

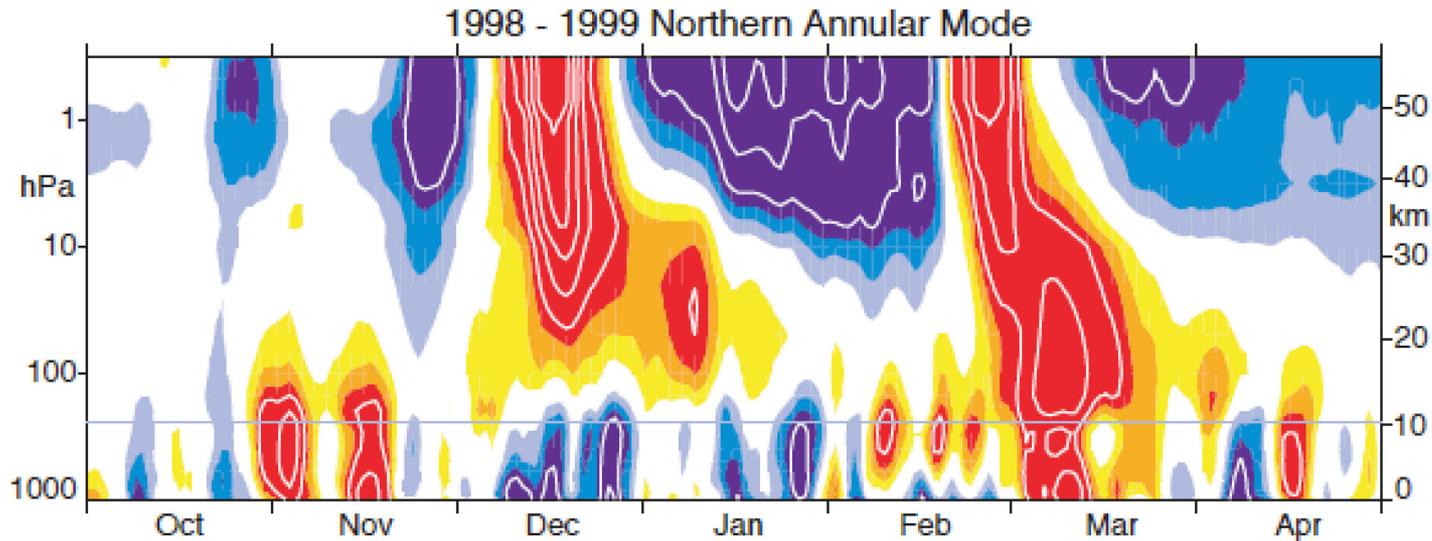


# Microwave Limb Sounder Observations of Polar Middle Atmosphere: Decadal and Inter-annual Variability

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# Motivation



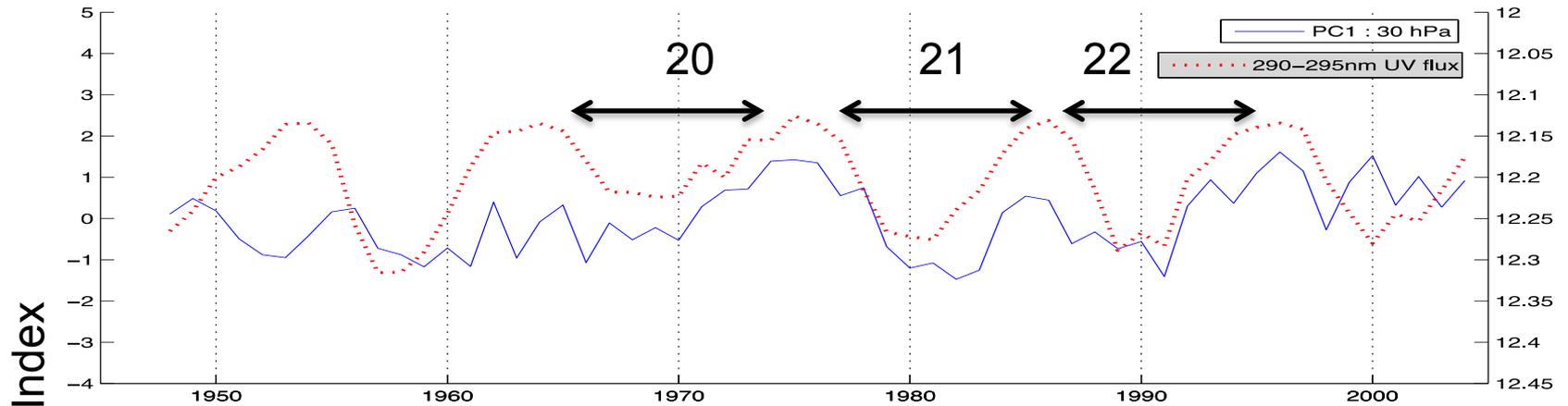
Baldwin and Dunkerton  
(2001)

- Why I care mode of variability?
  - to diagnose the dynamics and transport
  - to see the stratosphere-troposphere, mesosphere-stratosphere couplings

# Summer NAM at 30 hPa and UV

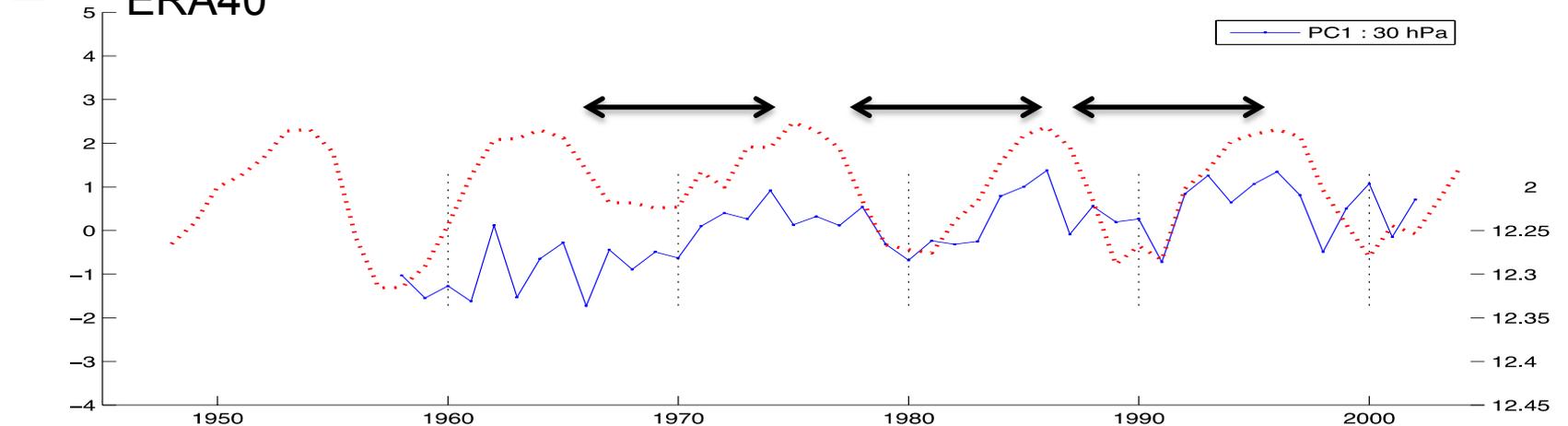
NCEP/NCAR

J. Lean's UV (w/m<sup>2</sup>)



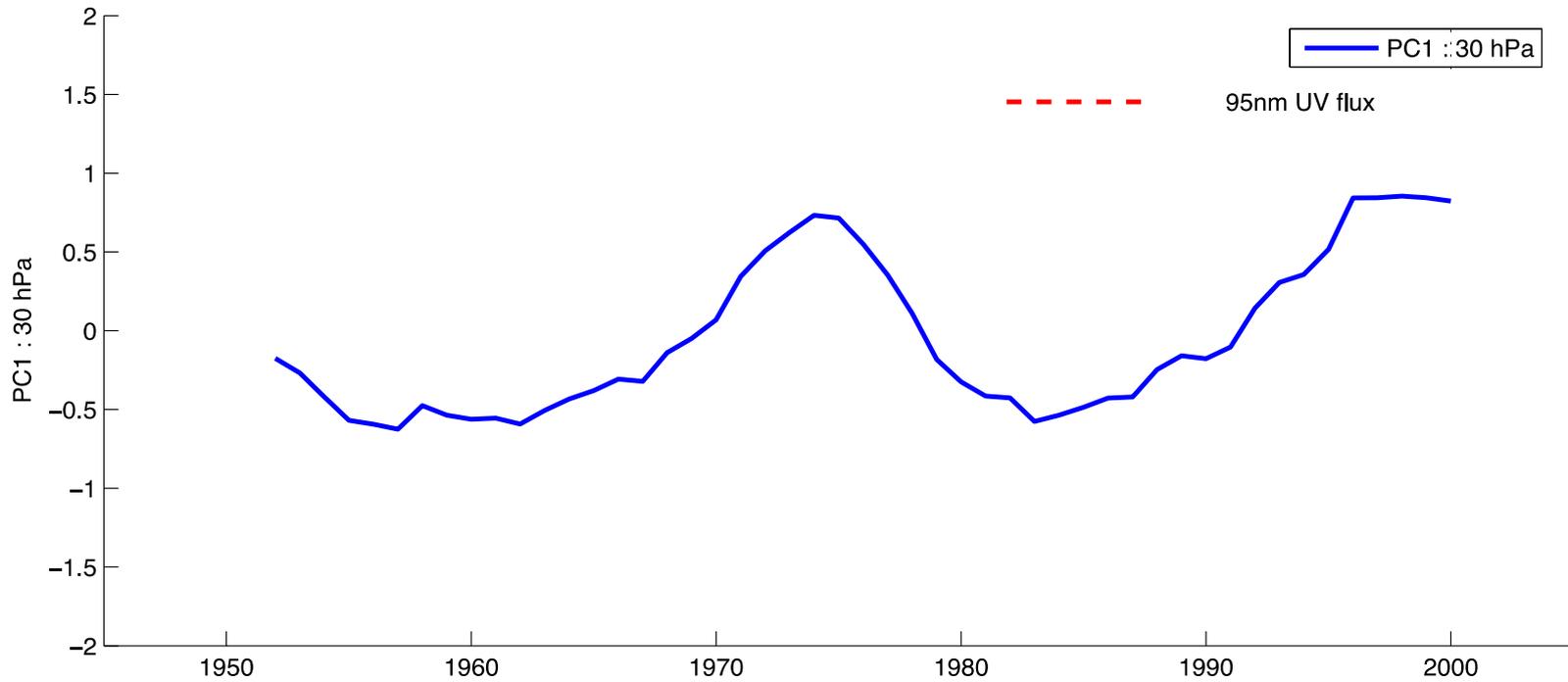
Lee and Hameed, 2007

ERA40



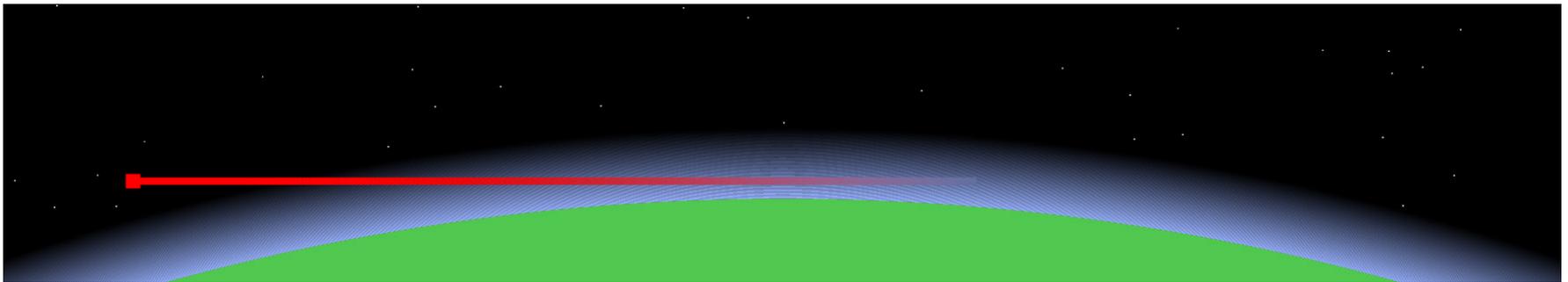
# Summer NAM and UV

NCEP/NCAR

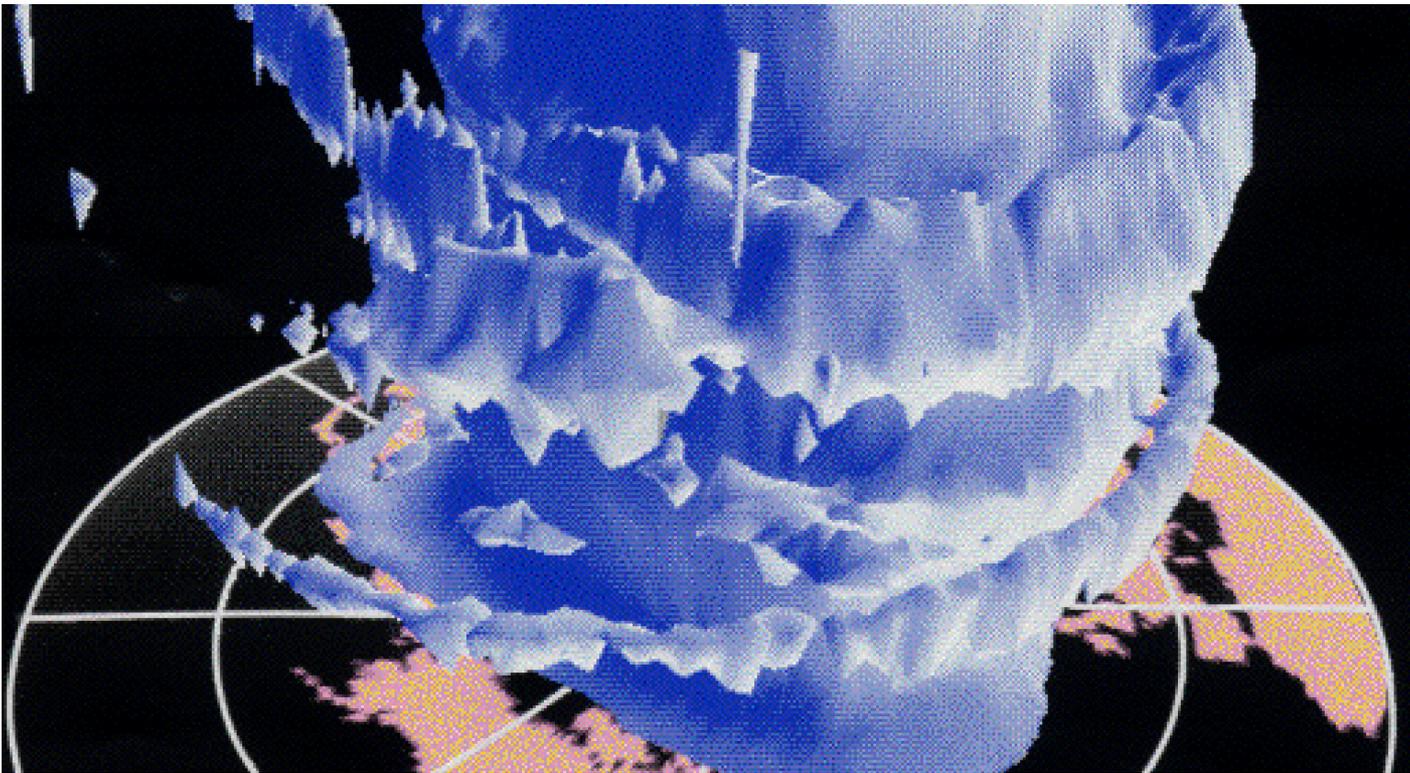


# Microwave Limb Sounder (since 2004- on going)

- The MLS observe thermal microwave emission from rotational lines of atmospheric molecules
- Rather than looking straight down (nadir), MLS instruments scan up and down in the limb geometry (edge onto the atmosphere)
  - This gives good vertical resolution.
  - The longer path length compared to nadir viewing gives a stronger signal for trace species ( $H_2O$ ,  $O_3$ ,  $CO$ ,  $OH$ , etc) and temperature.
  - Calibrations of MLS are less sensitive to solar variability than UV/VIS/IR instrument.
  - Microwave can measure the water vapor even in presence of the ice clouds and aerosols.



# 3-D visualization of polar vortex : GFDL model



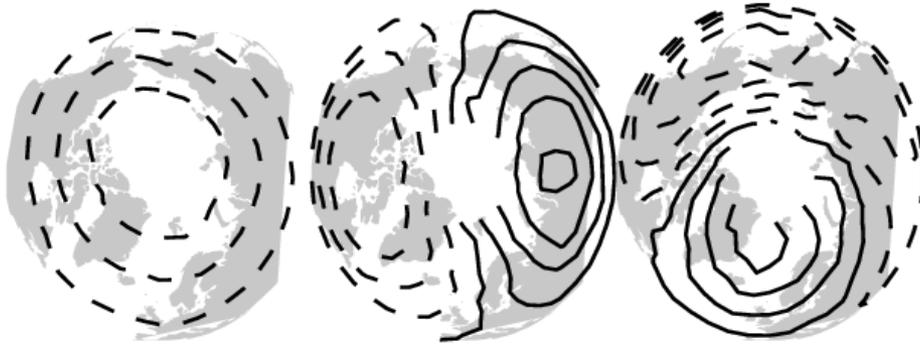
Courtesy : Kevin Hamilton

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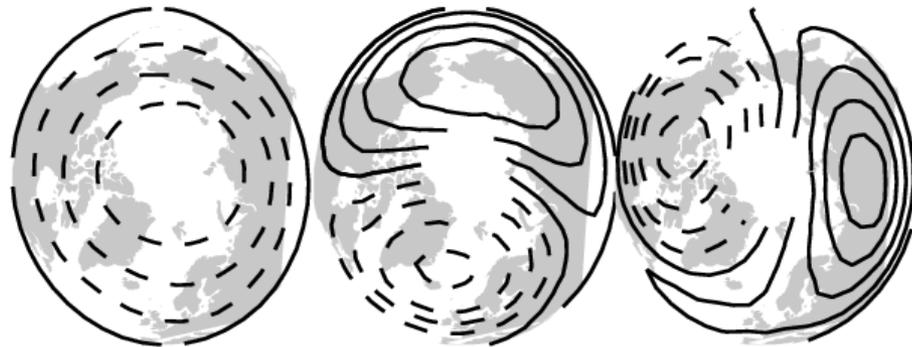
SORCE 2011, Sedona, Arizona

# The NAM patterns from the MLS (2005-2009)

MLS  
0.002hPa



1hPa

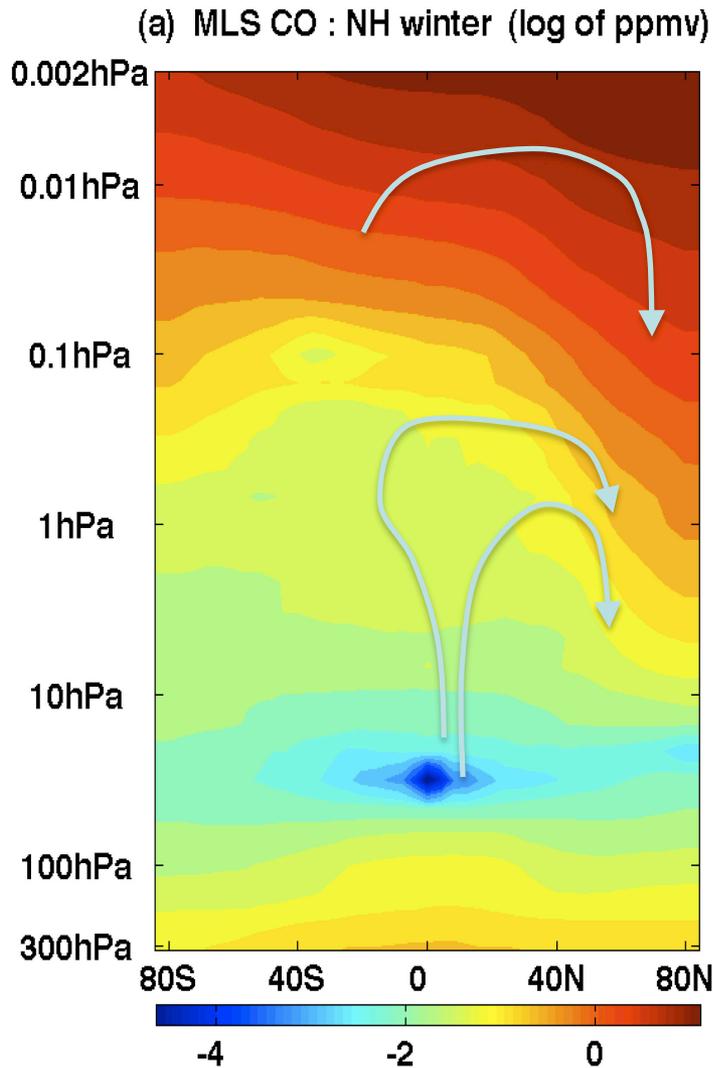


- NAM structure exists in the mesosphere.
- The NAM patterns derived from the MLS observations are consistent with those derived from long-term reanalysis below the middle stratosphere.

First annular mode : the same structure, with higher amplitudes at the pole

Second and third : orthogonal wave 1 patterns

# Aura MLS CO : DJF



- Vertical and horizontal gradients of zonal mean CO

- How does the polar descent shape up the tracer distribution?

- What is going to change during SSW? -> with strong perturbations.

- Besides geopotential height, the mode was constructed with CO field.

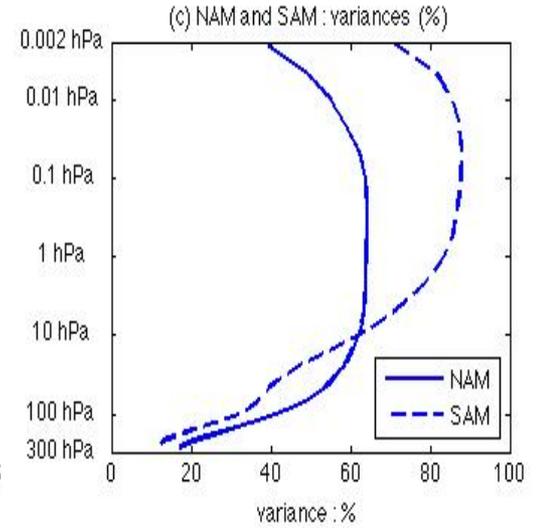
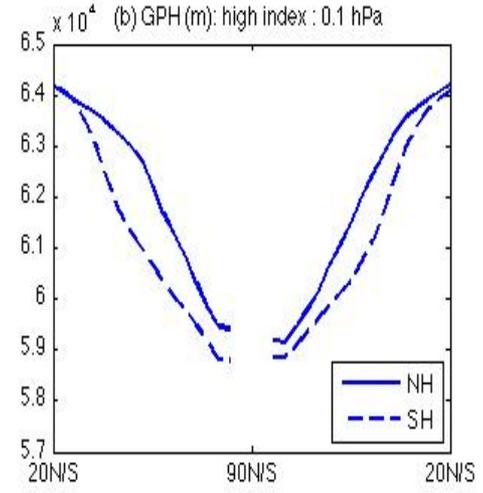
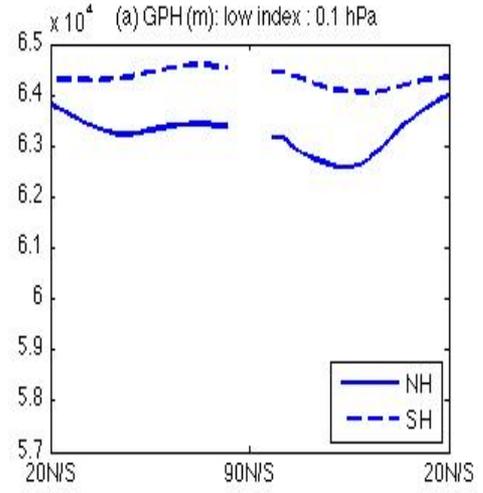
EOF1  
or  
NAM/SAM

Low Index

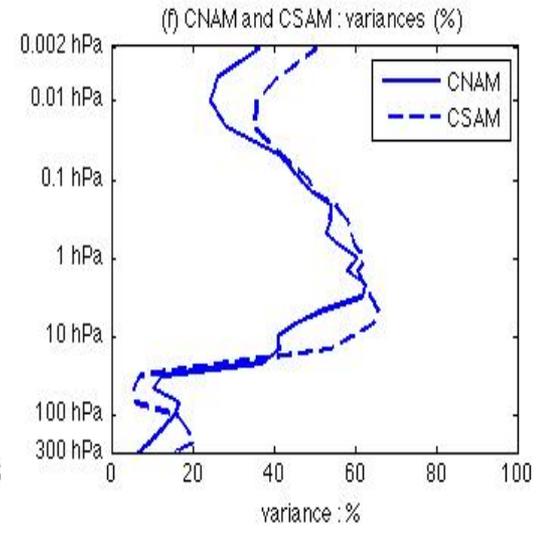
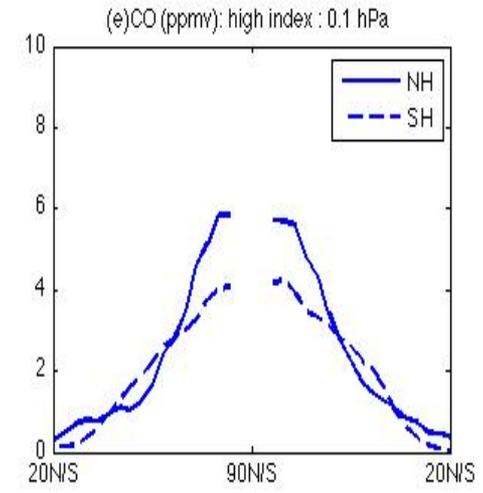
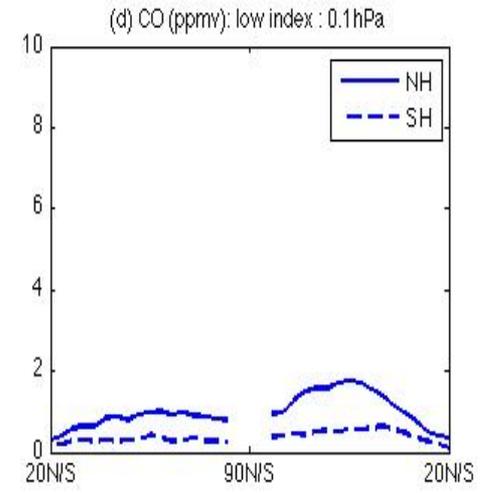
High Index

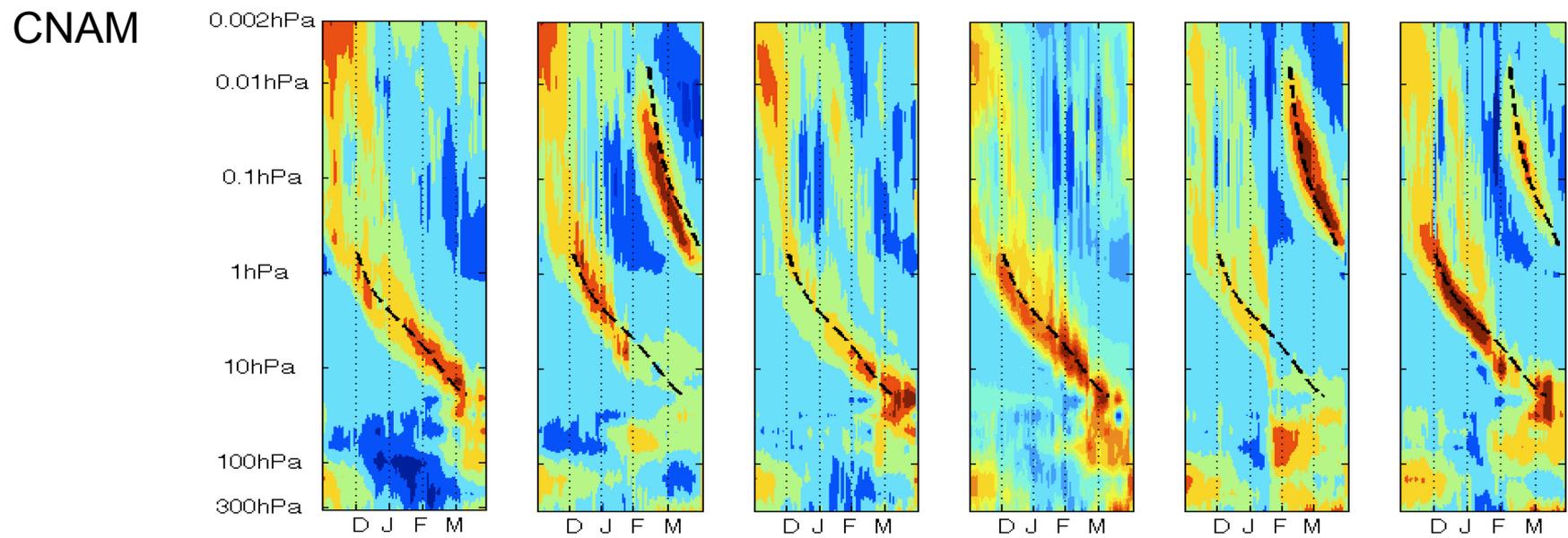
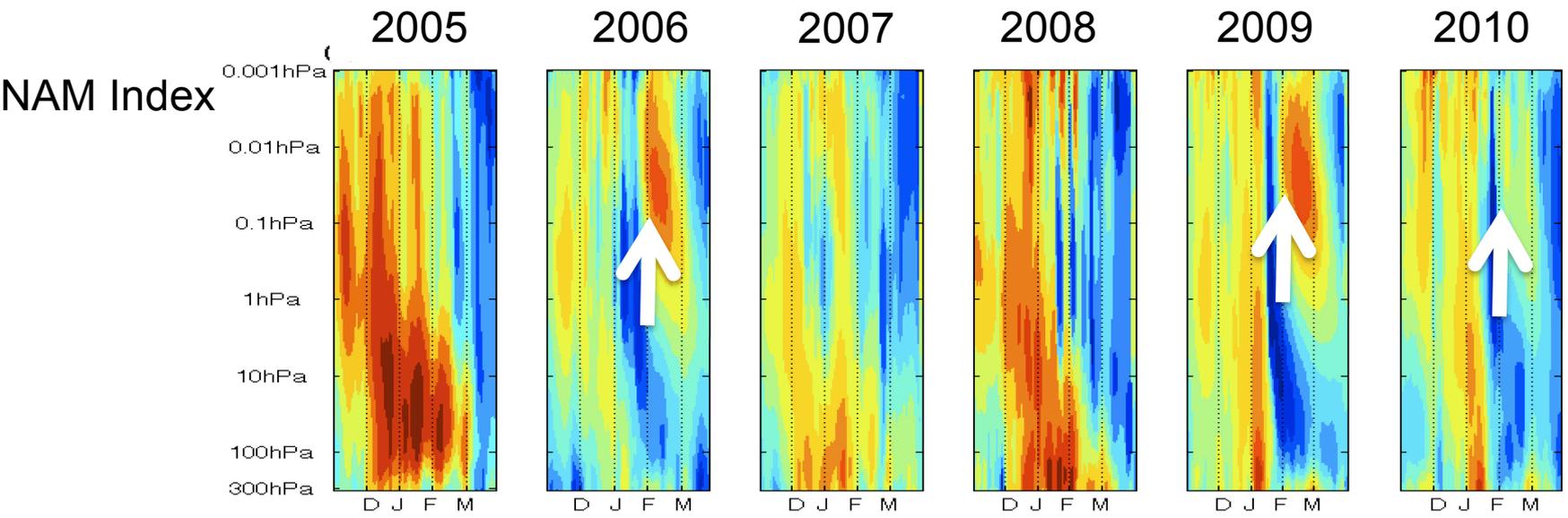
% of Variance

GPH



CO





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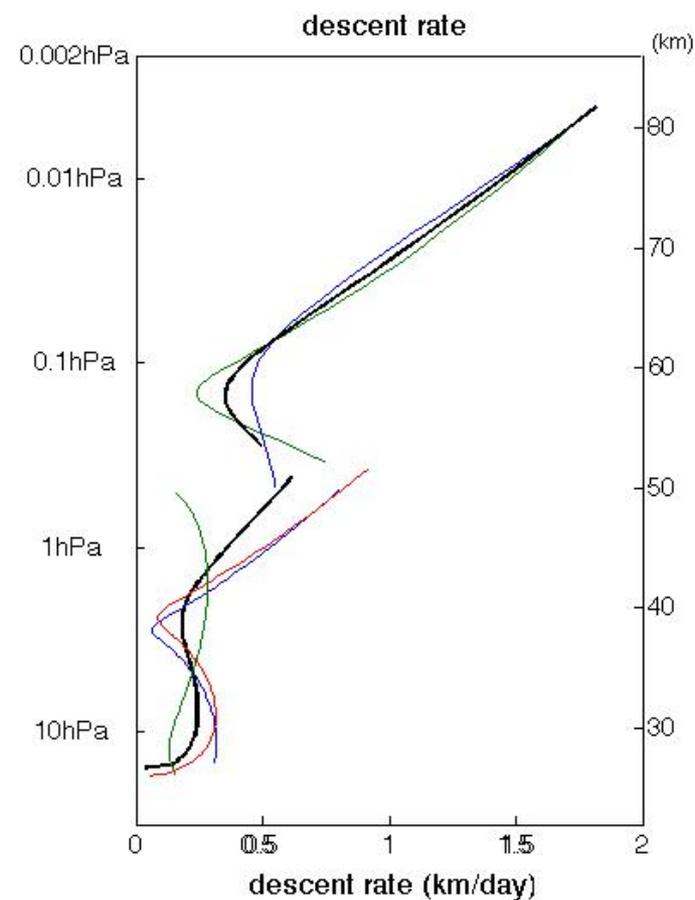
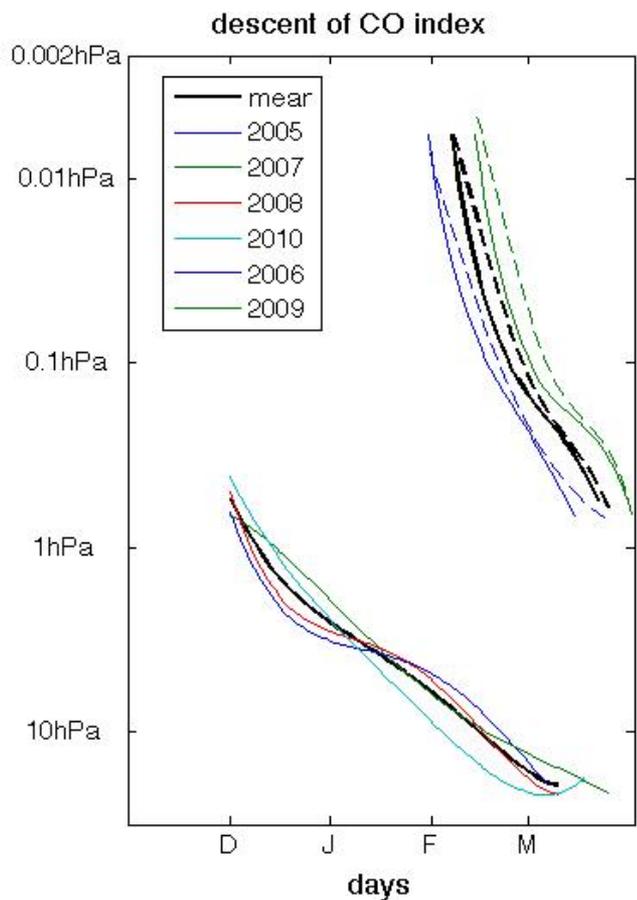


60° N-82° N zonal mean CO

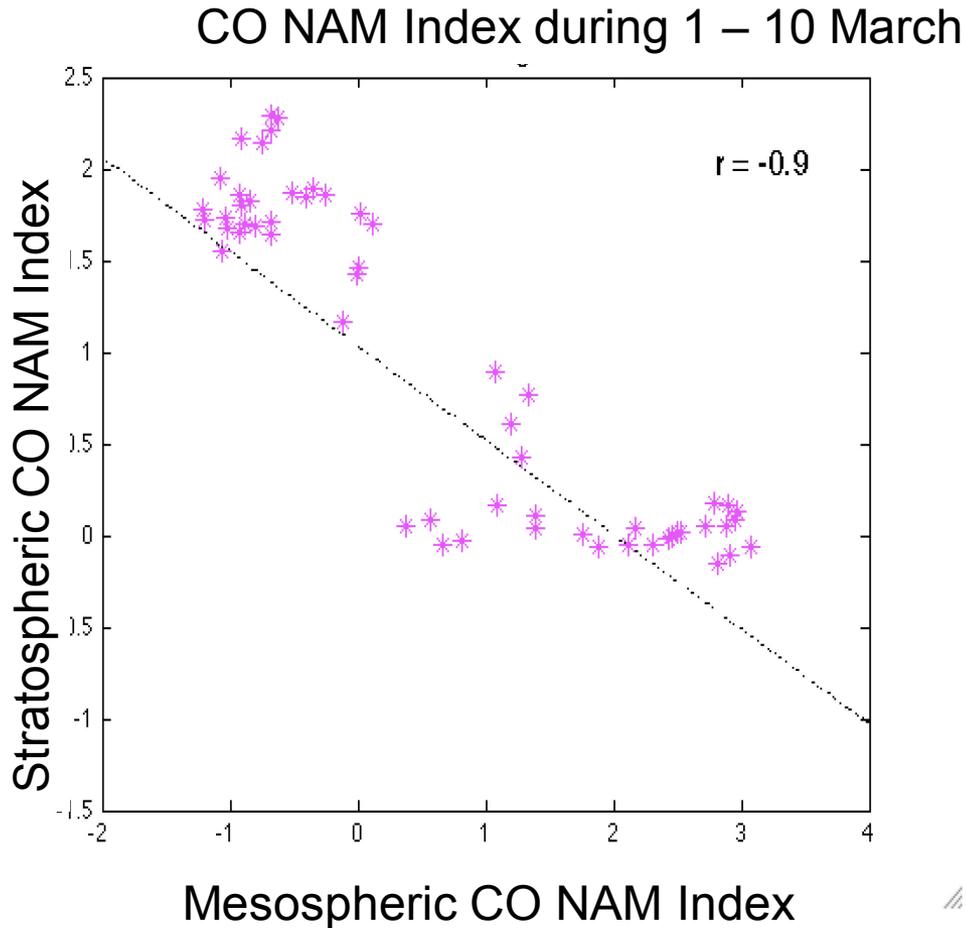
(log ppmv)



-CO NAM index line follows the maximum anomaly CO descent.



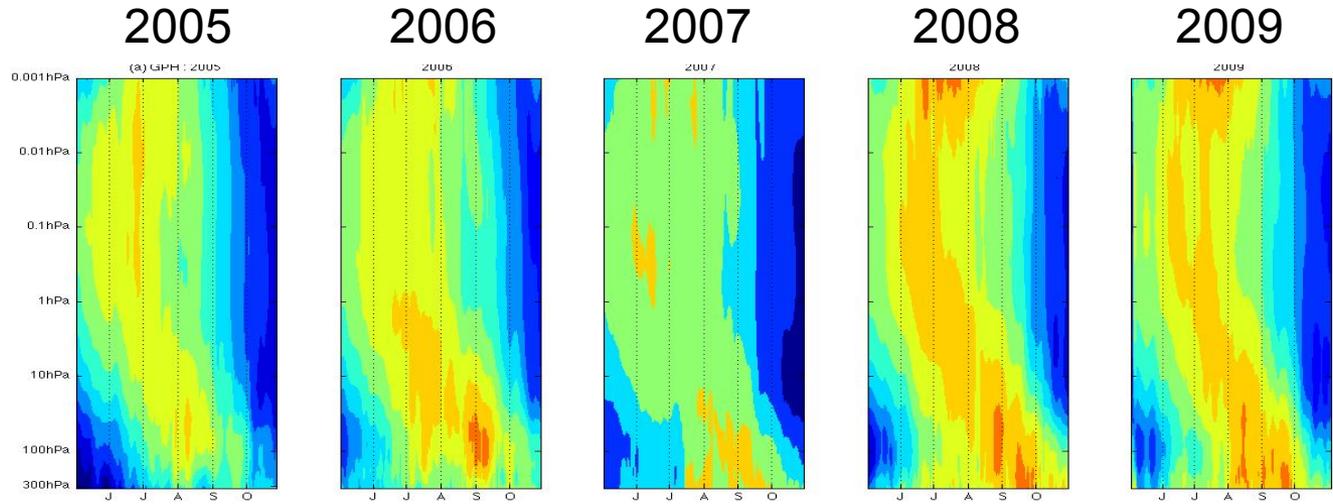
# Planetary and Gravity Wave Coupling



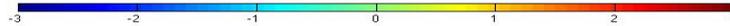
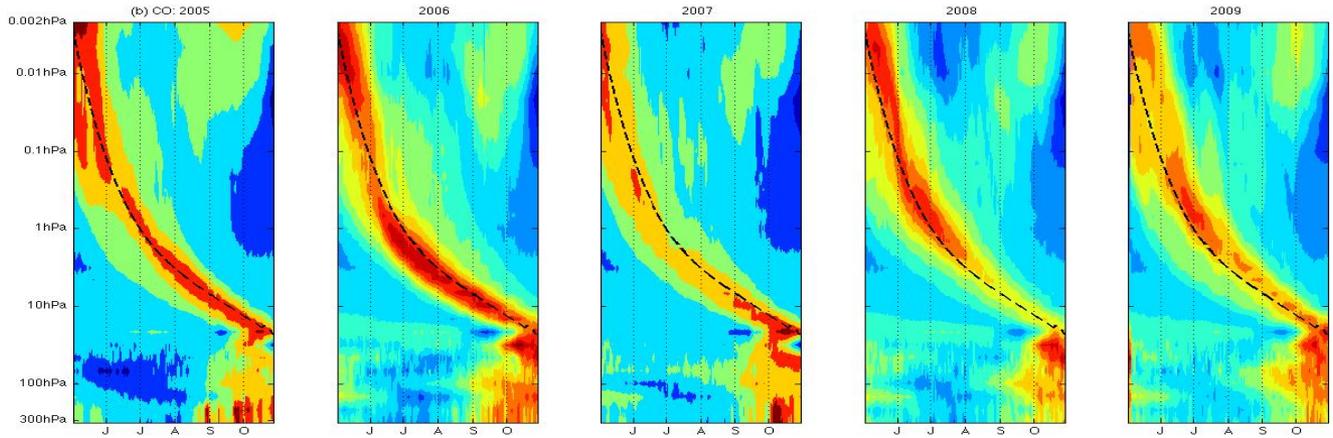
- Mesosphere and stratosphere CNAM anti-correlated
- Planetary and Gravity wave coupling
- weak vortices in the stratosphere (low index)
- Prevents gravity wave propagating upward
- forming strong vortex in the mesosphere
- Siskind et al. [2010]

# SAM Index

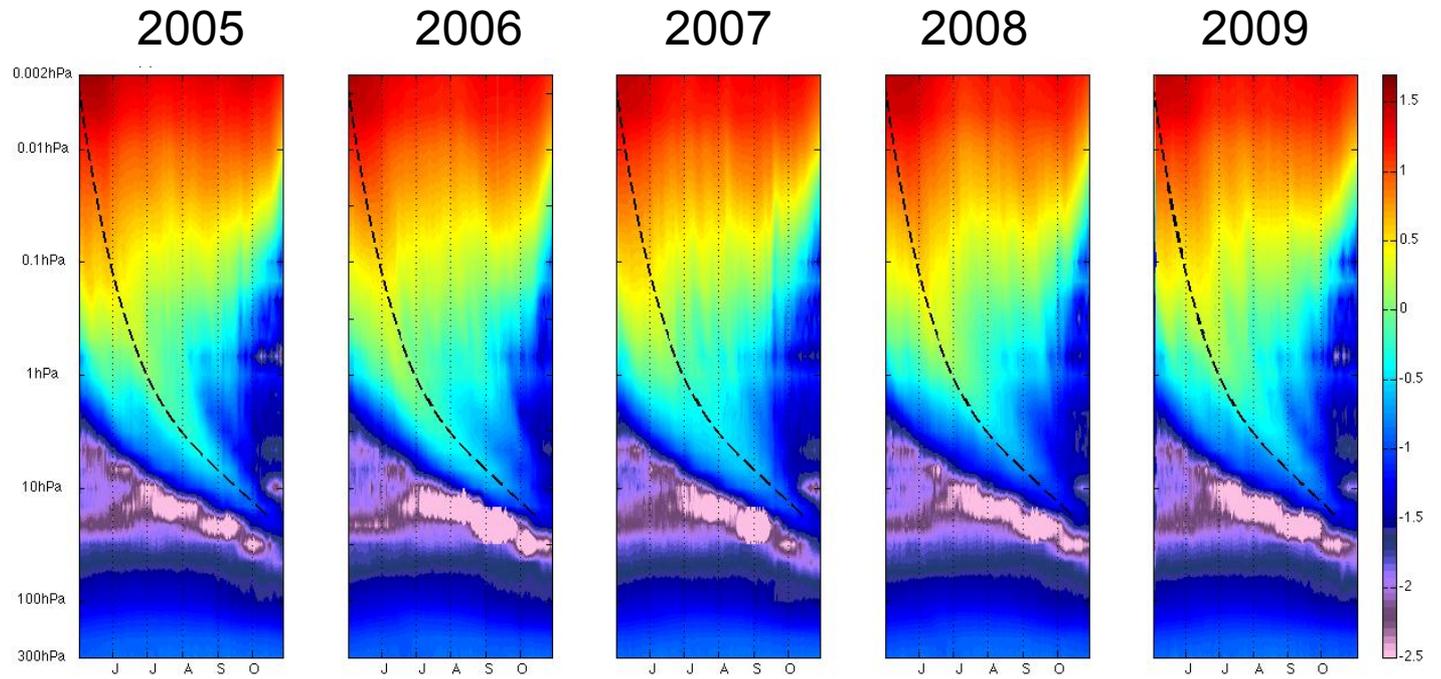
GPH

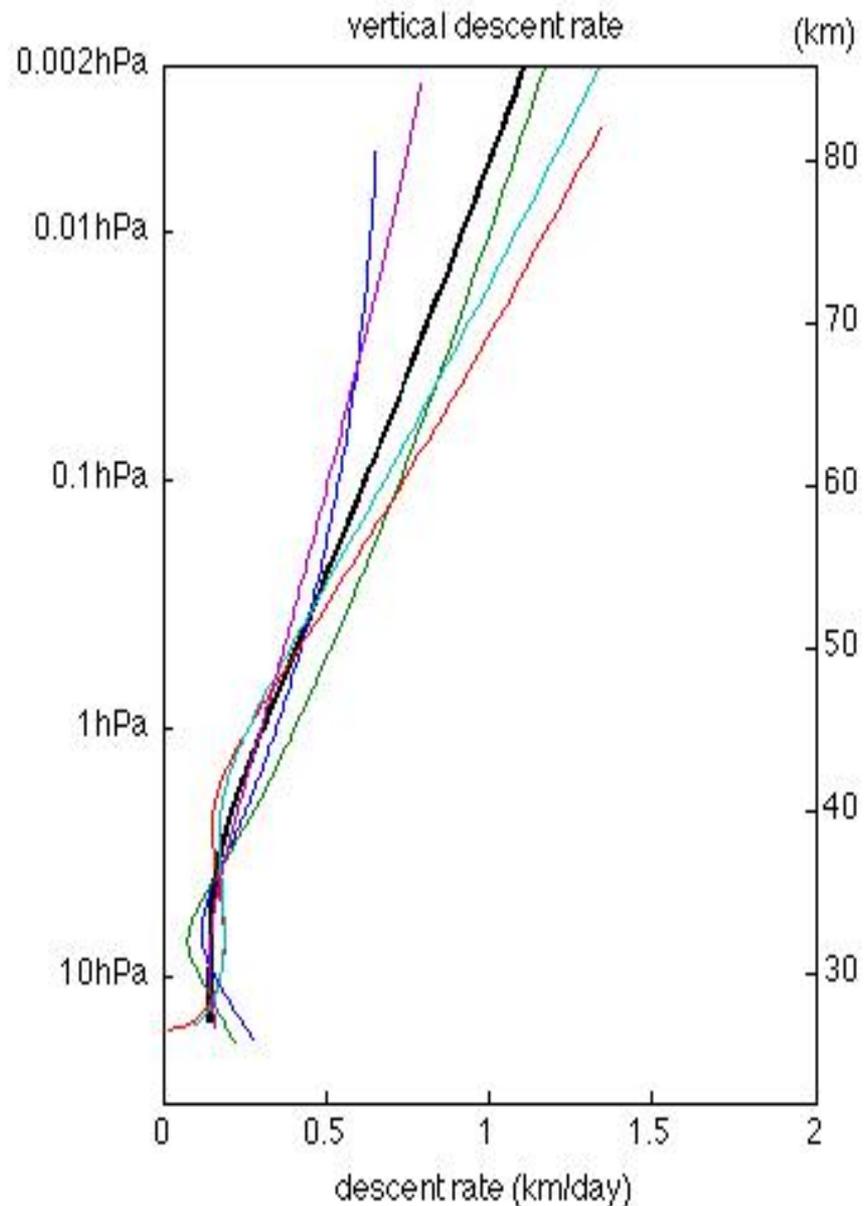
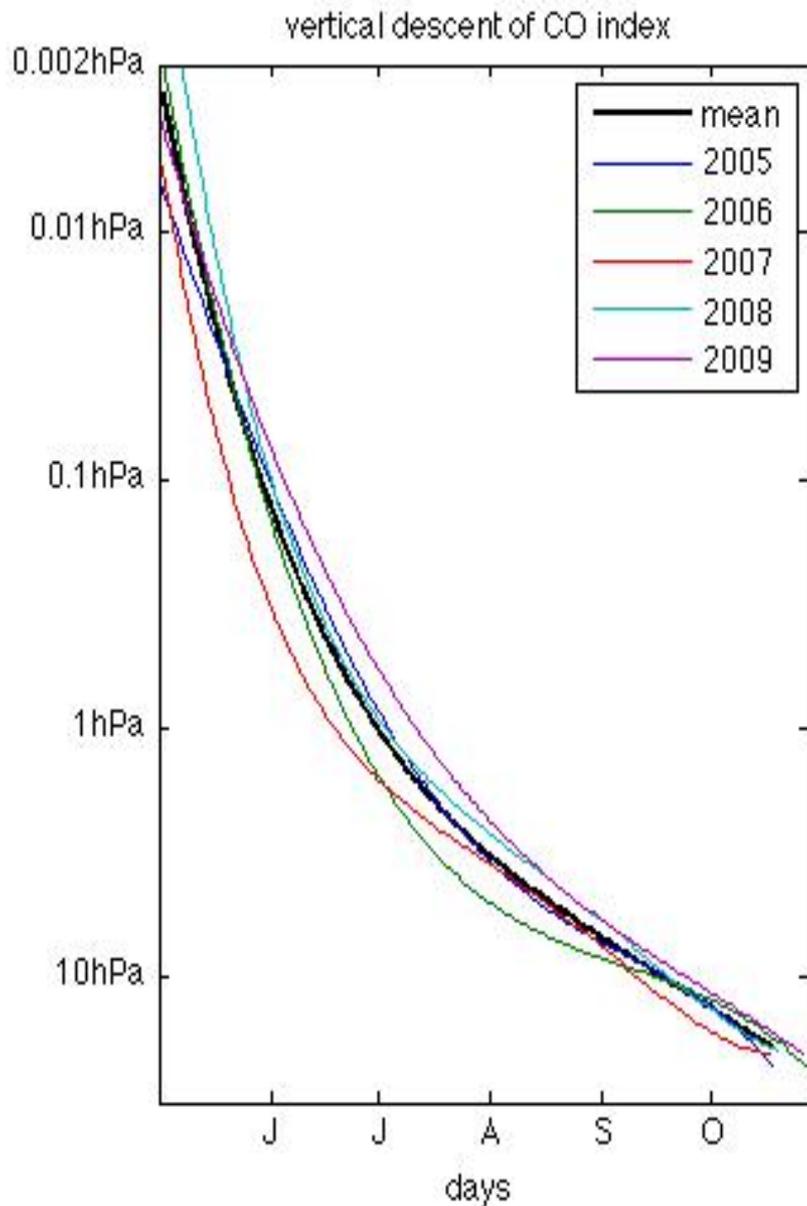


CO



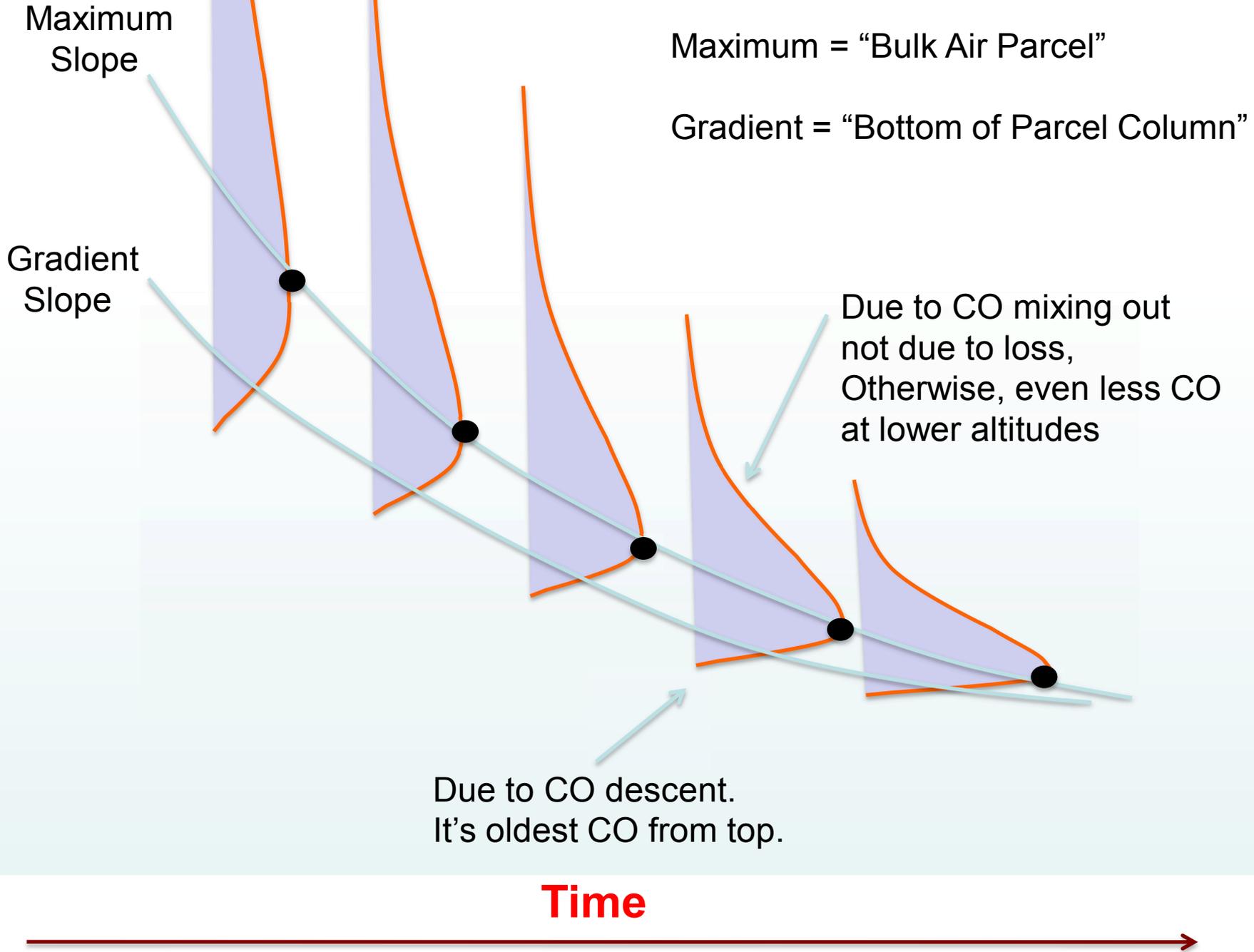
# 60° S-82° S zonal mean CO





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SORCE 2011, Sedona, Arizona



## Conclusions

- Strong coupling in the middle atmosphere
- Descent is fast ( $\sim 1\text{km/day}$ ) in the upper level ( $\sim 80\text{km}$ ) and slows down to  $\sim 0.2\text{km/day}$  at  $30\text{km}$
- Inter-annual variability in NH and SH

## Descent means?

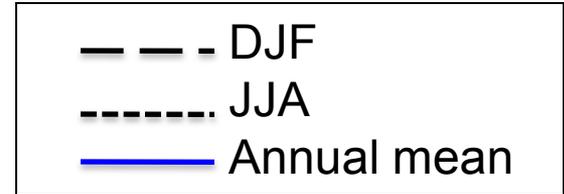
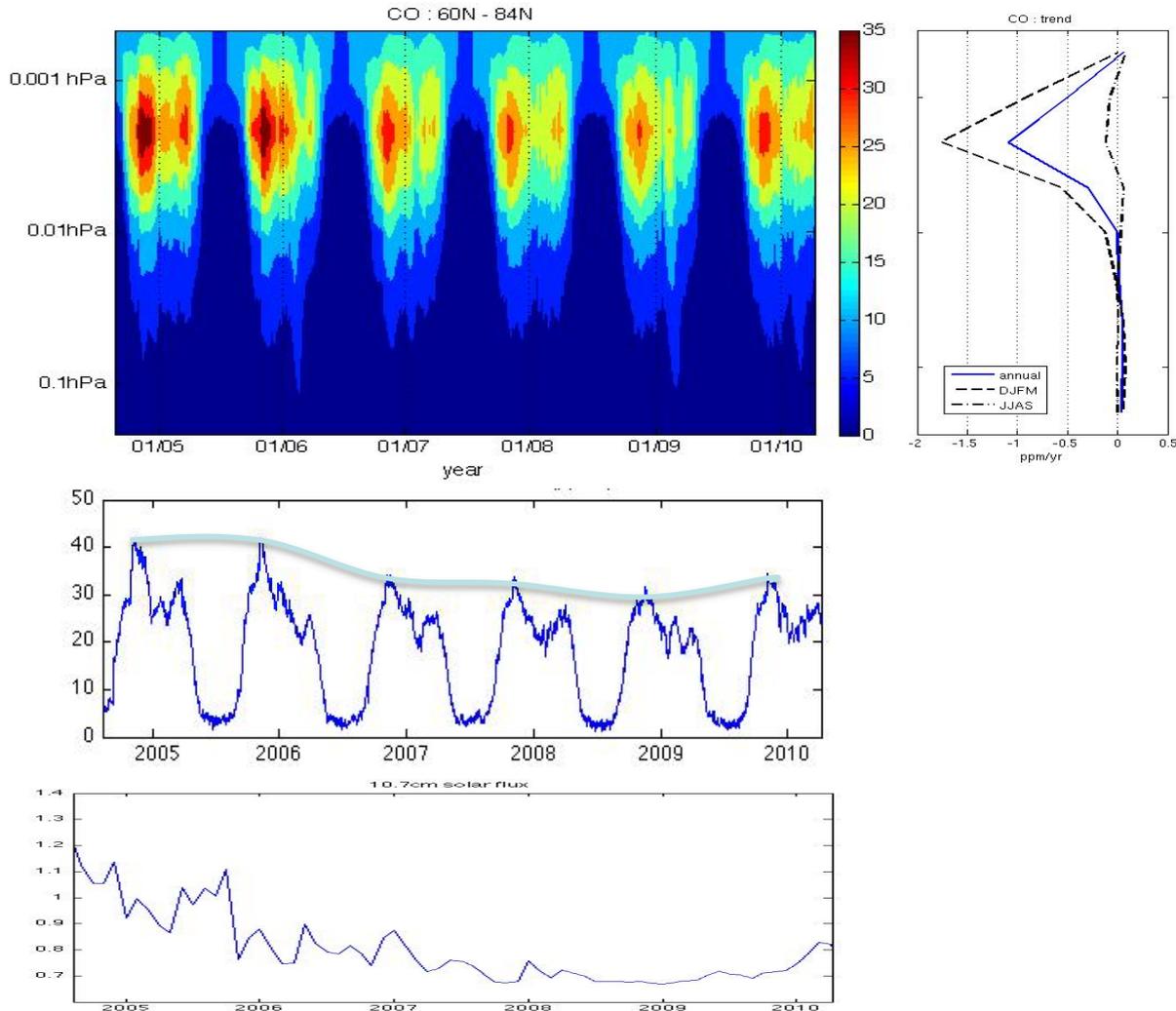
- Whole vortex descent?
- Constant mixing ratio gradient?
- Occurs at the center of air mass? Or at the bottom?
- Slope of the descent at the bottom of air mass is slower than that from the center of air mass.

# Future Plans

- Observational evidences of the solar signal in the middle and upper atmosphere
- Application of new remote sensing data to decadal variability
  - MLS and MEaSURES, Korean GEMS in conjunction with SORCE SIM
  - Stratosphere-troposphere coupling

# Zonal mean CO

## High Lat (60N-84N)

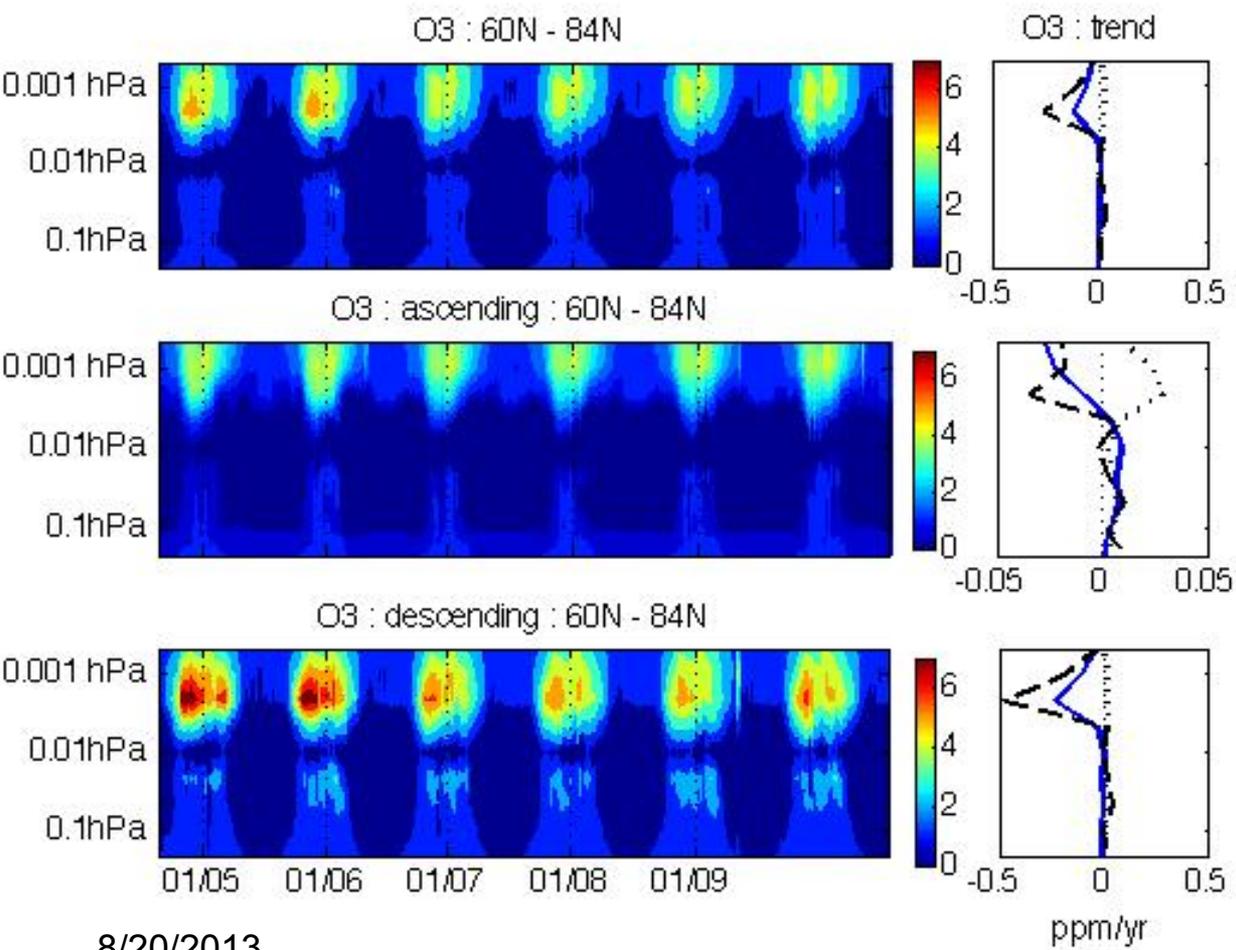
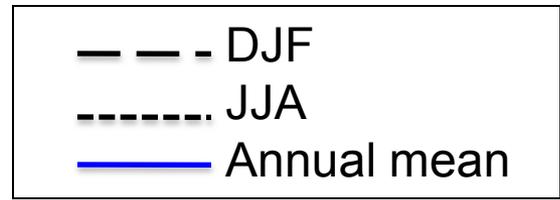


- CO decrease over 5 years during the winter ( 1.5ppmv/yr)



# Zonal mean O<sub>3</sub>

High Lat (60N-84N) in ppmv



- Decreasing trend since 2004 in the upper mesosphere (0.2ppmv/yr)

← day time

← night time

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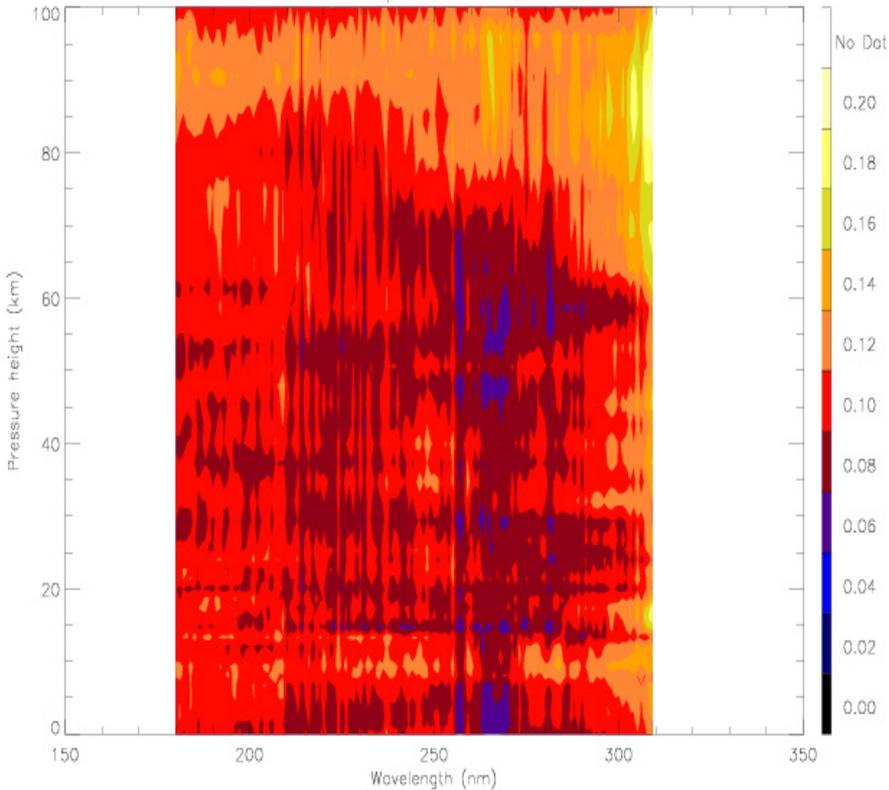
SORCE 2011, Sedona, Arizona

# Water Vapor : Correlation with spectral solar irradiance

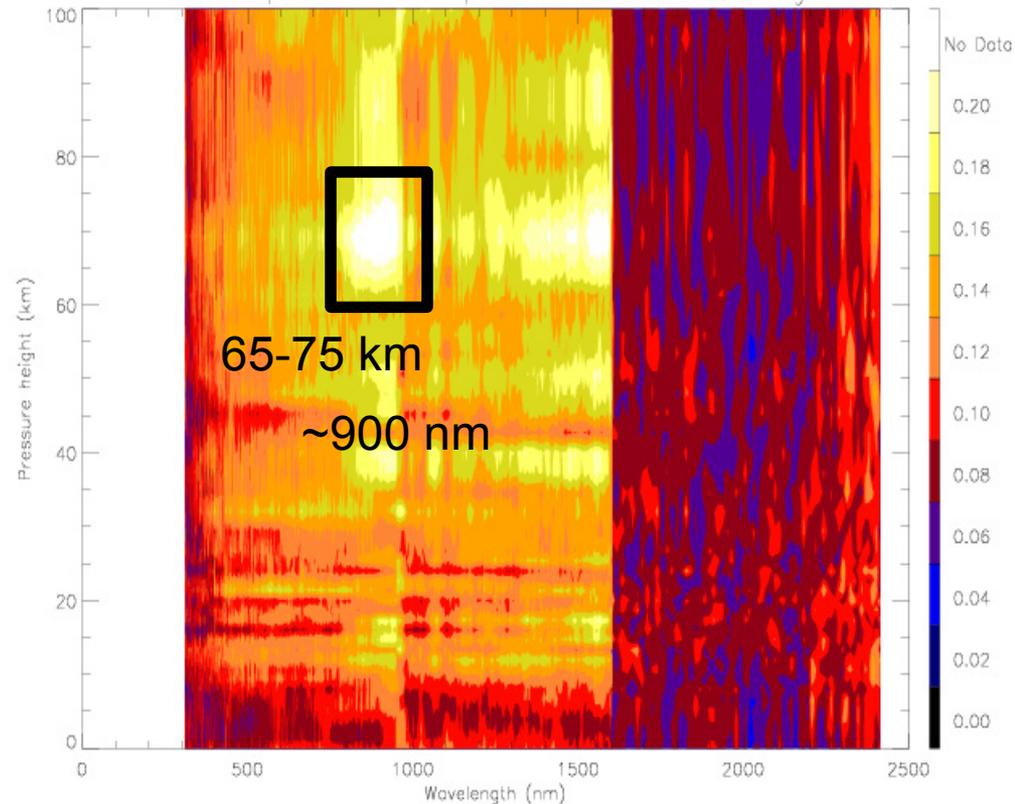
30N-60N H<sub>2</sub>O and SOLSTICE  
180-310 nm

30N-60N H<sub>2</sub>O and SIM  
310-2413 nm

Correlation Map between 30N and 60N



Water Vapor Correlation Map between 30N and 60N average



Courtesy : Greg Simonian

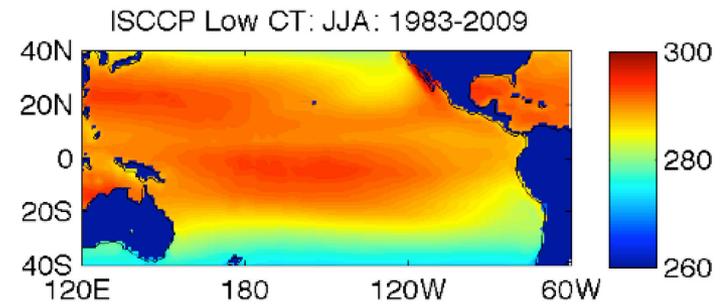
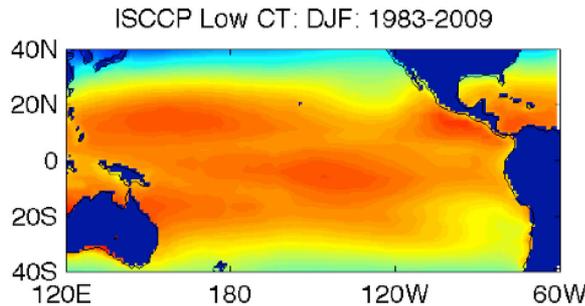
# Solar impact on cloud variability?

## ISCCP Low Cloud Top Temperature (1983-2009)

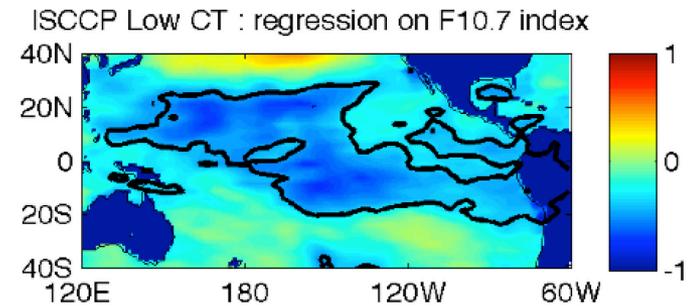
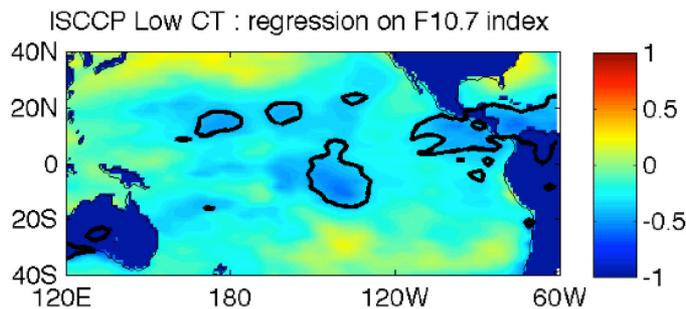
**DJF**

**JJA**

**mean**



**Regression on  
Solar F10.7  
index**



**Regression on  
Nino 3.4 index**

