



MRO: Continuing Discoveries at Mars

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Mars
Reconnaissance
Orbiter



AAS GN&C Conference
Breckenridge, CO
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4 MARS SCIENCE GOALS

1

WATER

LIFE

Determine if life ever arose on Mars

2

HABITABLE ZONES

CLIMATE

Understand Martian climate processes and history

3

SIGNS OF LIFE

GEOLOGY

Determine how the surface and interior of Mars evolved

4

HUMANS

Prepare for human exploration



EVOLVING THEMES

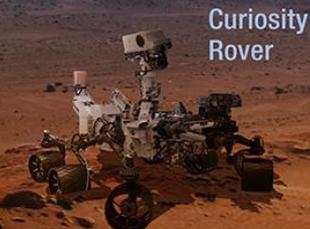
Operational 2001–2015

2016

2018

2020

2022



Follow the Water

Explore Habitability

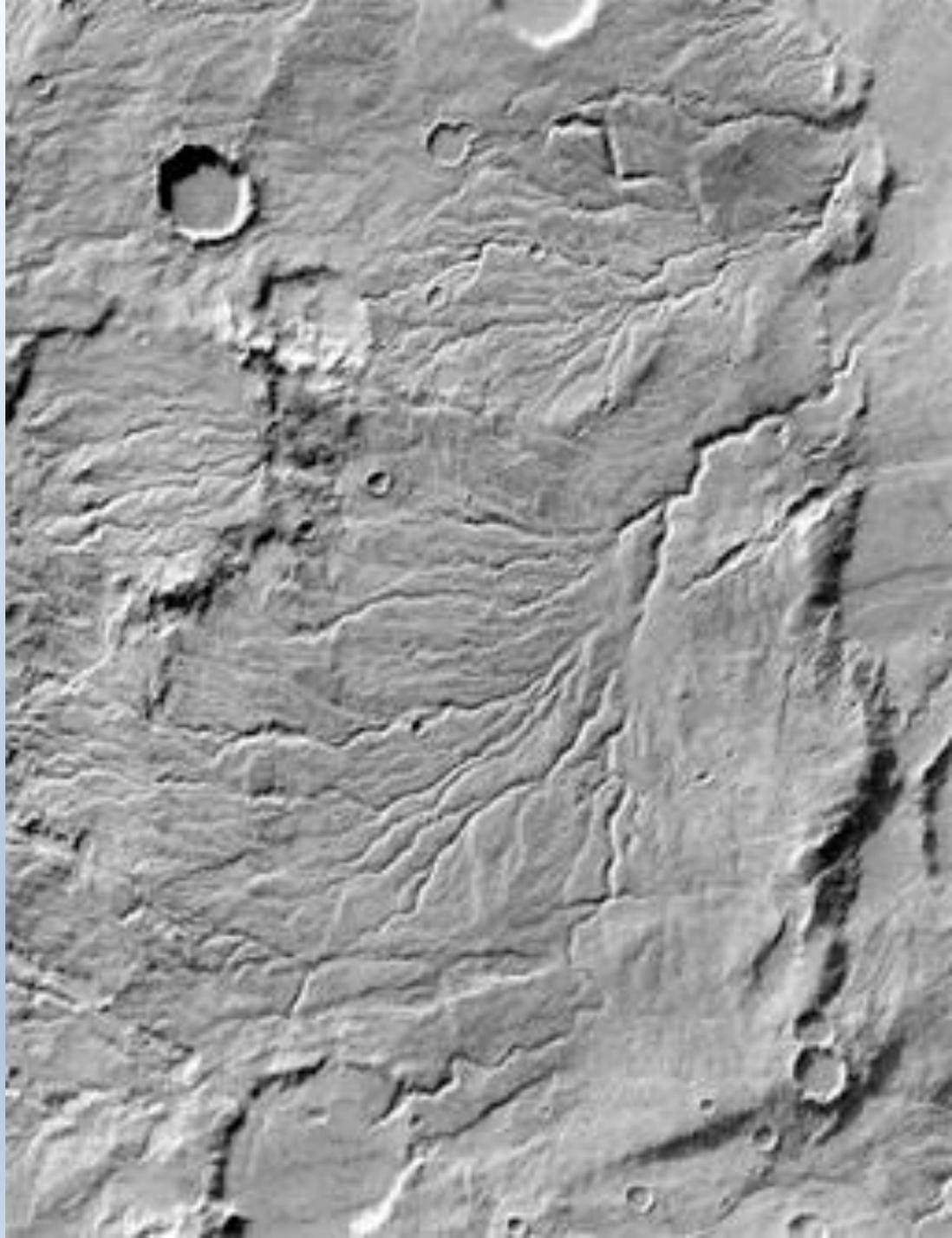
Seek Signs of Life

Prepare for Future Human Explorers



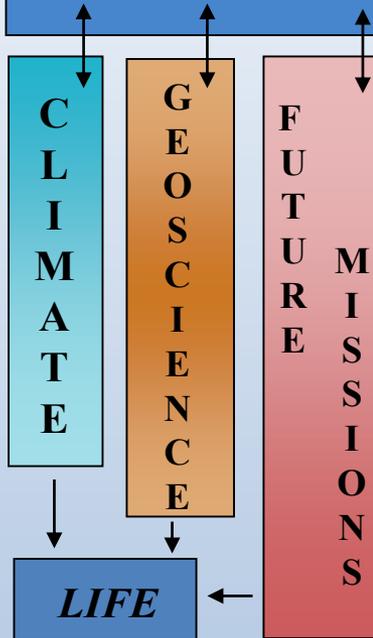
What did we know about Mars in 2005?

- Ancient Mars
 - Wet, but when & for how long, were the locations habitable?
- Geologically Recent Mars
 - Layering seen in polar cap edges, but composition, depth, formation unknown
 - Mid-latitude ice was a possibility, but not certain
- Present Mars
 - Dunes, but thought to be frozen from some ancient time
 - Gullies formed, but when?
 - Vertical distribution of dust and water-ice unknown
 - Was there CO₂ snowfall in polar regions (not just frost)?
 - No liquid water on Mars today



MRO Has Many Roles

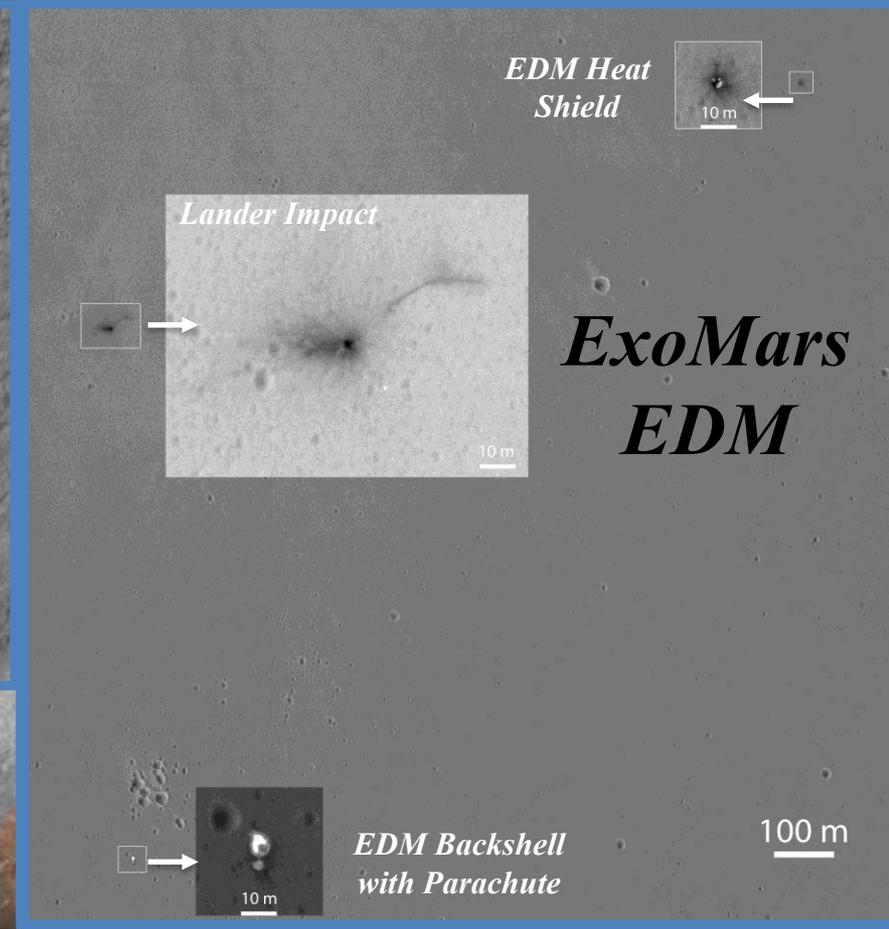
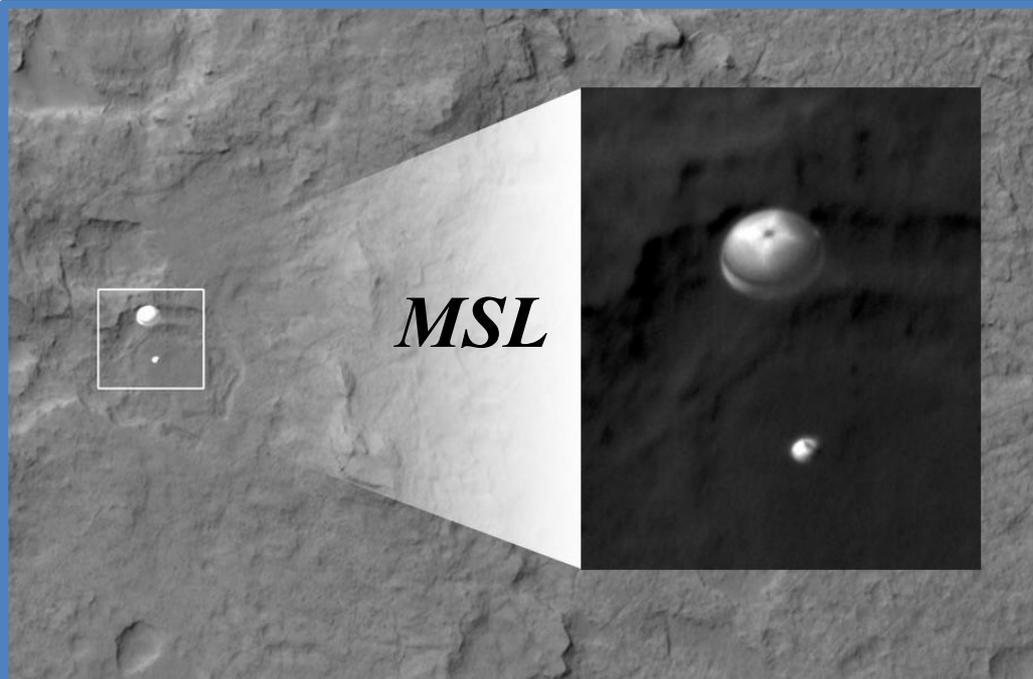
“Follow the Water” Theme



- ***Weather and Climate Satellite***
 - Monitor the present climate, seasonally and year-to-year
 - Alerts Mars Exploration Rovers to evolving dust storm events
 - Monitored atmosphere for landers as they approached Mars entry
- ***Geological Explorer***
 - Identified water-related landforms and aqueous surface deposits
 - Probe the subsurface looking for layering and water (ice)
- ***Site Finder***
 - MRO data led Phoenix to a new site, & characterized & certified the site
 - Gale Crater was selected for MSL-Curiosity based on MRO data
 - Observe hundreds of sites at very high-resolution to find the best places for future landed exploration (InSight, 2020 NASA & ExoMars Rovers, potential human exploration zones)
- ***Communications Satellite***
 - Provided critical event coverage (entry, descent and landing) and relay for both Phoenix and provides relay for Mars Science Laboratory (~450 Mb/sol)
- ***Technology Pathfinder***
 - Demonstrated optical navigation and use of new telecom frequencies (Ka-band) for future missions



MARS Program Support by MRO



MRO Science Investigations

<i>Instrument</i>	<i>Type</i>	<i>PI/TL, Institution</i>	<i>Attributes</i>
<i>CRISM</i>	Hyper-Spectral Imaging VIS-NIR Spectrometers	<i>Scott Murchie, PI</i> APL / Johns Hopkins University	Targeted Observing @ 18 m/pixel Regional Survey @ 100-200 m/pixel Very High Data Rate 85% of Mars surveyed in 72 channels
<i>HiRISE</i>	Very High Resolution Imaging	<i>Alfred McEwen, PI</i> University of Arizona	Targeted Imaging @ 30 cm/pixel Swath: 5.4 km w. 1.2 km 3-color strip Very High Data Rate / 2.7% of Mars
<i>SHARAD</i>	Shallow Subsurface RADAR (Provided by ASI)	<i>Roberto Seu, TL/PI</i> University of Rome <i>Roger Phillips, rDTL</i> <i>Nathaniel Putzig, aDTL</i>	Regional Radar Profiling of Subsurface Profiles to 0.5 km in regolith / 1.5 km in ice @ ~10 m vertical resolution High Data Rate
<i>CTX</i>	High-Resolution Context Imager	<i>Michael Malin, TL</i> Malin Space Science Systems	Targeted & Regional Survey 6 m/pixel, panchromatic in 30 km swath High Data Rate / 98% of Mars
<i>MARCI</i>	Mars Color Imager	<i>Michael Malin, PI</i> Malin Space Science Systems	Daily Global Mapping ~1 km/pixel in 7 color bands Moderate Data Rate / > 5.4 Mars yrs.
<i>MCS</i>	Mars Climate Sounder	<i>Daniel McCleese, PI</i> JPL / Caltech	Daily Global Sounding (T, p, aerosols) ~5 to 80 km Low-Data Rate / > 5.4 Mars yrs.
<i>ACCEL</i>	Facility Science Team Investigation	<i>Gerald Keating*, TL</i> GWU / LaRC	Profiled upper atmosphere using S/C Accelerometers during Aerobraking. (complete)
<i>Gravity Science</i>	Facility Science Team Investigation	<i>Maria Zuber, TL</i> MIT / GSFC	Data from DSN tracking using Spacecraft X Band Telecom



CRISM



HiRISE



SHARAD



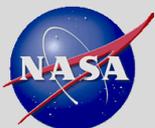
CTX



MARCI



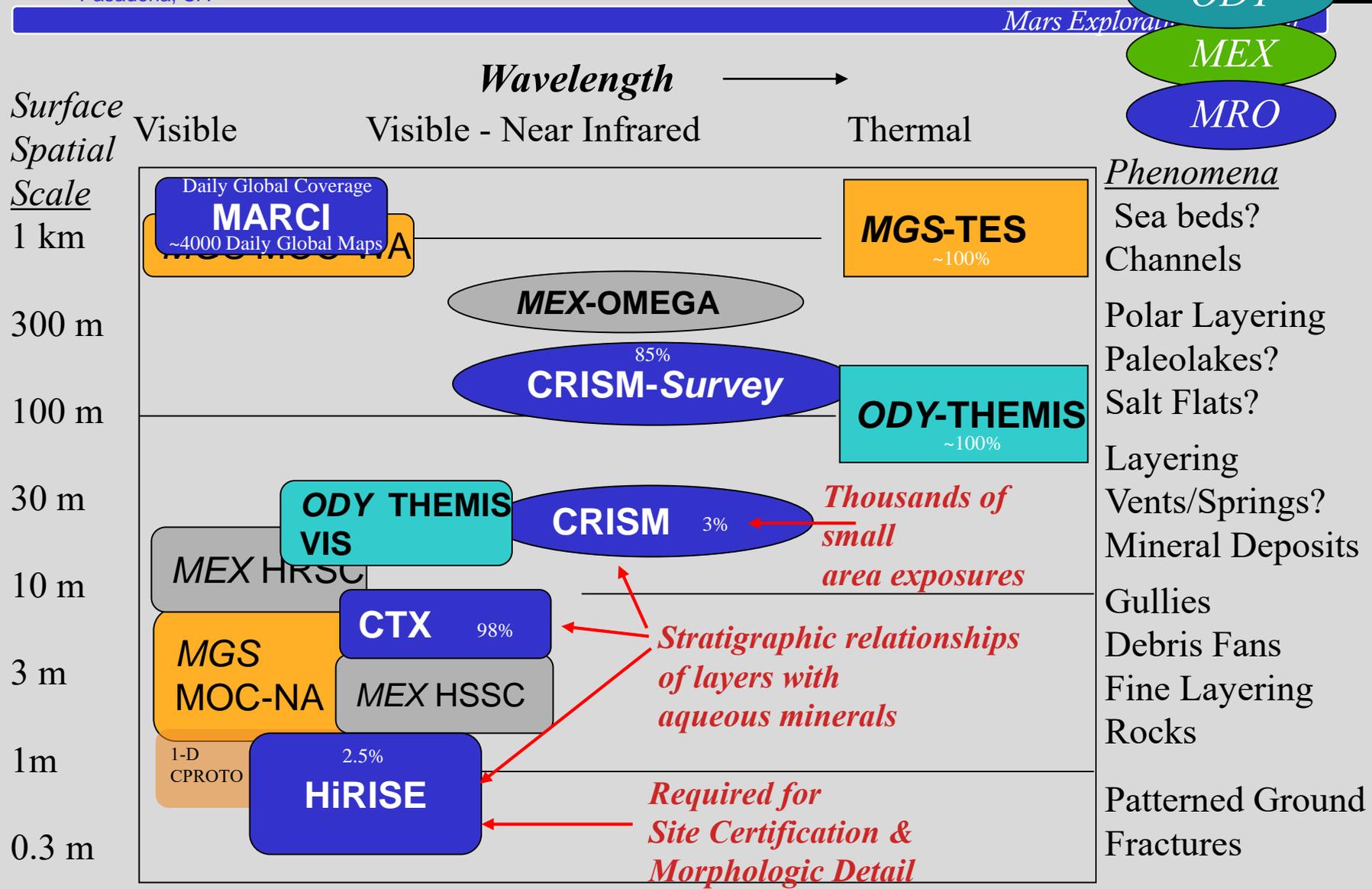
MCS



Gains in Resolution



- MGS
- ODY
- MEX
- MRO



Telecom:
X, Ka-Band & UHF
100 W X-Band TWTAs
3-m diameter High Gain Antenna

2180 kg Launch Mass (Atlas V)
2000W Array at Aphelion
500 kbps @max Mars-Earth Range
1551 m/s Delta-V

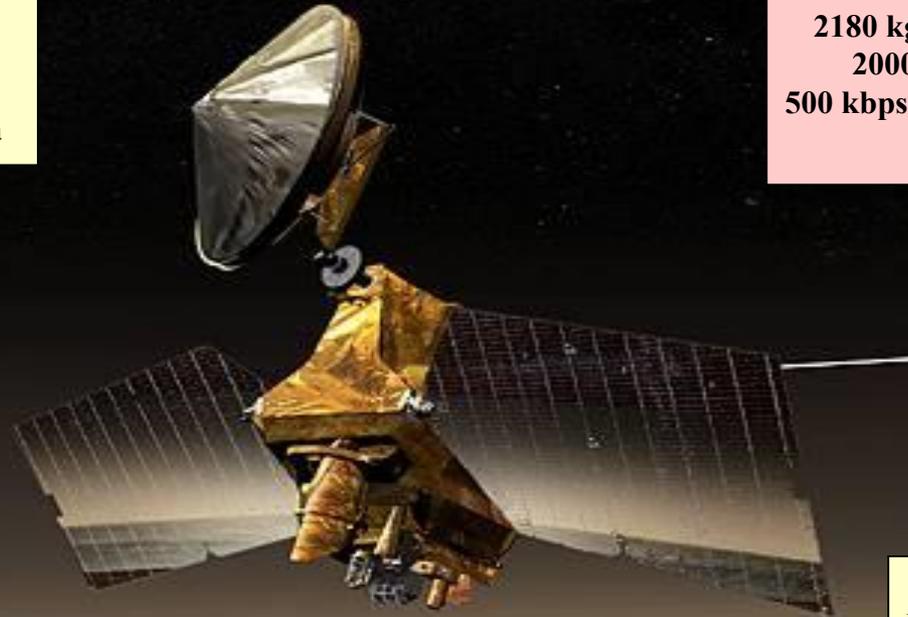
*~300 Tb of Science
Data Returned*

Power:
Dual 50 A-Hr NH₂ Batteries
20 m² of GaAs 3J Solar Cells

*Supports Continuous
Payload Operations*

Propulsion:
Single-Tank Mono Prop Design
20 Yrs Consumables

Fuel Not a Limit



Targeting:
Ephemeris-Based Targeting
Time-Tagged Sequence Fully Supported

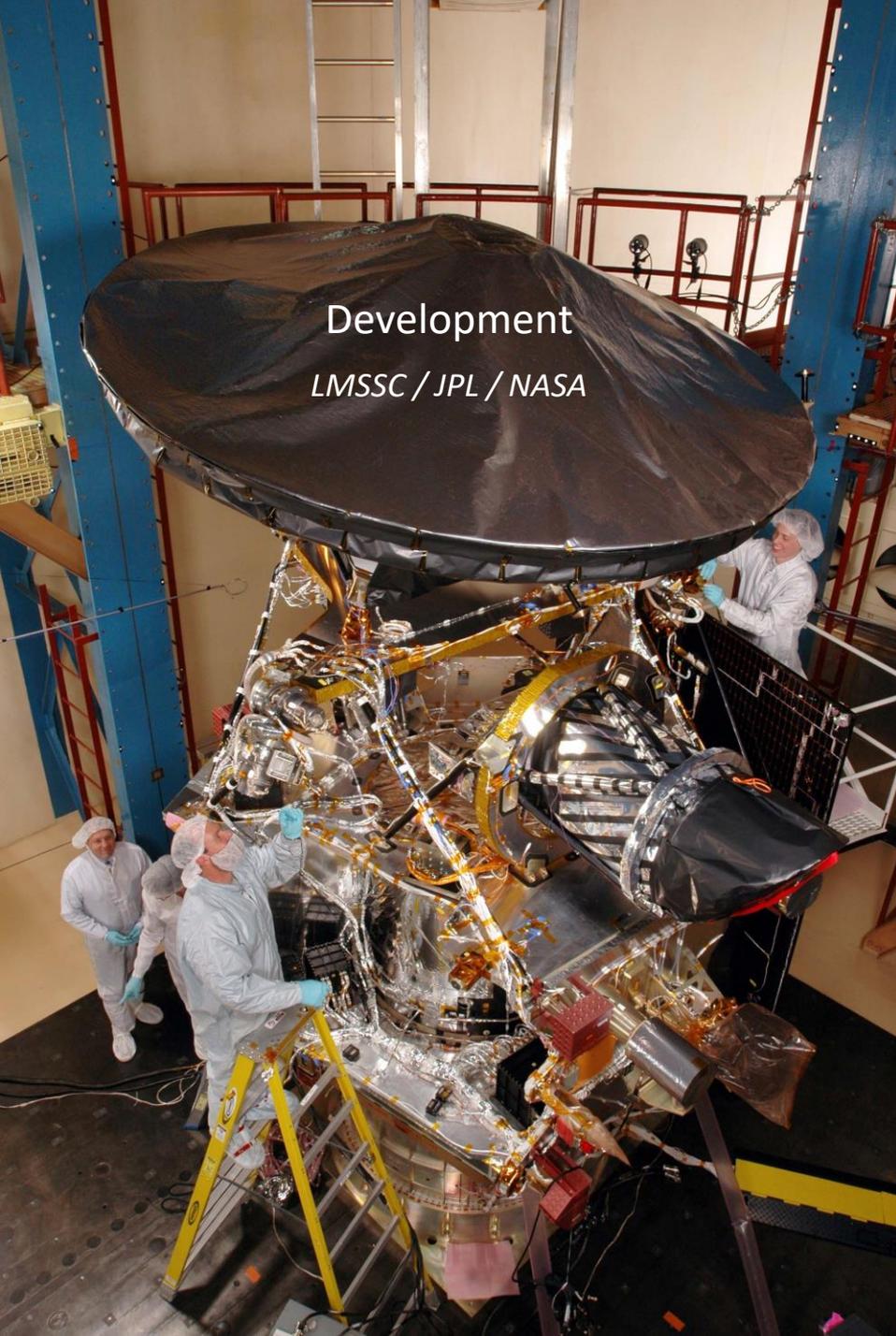
*Precise targeting of almost any
small target on Mars within a
2 week planning cycle*

10 m radar antenna

Payload:
6 Science Payloads
2 Eng Payload
Electra Eng SS
Simultaneous Operations
Nested Targeting

*All instruments still
operating after 10 yrs*

Command/Data Handling:
RAD750 FPC
160 Gbit SSR
100 Mbps Science I/F

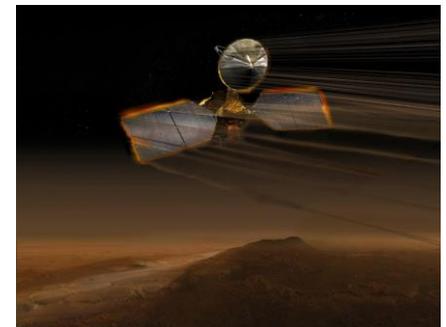


Launch
August 12, 2005

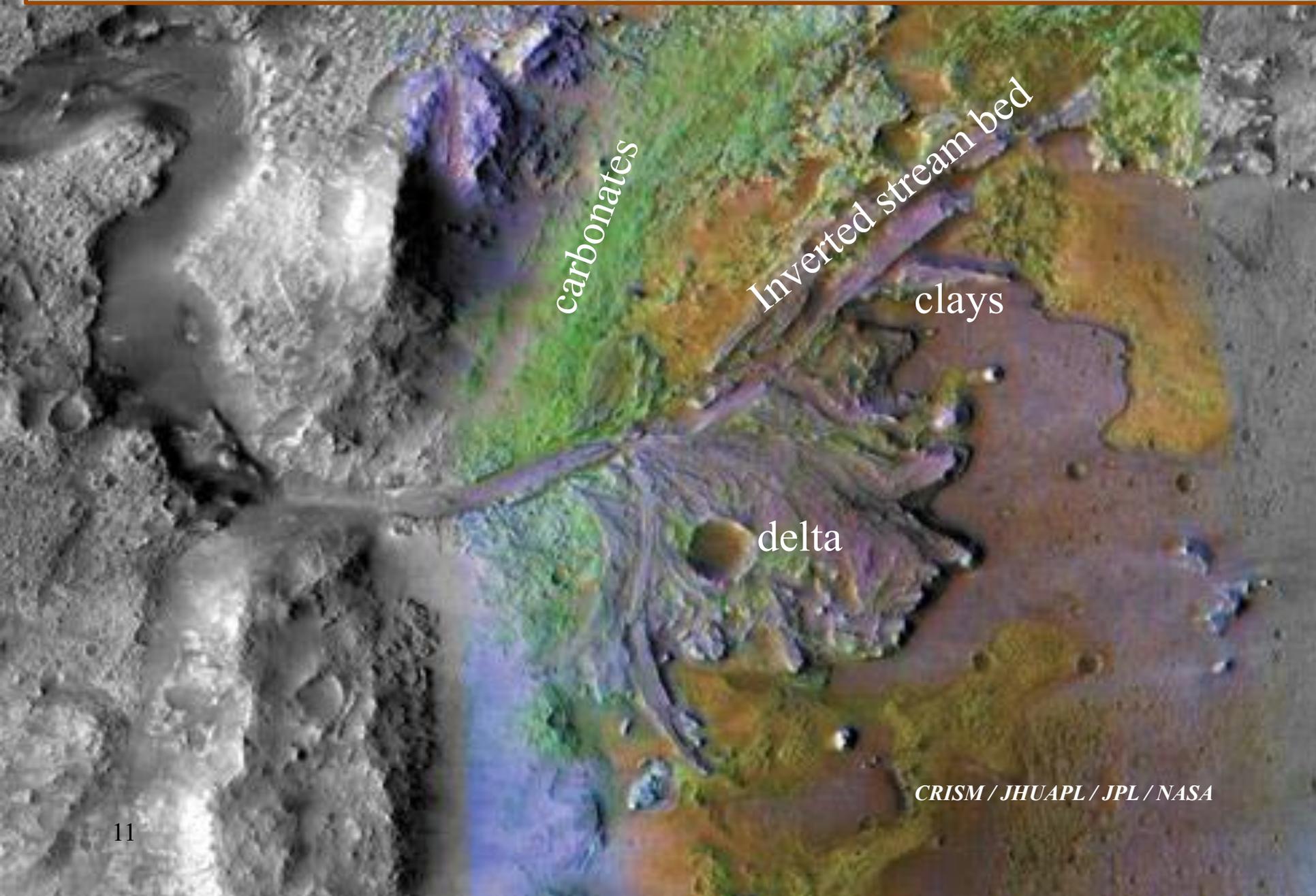


Orbit Insertion
March 10,
2006
(artist concept)

Aerobraking
March -August,
2006
(artist concept)

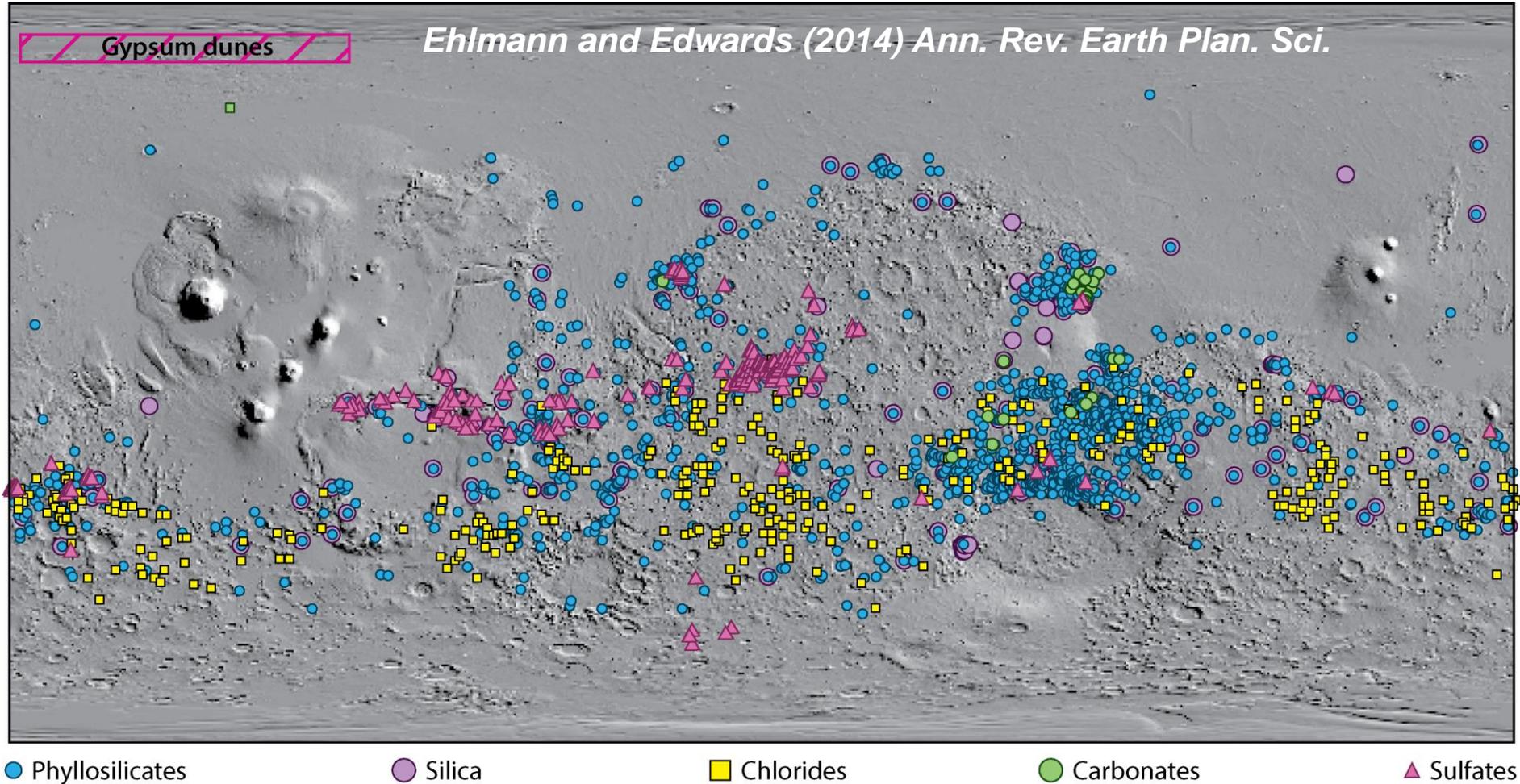


Ancient Mars: A Diversity of Water-Related Environment



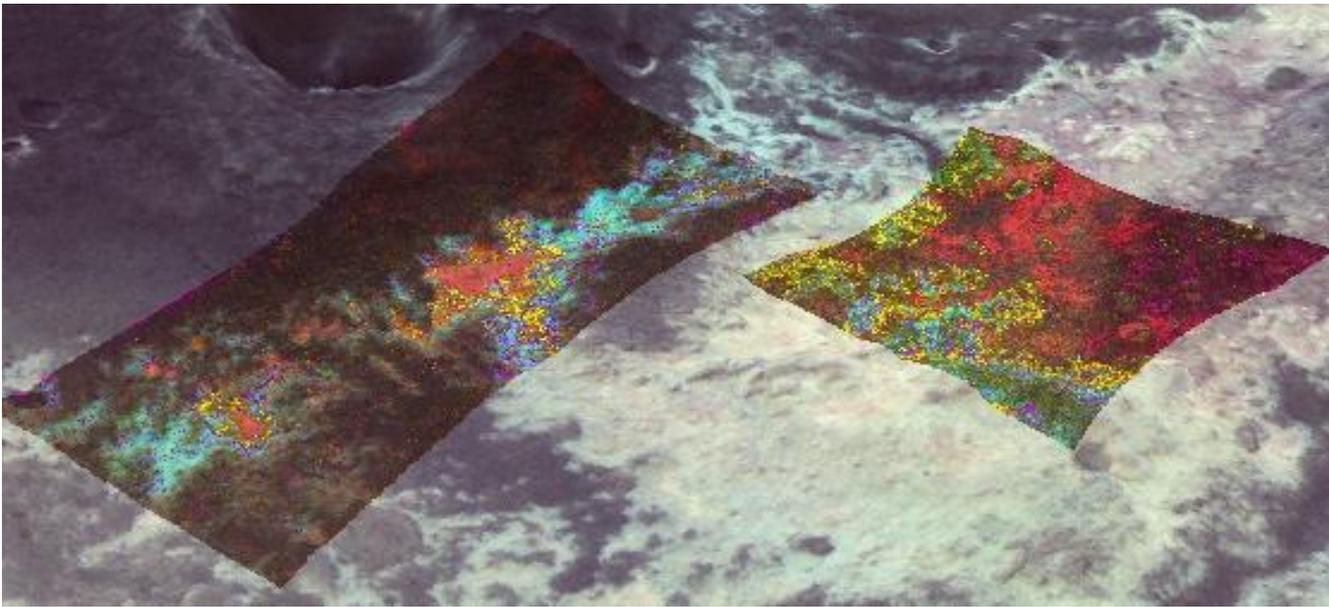
CRISM / JHUAPL / JPL / NASA

MRO CRISM, MEX OMEGA, and ODY THEMIS have mapped thousands of outcrops of minerals that must have formed in liquid water



Their sequence in time tells the story of Mars' habitability

Layered Phyllosilicate Deposits



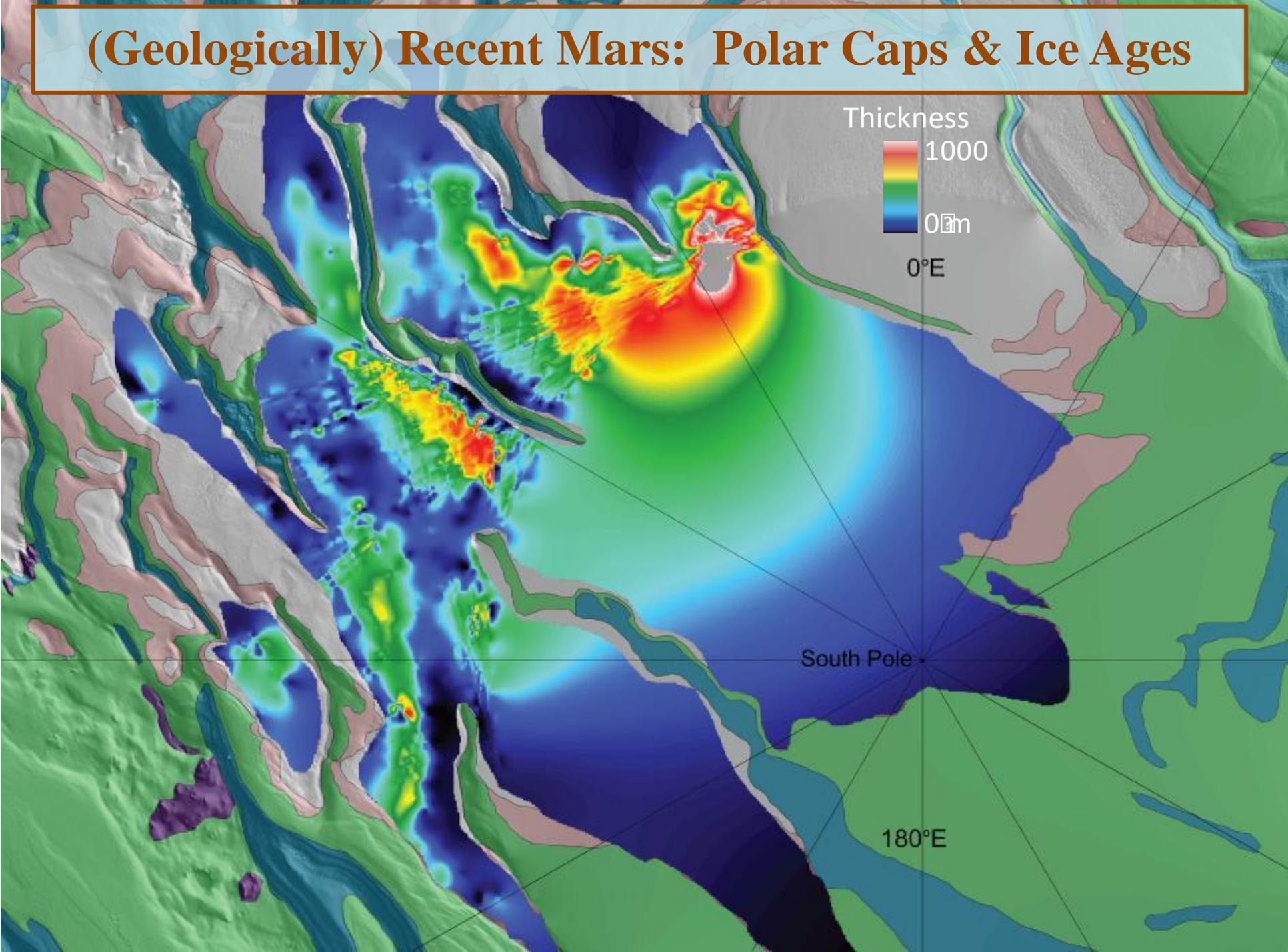
Vertical Sequence:

- ❖ Fe/Mg-clay in differing oxidation states
- ❖ Acid-altered clay
- ❖ Al-clay + silica
- ❖ Amorphous material



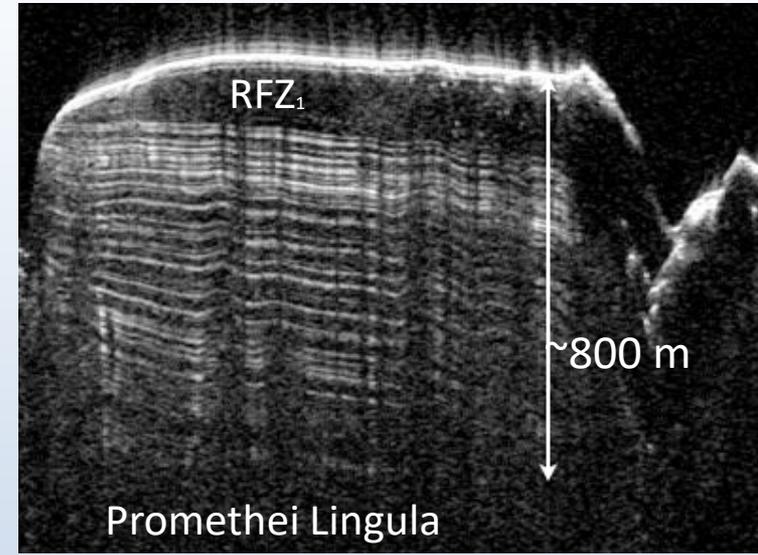
- This vertical sequence is consistent with weathering of basalt or ash to clay, then changing oxidation conditions and increasing acidity
- A lot of water is required implying a brief wet period – age coincides with valley networks and lakes
- ✓ Good candidate for habitability: redox gradients, water, pH moderate at least early on. Plus, clay minerals can serve as a substrate for pre-biotic chemistry and can fossilize microbes.

(Geologically) Recent Mars: Polar Caps & Ice Ages

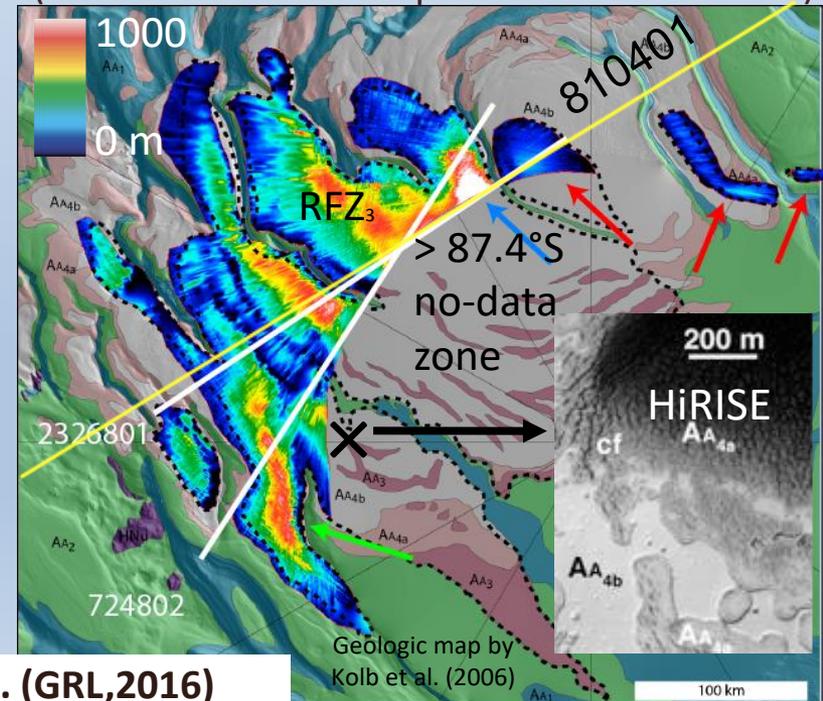
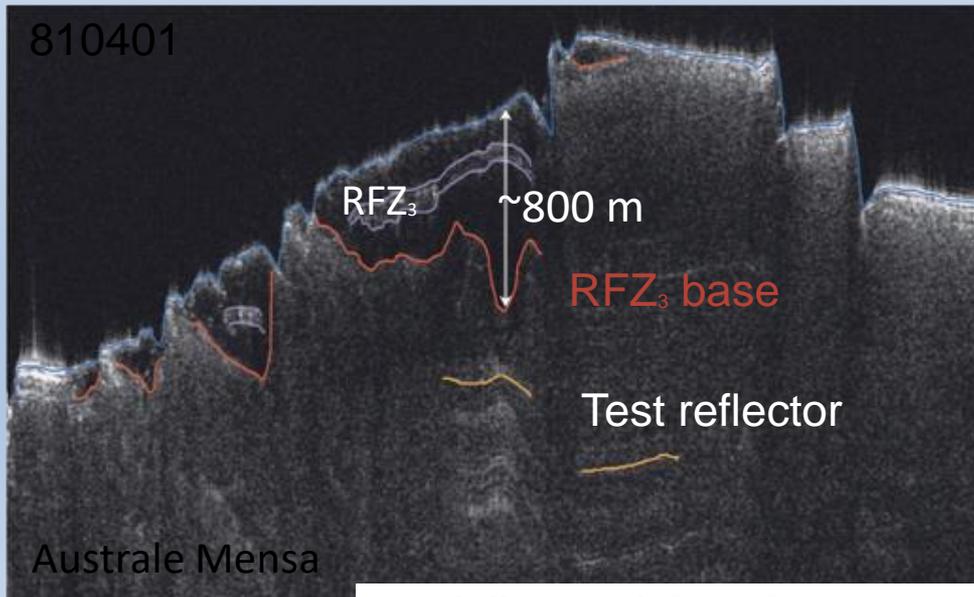


South Polar Buried CO₂

- Layering is discontinuous, sometimes truncated just below the surface.
 - » **Likely much older than NPLD.**
- Odd reflection-free zones (RFZs) occur near surface. In Australe Mensa, **RFZ₃ has dielectric properties of CO₂**, maps to a unit with CO₂ sublimation features.

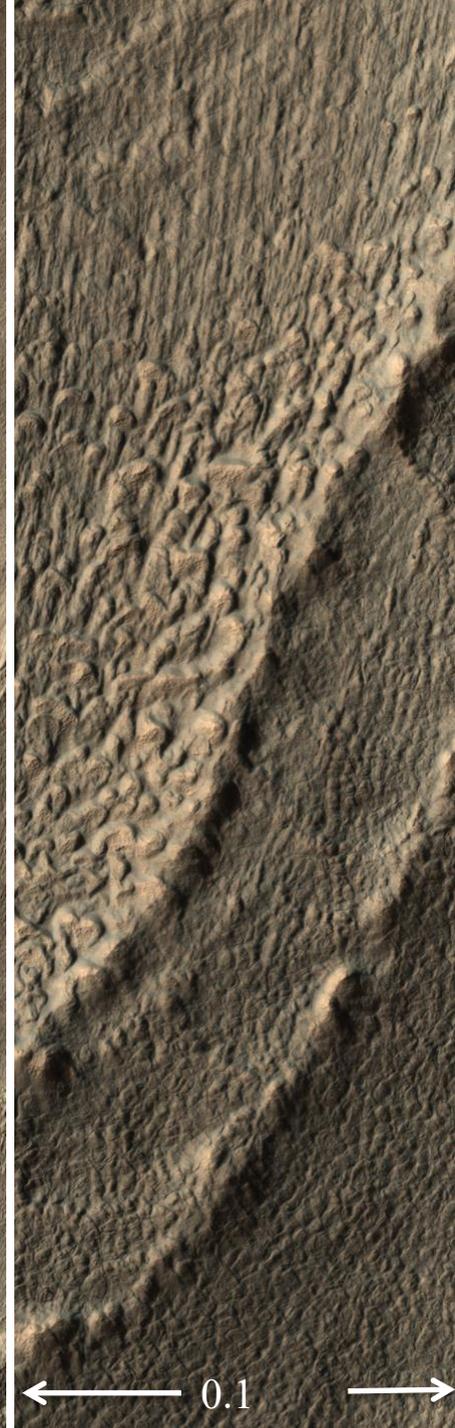
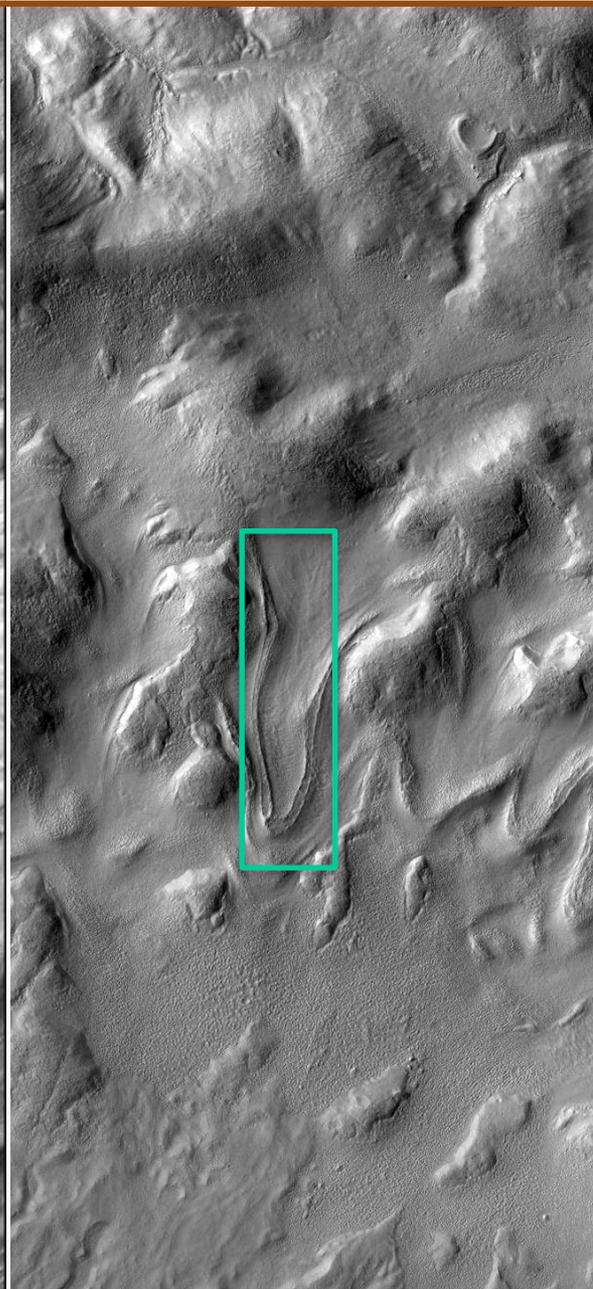


Amount > 6 mbar of CO₂!
(> 2 X current surface pressure if sublimated)



Phillips et al. (2011); Bierson et al. (GRL,2016)

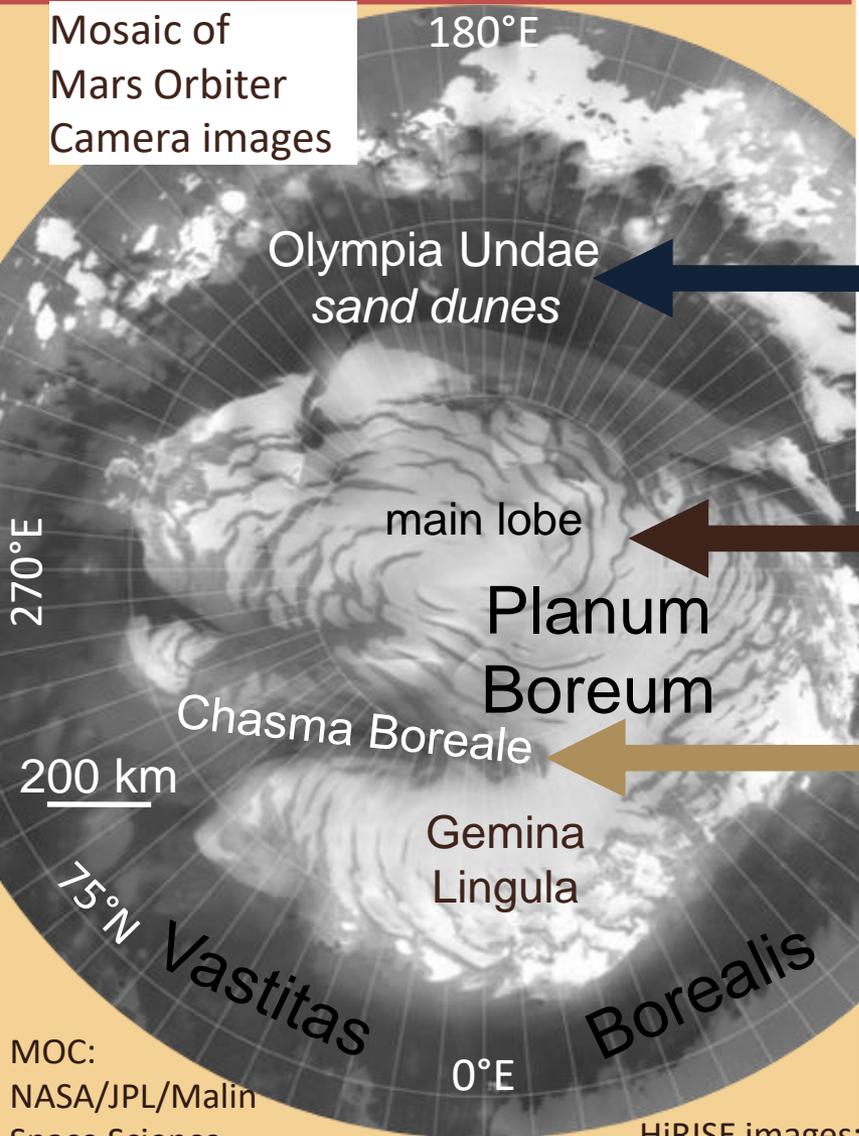
Recent Mars: Mid-latitude Ice



← 0.1 →

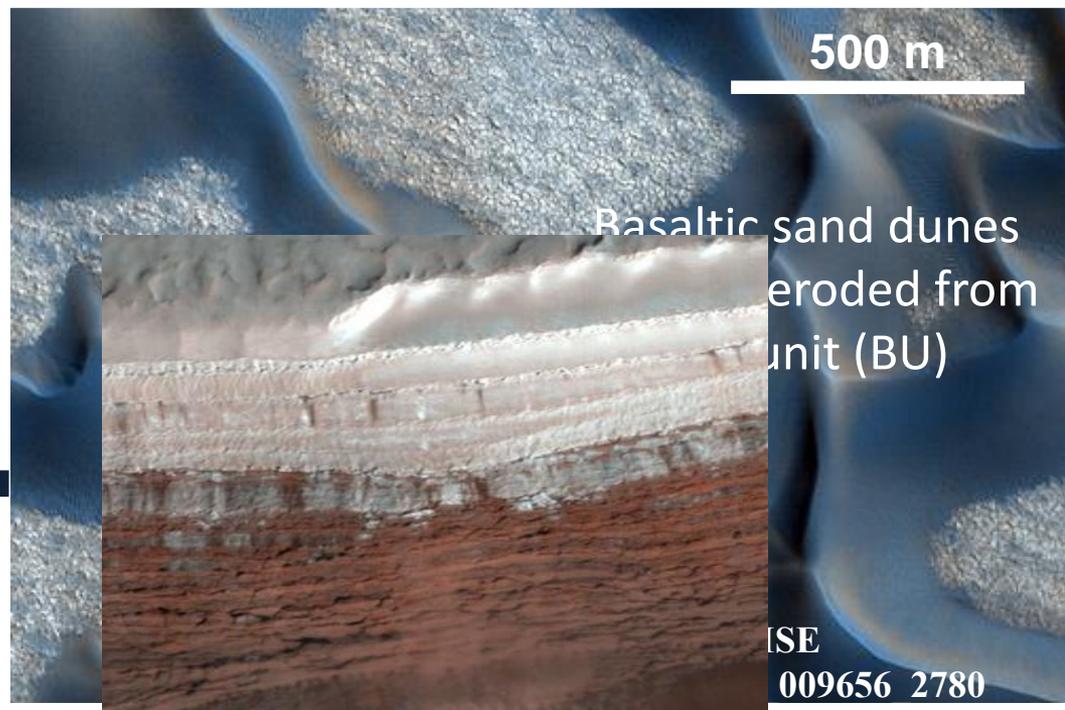
The North Polar Region of Mars

Mosaic of Mars Orbiter Camera images



MOC:
NASA/JPL/Malin
Space Science
Systems

HiRISE images:
NASA/JPL/University of Arizona



HiRISE / U. Arizona / JPL / NASA



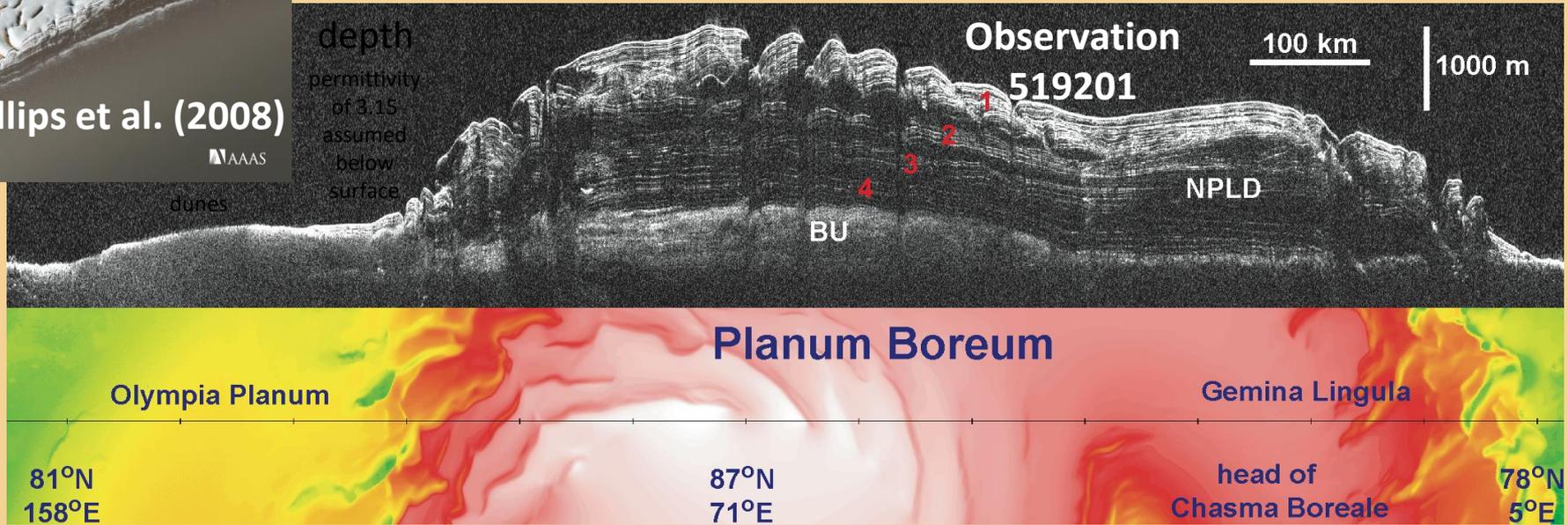
Phillips et al. (2008)

AAAS

dunes

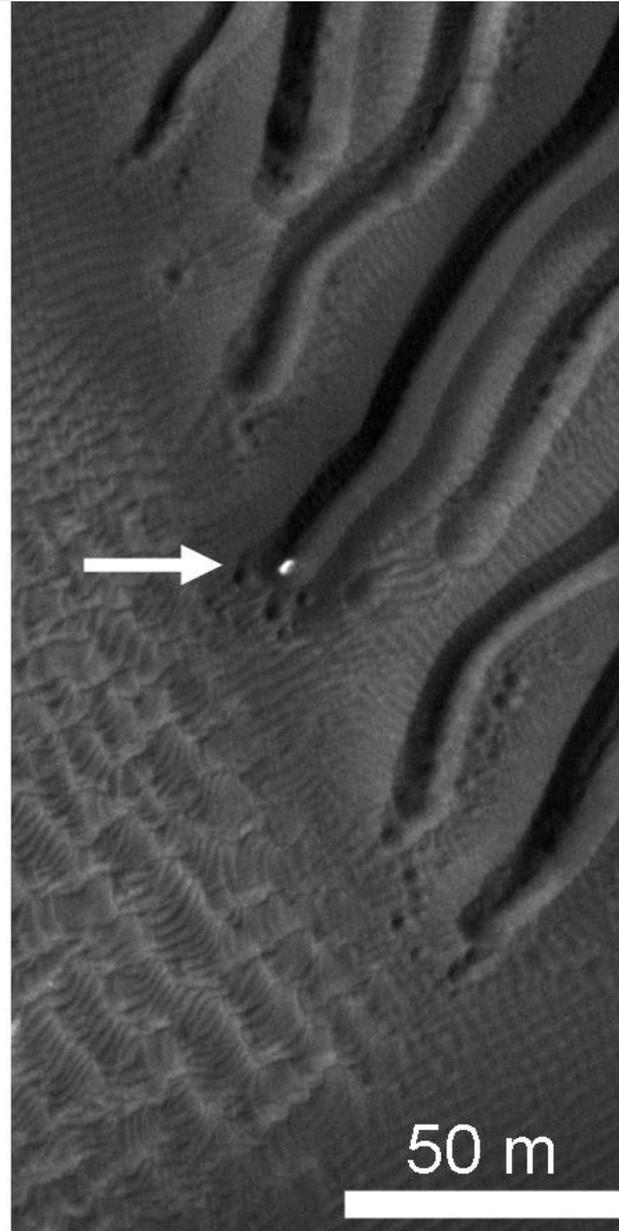
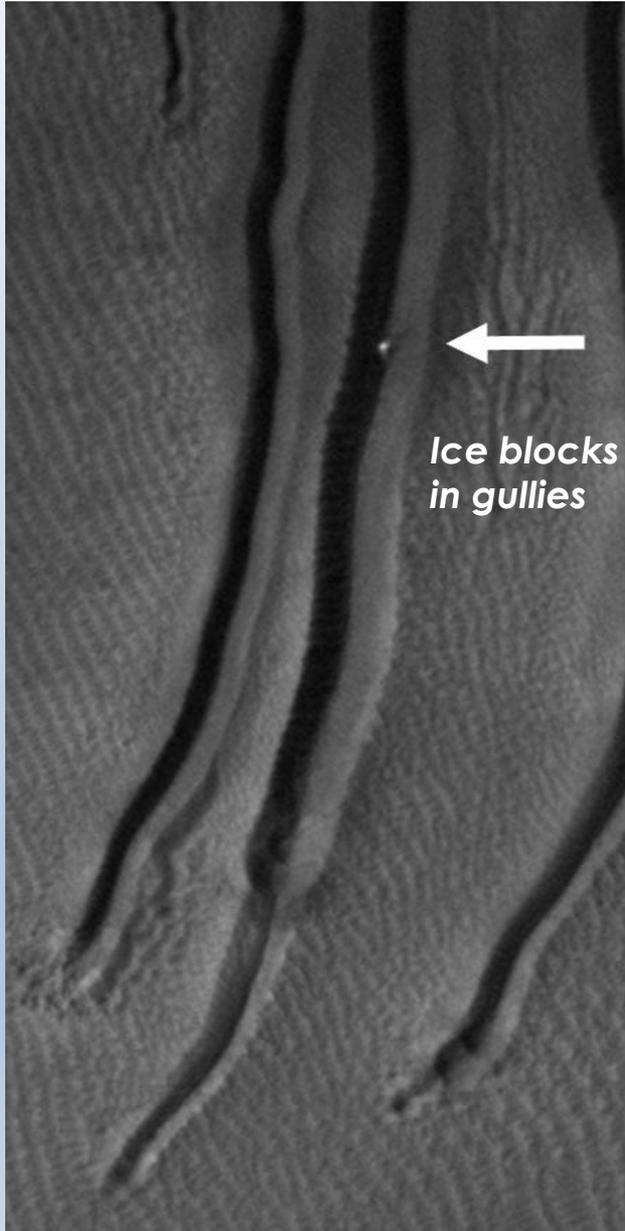
depth
permittivity
of 3.15
assumed
below
surface

In 2007, SHARAD began revealing the internal structure of the NPLD



- Strong basal returns imply relatively pure ice (<~5% rock/dust)
- NPLD layer-packet structure likely related to climate
- Older basal unit rarely layered, missing below >1/3 of NPLD
- **Flat basal boundary has ~ 0 flexure ⇒ very low heat flow**

Recent Mars: Changing Landforms

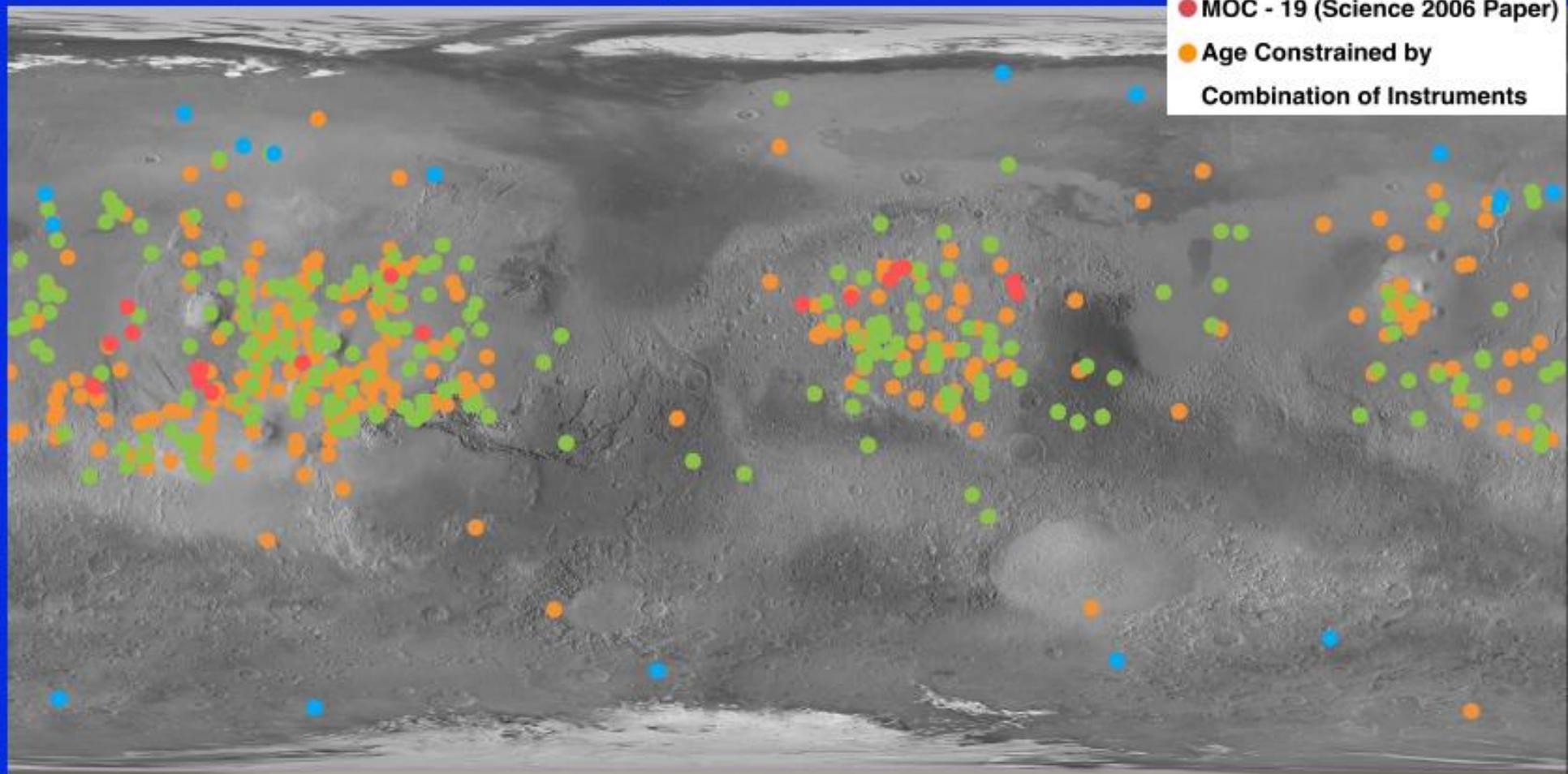


Linear
gullies on
sand dunes
from sliding
blocks of
CO₂ ice
(Dinięga et al.,
2013)

Key:

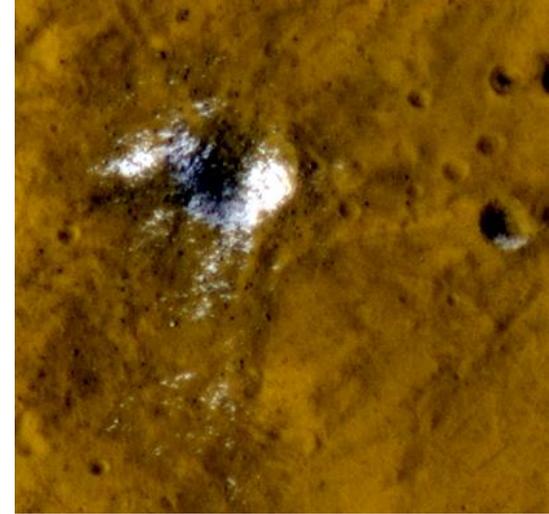
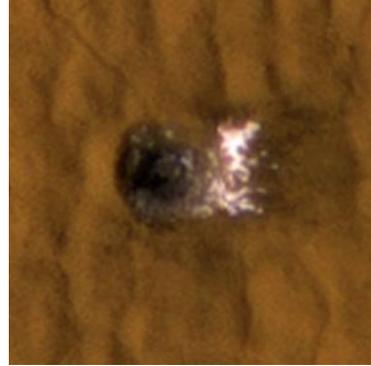
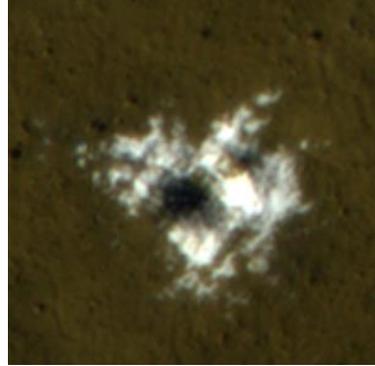
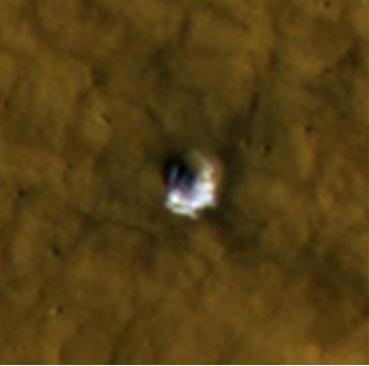
- Icy
- CTX - CTX Age Constrained
- MOC - 19 (Science 2006 Paper)
- Age Constrained by
Combination of Instruments

New Impact Site Distribution



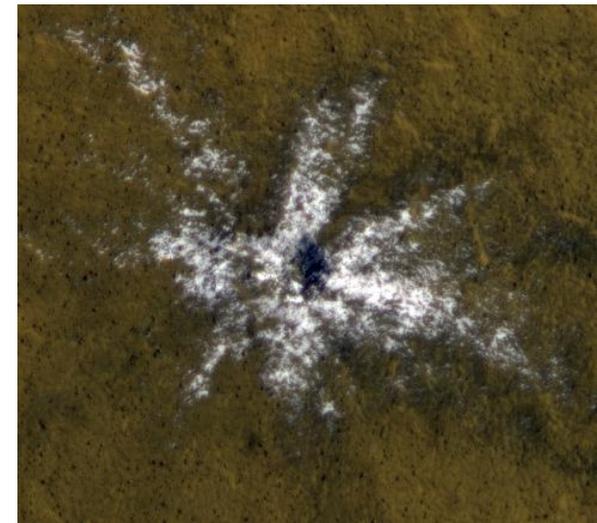
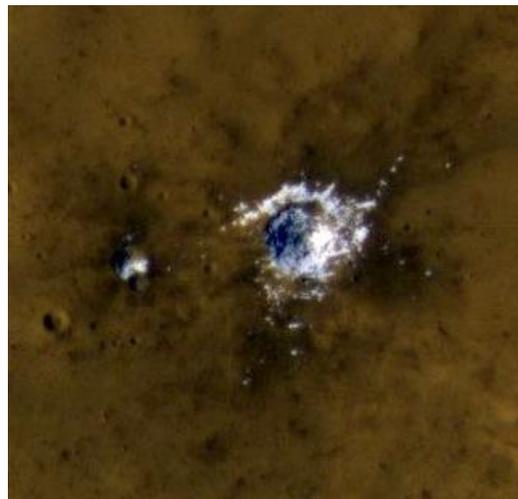
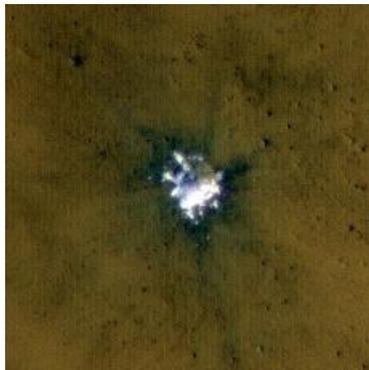
547 candidate impact sites found by CTX
489 impacts confirmed by HiRISE so far

CTX_10/1515_MCM: 13



Ice Exposed by New Impacts

- New impacts north of 39 N have distinct ice patches.
- Distribution and depth of ice is consistent with model calculations for an atmosphere slightly more humid than present.
- Image sequences show that the ice fades over time.
- Sublimation modeling indicate that mm of ice sublimate before it fades, indicating clean ice.
- Ice is a critical resource to possible human explorers of Mars.

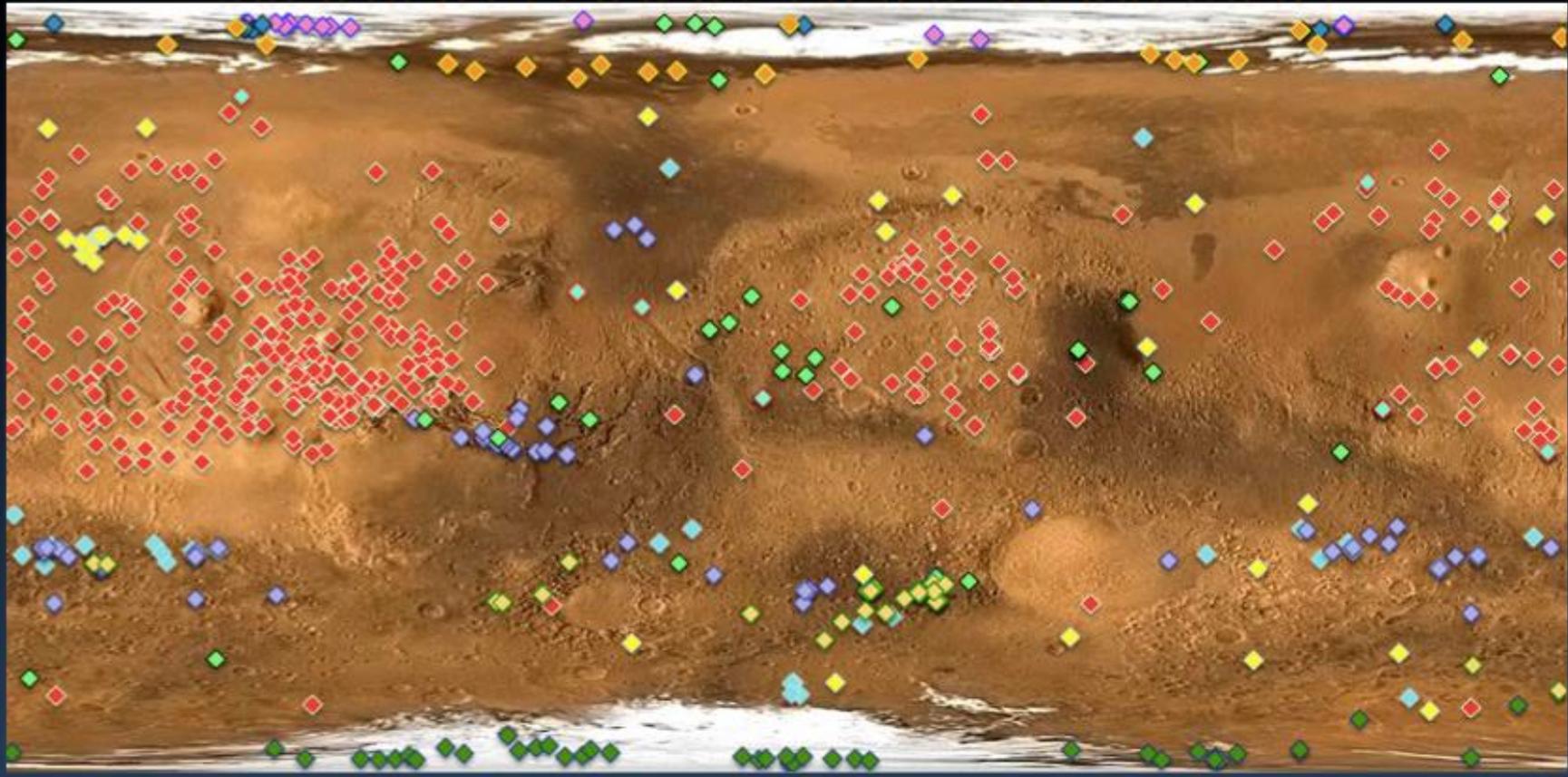


Many Martian Landscapes in Motion

(monitored by MRO/HiRISE)

Martian Landscapes in Motion

- Show All
- Active Gullies
- Avalanches
- Dunes
- Dune Gullies
- Dust Devils
- New Impacts
- Polar Pits
- RSL
- Robots
- North Polar
- South Polar
- Hide All

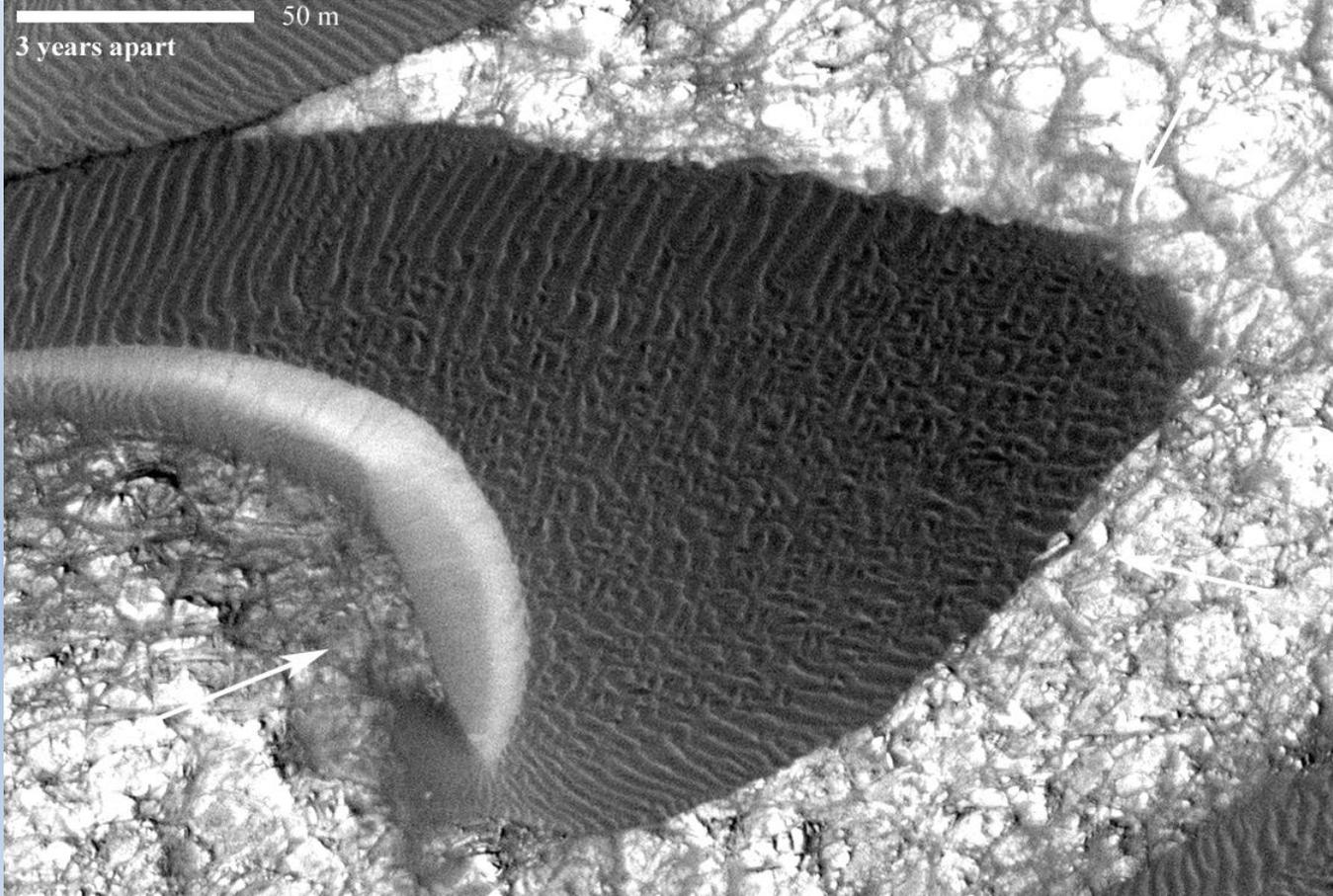
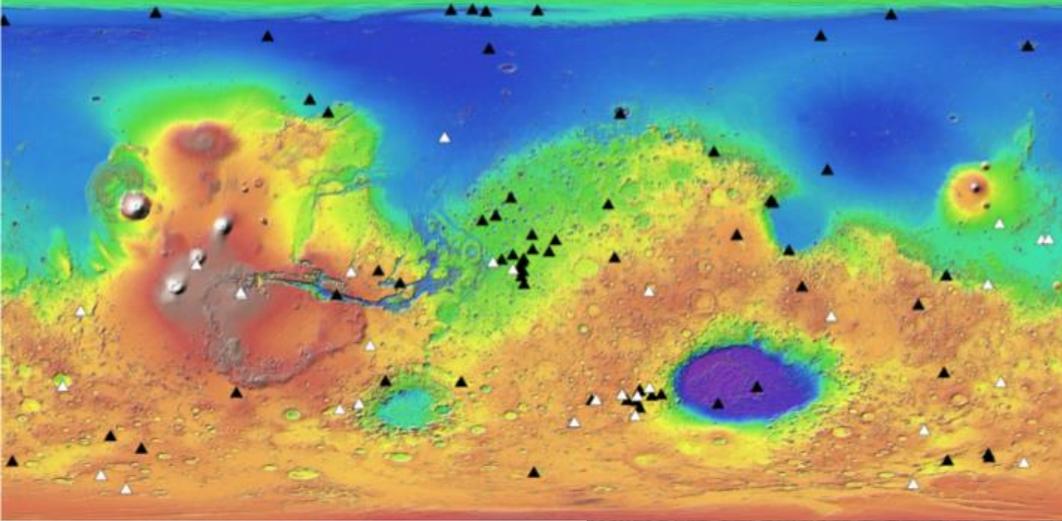


Windy Mars

Dune motion observed globally

Black = active

White = no activity observed (yet)



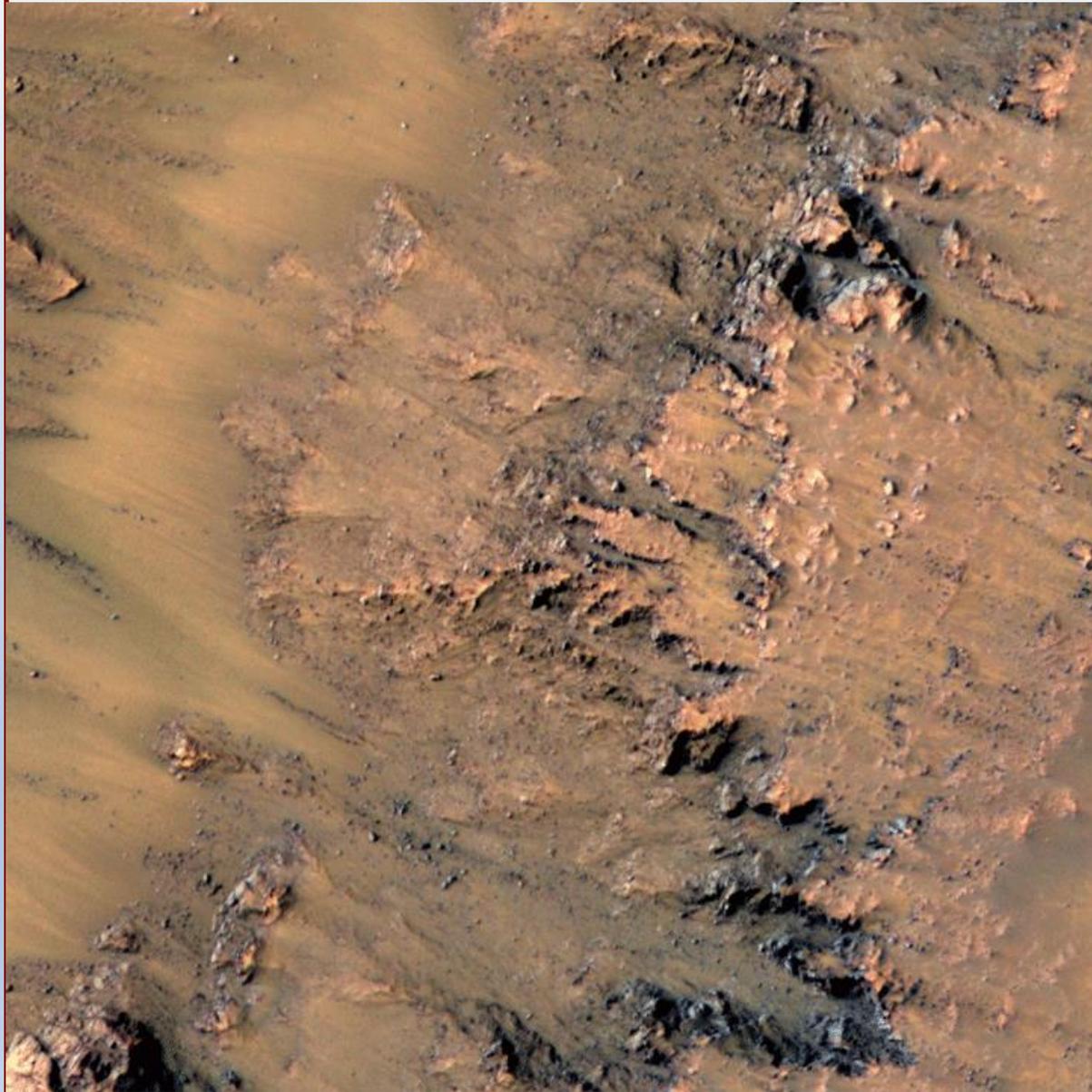
Dune
Movement

Recurring Slope Lineae (RSL)

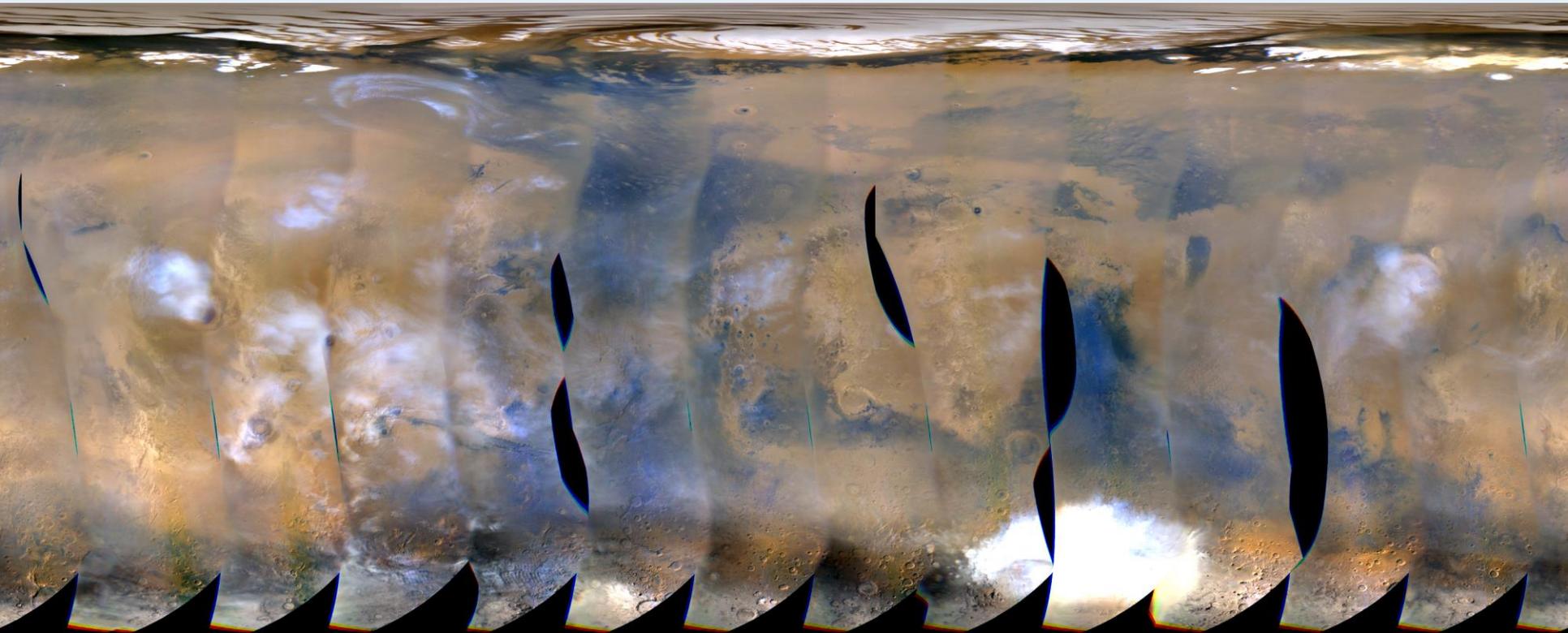
- RSL are narrow (0.5-5 m), dark markings on steep slopes ($>25^\circ$)
- Form and incrementally grow in late spring to summer, then fade or disappear in fall.
- Reform at nearly same locations in multiple Mars years.
- Extend downslope from bedrock outcrops or rocky areas; often associated with small channels.
- Form and grow at temperatures at which brines would be liquid.
- CRISM has detected hydrated perchlorates
 - Deliquescent -> easily takes up water from atmosphere
 - Depresses freezing point to -70 C
- However, Ody sees no effect on thermal inertia => little water
- Exact mechanism for RSL is not understood, but activity of brines is still current best model.

Seasonal Flows on Warm Martian Slopes

(McEwen et al., Science, 5 Aug 2011)



Recent MARCI global weather map

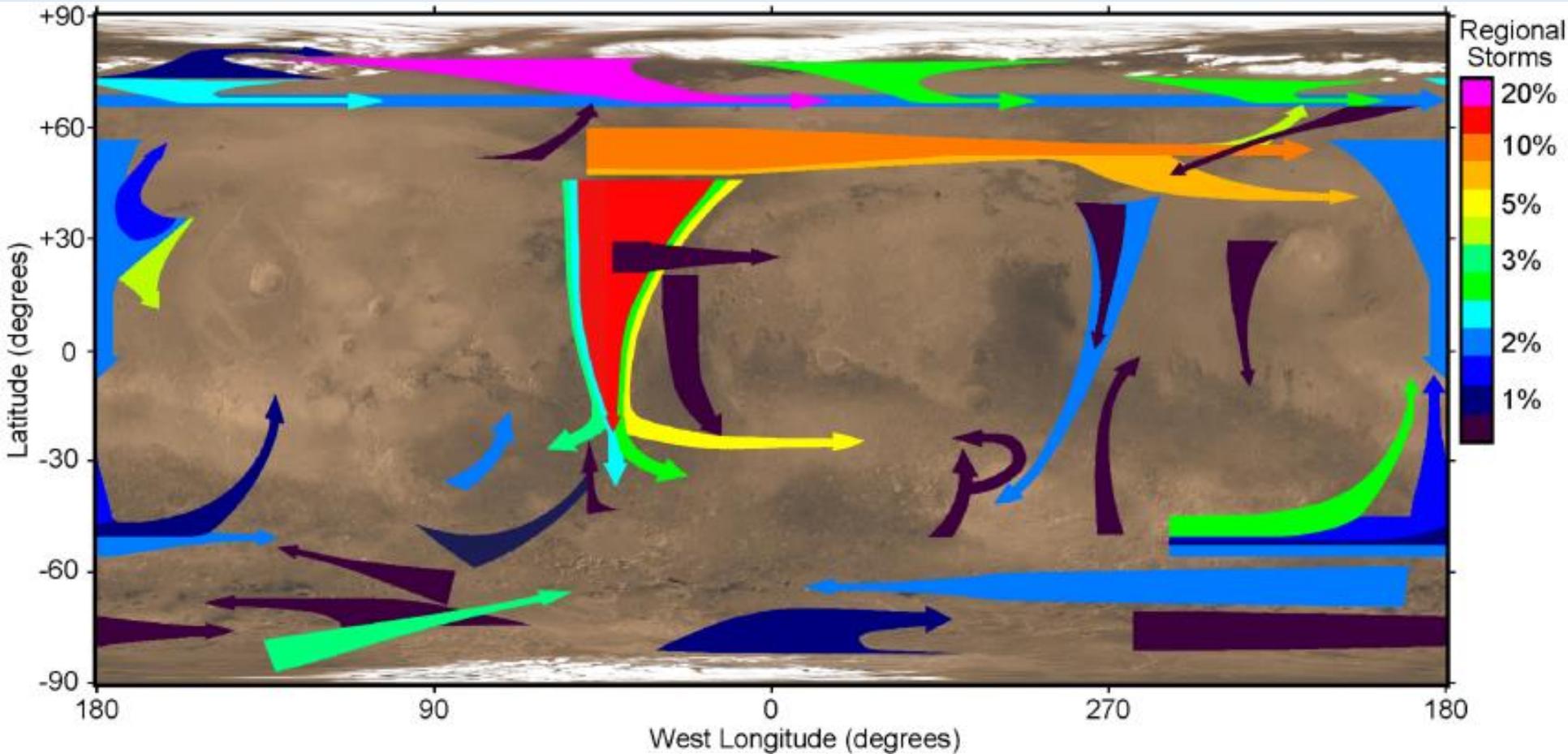


MARCI / MSSS / JPL / NASA

March 7, 2016 ($L_s = 119.1$)

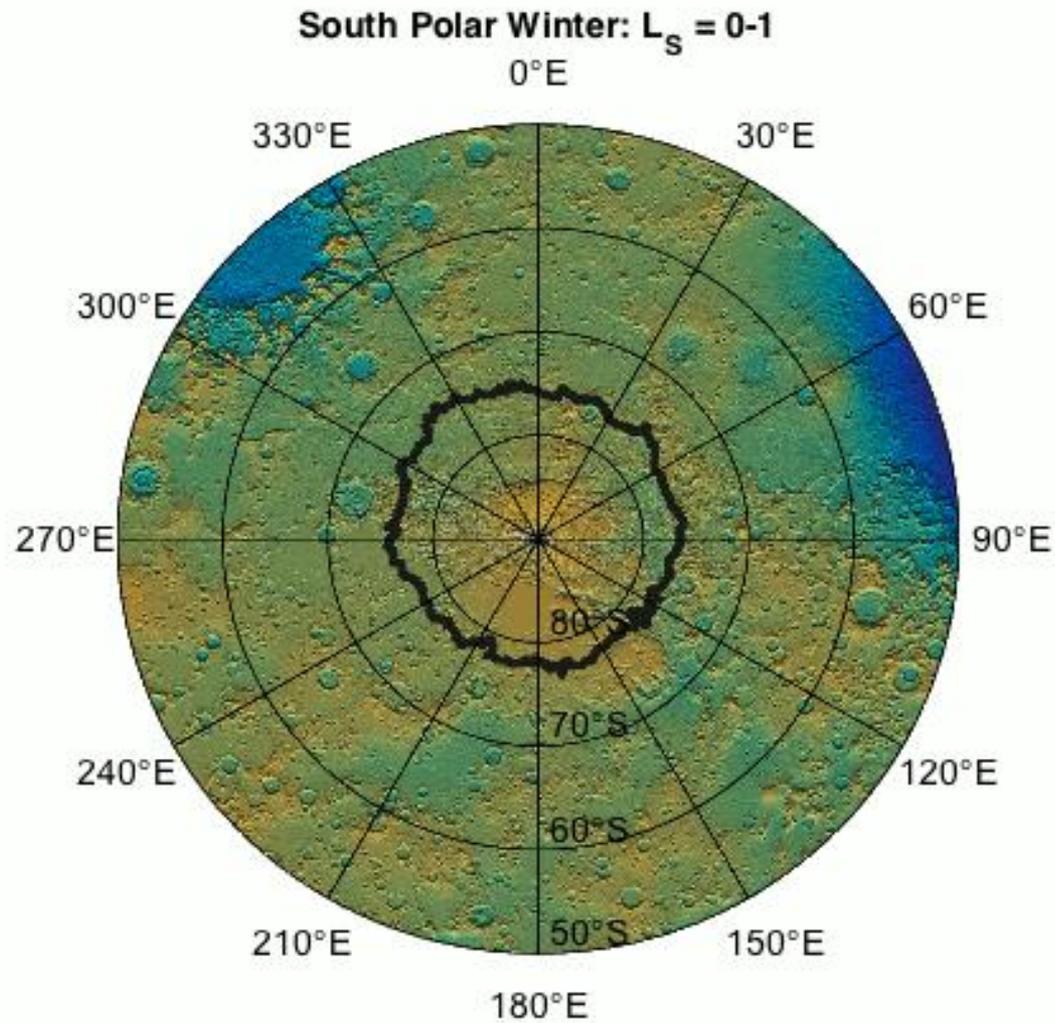
Primary Dust Storm Tracks

MARCI / MSSS / JPL / NASA



CO₂ Snow Clouds and Surface Snow: South Pole

Mars Climate Sounder / Caltech / JPL / NASA



11 Years in Orbit

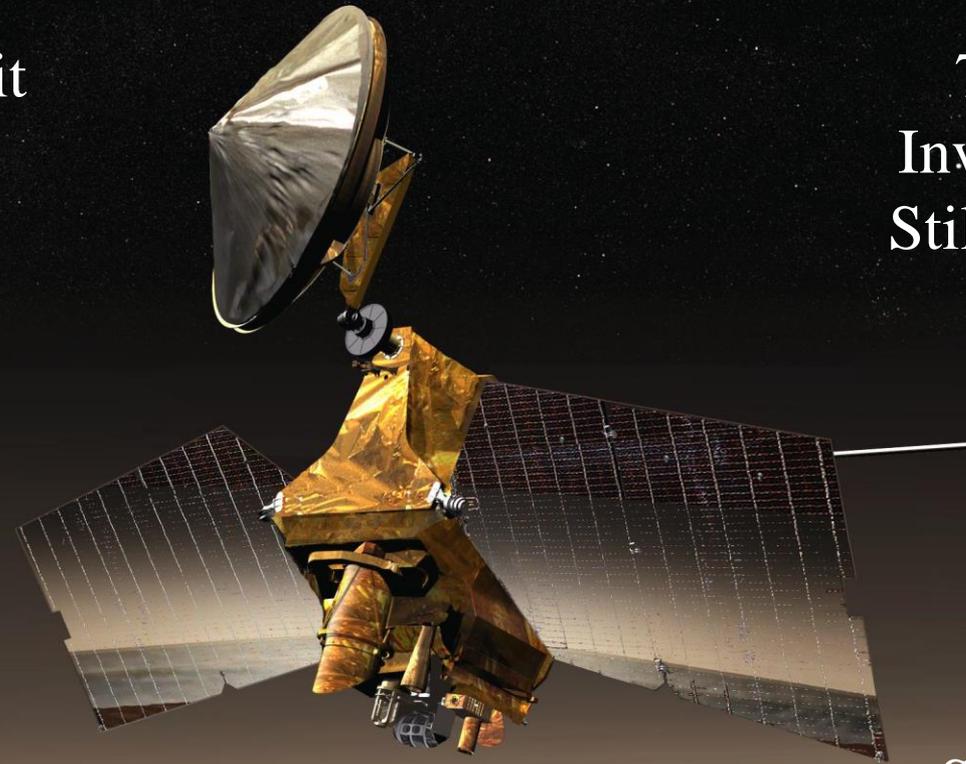
~50,000 orbits

300 Tb of
Science Data
Returned

~200 kg of
Usuable Fuel
still in the Tank

7 Science
Investigations
Still Returning
Data

More
Discoveries
Sure to Come!



**Mars
Reconnaissance
Orbiter**

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Government sponsorship acknowledged.