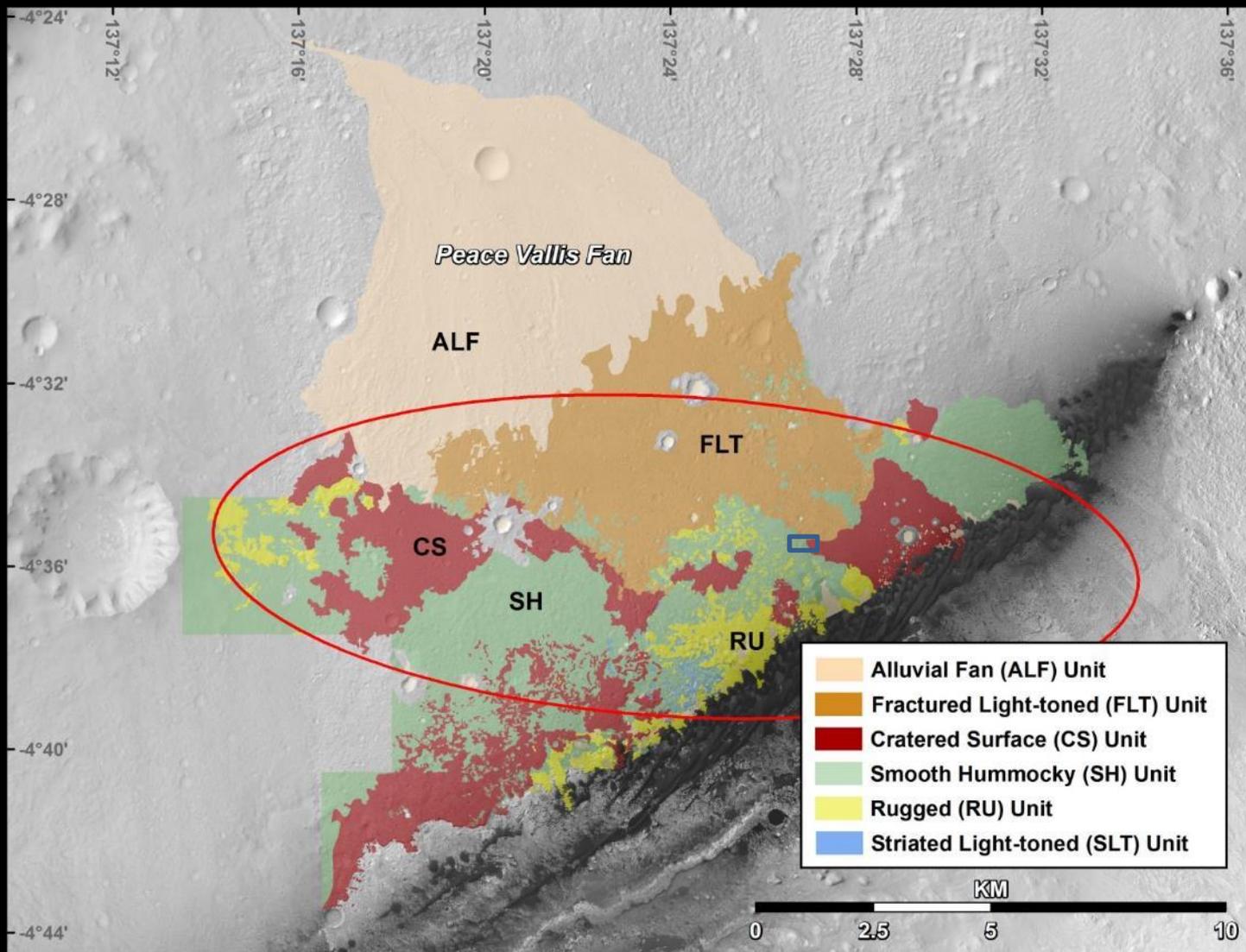
Curiosity's Rover Environmental Monitoring Station is taking weather readings continually

Multiple Co-Registered Datasets in Landing Zone: Slope, Rock Abundance, Hazards, and more



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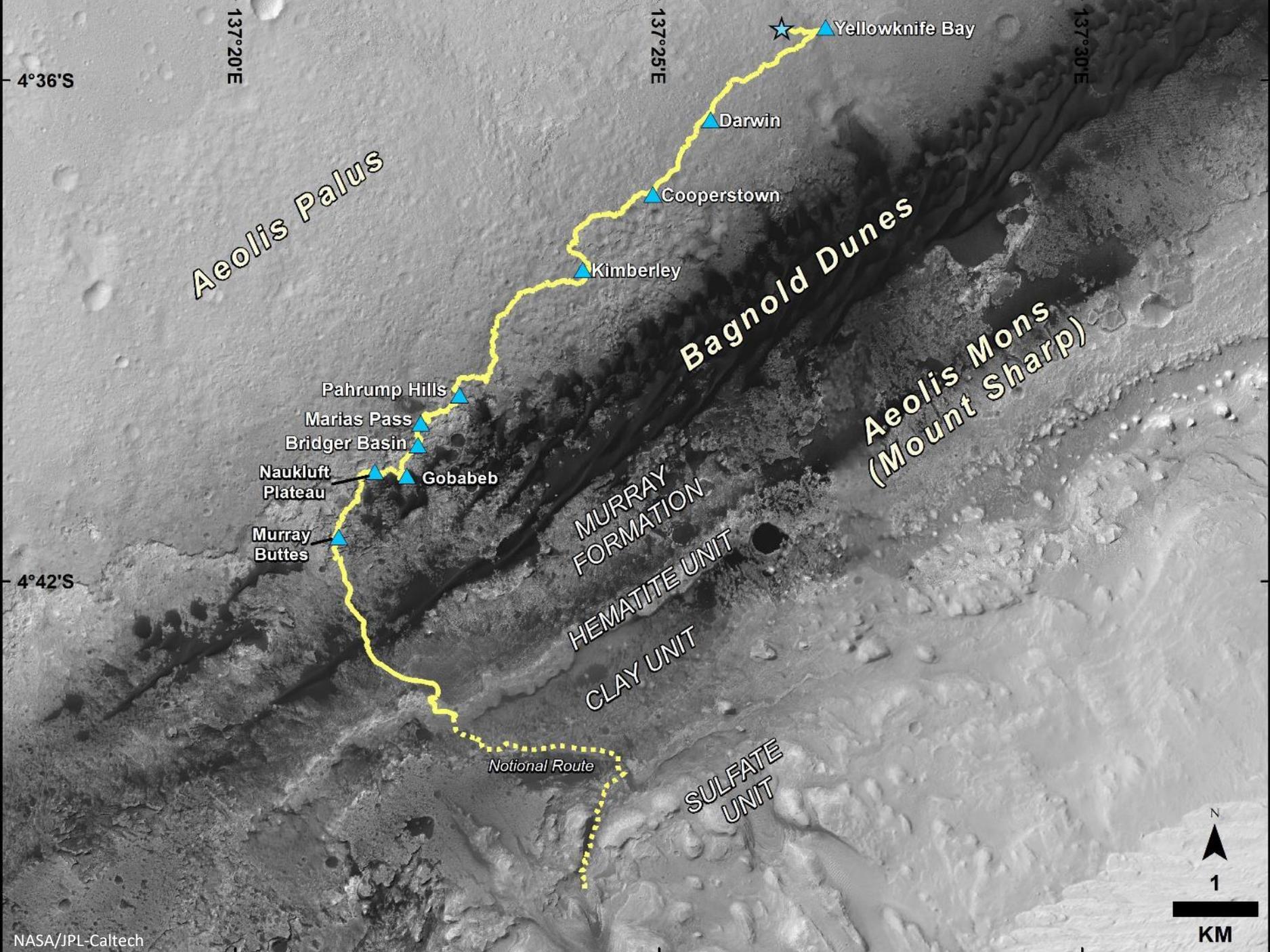


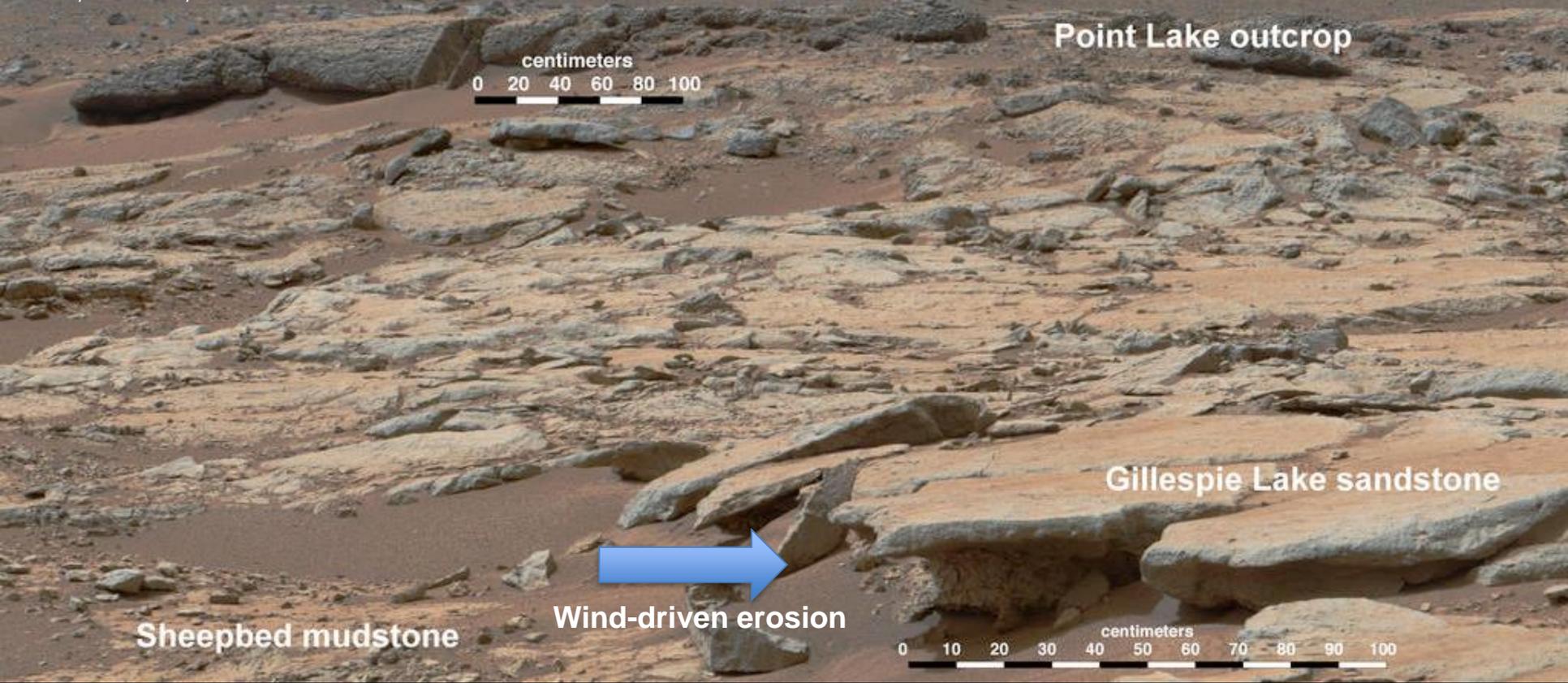


NASA/JPL-Caltech/Univ. of Arizona

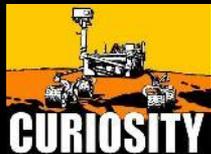
Orbital data used to derive first science product (geologic map) & strategic traverse decisions (Where do we go first? Why?)



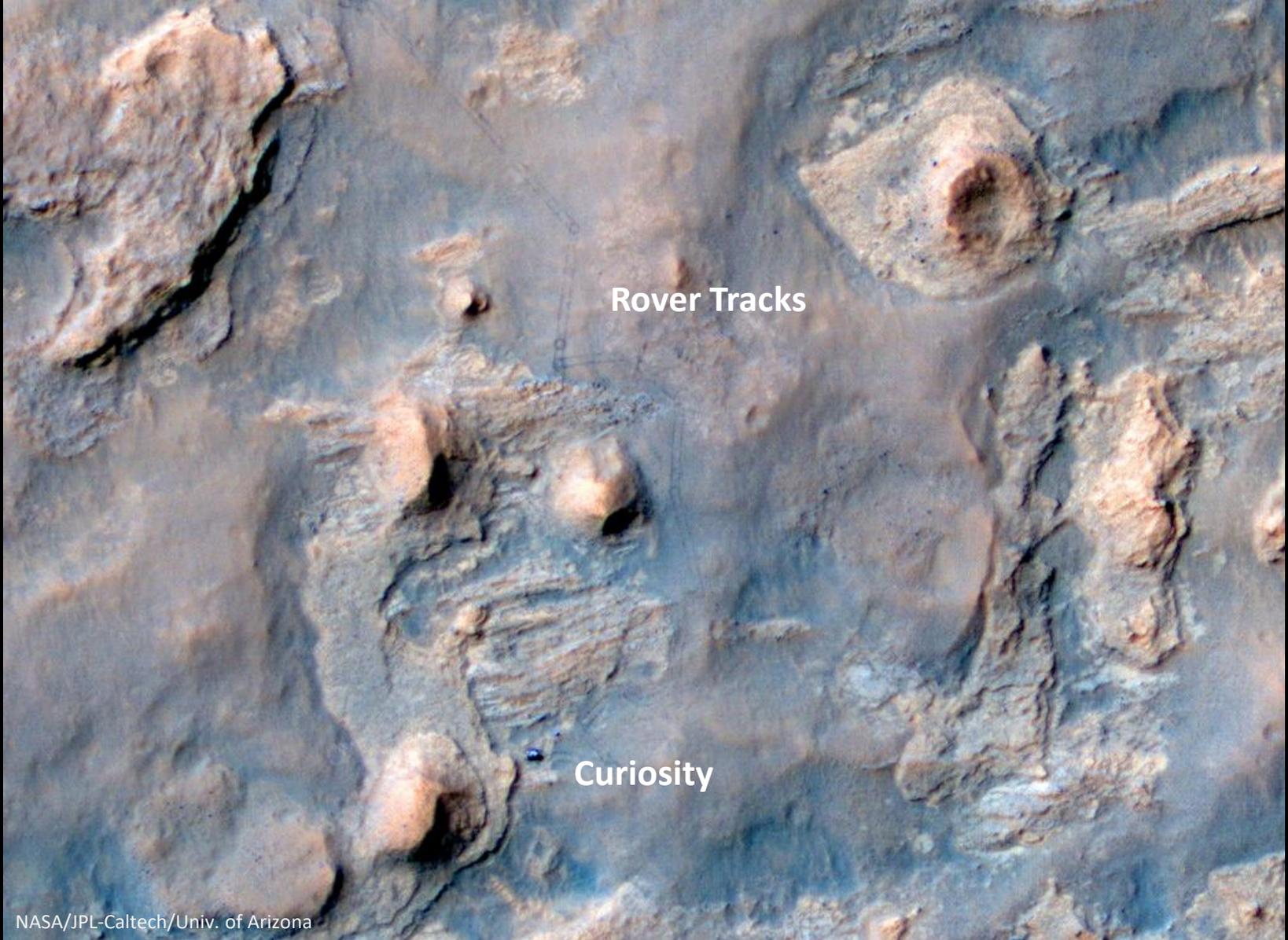




Curiosity carried out the first radiometric (K-Ar) and exposure age dating on the surface of another planet. The “Sheepbed” mudstone has been exposed for only 78 ± 30 million years, based on ^{36}Ar , ^3He , and ^{21}Ne cosmogenic isotopes.



These results suggest that natural erosion can be used to find sites where organic materials would have experienced less degradation.



Curiosity at the Kimberley, where four rock types typical of Gale Crater's plains come together

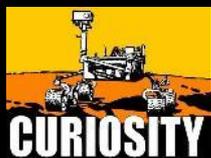
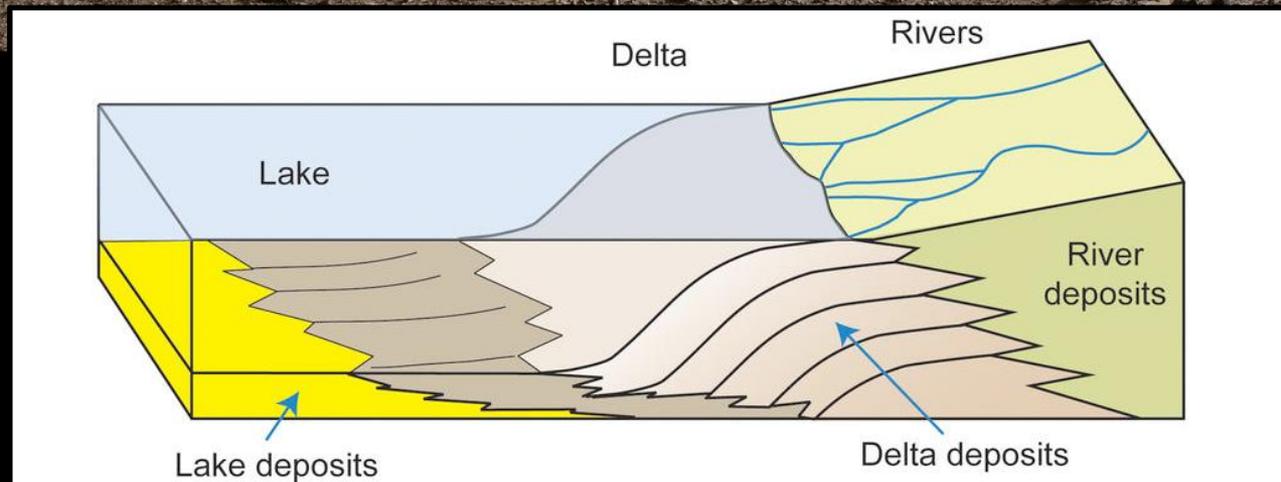


← Mount Sharp



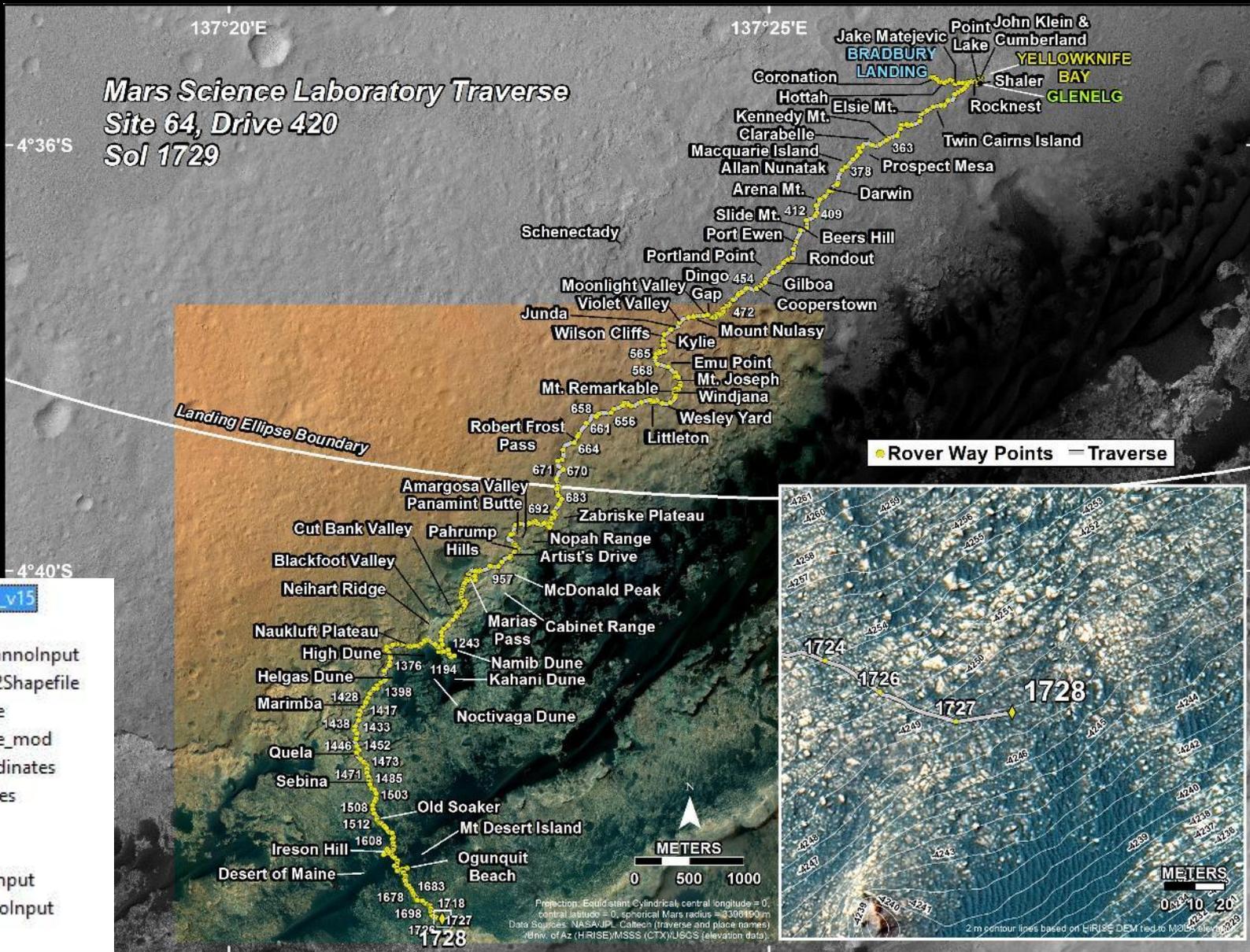
centimeters
0 20 40 60 80 100

Possible lake deposits at the base of Mount Sharp



Curiosity discovered south-tilting sandstone beds on Gale Crater's plains that indicate water-driven transport of sediment, building up lower Mount Sharp from lake deposits

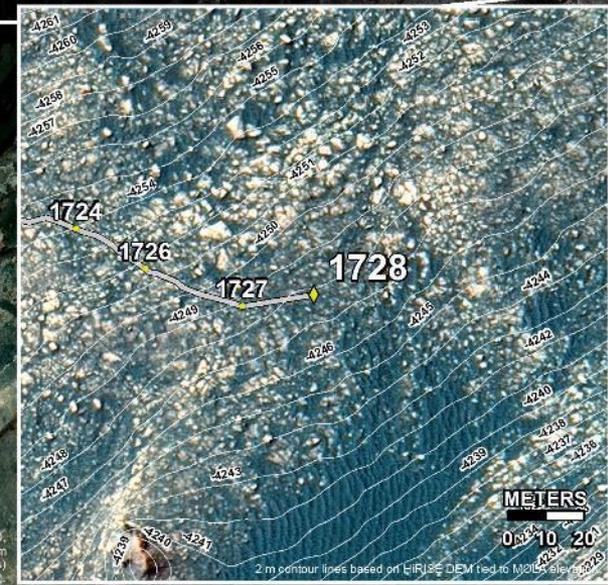
Mars Science Laboratory Traverse
 Site 64, Drive 420
 Sol 1729



● Rover Way Points — Traverse

Custom Scripts

- ☑ Mobility_Localization_v15
- ☑ Append_Traverse
- ☑ ExportFeature2rpannoInput
- ☑ LocoSpreadsheet2Shapefile
- ☑ mobility2shapefile
- ☑ mobility2shapefile_mod
- ☑ Output GEO coordinates
- ☑ Output webGIS files
- ☑ Strike-Dip
- ☑ toGeoJSON
- ☑ Traverse2rpannoInput
- ☑ Waypoints2rpannoInput
- ☑ ZipShapefile



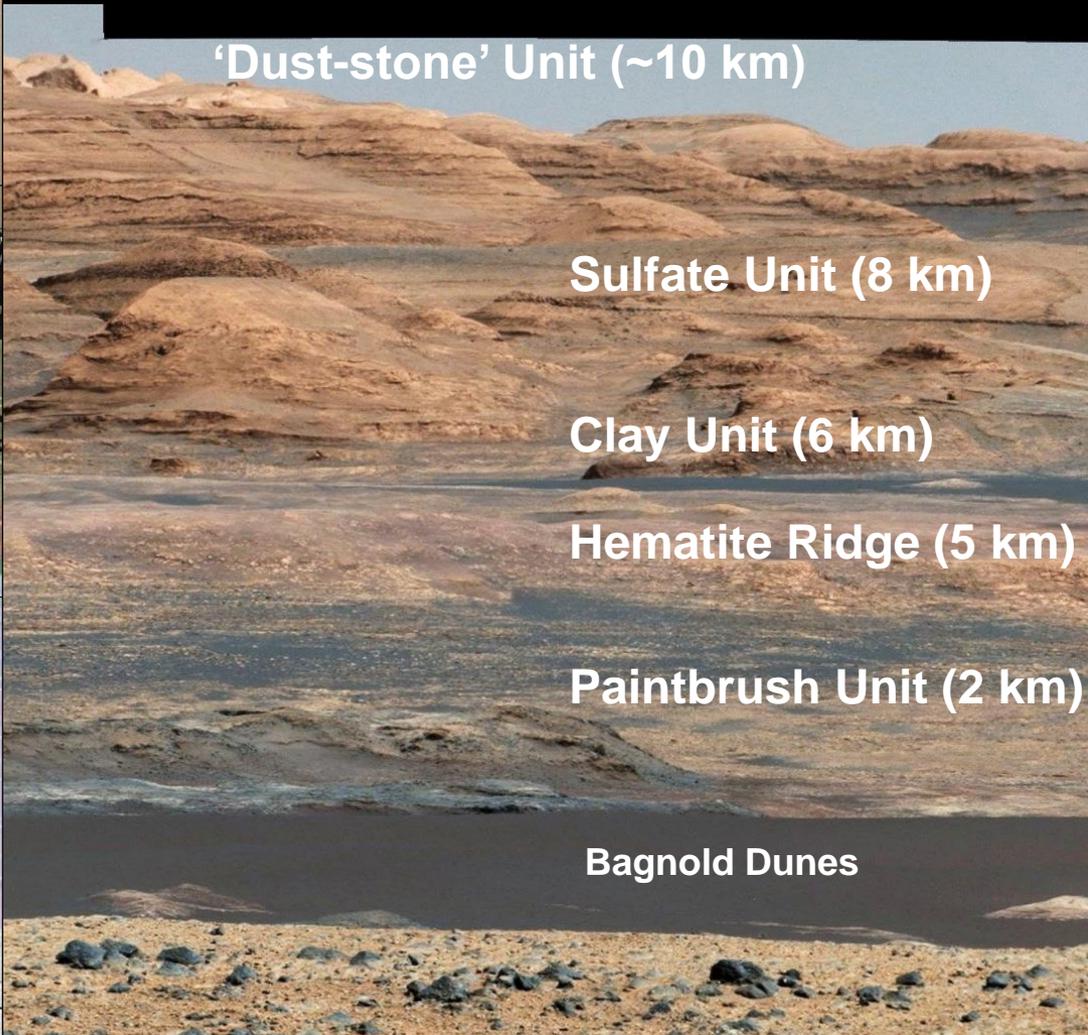
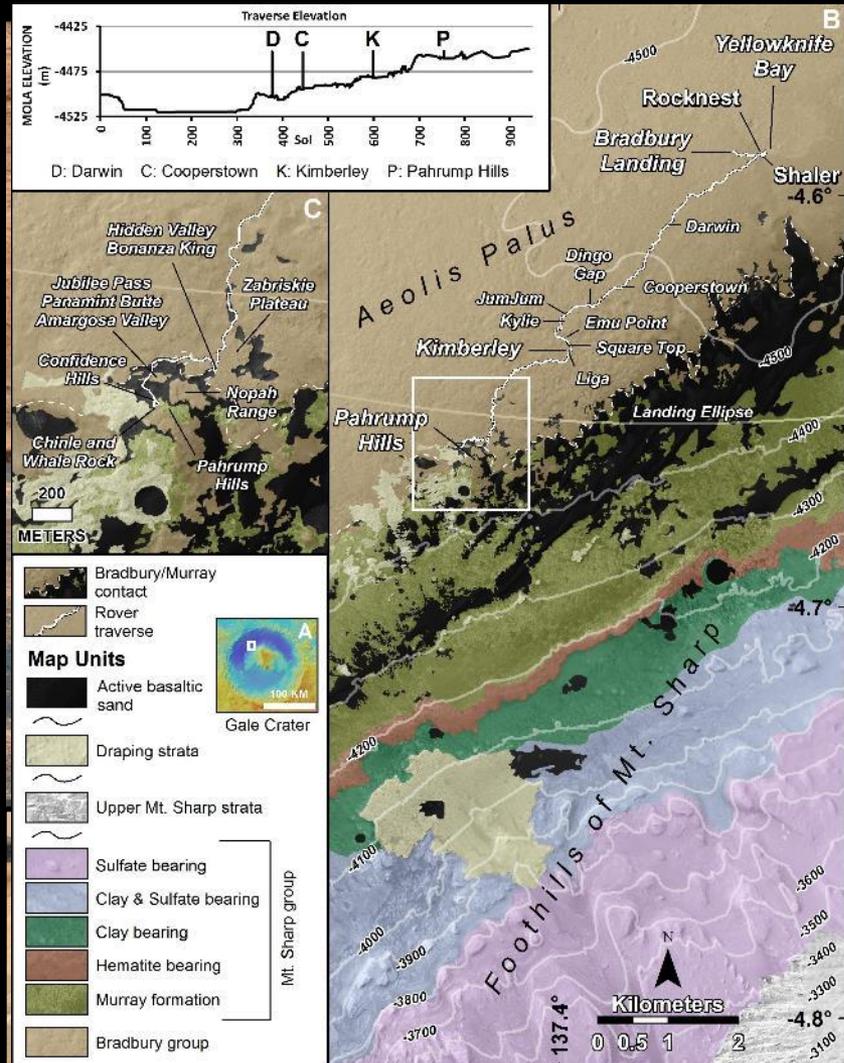
Automated Traverse Map for MSL/Curiosity Rover.

<http://mars.jpl.nasa.gov/msl/mission/whereistherovernow>

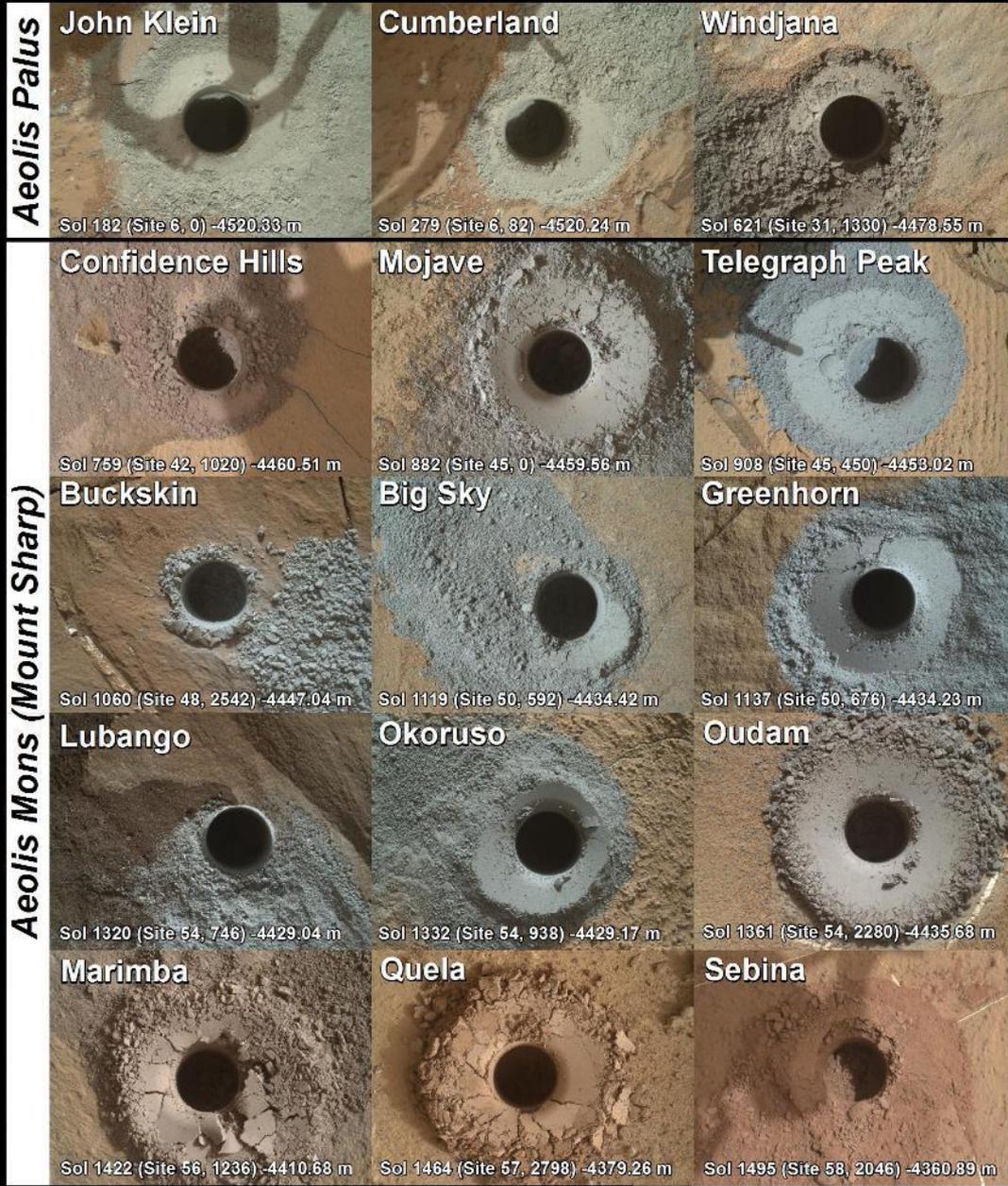
NASA/JPL-Caltech/Univ. of Arizona



Aeolis Mons (Mt. Sharp) Geology: Transition from 'Wet' to 'Dry' Mars



Left figure from Grotzinger et al., Science, 2015

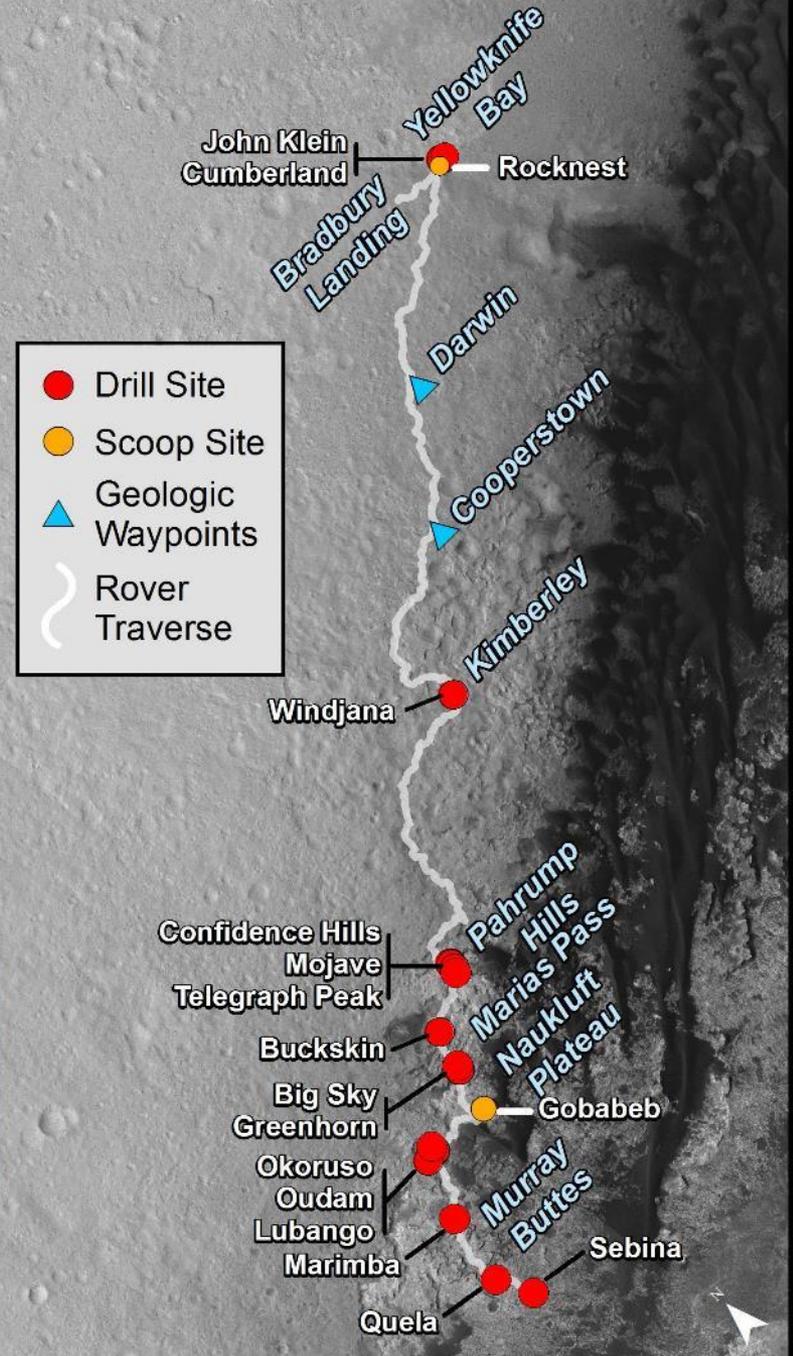


Drill hole diameter = ~1.6 cm.

Map Produced by NASA/JPL-Caltech, 2016
 MAHLI and basemap images courtesy
 NASA/JPL-Caltech/MSSS/UofA/USGS-Flagstaff



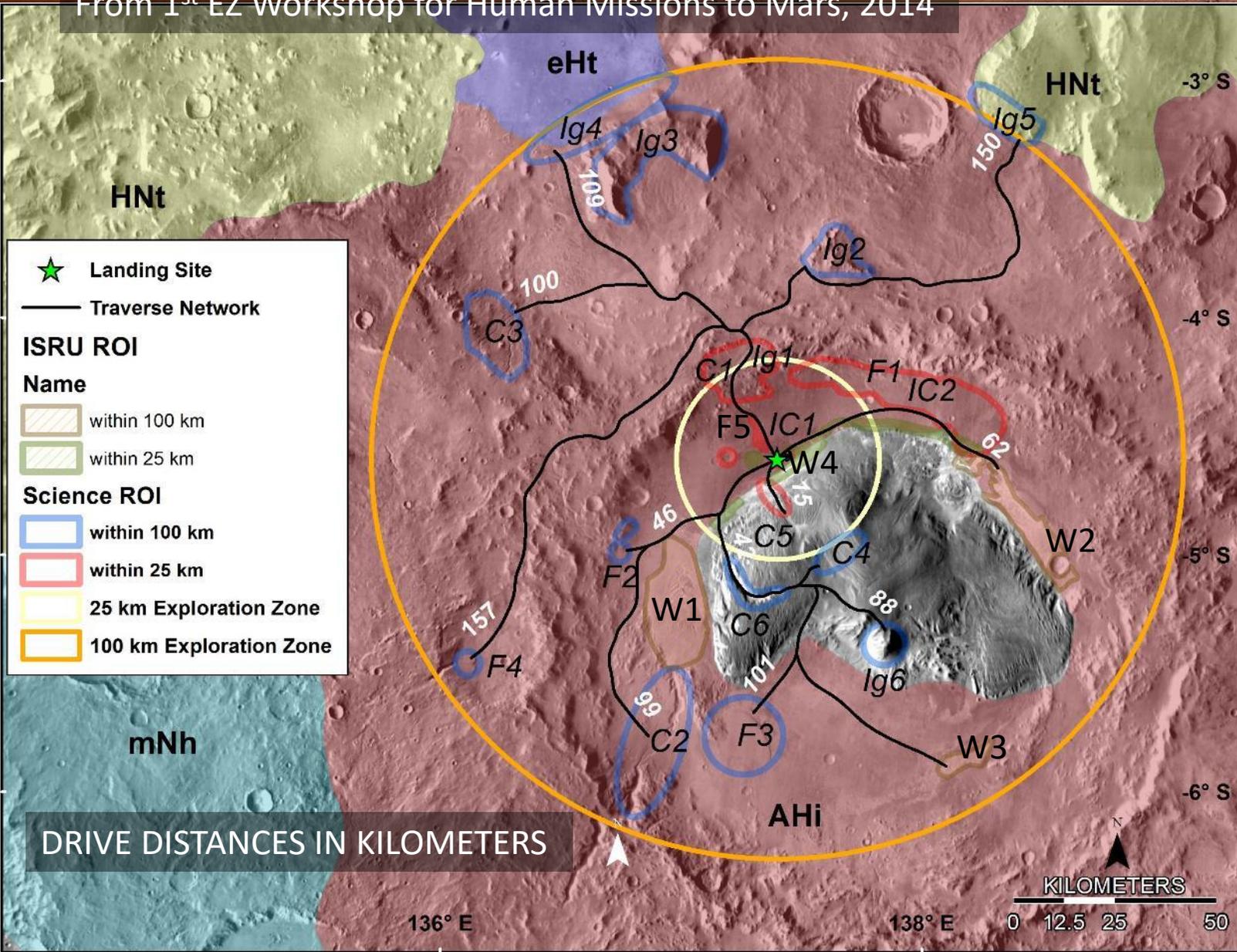
DRILL SITES AT GALE CRATER



Gale Crater as a Human Exploration Zone

From 1st EZ Workshop for Human Missions to Mars, 2014

Human Missions to Mars



★ Landing Site

— Traverse Network

ISRU ROI

Name

within 100 km

within 25 km

Science ROI

within 100 km

within 25 km

25 km Exploration Zone

100 km Exploration Zone

LZ Coordinates

Longitude (E)
137.42009295°

Latitude (S)
4.59310427°

Elevation (MOLA)
-4497.77

Easting
8145534.27 m

Northing -
272254.70 m

Science ROI

C = Channel
F = Fan
Ig = Igneous
IC = Inverted Channels

ISRU ROI

W = Water ROI

DRIVE DISTANCES IN KILOMETERS



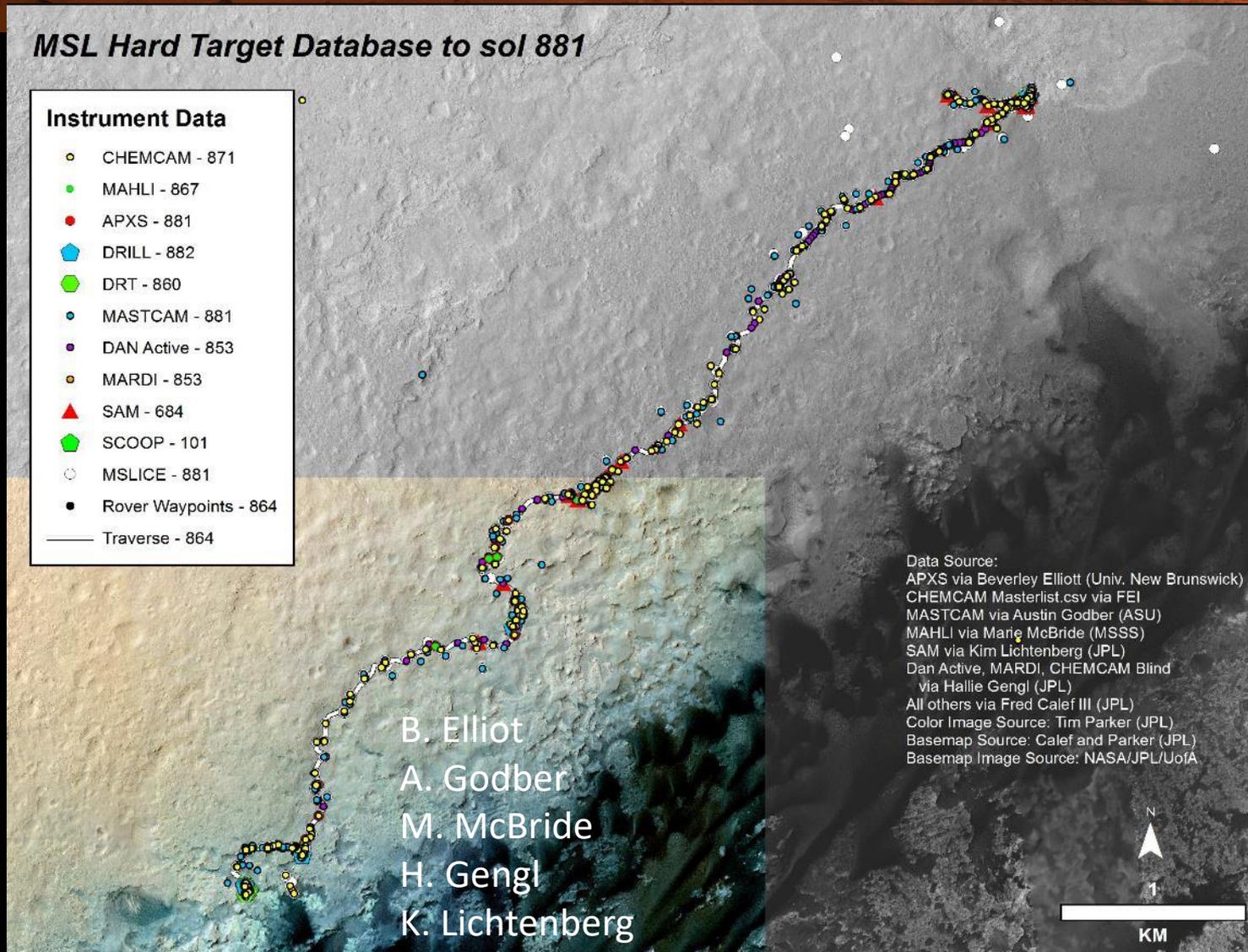
Tracking Science Observations: thousands of potential sample locations known to cm scale



MSL Hard Target Database to sol 881

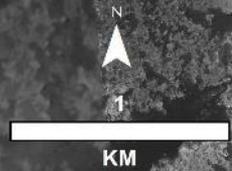
Instrument Data

- CHEMCAM - 871
- MAHLI - 867
- APXS - 881
- DRILL - 882
- DRT - 860
- MASTCAM - 881
- DAN Active - 853
- MARDI - 853
- ▲ SAM - 684
- SCOOP - 101
- MSLICE - 881
- Rover Waypoints - 864
- Traverse - 864



B. Elliot
A. Godber
M. McBride
H. Gengl
K. Lichtenberg

Data Source:
APXS via Beverley Elliott (Univ. New Brunswick)
CHEMCAM Masterlist.csv via FEI
MASTCAM via Austin Godber (ASU)
MAHLI via Marie McBride (MSSS)
SAM via Kim Lichtenberg (JPL)
Dan Active, MARDI, CHEMCAM Blind
via Hallie Gengl (JPL)
All others via Fred Calef III (JPL)
Color Image Source: Tim Parker (JPL)
Basemap Source: Calef and Parker (JPL)
Basemap Image Source: NASA/JPL/UofA

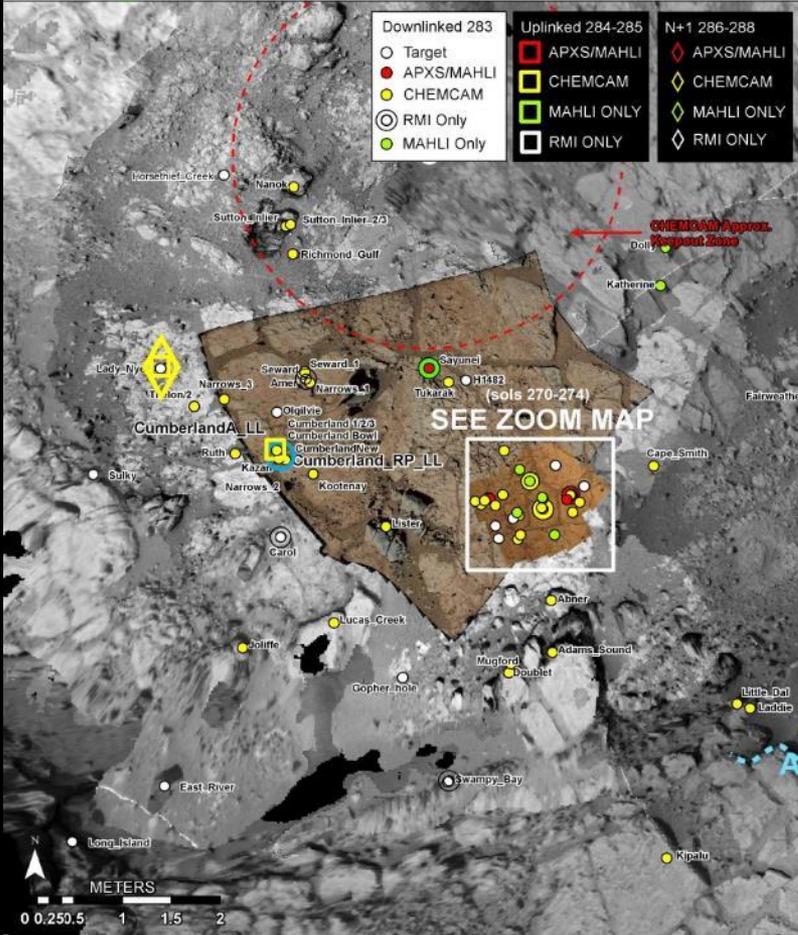


ons to Mars

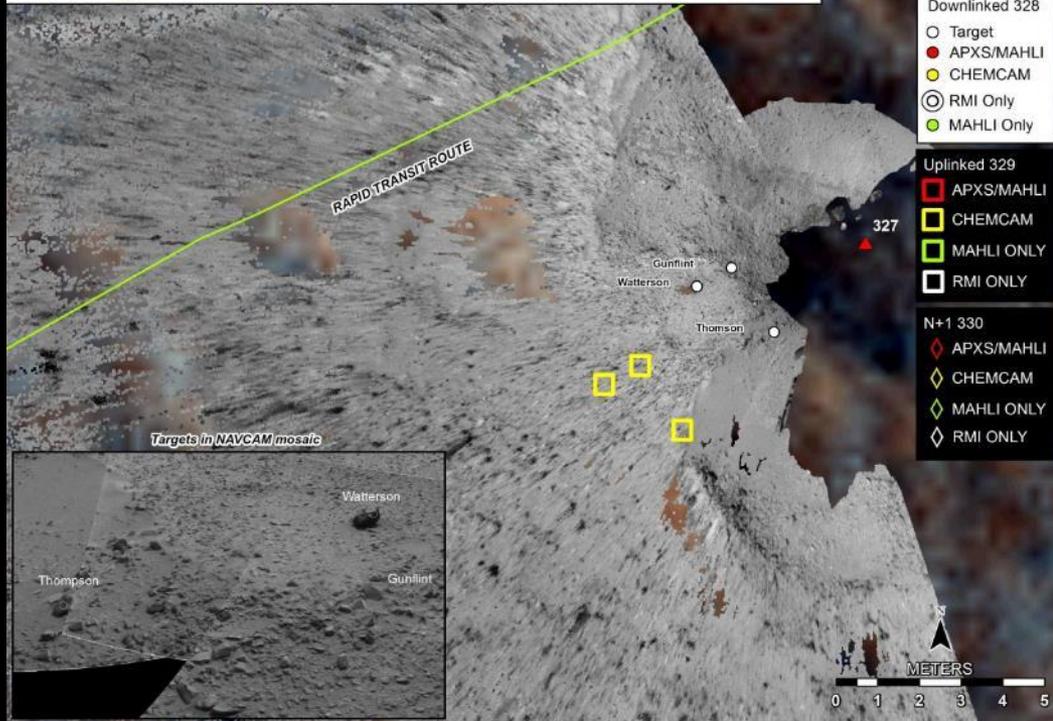
Example Target Map from MSL Operations: known locations for future science opportunities maximize science return during field excursions



Target Data: Sol 283-288



Target Data: Sol 328-330



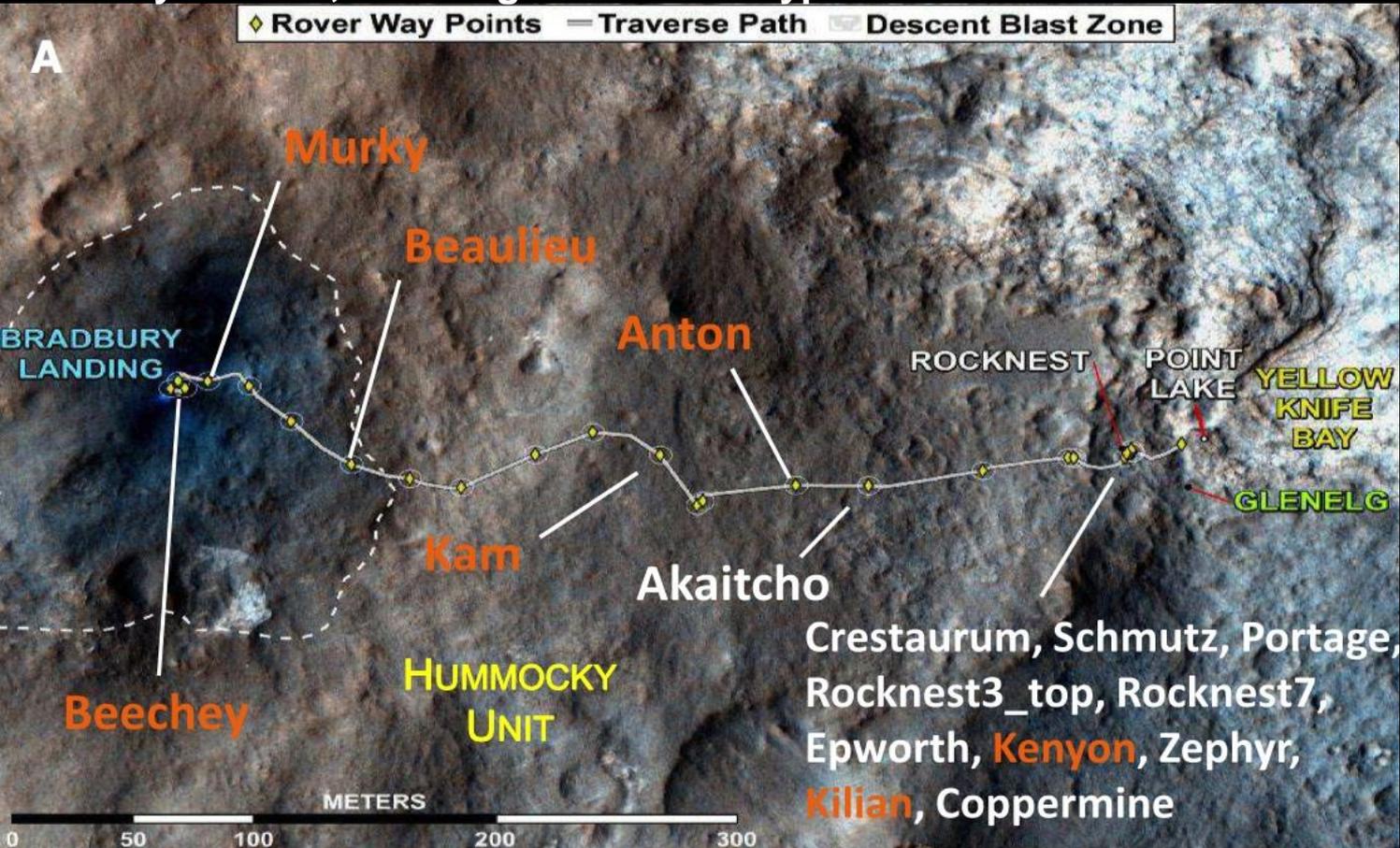
Felsic vs. Mafic ISRU Source Areas



1st EZ Workshop for Human Missions to Mars

ChemCam identified two principal soil types along the traverse to Yellowknife Bay: a fine-grained, mafic type similar to other soils, and a locally derived, coarse-grained felsic type.

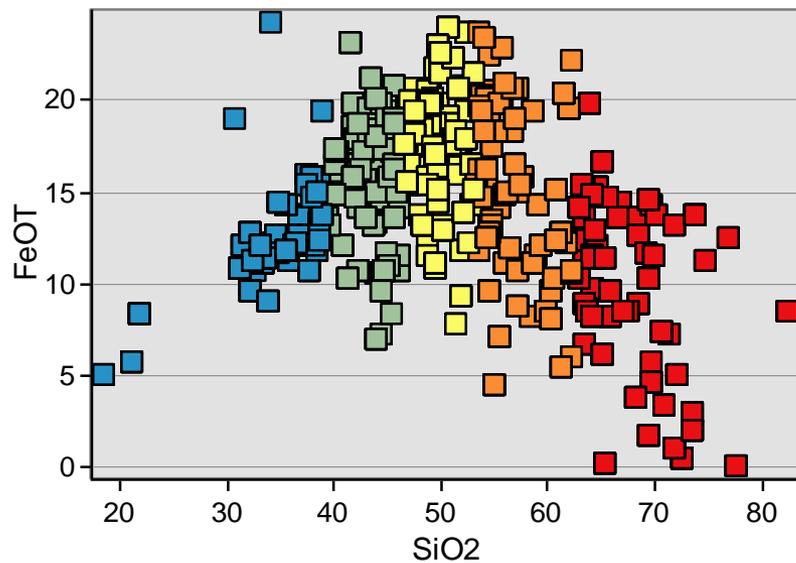
Mafic soil component has hydration signature, corresponding to the X-ray amorphous component sampled by CheMin and SAM.



Meslin et al. [2013]

Mafic
Felsic

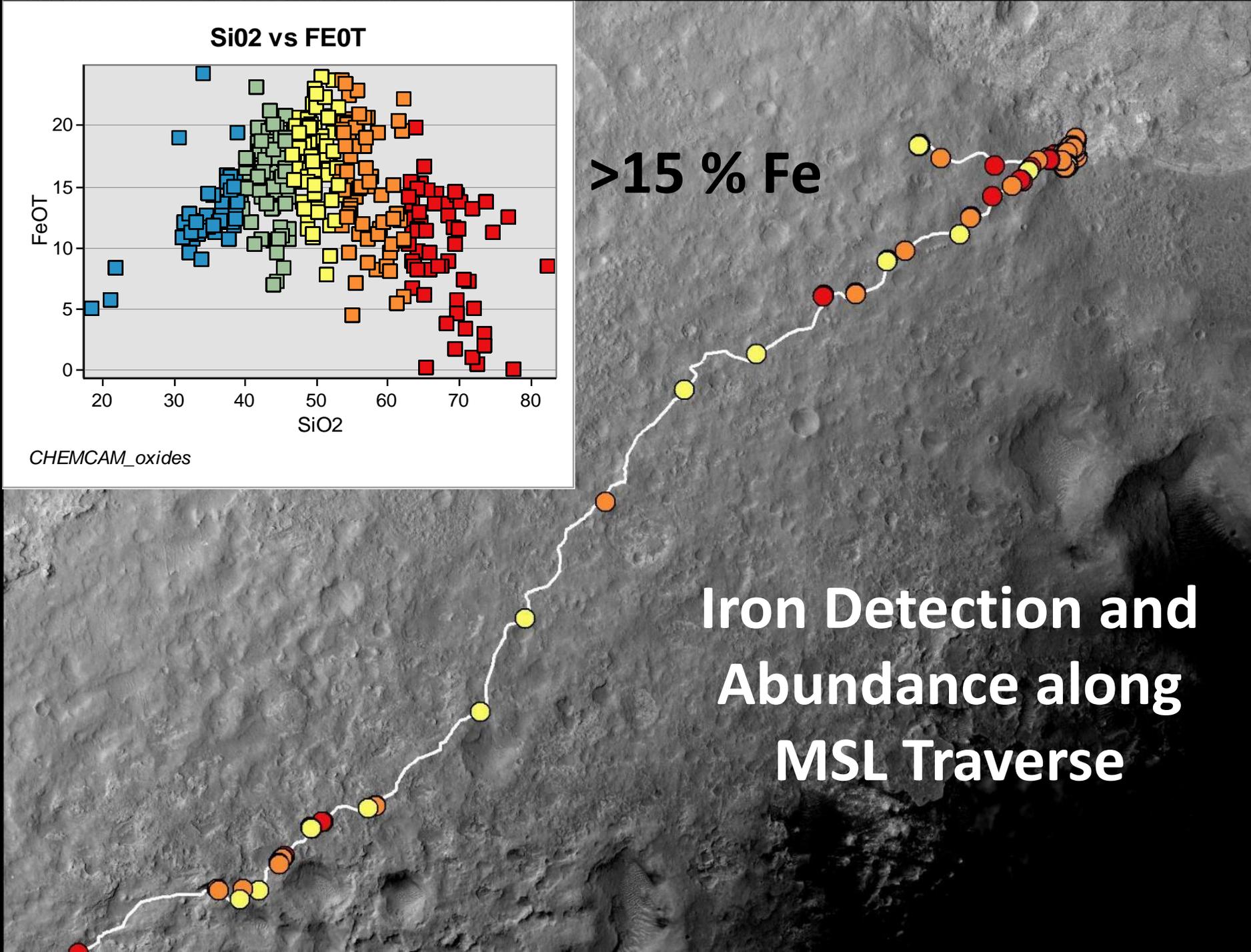
SiO₂ vs FeOT



CHEMCAM_oxides

>15 % Fe

Iron Detection and
Abundance along
MSL Traverse

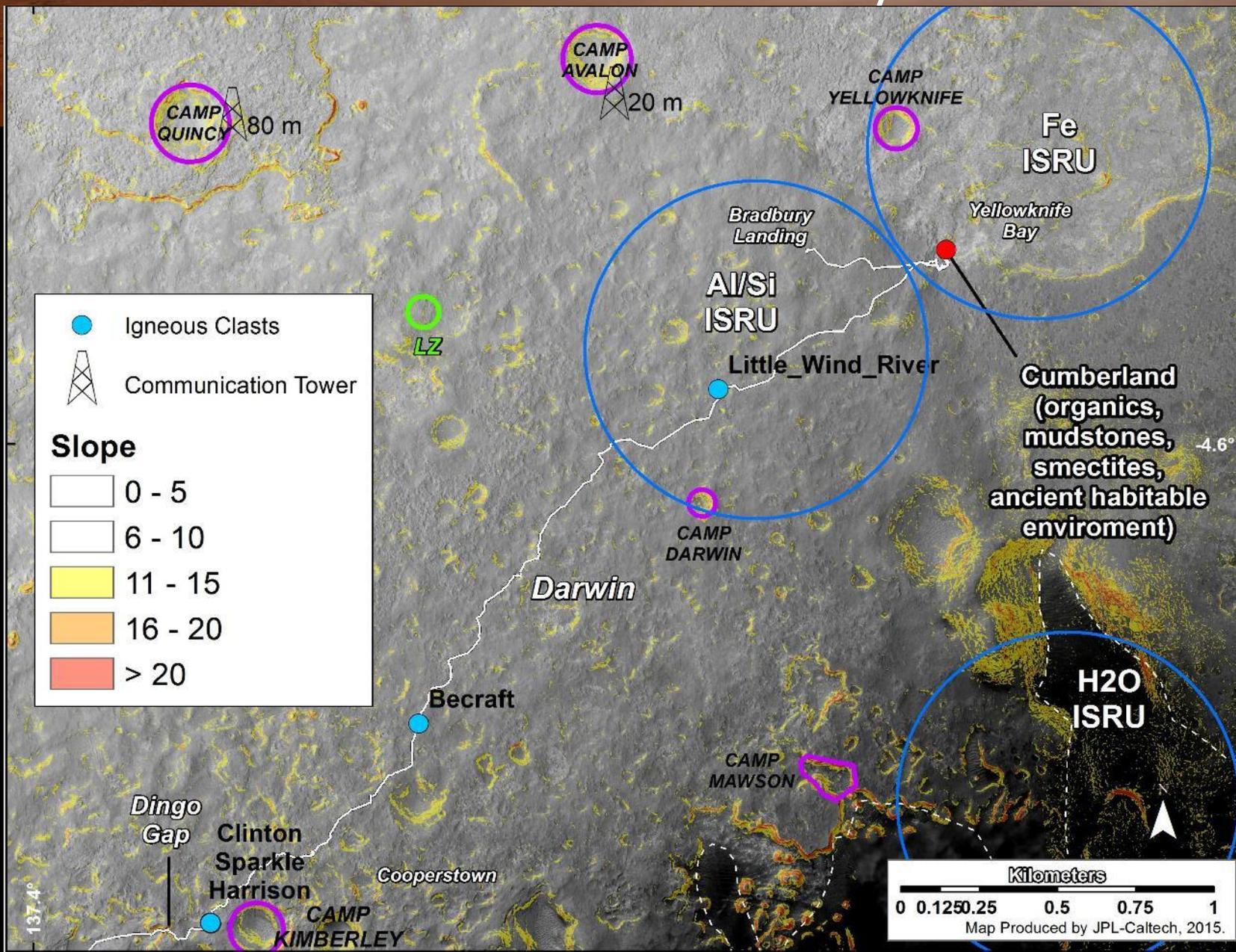


Exploration Zone:

Potential Habitation Sites and nearby ISRU Zones



ns to Mars



ISRU Resource Numbers

– Water

- Based on Leshin et al. (2013), Rocknest soil contained up to 3 wt% H₂O via SAM, Bagnold Dune field contains 10⁴-10⁶ MT adsorbed water (were closer to 1%)
- Easy to process. Potentially 'reusable'.
- Rocknest sand has a typical Mars basalt composition, but also 1.5-3% bound water.
- Both a water and food growth medium.

• Aluminum and Silicon

- ISRU01 rocks on average have ~12 wt% Al and ~55-60 wt% Si. ISRU02 rocks, ~18 wt% Fe.
- Assuming rock density ~2.5 gm/cc:
 - 300 kg/m³ Al (ISRU01)
 - 1375-1500 kg/m³ Si (ISRU01)
 - 450 kg/m³ Fe (ISRU02)



Resource ROI(s) Rubric



Site Factors			RW01-04	RISUR1	LZ01	ISR02							EZ SUM		
ISRU and Civil Engineering Criteria	Water Resource	Engineering	Meets First Order Criteria (Latitude, Elevation, Thermal Inertia)										1		
		Threshold	AND/OR	Potential for ice or ice/regolith mix			●								1
			Potential for hydrated minerals	●											4
			Quantity for substantial production	●											4
			Potential to be minable by highly automated systems	●											4
			Located less than 3 km from processing equipment site	●											1
			Located no more than 3 meters below the surface	●											4
			Accessible by automated systems	●											4
		Qualifying	Potential for multiple sources of ice, ice/regolith mix and hydrated minerals												
			Distance to resource location can be >5 km	●											3
	Route to resource location must be (plausibly) traversable		●											3	
	Civil Engineering	Threshold	~50 sq km region of flat and stable terrain with sparse rock distribution			●								1	
			1-10 km length scale: <10°			●								1	
			Located within 5 km of landing site location			●								1	
		Qualifying	Located in the northern hemisphere			○								0, 1	
	Evidence of abundant cobble sized or smaller rocks and bulk, loose regolith				●									1	
	Food Production	Qualifying	Utilitarian terrain features			●								1	
			Low latitude			●								1	
			No local terrain feature(s) that could shadow light collection facilities			●									1
			Access to water			●									1
Access to dark, minimally altered basaltic sands					●									1	
Metal/Silicon Resource	Threshold	Potential for metal/silicon		●		●							2		
		Potential to be minable by highly automated systems		●		●								2	
		Located less than 3 km from processing equipment site		●		●								2	
		Located no more than 3 meters below the surface		●		●								2	
		Accessible by automated systems													
	Qualifying	Potential for multiple sources of metals/silicon		●		●								2	
		Distance to resource location can be >5 km		●		●								2	
		Route to resource location must be (plausibly) traversable		●		●								2	

Key

- Yes
- Partial Support or Debated
- No
- ? Indeterminate