



# Terahertz Low Background Photon Noise Limited Performance Quantum Capacitance Detector

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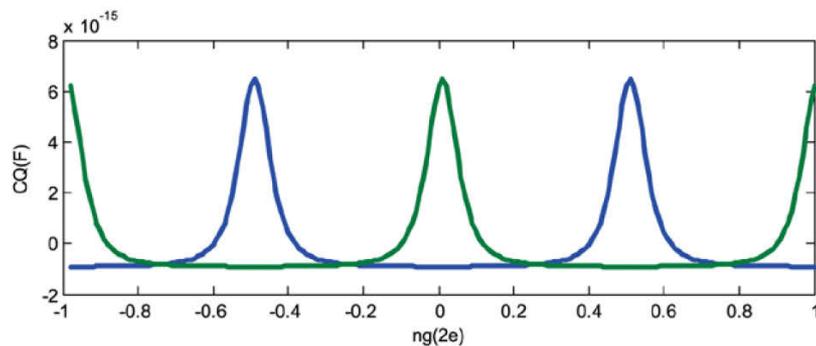
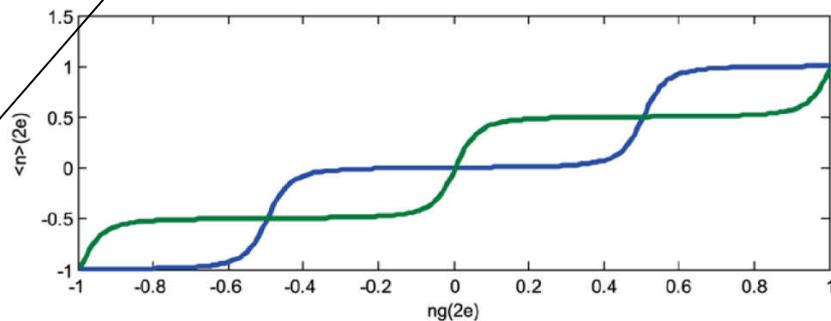
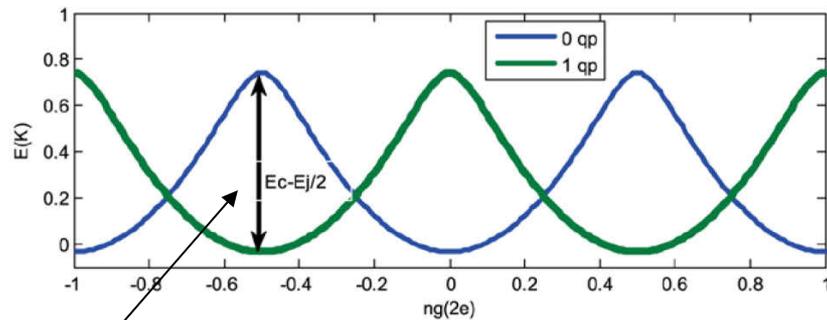
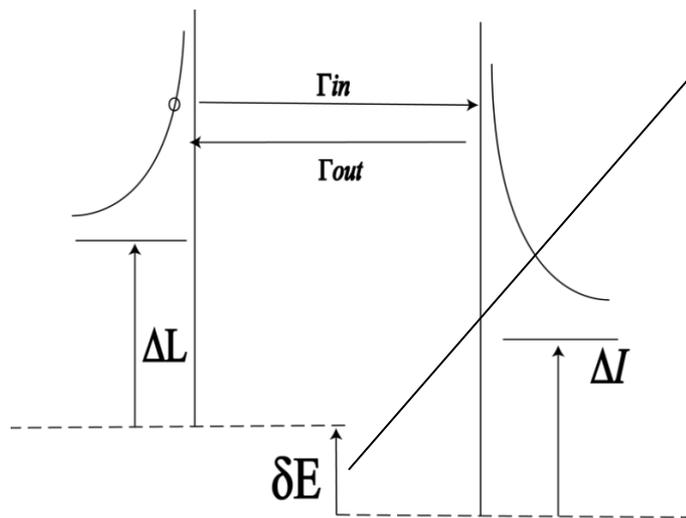
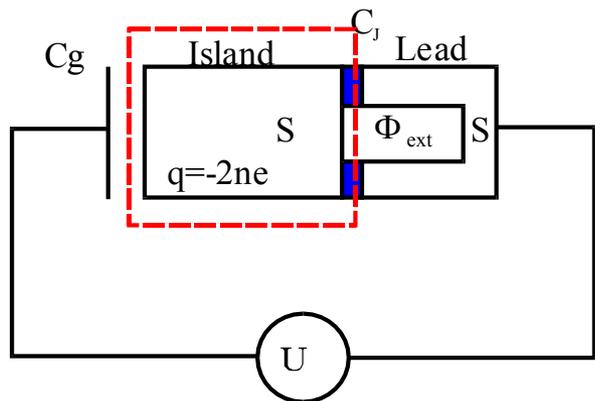
Electron Beam Lithography by  
Richard E. Muller

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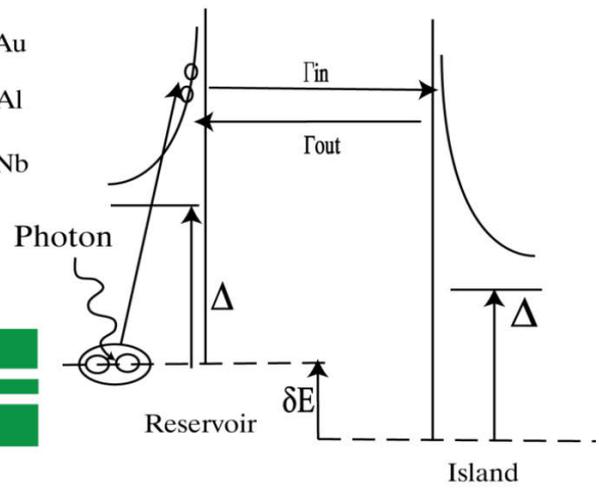
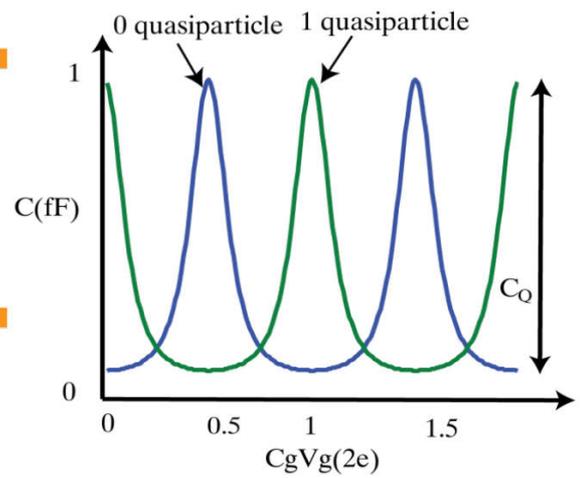
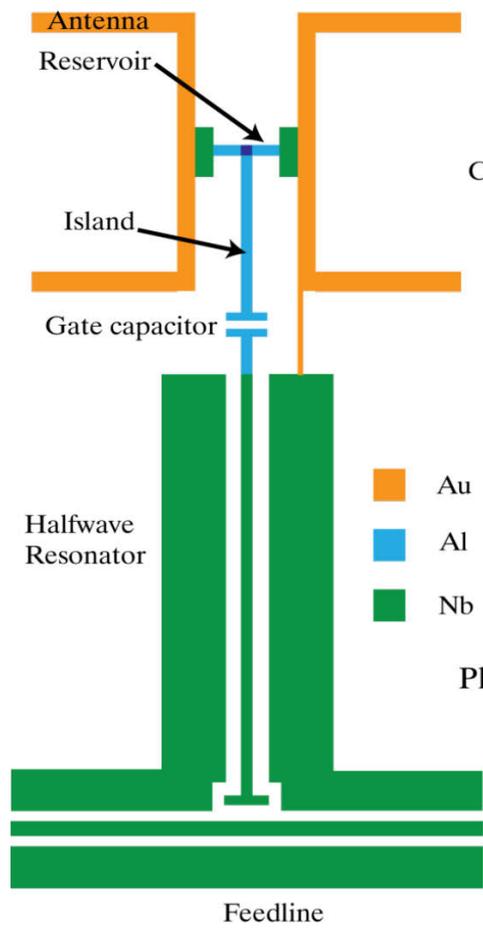
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# Single Cooper-pair Box (SCB)



# The Quantum Capacitance Detector

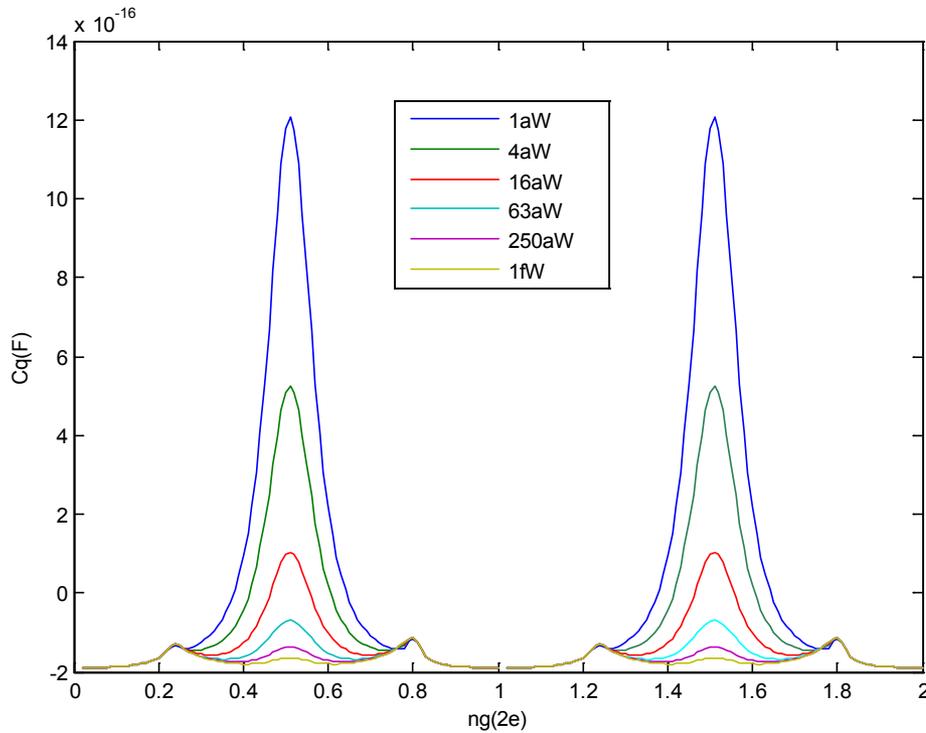


- Radiation coupled by an antenna breaks Cooper pairs in the reservoir (absorber)
- Quasiparticles tunnel onto the island with a rate  $\Gamma_{in}$  proportional to the quasiparticle density in the reservoir
- Quasiparticles tunnel out of the island with a rate  $\Gamma_{out}$  independent of the number of quasiparticles in the reservoir
- At steady state the probability of a quasiparticle being present in the island is given by  

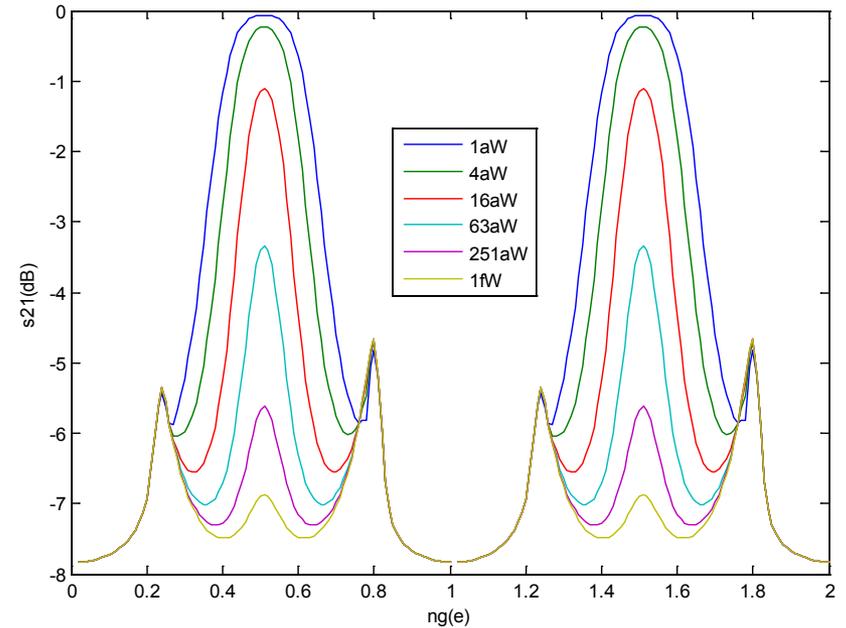
$$P_0(N_{qp}) = \Gamma_{in} / (\Gamma_{in} + \Gamma_{out})$$
- The resulting change in the average capacitance will be  $C_Q = (4E_C/E_J)(C_g^2/C_J)P_0(N_{qp})$
- This change in capacitance will produce a phase shift  $\delta\Phi \sim 2C_Q / (\omega_0 Z_0 C_c^2)$



# Simulated response



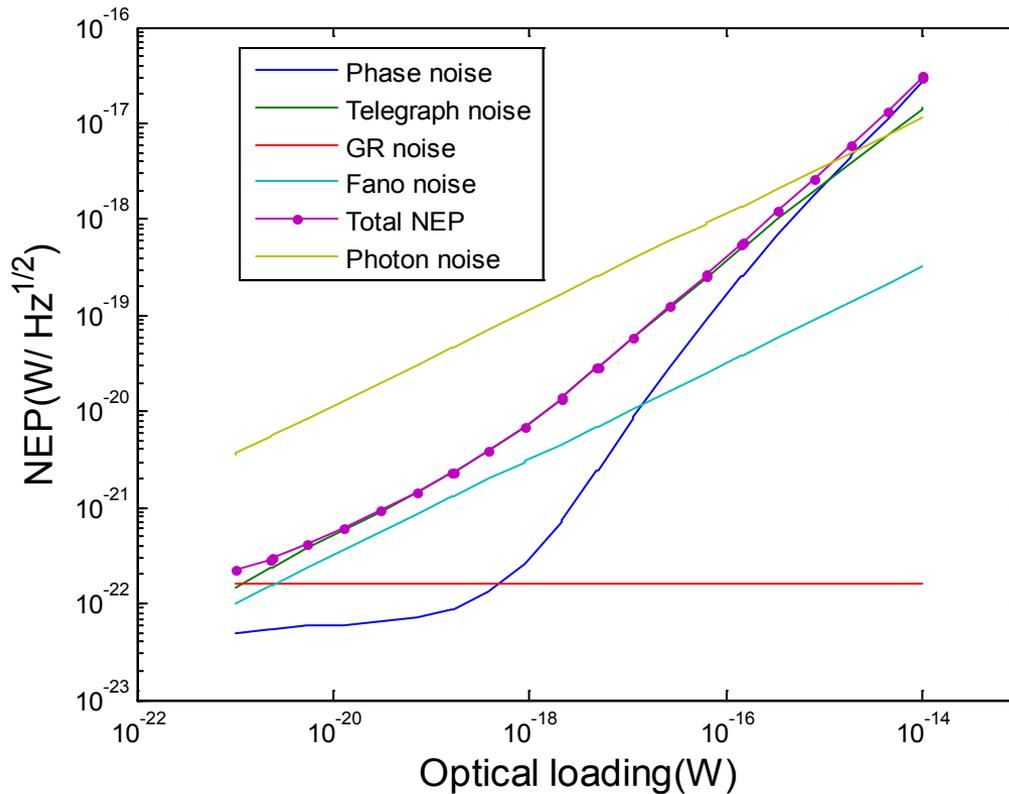
- SCB capacitance x gate voltage (in units of Cooper Pair charge) for different coupled optical signal power



- transmission through feedline x gate voltage (in units of Cooper Pair charge) for different coupled optical signal power

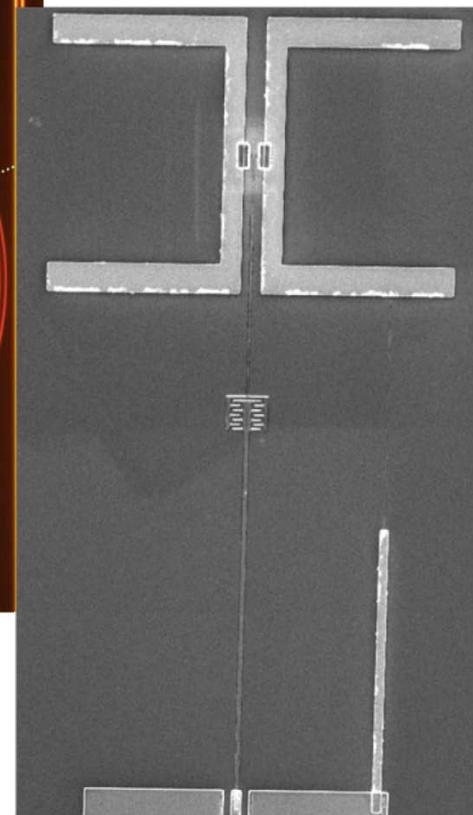
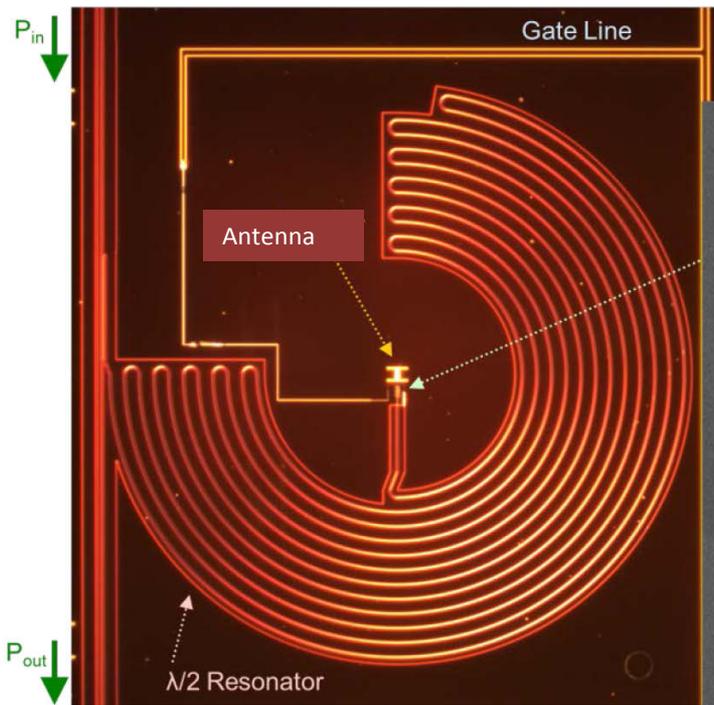
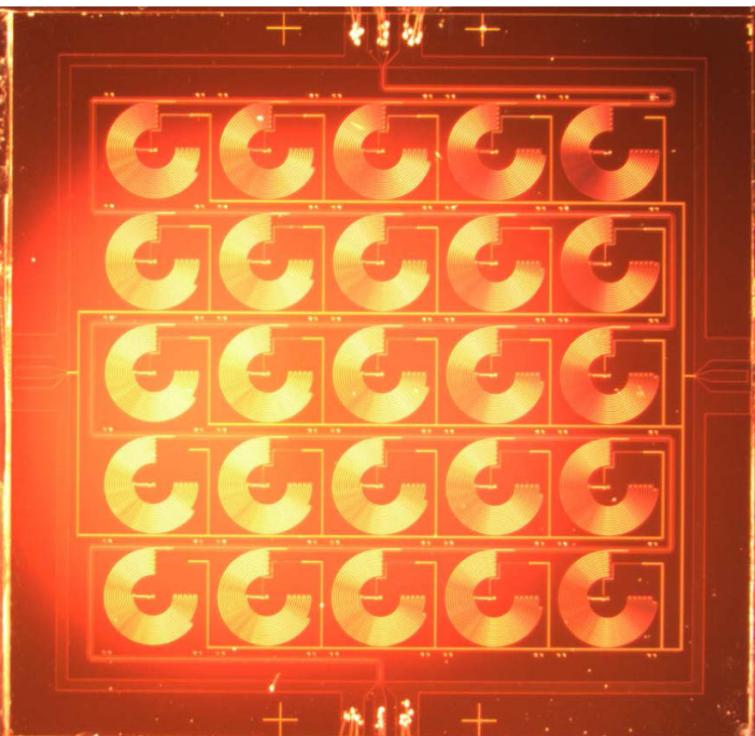


# Theoretical Sensitivity vs. Signal Power

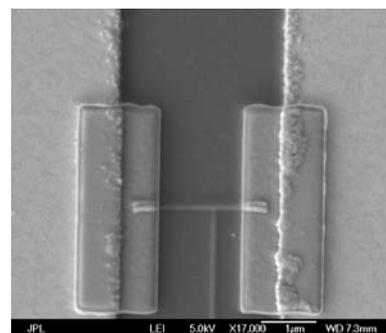


- *Detector is background limited over a wide range of operation*

# Quantum Capacitance Detector: 5x5 Array

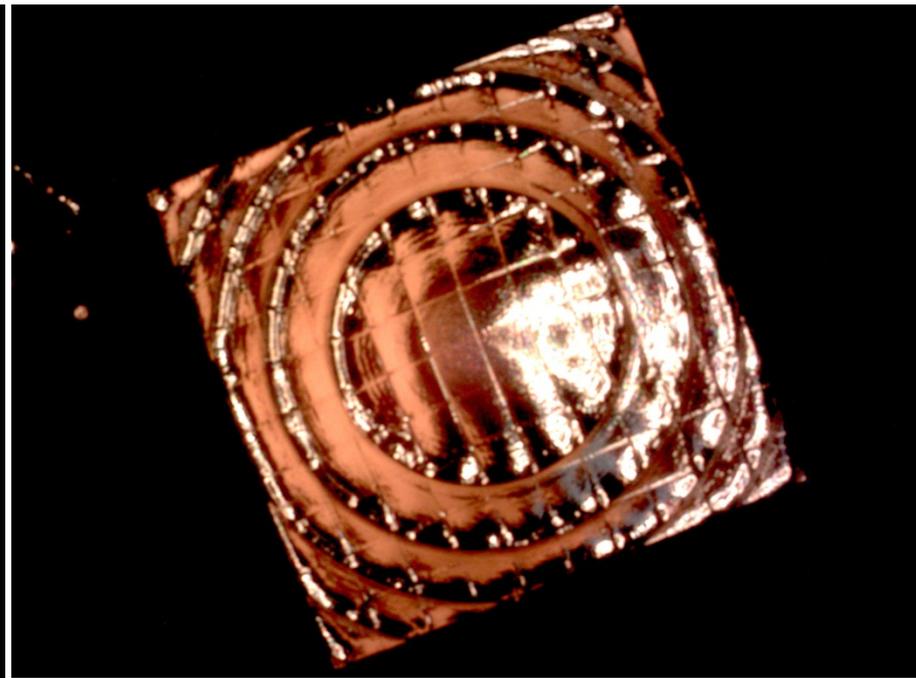
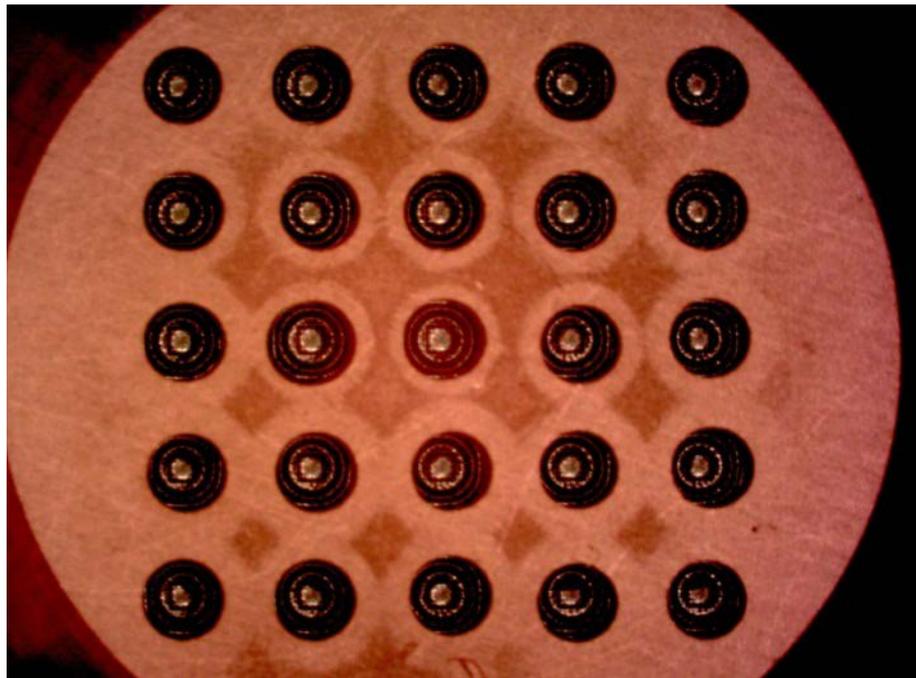


Only center device  
Illuminated by lens.  
Each device has a slightly  
Different resonance frequency.



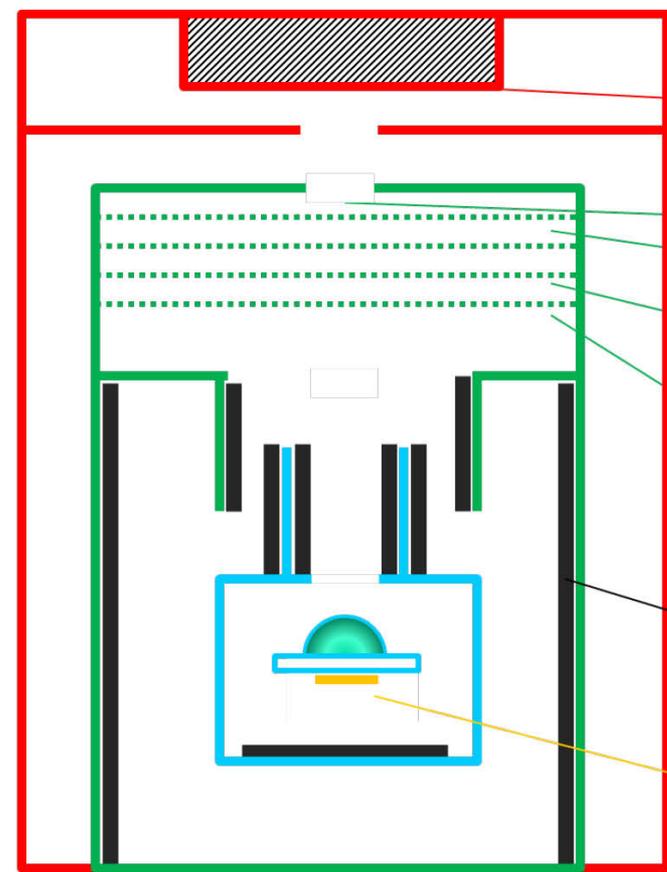
## Fresnel lens array

Lens array made by Dan Wilson using an electron beam lithography technique developed by Paul Maker, Dan Wilson and Rich Muller at JPL





# Experimental Setup



Blackbody source

Aperture  
2 mm diameter  
1.5THz band pass filter  
10% band  
3THz low pass filter  
0.03" teflon

Bock Black  
Radiation absorber

Sample

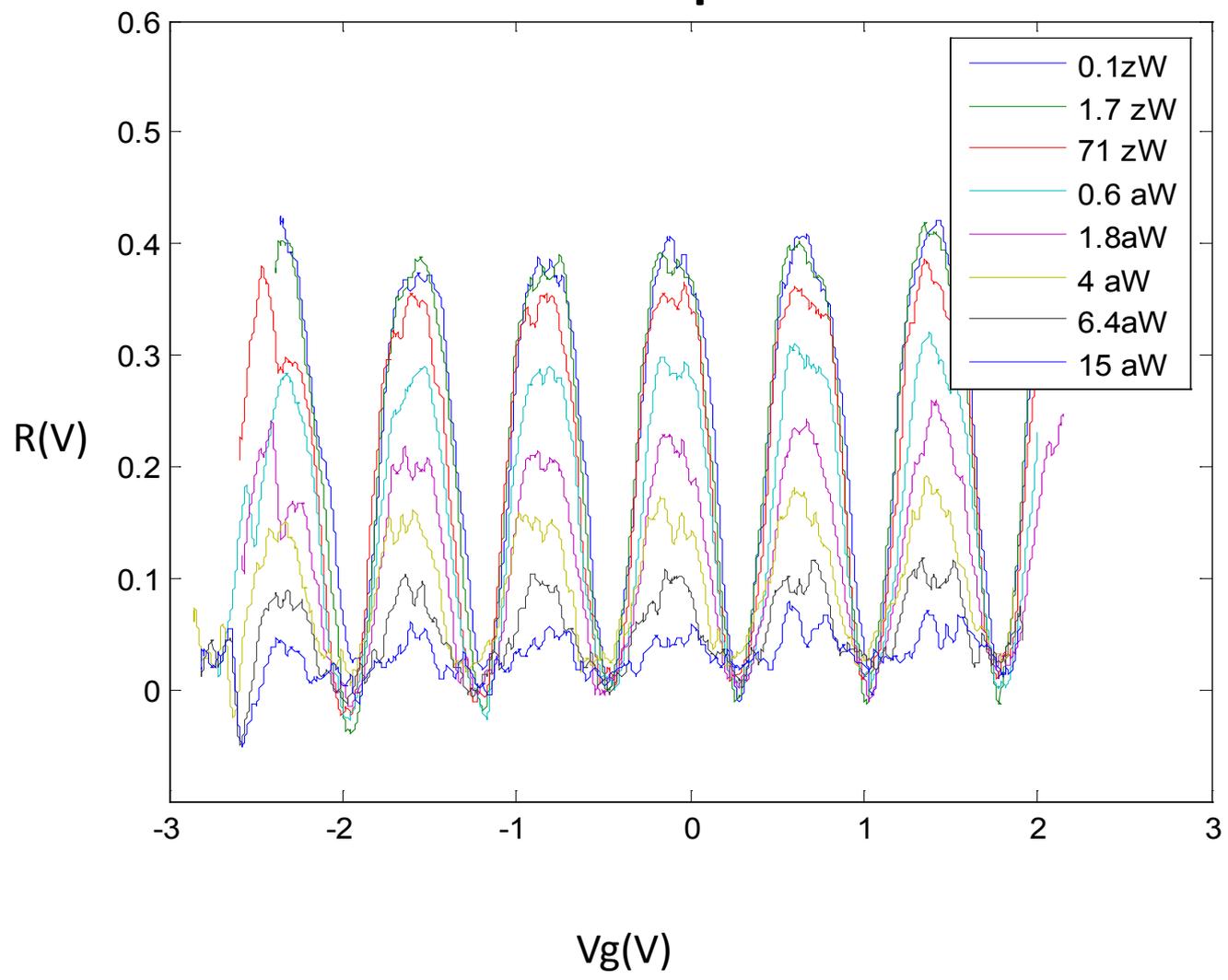
4K    Still Temp.    MC Temp.



- Black body source and filters provide 1.5THz radiation from 4.2 – 40 K. Bock Black absorbs stray 4K radiation

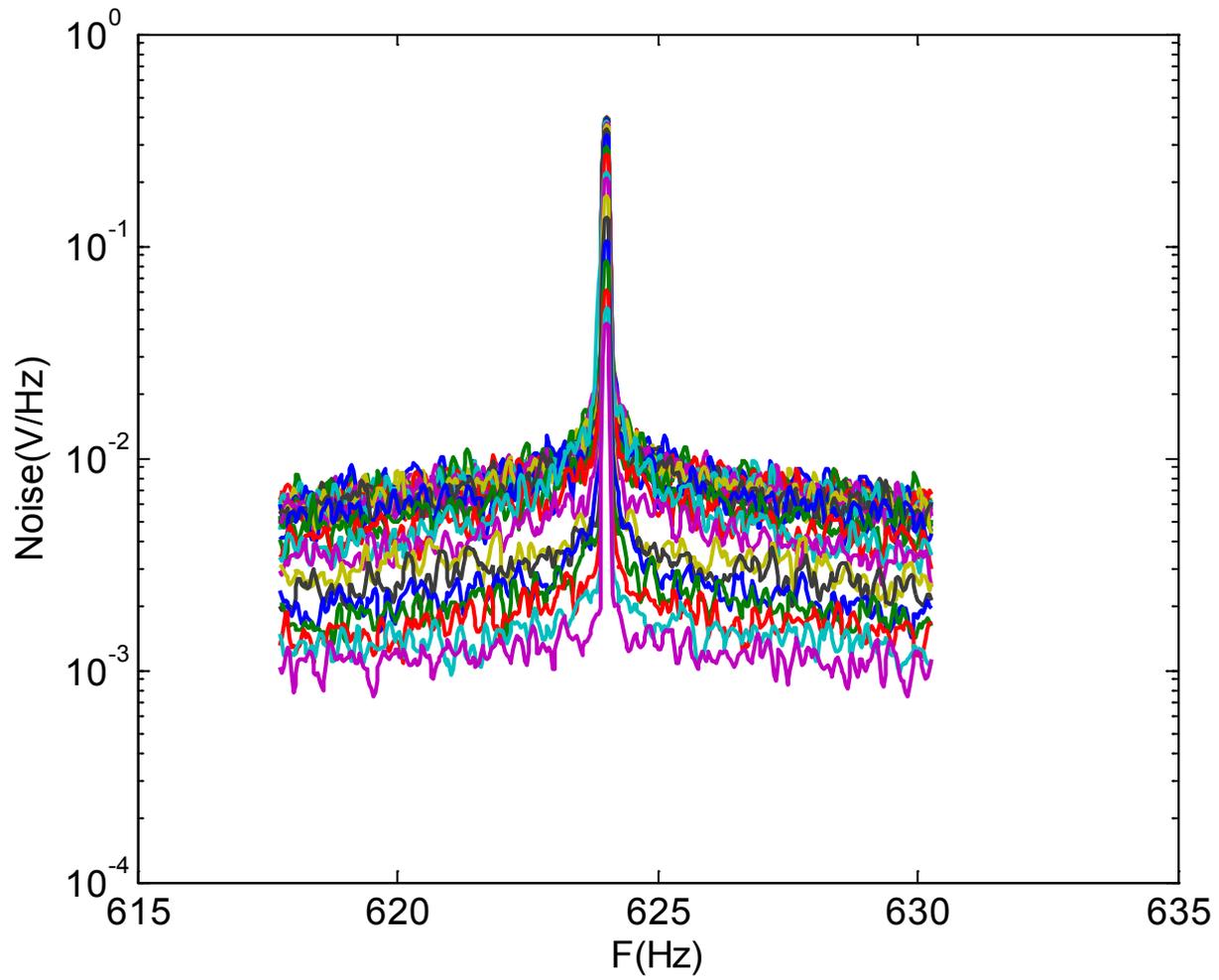


# Response x gate voltage as a function of black body source temperature



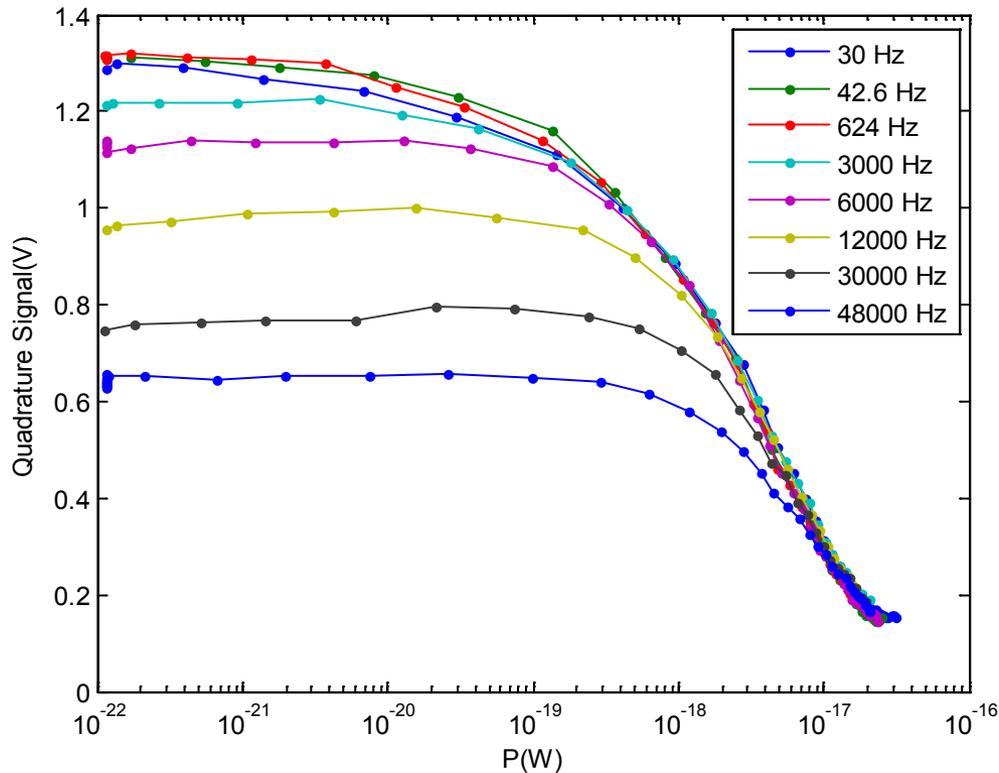


# Noise measurements



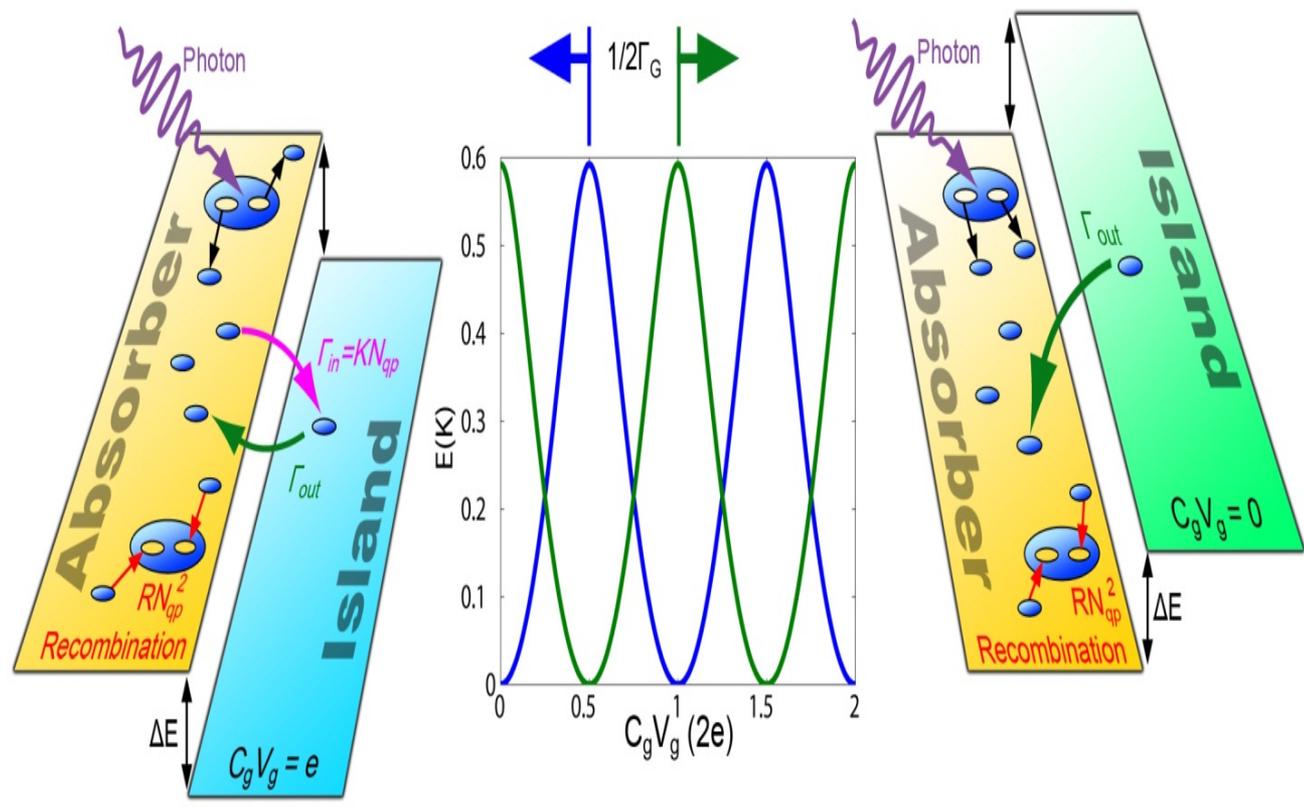


# Measurements of Response versus optical power for various gate sweep rates



- Aperture 500um
- Diffraction causes illumination to be about the same for all pixels
- Response similar from pixel to pixel
- What causes sweep rate dependence?

# Sweep rate dependence model





# Sweep rate dependence model

Detailed balance equation

$$dN_{qp} / dt = \eta P_S / \Delta - RN_{qp} (N_{qp} - 1) - \Gamma_{in} N_{qp} + \Gamma_{eff} = \eta P_S / \Delta - (R + K) N_{qp}^2 + RN_{qp} + \Gamma_{eff}$$

$N_{qp}$  = number of quasiparticles

$\Delta$  = superconducting gap energy

$P$  = optical signal power

$\eta$  = conversion efficiency from photon energy to quasiparticles  $\sim 0.57$

$R$  = recombination rate

$\Gamma_{in}$  = tunneling rate from absorber to island =  $KN_{qp}$

$\Gamma_{eff}$  = effective tunneling rate from island to absorber

Model for  $\Gamma_{eff}$  ->  $\Gamma_{eff} = \sqrt{\Gamma_{out}^2 + SR^2}$

At the end of each sweep quasiparticles are dumped back into reservoir. If the sweep rate ( $SR$ ) is faster than the intrinsic tunneling out time  $\Gamma_{out}$ , then  $\Gamma_{eff} \sim SR$ .

If  $SR$  is much slower,  $\Gamma_{eff} \sim \Gamma_{out}$



## Sweep rate dependence

Steady state solution

$$N_{qp} = (R + \sqrt{R^2 + 4(\eta P_S / \Delta + \Gamma_{eff})(R + K)}) / 2(R + K)$$

PR is the photon arrival rate,  $h\nu$  the photon energy,  $\Delta$  the absorber superconducting gap

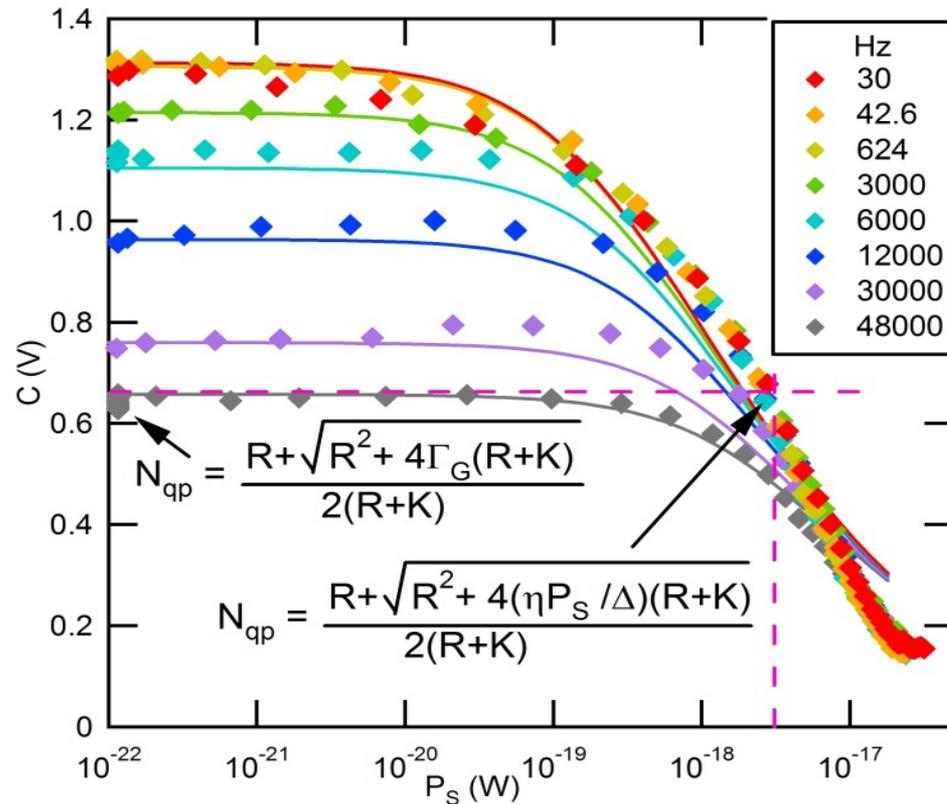
Response is a single value function of  $N_{qp}$

$$C = A\Gamma_{out} / (\Gamma_{out} + \Gamma_{in}) = A / (1 + KN_{qp} / \Gamma_{out})$$

A is a constant (depending on the electronics)



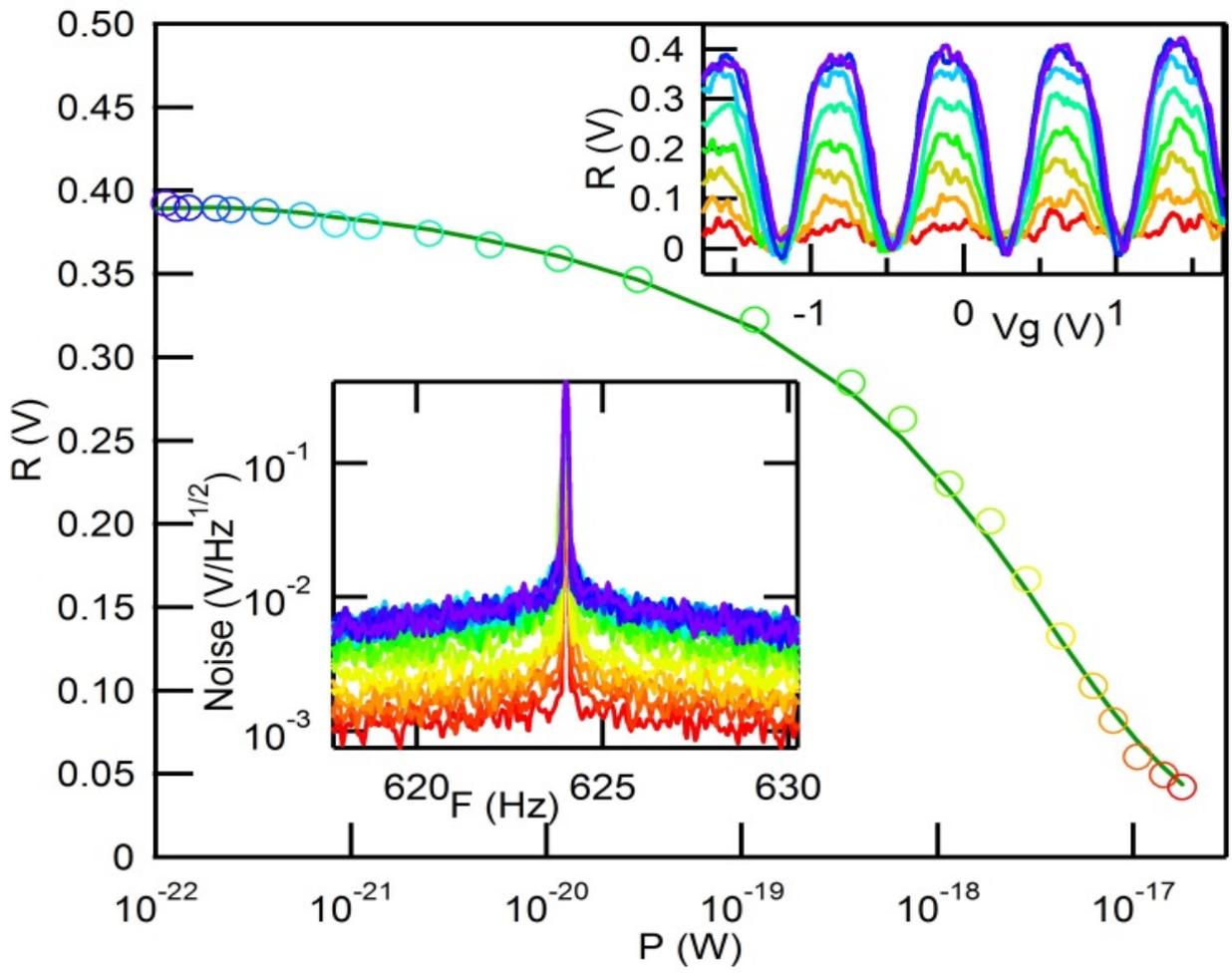
# Sweep rate dependence = calibration of photon arrival rate!



By comparing the response at low power and high sweep rate with the same response at high power and low sweep rate we arrive at  $P_s = \Gamma_G \Delta / \eta$



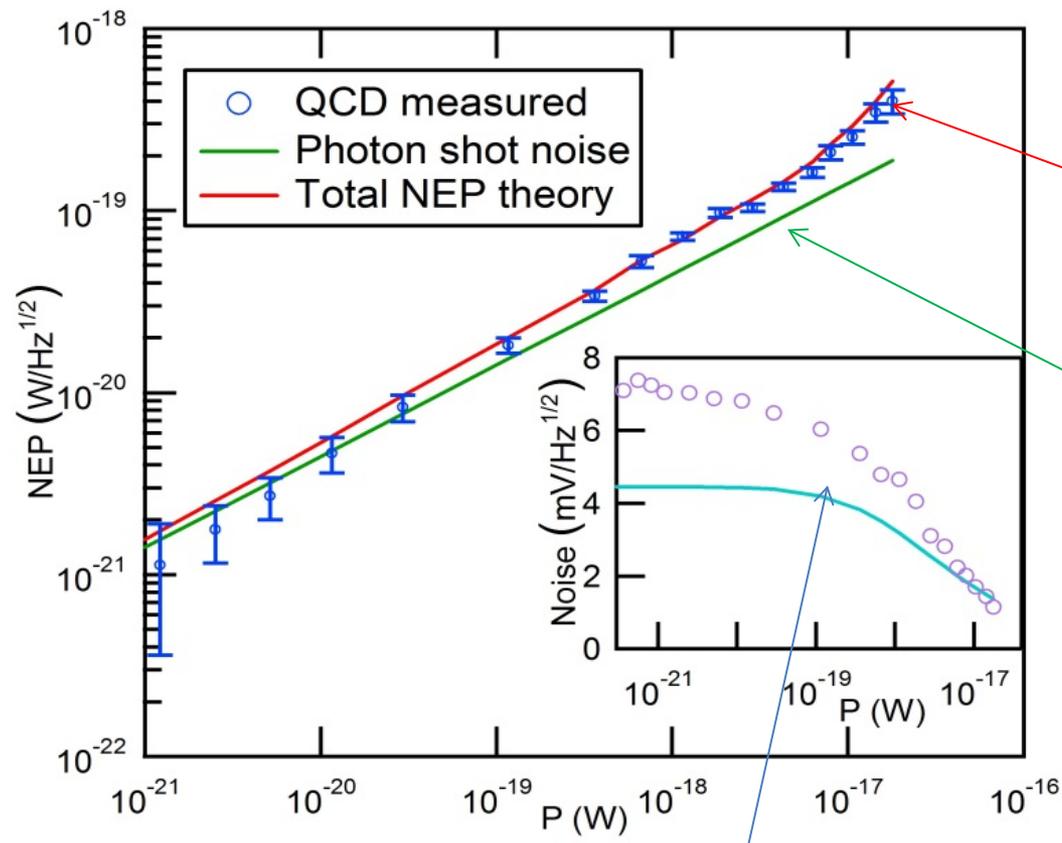
# Response





# NEP measured with power calibration

Photon shot noise limited!



$$NEP_{tot} = \sqrt{NEP_{ph}^2 + NEP_{sn}^2}$$

$$NEP_{ph} = \sqrt{2h\nu P_s}$$

$$NEP_{sn} = S_{sn}(f) / (dC / dP)$$

$$S_{sn}(f) = \sqrt{2A^2 (\Gamma_{in} \Gamma_{out} / \Gamma_{\Sigma}) / (\Gamma_{\Sigma}^2 + (2\pi f)^2)}$$



## Conclusion

- Achieved photon shot noise limited NEP at 200 $\mu$ m wavelength in a 5x5 array
- Novel way of calibrating absorbed optical power
- Fresnel lenses working
- Introduced special filters to cut down noise through coaxes
- Next
  - Redesign antenna for better efficiency
  - Tweak Fresnel lens fabrication for better lens profile



# Noise Equivalent Power

