



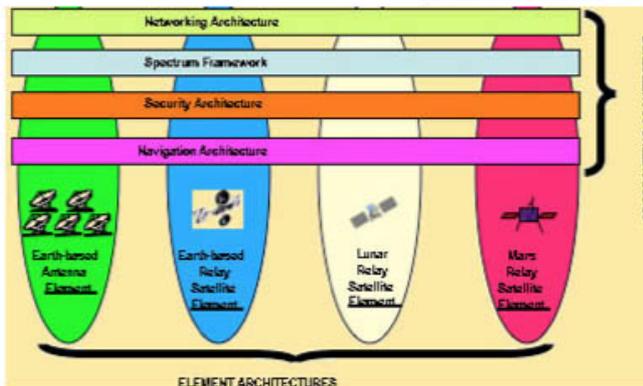
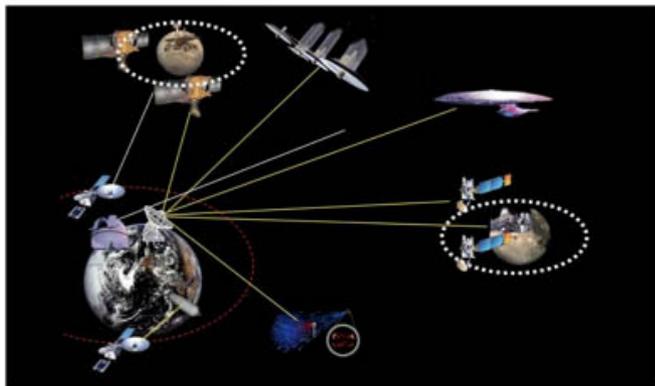
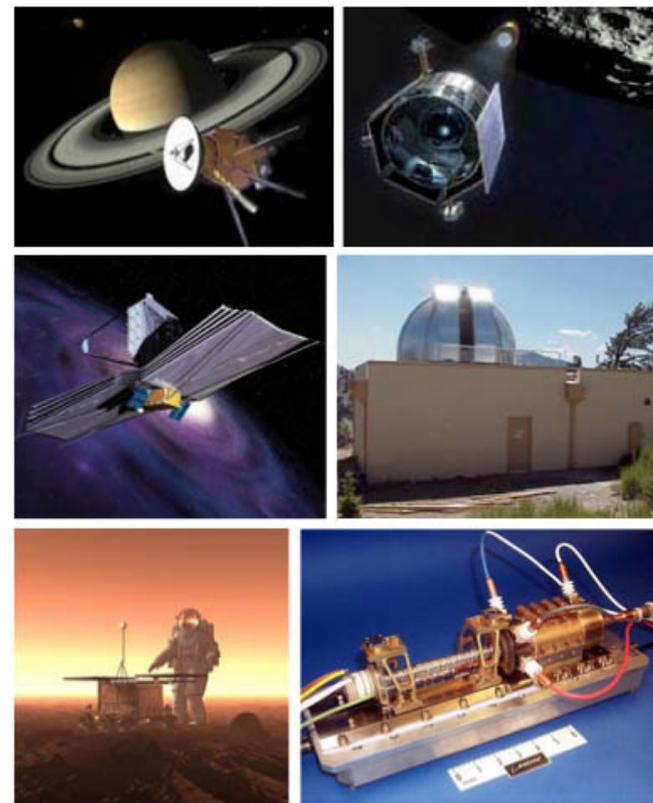
Advanced Communications, Navigation and Technology Concepts for a Mars 2018 Orbiter

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California Institute of Technology

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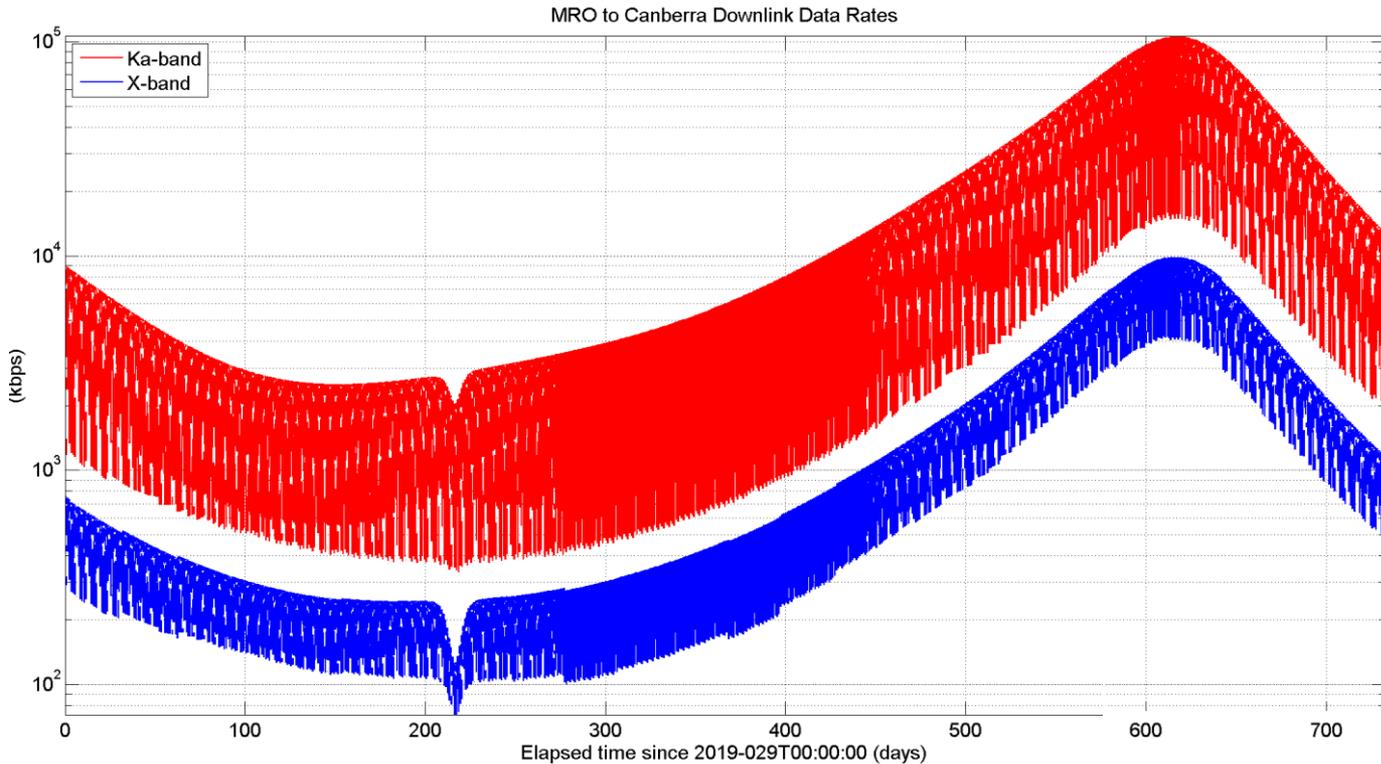
Overview



- **High Rate RF Communications**
 - **Spacecraft System Bottlenecks**
 - **DSN Bottlenecks**
- **Optical Communications Technology Demo**
- **Navigation Technology Demo**
 - **Rendezvous in Mars orbit and related technologies**
 - **Autonomous aerobraking**
 - **Δ DOR at Ka-band**
 - **One-way radio metric tracking/Deep Space Atomic Clock**
 - **Ranging at optical frequencies**
- **Disruption Tolerant Networking (DTN) Demo & Avionic Requirements**
 - **Flight segment**
 - **Ground segment**
- **Summary**



Orbiter to Canberra Data Rate Spans (2 Years) (X-band and Ka-band)



- Assumptions:**
- Science operations from 01/29/19 – 01/29/21.
 - One 8-hour pass / day
 - DSN 34-m Beam Wave Guide (BWG) antenna
 - Coverage model for celestial body dynamics, spacecraft orbits and station locations
 - RF model for signal / noise environment (link quality)
 - End-to-end model for protocols and routing of network topology.

- **Data rates are a strong function of the Mars–Earth range, which varies widely over the two-year synodic period.**
- **Downlink data rates at Ka-band are typically 10 times those at X-band**

Average Data Rate (kbps)

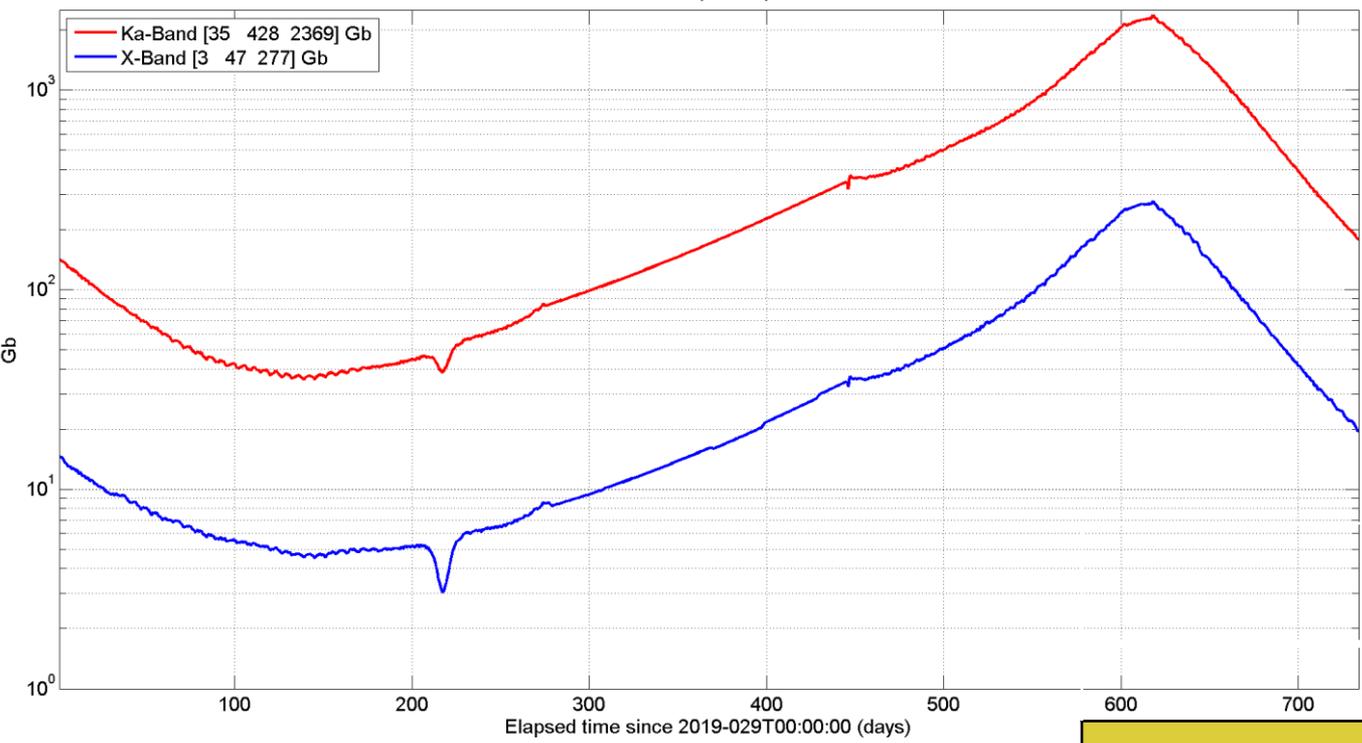
	Min	Mean	Max
X-Band D/L	115.5	1756.5	9060.3
X-Band U/L	315	2604	13169.7
Ka-Band D/L	2127	22225	110505



Orbiter to Canberra Data Volumes (2 Years) (X-band and Ka-band)



MRO to Canberra Day-to-Day Data Volumes



Assumptions:

- Same as previous.

- **Data volumes are, again, a strong function of the Mars–Earth range.**
- **Downlink data volumes at Ka-band are typically ten times those at X-band**

Daily Data Volume (Gb)		
Min	Mean	Max
X-Band D/L	3.8	344
X-Band U/L	9.9	505.0
Ka-Band D/L	68.7	4140



High-Rate RF Communications & Spacecraft System Bottlenecks



- **Current Command & Data Handling (C&DH) avionics limited to ~6 Msps throughput**
 - Limitations will have to be relieved to realize Ka-band data rates
- **Current Solid State Recorders (SSR) have a capacity of ~160 Gb.**
 - Will have to increase to accommodate daily Ka-band data volumes
- **Current radio of choice is the *Small Deep Space Transponder (SDST)*:**
 - Frequencies: Deep space S- or X-band, Ka-band
 - Uplink: 4 kbps (uncoded)
 - Downlink: 6 Msps (typically coded)
 - Encryption/decryption: None
 - Operates at the required frequencies
 - Can support lower downlink rates for links at large Mars-Earth range
 - Becomes a flight-side bottleneck at closer ranges on the downlink
 - Can only receive a very limited uplink rate (<< X-band link can support)
- **Next-Gen radio (albeit notional) is the *Universal Space Transponder (UST)*:**
 - Frequencies: Deep space S- or X-band, Ka-band
 - Uplink: 26 kbps (uncoded) up to 6 Msps (coded)
 - Downlink: ≥ 80 Msps (typically coded)
 - Bandwidth Efficient Modulation: Quadrature Phase Shift Keying (QPSK); Offset-QPSK or Gaussian Minimum Shift Keying (GMSK)
 - Encryption/decryption: Advanced Encryption Standard (AES)



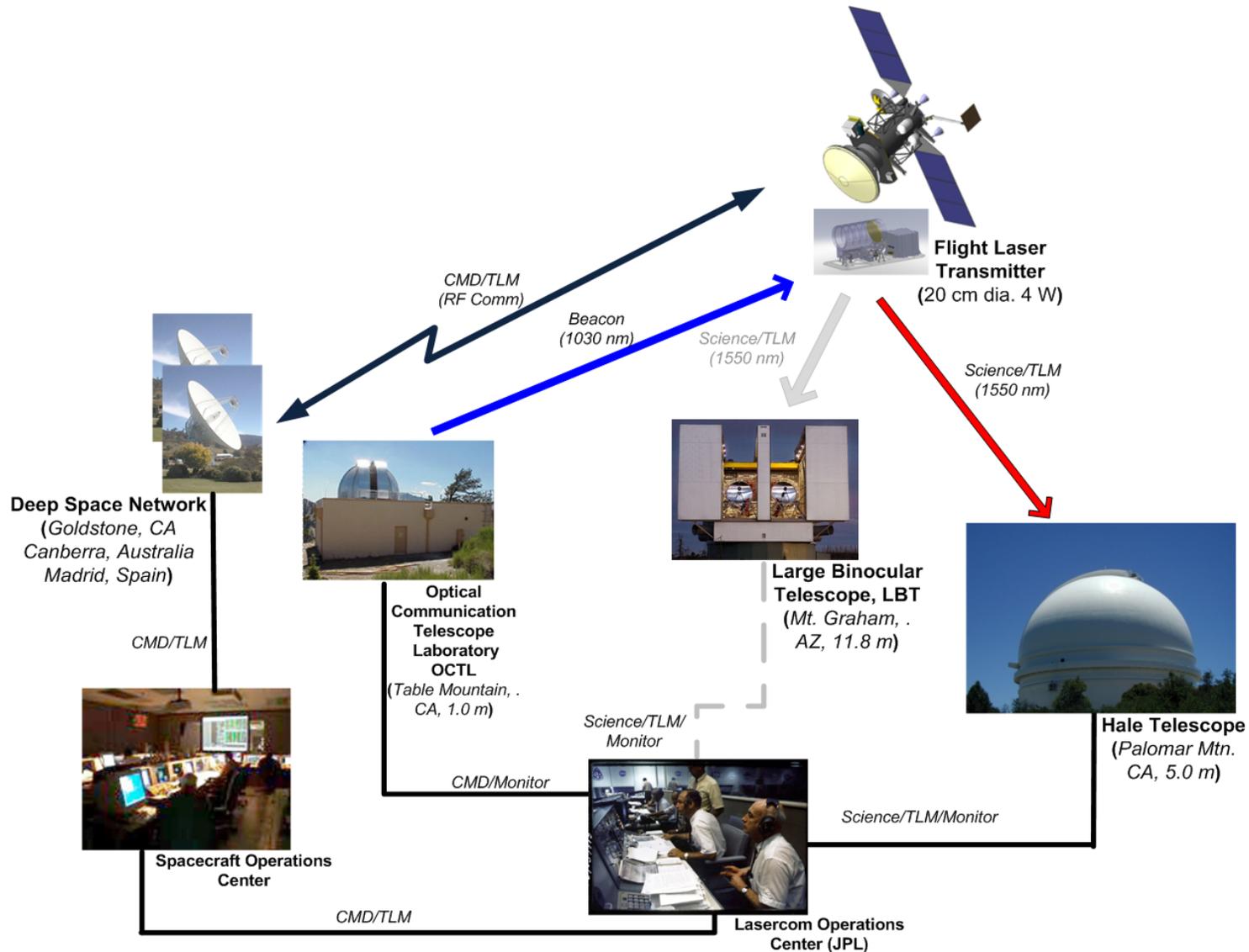
High-Rate RF Communications & Ground System Bottlenecks



- For downlinks ≤ 26 Msps, current system throughput is sufficient
 - To triple the downlink rates (as seems likely) augmentation is in order
- Deep Space Receiver/Telemetry Processor will need increased capacity
 - High-rate commercial receiver similar to the Near Earth Receiver/Telemetry Processor
 - Next generation wideband high-rate receiver being prototyped at JPL
- Demodulators and decoders
 - Missions currently utilize BPSK/QPSK with convolutional, Reed–Solomon or Turbo codes
 - DSN currently supports convolutional and Reed Solomon codes at high data rates
 - Turbo codes have more gain – but are complex and hard to decode (a few Mbps limit)
 - Low Density Parity Check (LDPC) codes enable high rate decoding with high coding gain
 - At higher rates, channel allocations become filled requiring bandwidth-efficient modulation.
 - A likely option is Gaussian Minimum Shift Keying (GMSK)
 - New modulation and decoding techniques are recommended by CCSDS
 - Enhance performance as well as interoperability with other missions or service providers
- Data received at DSN complexes must be sent back to JPL or some other specified destination
 - Commercial leased lines provided via NASA Integrated Services Network (NISN)
 - As data volumes increase, so must the leased line capacity – but no technology challenge
- Above upgrades assume a link between a Mars 2018 orbiter and a single DSN 34m antenna
 - If two antennas are arrayed, then all numbers increase by 2x, and so on . . .
 - Remote sensing spacecraft, with modern instruments, can far exceed downlink capacity
 - Project requirements for arrayed passes can significantly enhance downlink performance



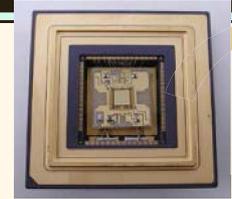
Lasercom Demonstration Architecture



Key Technologies for Deep-Space Demo Under OCT Development

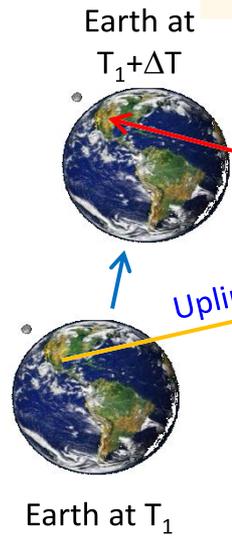


Photon Counting Space Receiver

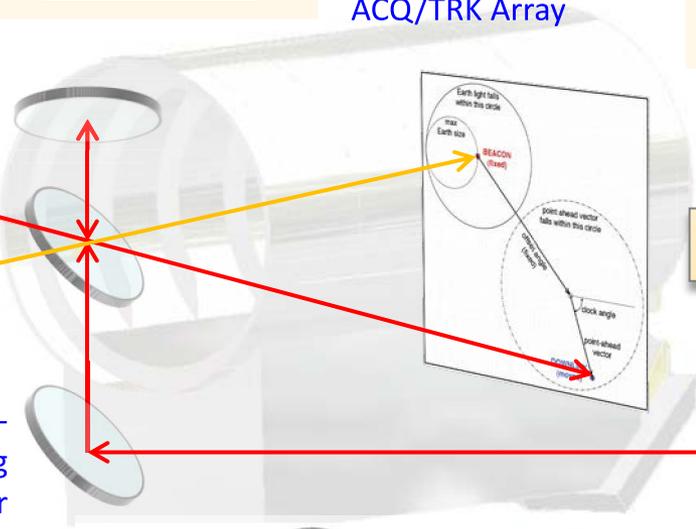


ACQ/TRK Array

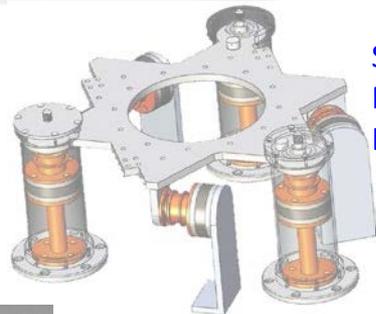
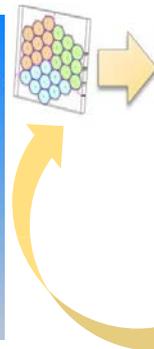
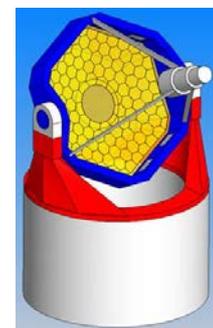
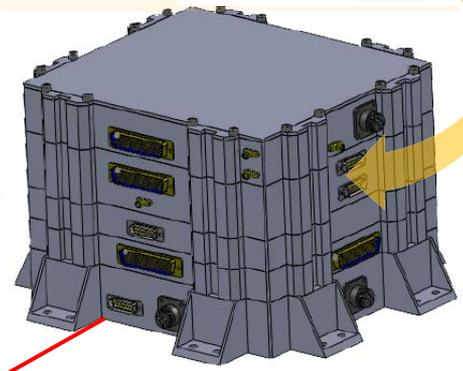
Photon Counting Space Receiver



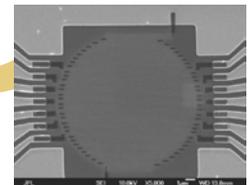
Fine-Pointing Mirror



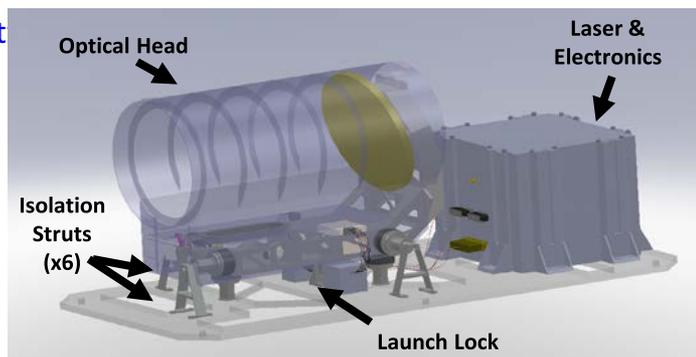
Laser & Electronics Assembly



Spacecraft Disturbance Rejection Plate



Ground Receiver Photon-Counting Detector Array





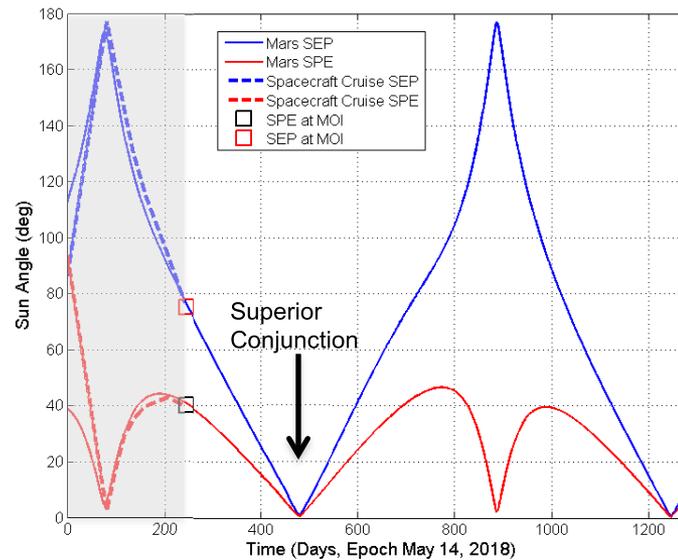
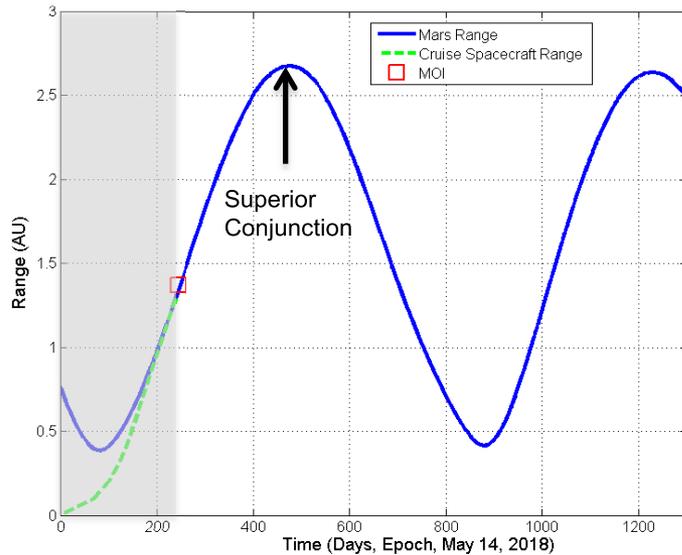
Mars 2018 Optical Comm Demo



- **Wide variety of ranges and deep-space link conditions during cruise (5-months) and following MOI (> 4 years)**
 - Ranges from 0.2-1.37 AU during cruise and 0.42-2.64 AU following MOI
 - **Sun angles**
 - Minimum SEP angle supportable using the Hale telescope is 12-20°
 - Need new build to support operations at 3-5° SEP
 - Flight terminal designed for 2-deg SPE operation

Link Outage Summary

Min SEP	Days of Outage	Min SPE	Days of Outage
20	120	2	19
12	72		
5	30		
3	17		



Diverse link conditions encountered during cruise and orbit phases of mission



Rendezvous in Mars Orbit (and Related Technologies)



- **2018 Mars opportunity could demonstrate proximity operations, as well as rendezvous and docking in Mars orbit, all likely required for eventual Mars Sample Return**
 - Updated version of MRO optical navigation camera could demonstrate autonomous navigation in Mars orbit by imaging Phobos or Deimos against background stars or by tracking landmarks on Martian surface
 - Optical navigation camera might also be able to produce science-quality mapping images of Phobos and Deimos
- **Orbiter could deploy pop-out (university?) cubesat with COTS release mechanism**
 - Orbiter spacecraft would locate cubesat with optical navigation camera, to simulate search for sample capsule; rendezvous & docking would follow
 - Cubesat could demonstrate FEEP (field-effect electric propulsion)



Autonomous Aerobraking



- **Aerobraking atmospheric passes must occur at altitudes such that**
 - Aerodynamic forces or heating rates are within spacecraft design limits
 - Aerodynamic effects are still sufficient to modify the orbit in a timely fashion
- **Given periapsis orbit accuracy requirements, along with the duration of the aerobraking process, an on board means of automating orbit determination and periapsis altitude control is desired.**
 - Spacecraft accelerometer data will be an enabling capability
- **NASA activity, led by LaRC with JPL participation, is developing autonomous aerobraking techniques**



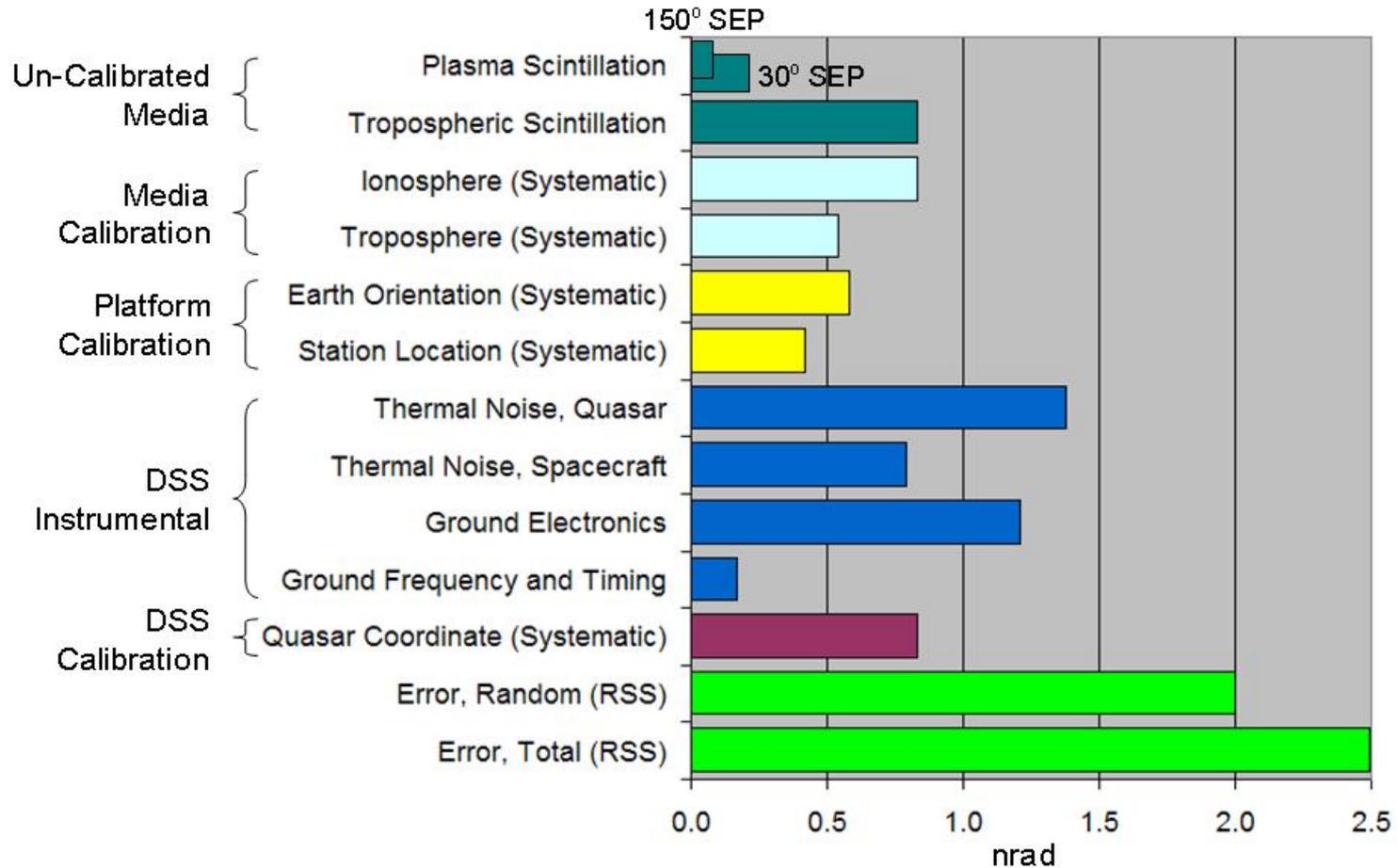
Ka-Band Δ DOR Data



- **Δ DOR performance has likely been pushed as far as possible at X-band**
- **Next significant advance will come by transitioning to Ka-band**
- **Important spacecraft feature is Ka-band DOR tone at 120 to 160 MHz**
 - **MRO Ka-band downlink demo (2005) used DOR tone of 76 MHz**
 - **Results indicated that higher frequency DOR tone was needed to get Ka-band performance that surpasses X-band**
- **Improved Δ DOR accuracy using Ka-band should be validated**
 - **When spacecraft enters Mars orbit, absolute accuracy of Δ DOR can be validated, using ephemeris of Mars as truth model**
 - **Future projects can then plan to use newly validated capability**
- **S/C-S/C Δ DOR demo opportunities may exist, for spacecraft near Mars**
- **2018 orbiter might provide a Ka-band beacon for future Mars missions**
- **Ka-band Δ DOR reduces key error sources of X-band Δ DOR**
 - **Reduce charged-particle errors by 15x ($1/f^2$ frequency dependence)**
 - **Reduce thermal noise and ground electronics errors by 4x (increased B/W)**
 - **Reduce quasar position errors by 4x (more compact radio sources)**
- **Δ DOR and telemetry should migrate to Ka-Band contemporaneously**



Typical DSN Δ DOR Error Budget at X-Band



From Border, J. S., Lanyi, G. E., and Shin, D. K., "Radiometric Tracking for Deep Space Navigation," in *Advances in the Astronautical Sciences: Guidance and Control 2008*, Vol. 131, edited by M. E. Drews and R. D. Culp, Univelt, San Diego, 2008, pp. 309-328.



Deep Space Atomic Clock (DSAC)



- **Scheduled for Earth orbit flight (2015) as NASA Tech Demo Mission**
- **A deep-space follow-on demo could be done with 2018 Mars orbiter**
 - Mass and power would be reduced for more fully flight-ready instrument
 - Long-haul, as well as in situ, links could be demonstrated
- **Use of DSAC on Mars approach may reduce demand on DSN antennas**
 - Tracking is typically continuous during last 45 days of Mars approach
 - Tracking in 2-way mode can accommodate single spacecraft per antenna
 - Tracking in 1-way (downlink only) mode can accommodate multiple spacecraft per antenna (MSPA), allowing antenna sharing among missions
- **Use of DSAC in Mars orbit can take full advantage of MSPA, allowing more extensive tracking coverage of each orbiting spacecraft**
 - Enables improved orbit knowledge, better gravity field determination, improved radio science, and more accurate and robust aerobraking
- **1-way Ka-band downlink (with DSAC) improves accuracy over 2-way, i.e., X-up/Ka-down tracking (eliminates higher intrinsic X-band error levels)**
- **1-way uplink (with DSAC) allows on-board radio-metric-based ‘*auto-nav*’**
 - Useful in various phases of general planetary missions
 - Radio-metrics, along with optical and inertial, would have broad applicability, desirable accuracy and robustness-to-mismodeling



Laser Ranging



- **Metric tracking for navigation should be derivable from deep space communication at optical frequencies**
- **Lasercom on 2018 Mars Orbiter could allow demonstration of S/C-Earth terminal ranging at optical frequencies**
 - Ranging capability can be added to current lasercom development for a modest incremental cost
- **S/C-S/C ranging demonstration could be done relative to other Mars orbiting S/C, if any are suitably equipped**
 - Secondary spacecraft (e.g., cubesat) could be deployed for this purpose, with corner reflector
- **Scientific experiments could be carried out also to obtain**
 - Improved estimate of post-Newtonian parameter gamma
 - Improved estimate of Martian ephemeris and 10-20 asteroidal masses
 - Improved Martian gravity field, yielding information on crustal structure and seasonal CO₂ flow, from
 - 2-way laser ranging or 1-way laser range linked to DSAC
 - Note that latter gravity experiment could also be carried out with
 - 2-way Ka-band Doppler or 1-way Ka-band Doppler linked to DSAC
 - Inclusion of accelerometer could further improve gravity field determination



Flight Segment DTN Implementation



- Integrate *Interplanetary Overlay Network (ION) S/W* with the avionic system (C&DH)
 - Re-use telecomm payload design
 - (no change from MRO and MAVEN)
 - Leverage existing ION software demonstrated on *Deep Impact Network Experiment (DINET)*
 - Require further integration with avionic S/W
 - Increased processing complexity and load on C&DH
 - *Performance can be improved by implementing better storage management and hardware/firmware acceleration on the most tasked processes*
- Demonstrate DTN integration with avionic system

Acronym Definitions:

BP = Bundle Protocol

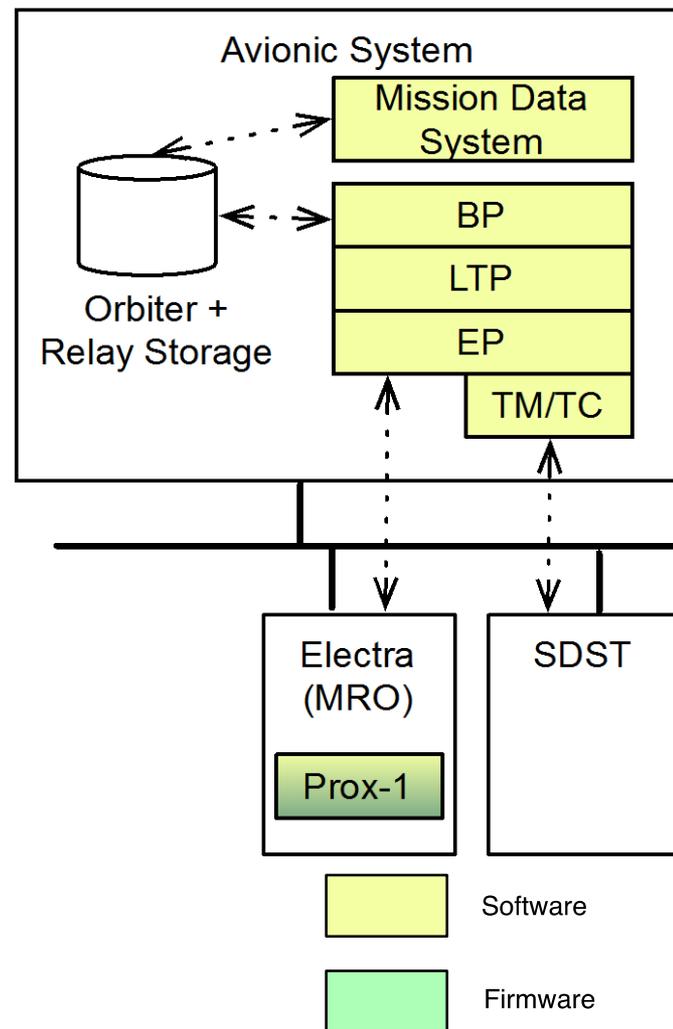
LTP = Licklider Transmission Protocol

EP = Encapsulated Protocol

TM/TC = Telemetry/Telecommand

C&DH = Command & Data Handling

Option 1: DTN in the avionic system





Ground Segment DTN Implementation

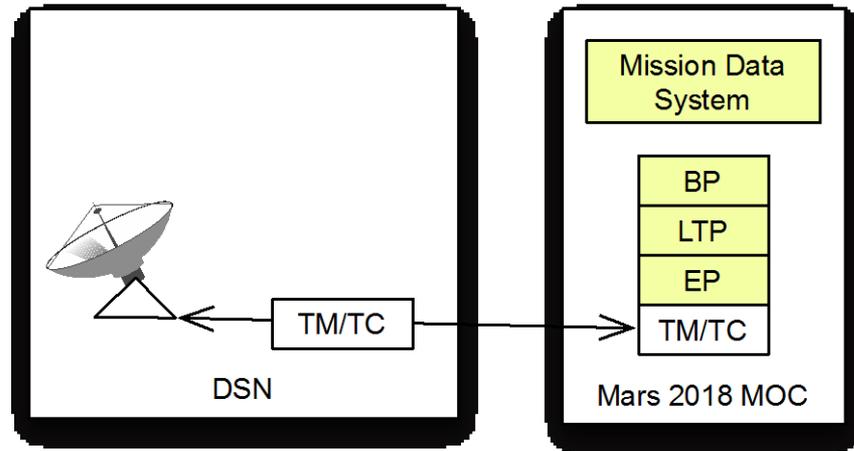


- **Option 1: At the MOC only**
 - EP, LTP and BP process done by MOC
 - LTP ARQ acknowledgement is generated at the MOC

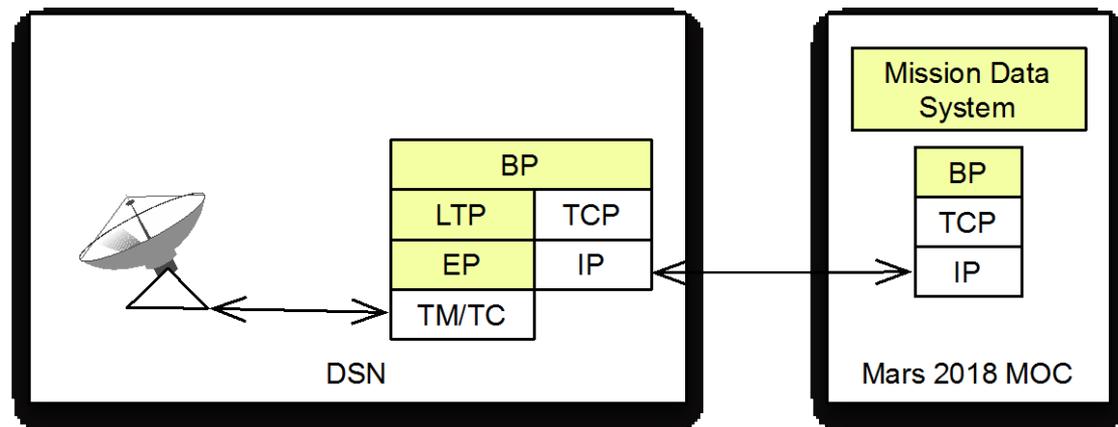
Acronym Definitions:

MOC = Mission Operations Center
ARQ = Automatic Repeat Request

OPTION 1: DSN PASS-THROUGH



OPTION 2: DTN in DSN



- **Option 2: LTP/BP at DSN; BP only at MOC**
 - MOC processes Bundles that contain mission data
 - DSN performs framing, encapsulation, and LTP retransmission



Summary



- **The paper surveys communications and navigation technology advancements relevant to the design and flight of a 2018 Mars orbiter.**
 - **Advancements in RF communications, for both the spacecraft and ground ends (i.e., DSN) of the link.**
 - **Importance of a deep space optical communications demonstration.**
 - **Advances in navigation techniques, both Earth-based and applicable to proximity operations.**
 - **Validation of space networking benefits, within the context of DTN.**
- **A 2018 mission is ‘made to order’ as a platform for advances described.**
- **These developments would materially contribute to NASA’s long-term goals in science, exploration and technology development.**
- **A 2018 Mars orbiter could reinvigorate the Mars program as well as provide a rationale for implementation of new capabilities that are important to future spacecraft as well as the future DSN.**