

# The Quest for Other Worlds Like Earth - The Kepler Mission

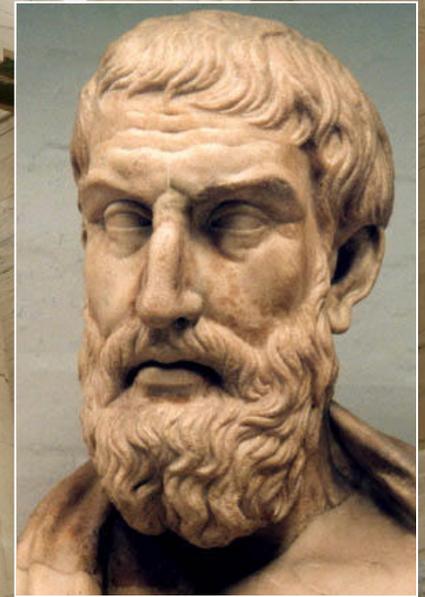
A large, bright yellow star with a small blue planet in the foreground, set against a dark background of stars.

Dr. Thomas Gautier  
Kepler Project Scientist  
Jet Propulsion Laboratory,  
California Institute of Technology.

# Speculation on the existence of other habitable worlds is age-old

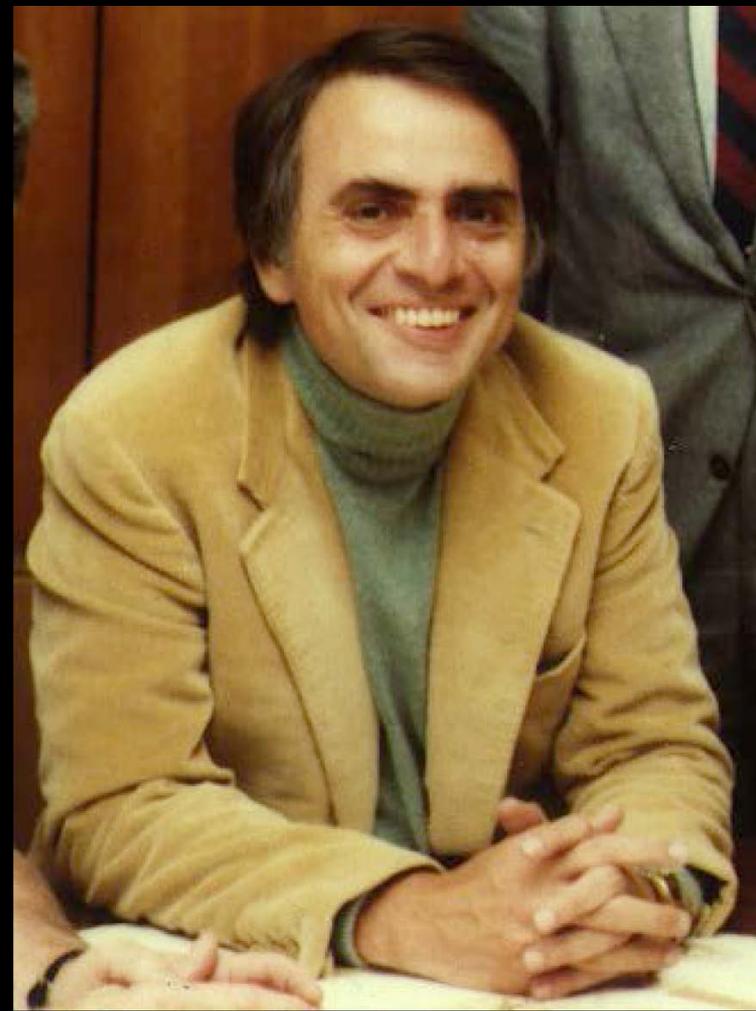
*There are infinite worlds both like and unlike this world of ours ... we must believe that in all worlds there are living creatures and plants and other things we see in this world....*

**Epicurus (341-270 BC)**



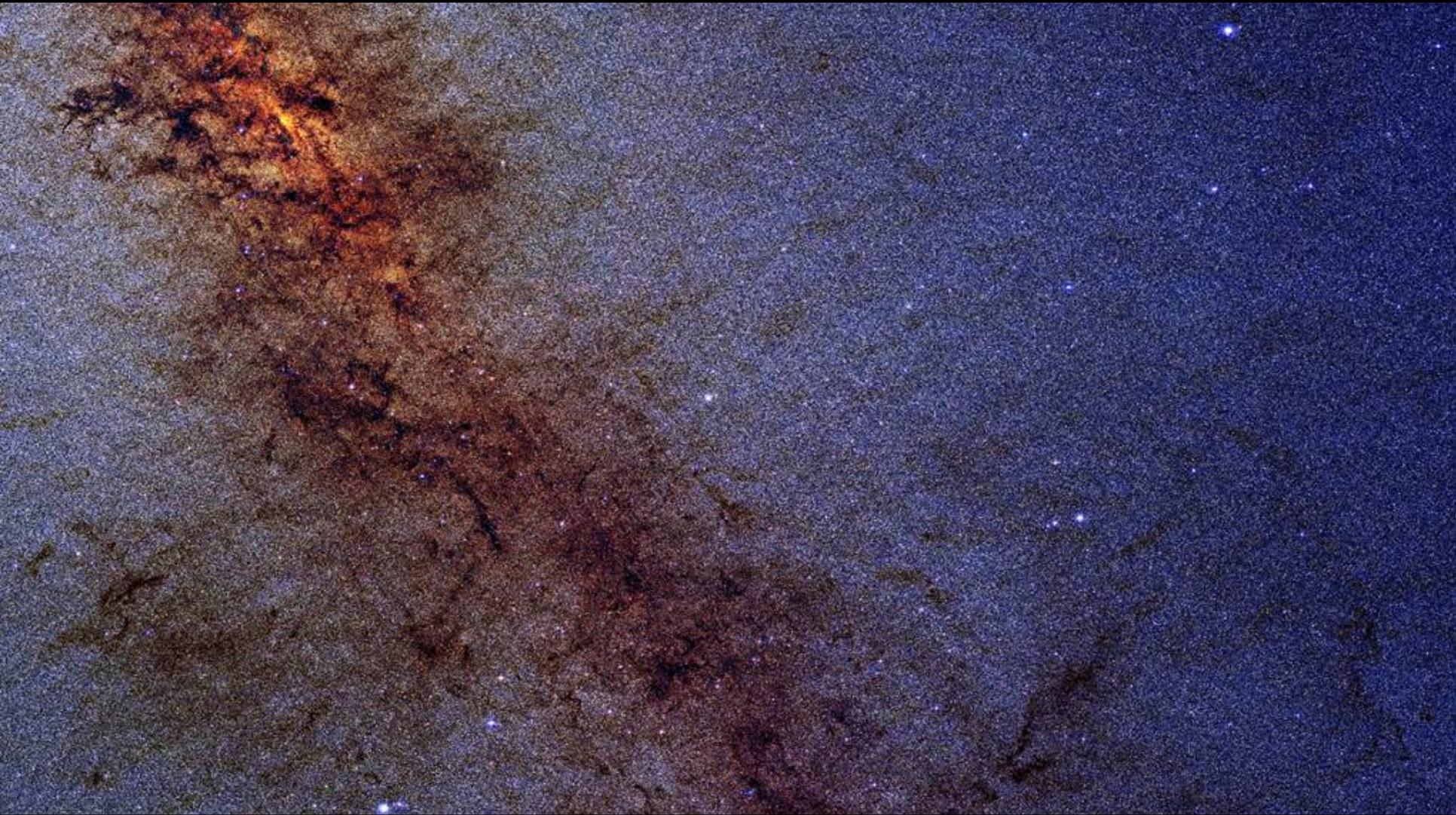
## And modern ...

- "There are perhaps as many as ten billion trillion [planets]. In the face of such overpowering numbers, what is the likelihood that only one ordinary star, the Sun, is accompanied by an inhabited planet?"
- "Why should we, tucked away in some forgotten corner of the Cosmos, be so fortunate? To me, it seems far more likely that the universe is brimming over with life."



*Carl Sagan - Cosmos (1980)*

*There are 200 billion stars in our galaxy...*



*There ought to be some other Earths.*

# How do we look for them?

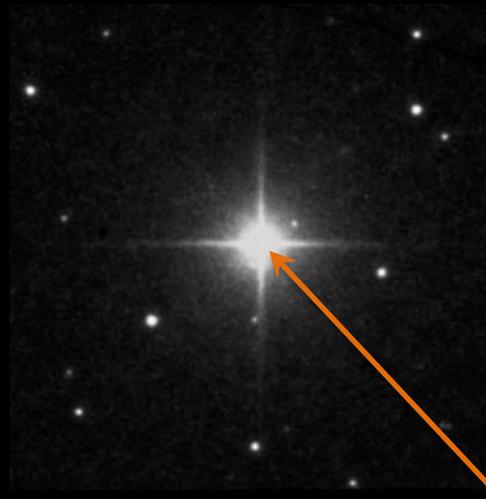
Take their pictures?

Problem: Stars are far away and REALLY bright.  
Planets are dim and close to their star.

Taking pictures of extrasolar planets is hard.

Even then can't see close in ones like Earth.

What about other ways?



Solar type star  
24 Sextans

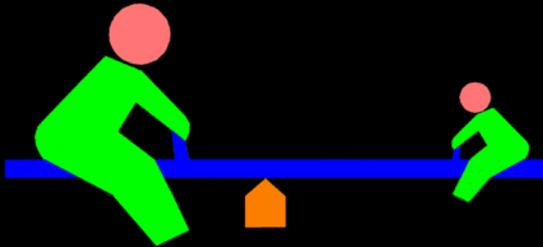
2 known planets  
About the same distance as Mars and Jupiter from our Sun

# Planets Move Stars

Planets and stars orbit each other, both circling their common center of gravity.

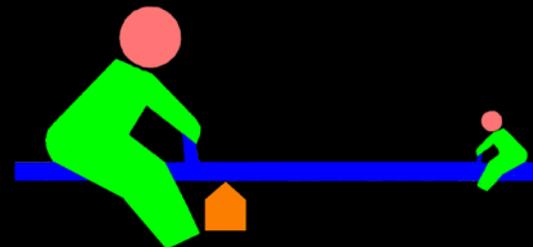
Larger planets pull the center of gravity farther from the star and make the star move in a larger orbit

Like kids on a see-saw:



Big kid (star) sits farther from balance point with medium size partner (planet).

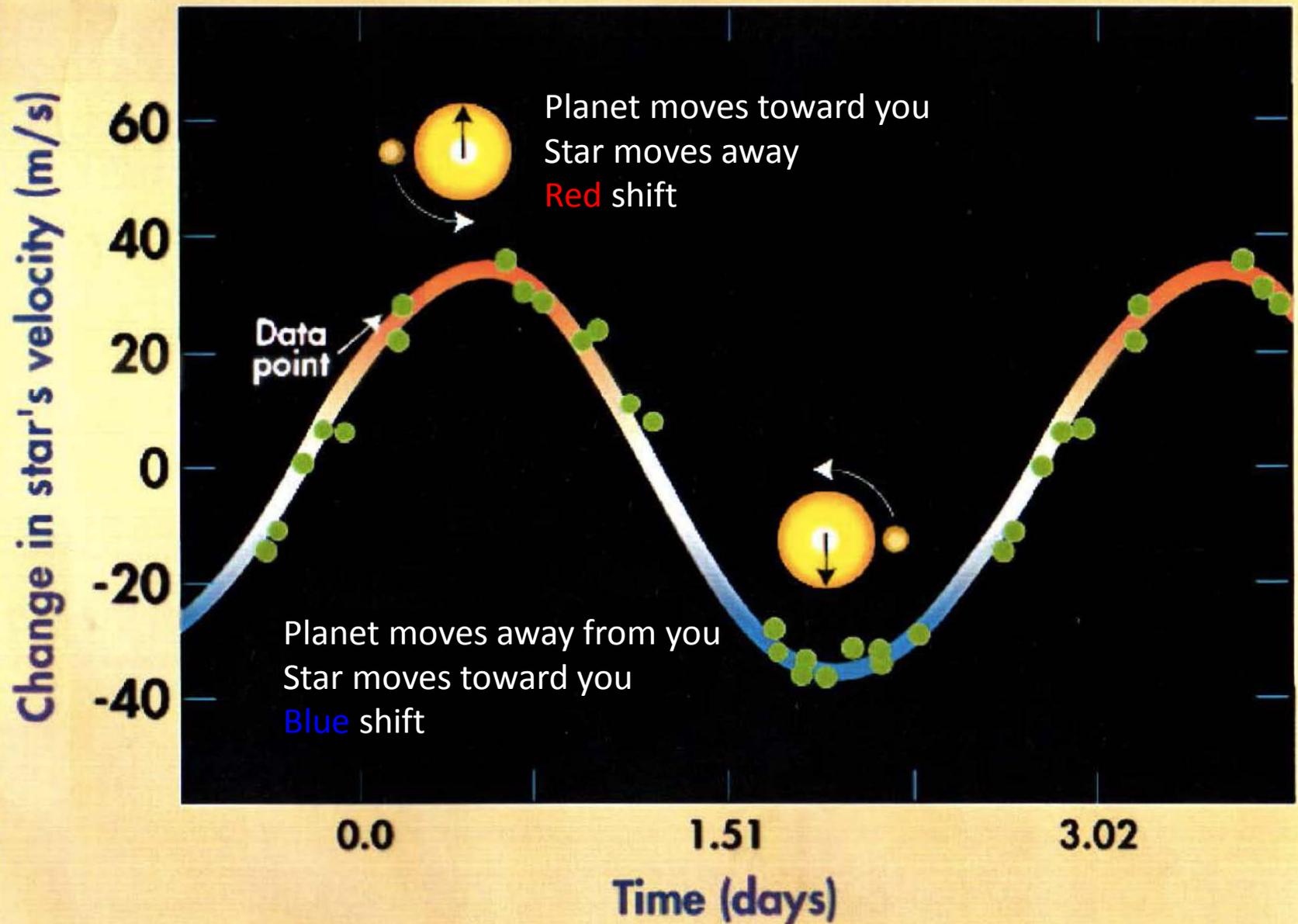
Moves more



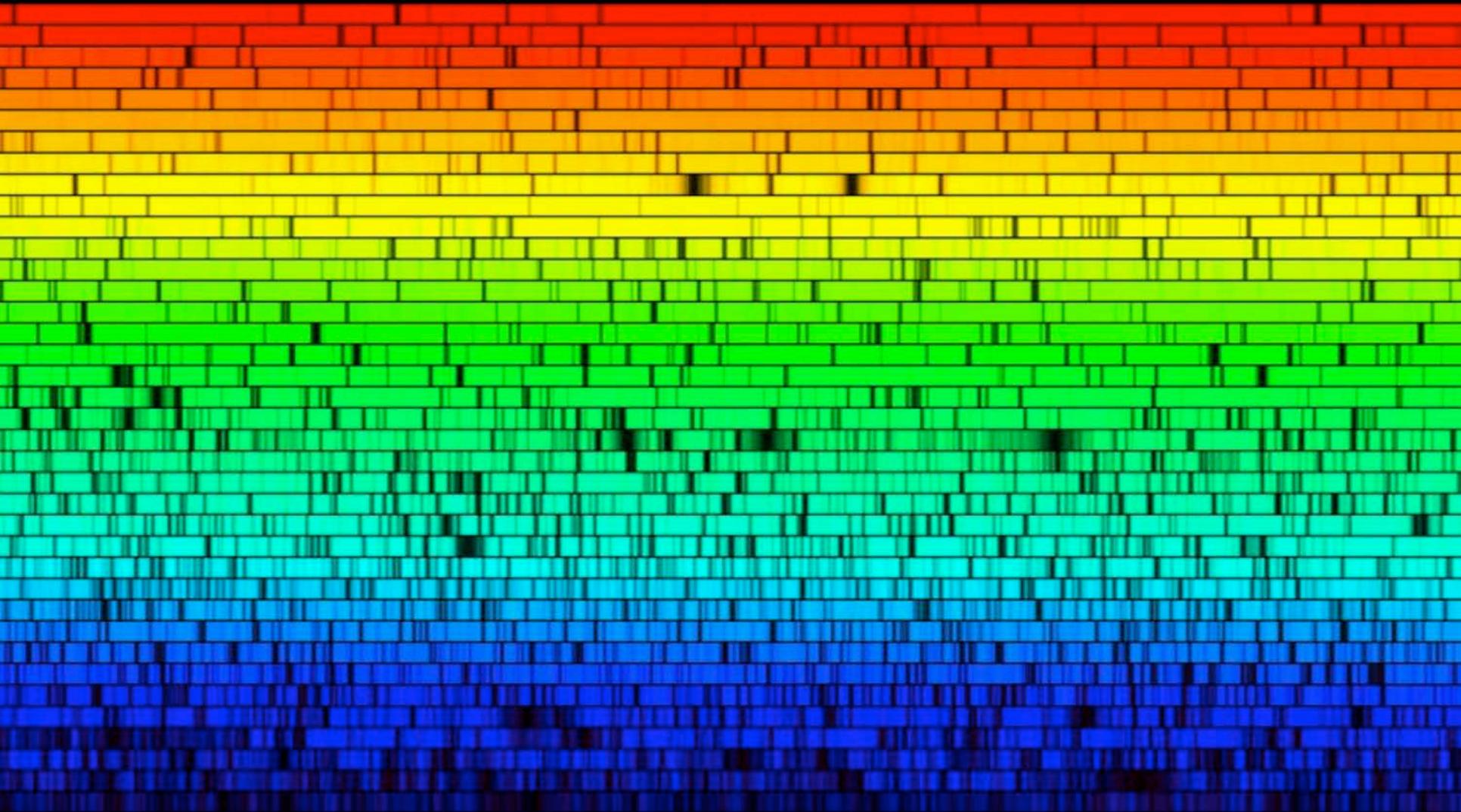
Big kid (star) sits closer to balance point with small size partner (planet).

Moves less

# We See this Motion in the Star's Spectrum



# Stellar Spectrum



**Radial Velocity measurement can find planet's MASS**

And distance from its star but not its size

**RV measurements good for finding large planets**

Also Earth-size around small, cool stars

Not good enough to find Earths around Sun-like stars

**Also too slow.**

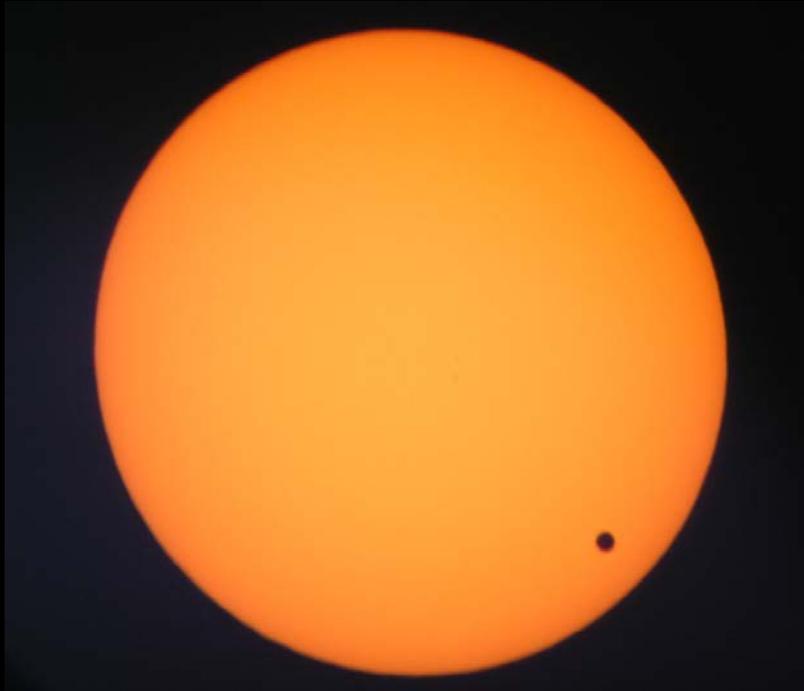
Can only do 1 star at a time.

Need many measurements totaling, maybe, hundreds, of hours

**Need another way to find Earths.**

But remember the RV method. It's important later.

# Planets Cast Shadows



## Planetary Transits

Sometimes a planet's orbit takes it in front of its star as seen from Earth.

For a few hours each orbit, it blocks a small amount of the star and making the star appear dimmer.

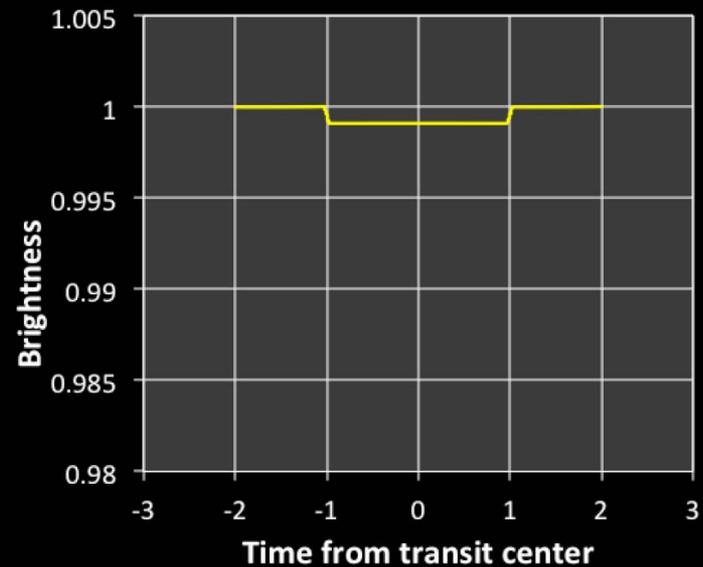
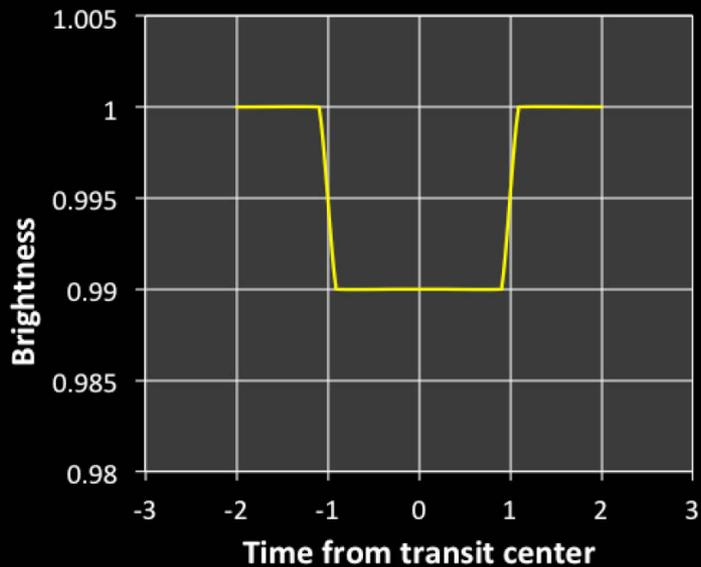
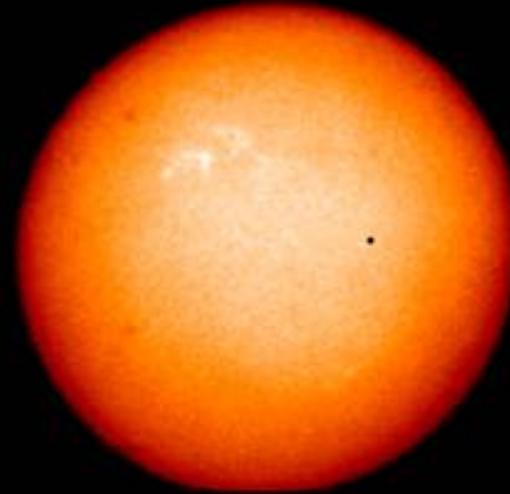
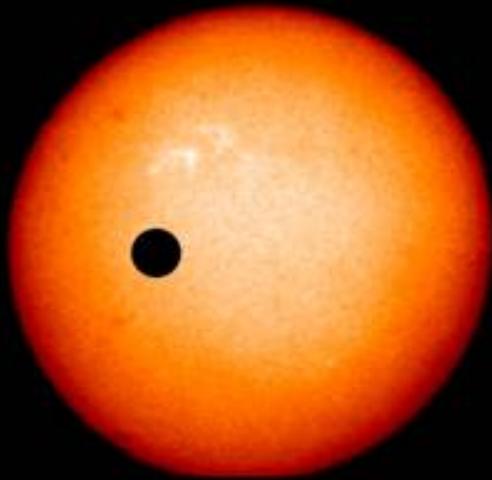
Transit measurements can find the size of the planet and distance from its star but not its mass.

# Transit Photometry





# Transits Have Different Depths



Simulated light curves

# Transits Won't Find ALL Planets

Not all planetary orbits are aligned to produce a transit

$$\text{) Range of Pole Positions} = \frac{d^*}{R_p}$$

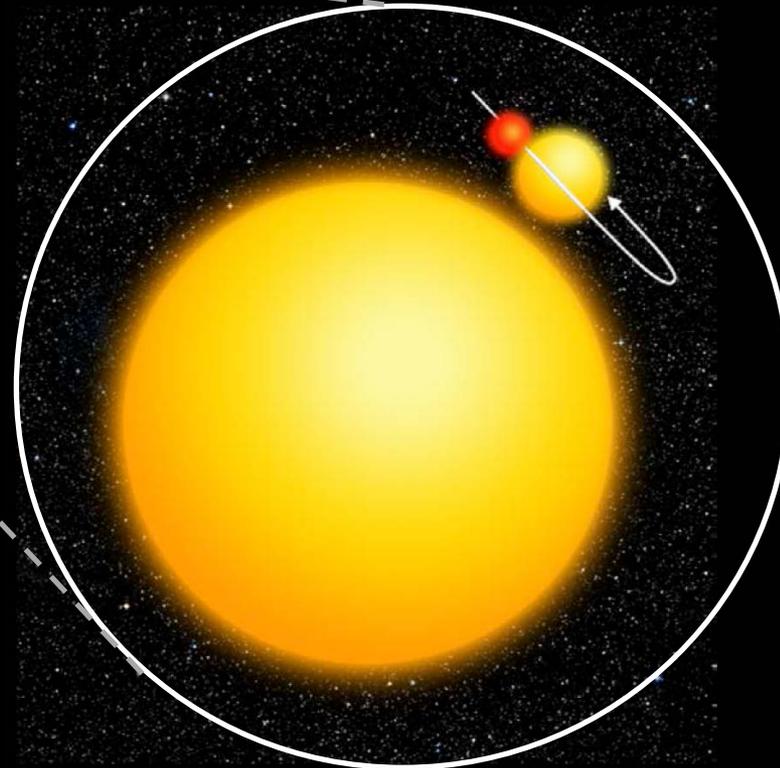
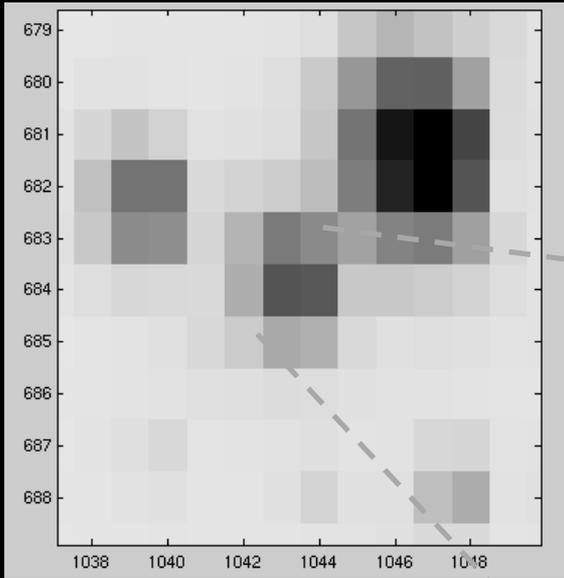
Only a few percent chance for a given planet to transit

Earth and Sun have only 0.5% chance

Transits last only a few hours each orbit

Must stare constantly at many thousands of stars to find many planets.

# A pitfall: False Positives



Other things can look like transiting planets.

Eclipsing binary stars behind target stars

Candidate transiting planets need to be carefully examined to look for false positives.

# Ways to See Exoplanets

**Star wobble (Radial Velocity)**  
distance from star

Mass, orbital period and shape:

**Star dimming (Transit photometry)**  
from star

Size, orbital period: distance

With both mass and size get planet  
density. Then can tell if the planet is:

Gas : Low density

Ice : Medium density

Rock : High density

How do we know it when we find  
another Earth?

# Perhaps a Habitable Planet?

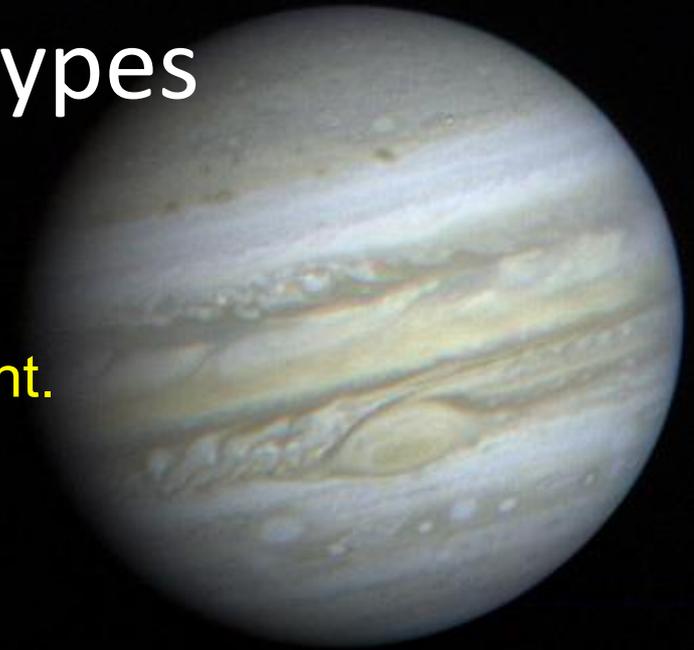
- Right temperature
- Liquid water
- Atmosphere



# Planets come in several types

## Too big (about $>10 M_{\text{earth}}$ )

Holds onto light gases (Hydrogen and Helium) too well and turns into a gas giant.  
(Jupiter, Saturn, Uranus, Neptune)



## Just right ( $0.5 - 10 M_{\text{earth}}$ )

Holds on to heavier gases (Nitrogen, Oxygen) long enough to have a potentially habitable atmosphere (Earth, Venus)



## Too small (about $<0.5 M_{\text{earth}}$ )

Can't hold onto a life sustaining atmosphere.  
(Moon, Mercury)

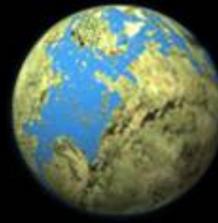


# And a range of temperatures



Too hot!

(too close)



Just right



Too cold!

(too far)

# The Habitable Zone

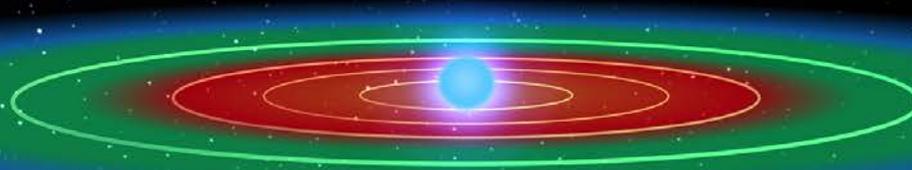
The region around a star where liquid water might exist on a planet's surface

The Goldilocks Zone

## Hotter Stars

F0

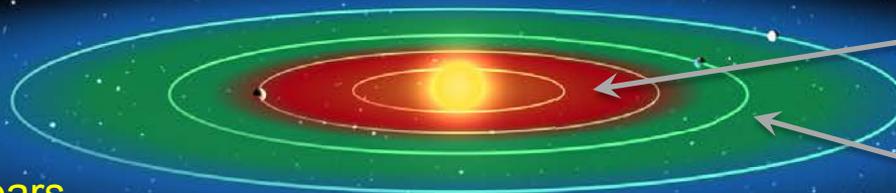
2 to 10 year orbits



## Sunlike Stars

G2

9 months to 3.5 years



Too hot

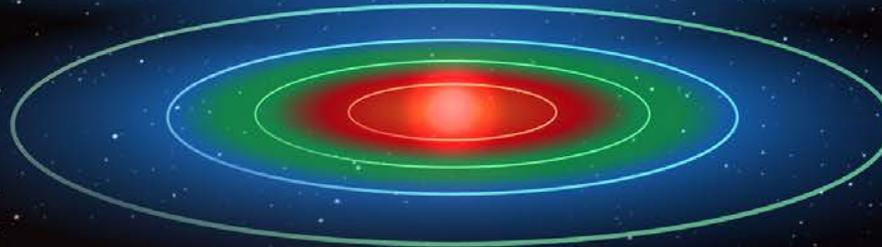
Too cold

Just right

## Cooler Stars

K5

3 to 13 months



# History of Exoplanet Discovery

- **First surveys looked for wobble with RV**
  - 1989: First exoplanet discovered
    - Very hot gas giant in close orbit around solar-like star
- **Many more discoveries followed**
  - More RV discoveries, ground based transit surveys, direct imaging, pulsar timing
  - By **2009**: **332** exoplanets known, **61** with measured size and density
  - **None like Earth**. Mostly too big and too hot. Some too cold.



# Why no Earths?

- **Low sensitivity**
  - RV wobble measurements not sensitive enough for low mass planets
  - Ground based transit surveys can't see small enough planets
- **Haven't looked at enough stars**
  - RV surveys are slow: one star at a time
  - Long periods required for habitable zones make planets hard to catch and make transit line-up probability low

) Range of Pole Positions =  $\frac{d^*}{R_p}$

# So, what to do?

Build a special transit detecting telescope with high sensitivity that can look at lots of stars at once –

Kepler

# Kepler Telescope-Photometer

Wide field for lots of stars

15° diameter field

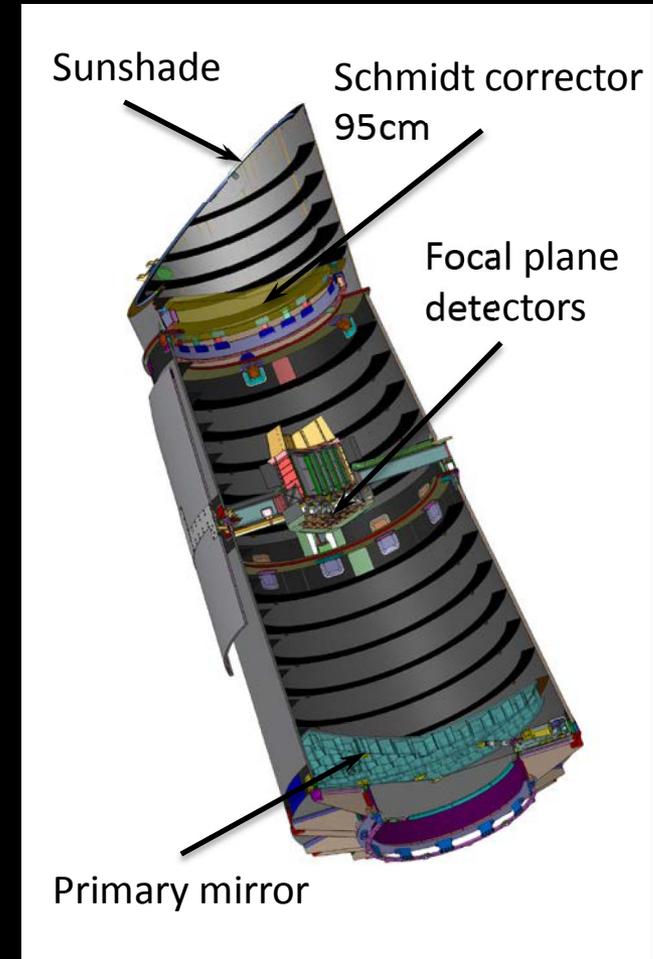
Monitor 150,000 stars at once

Big, to get enough light

95 cm aperture

Build it for use in space

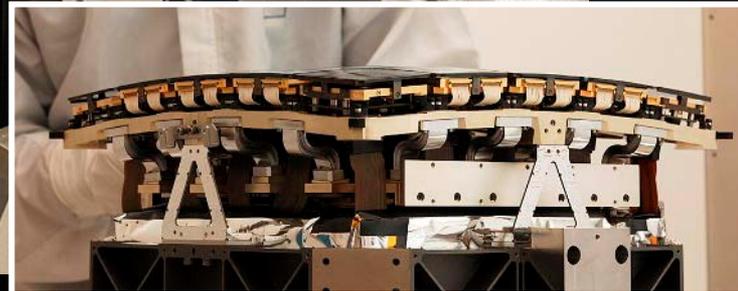
No variability from atmosphere to mask small transits



# Kepler's 95 Megapixel Camera



>100 square degree  
field of view



# Kepler in final assembly



# Kepler Launched March 6, 2009



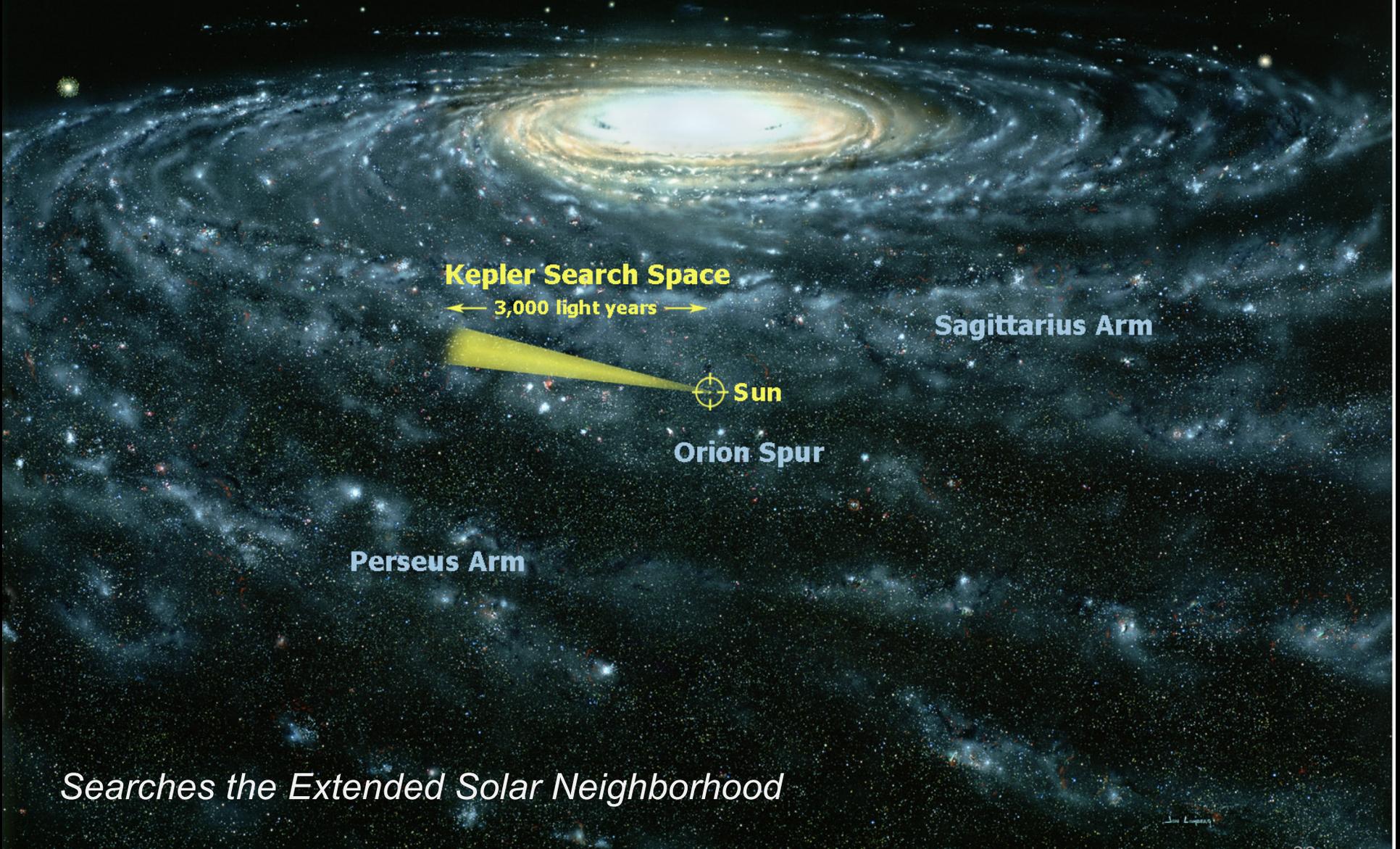
©2009 Ben Cooper  
<http://www.launchphotography.com>



Kepler  
Field of  
View



# Milky Way Galaxy



**Kepler Search Space**

← 3,000 light years →

**Sagittarius Arm**

**Sun**

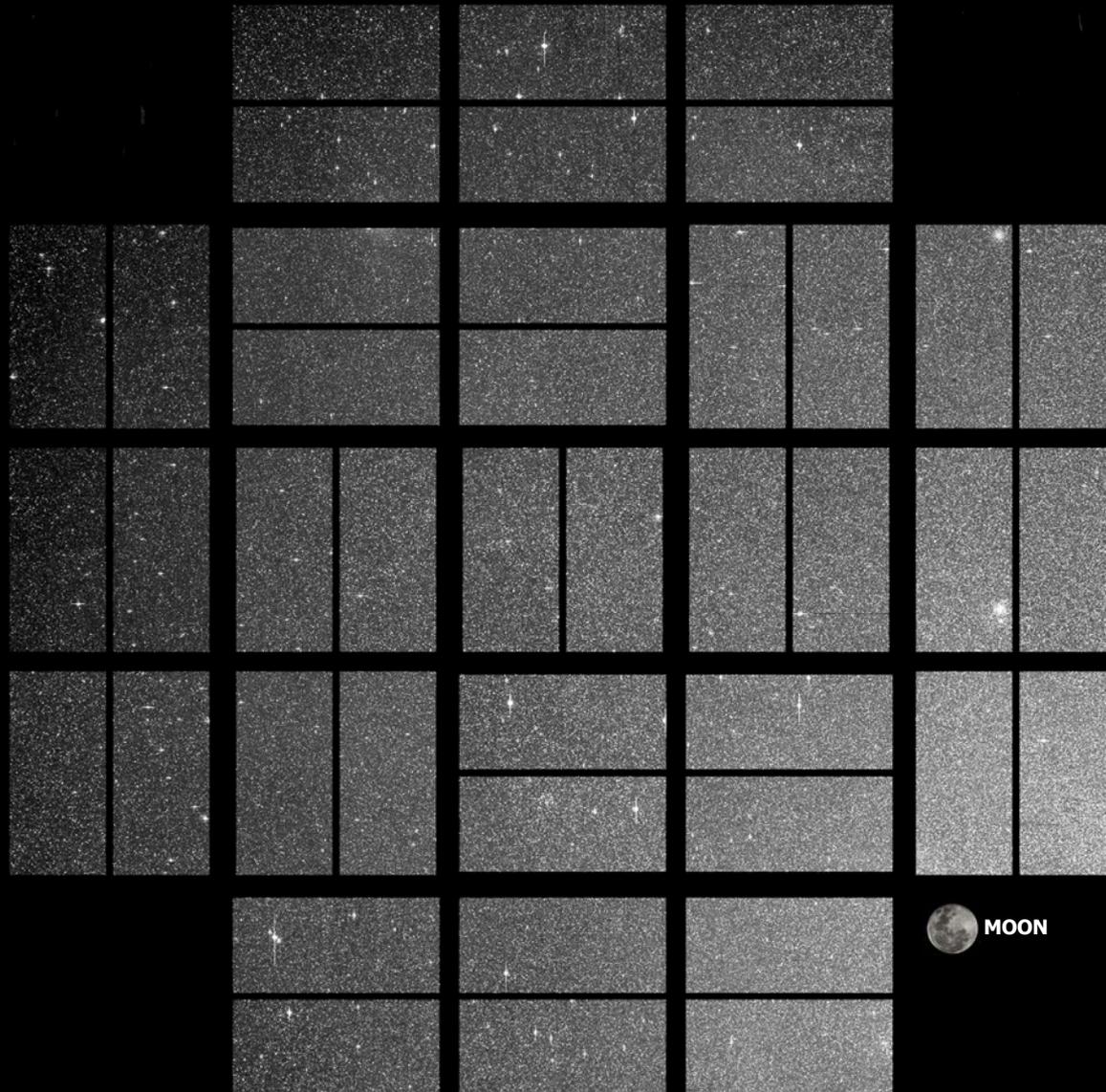
**Orion Spur**

**Perseus Arm**

*Searches the Extended Solar Neighborhood*

Jon Lomberg

# Kepler Full Field Image



# What has Kepler found?

Yesterday 763 exoplanets were known. 70 of those from Kepler observations

Still no habitable zone Earth-size planets

But, Kepler has found LOTS more planet candidates: Over 3000 with more coming

90% of candidates are expected to be real planets

Expect Kepler to announce ~800 new confirmed planets by fall.

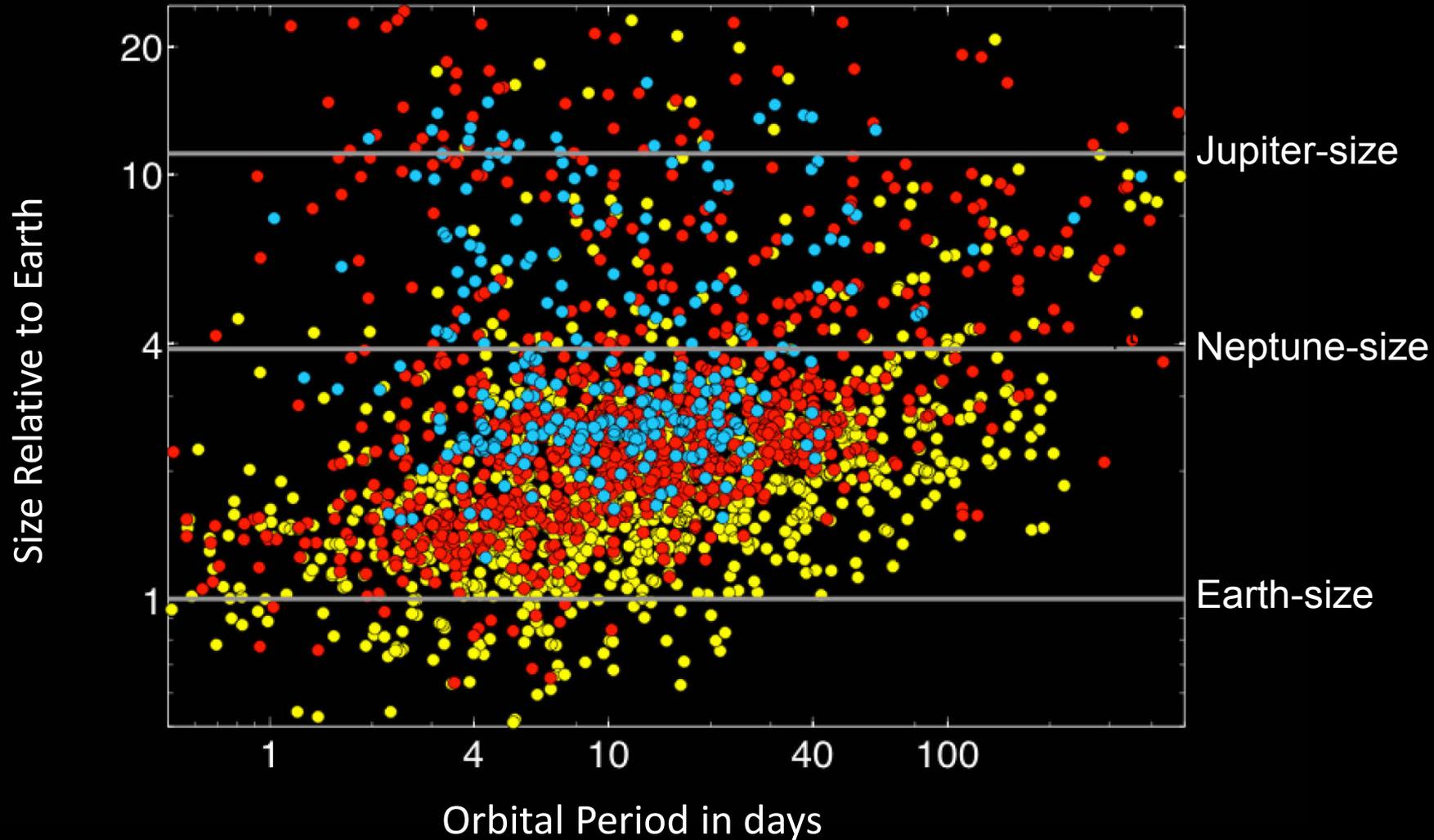
# Candidates as of Feb 2012

2312 total

● Jun 2010

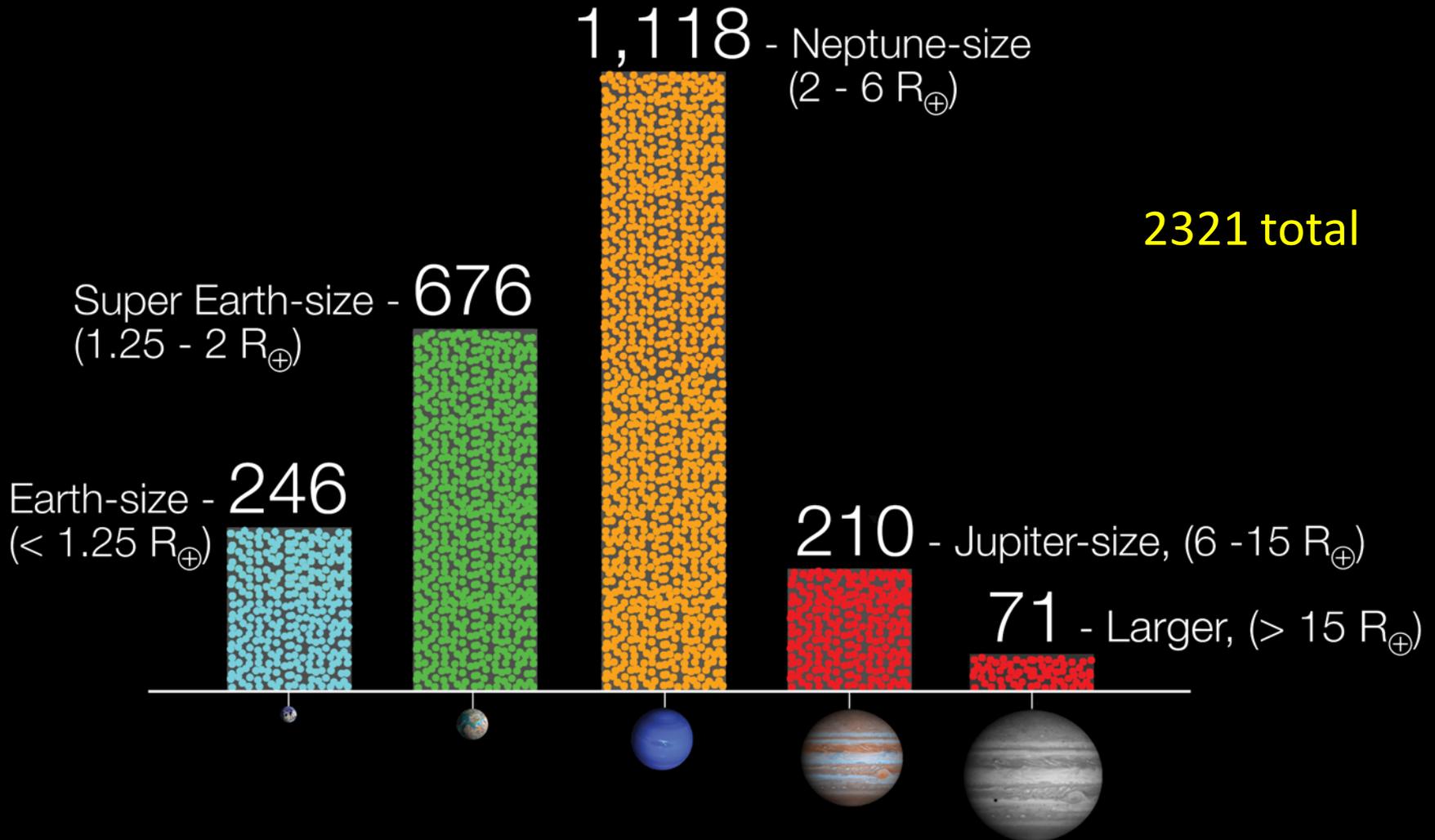
● Feb 2011

● Feb 2012



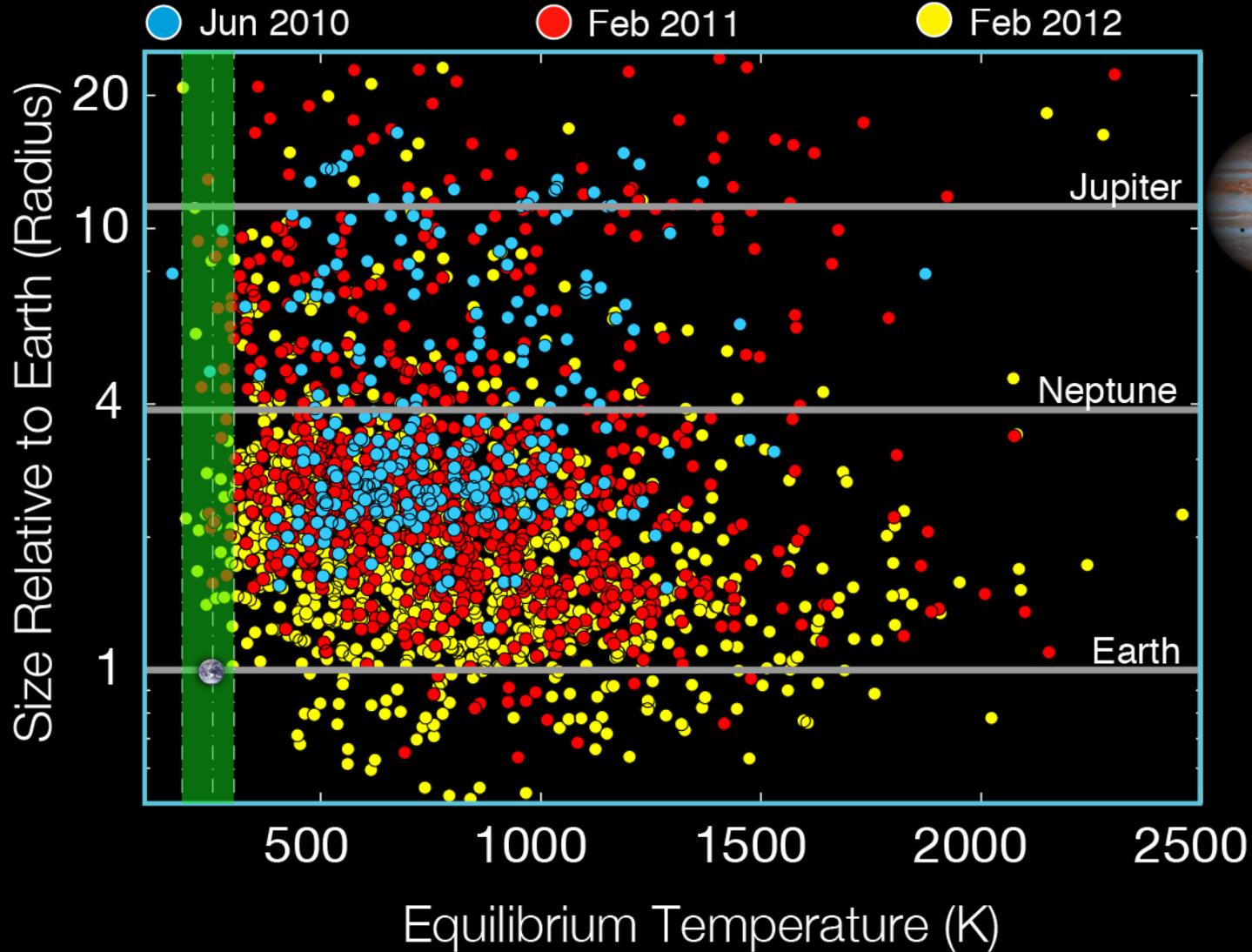
# Numbers of Planet Candidates

*As of February 27, 2012*



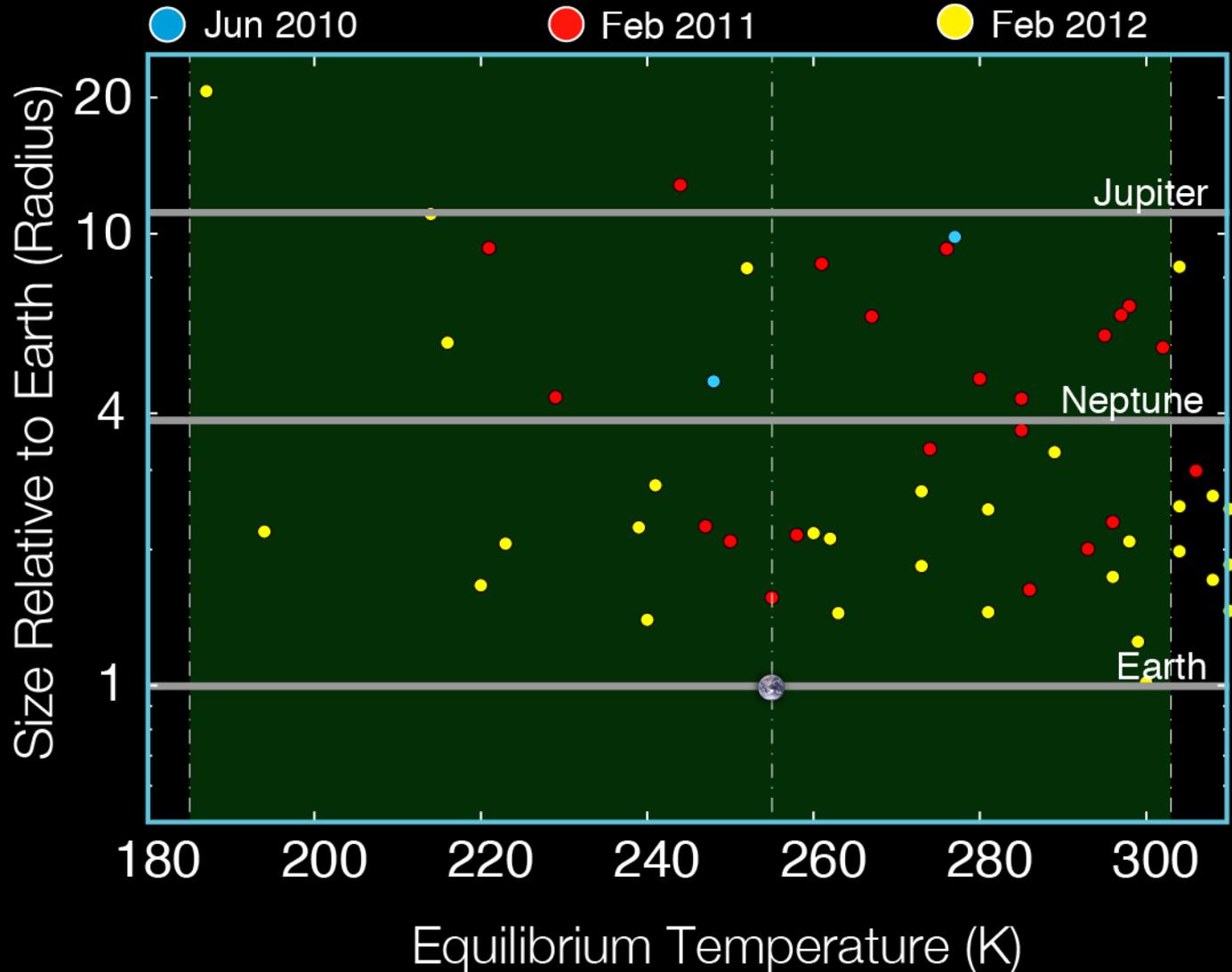
# Candidates in the Habitable Zone

*As of February 27, 2012*



# HZ Candidates

48 with  $T_{eq}$  between 185 and 303 K



# Earth: Our Reference Planet

## Known habitable

Size:  $1.0 R_{\text{Earth}}$   
Density:  $5.5 \text{ g/cm}^3$   
Surface: rocky/water  
Atmosphere: thin  
Location: habitable zone  
Temperature: 57F (287K)



# Kepler Planets

As of February 27, 2012

Kepler-14b



12.7 R<sub>E</sub>

Kepler-30c



14.4 R<sub>E</sub>

Kepler-17b



14.7 R<sub>E</sub>

Kepler-6b



14.79 R<sub>E</sub>

Kepler-8b



15.86 R<sub>E</sub>

Kepler-5b



16.00 R<sub>E</sub>

Kepler-7b



16.52 R<sub>E</sub>

Kepler-12b



19.0 R<sub>E</sub>

Kepler-27c



4.9 R<sub>E</sub>

Kepler-18c



5.49 R<sub>E</sub>

Kepler-18d



6.98 R<sub>E</sub>

Kepler-35b



8.16 R<sub>E</sub>

Kepler-16b



8.45 R<sub>E</sub>

Kepler-34b



8.56 R<sub>E</sub>

Kepler-9c



9.2 R<sub>E</sub>

Kepler-9b



9.4 R<sub>E</sub>

Kepler-30d



10.7 R<sub>E</sub>

Kepler-15b



10.8 R<sub>E</sub>

Jupiter



11.2 R<sub>E</sub>

Kepler-32c



3.7 R<sub>E</sub>

Kepler-33f



3.83 R<sub>E</sub>

Neptune



3.88 R<sub>E</sub>

Kepler-4b



3.99 R<sub>E</sub>

Kepler-27b



4.0 R<sub>E</sub>

Kepler-32b



4.1 R<sub>E</sub>

Kepler-31c



4.2 R<sub>E</sub>

Kepler-31b



4.3 R<sub>E</sub>

Kepler-25c



4.5 R<sub>E</sub>

Kepler-11e



4.52 R<sub>E</sub>

Kepler-33d



4.56 R<sub>E</sub>

Kepler-11c



3.15 R<sub>E</sub>

Kepler-23c



3.2 R<sub>E</sub>

Kepler-28c



3.4 R<sub>E</sub>

Kepler-11d



3.43 R<sub>E</sub>

Kepler-33e



3.45 R<sub>E</sub>

Kepler-26b



3.6 R<sub>E</sub>

Kepler-26c



3.6 R<sub>E</sub>

Kepler-28b



3.6 R<sub>E</sub>

Kepler-29b



3.6 R<sub>E</sub>

Kepler-11g



3.66 R<sub>E</sub>

Kepler-30b



3.7 R<sub>E</sub>

Kepler-19b



2.21 R<sub>E</sub>

Kepler-10c



2.23 R<sub>E</sub>

Kepler-22b



2.38 R<sub>E</sub>

Kepler-24b



2.4 R<sub>E</sub>

Kepler-25b



2.6 R<sub>E</sub>

Kepler-11f



2.61 R<sub>E</sub>

Kepler-33c



2.75 R<sub>E</sub>

Kepler-20d



2.75 R<sub>E</sub>

Kepler-24c



2.8 R<sub>E</sub>

Kepler-29c



2.9 R<sub>E</sub>

Kepler-20c



3.07 R<sub>E</sub>

Kepler-20e



.87 R<sub>E</sub>

Earth



Kepler-20f



1.03 R<sub>E</sub>

Kepler-10b



1.42 R<sub>E</sub>

Kepler-33b



1.5 R<sub>E</sub>

Kepler-21b



1.64 R<sub>E</sub>

Kepler-9d



1.64 R<sub>E</sub>

Kepler-23b



1.9 R<sub>E</sub>

Kepler-20b



1.91 R<sub>E</sub>

Kepler-11b



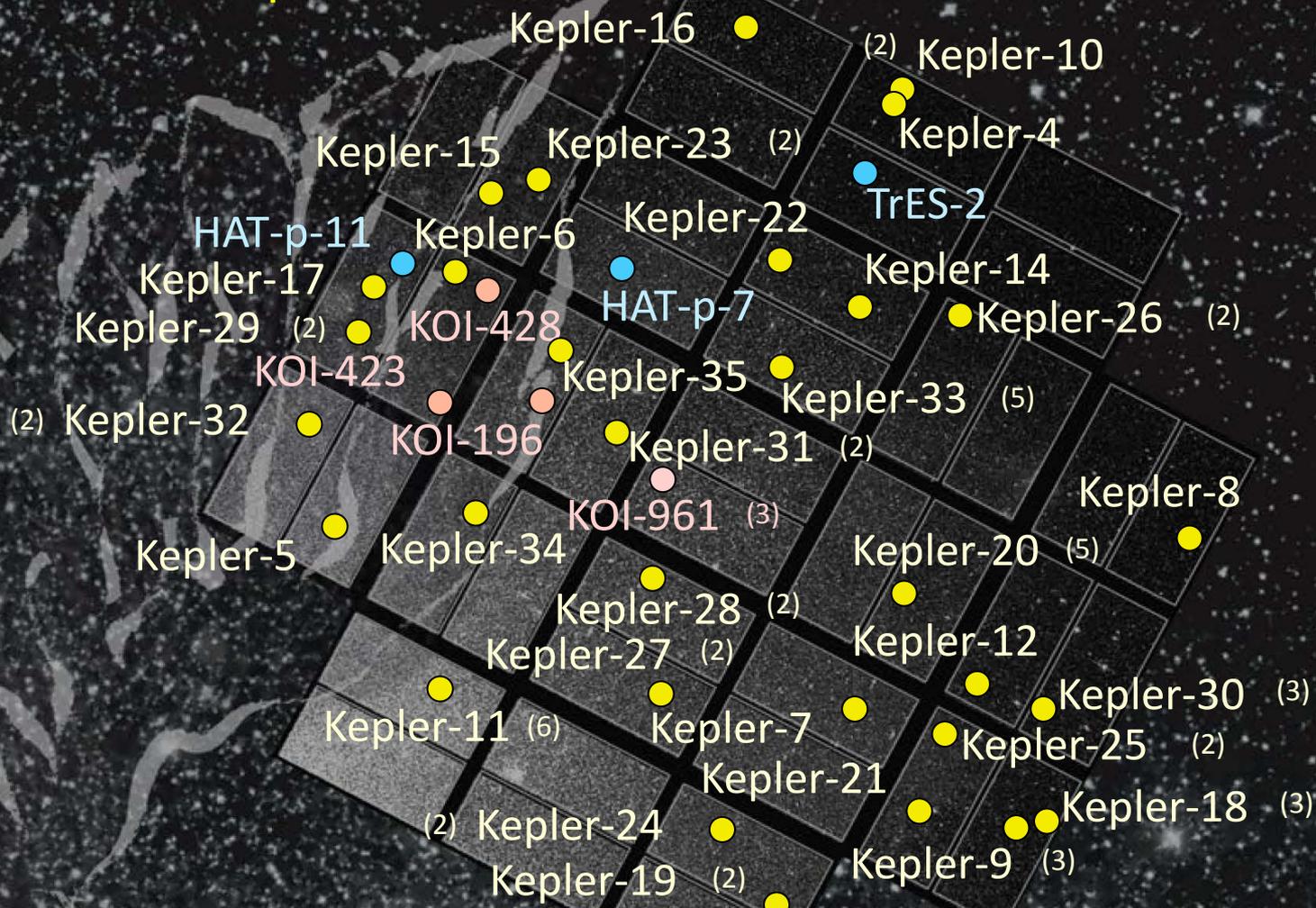
1.97 R<sub>E</sub>

Kepler-18b



2.0 R<sub>E</sub>

# 73 confirmed planets



61 from the Kepler Team  
9 from other astronomers  
3 known before Kepler

Kepler-16



(2)

Kepler-10



## Kepler-10b

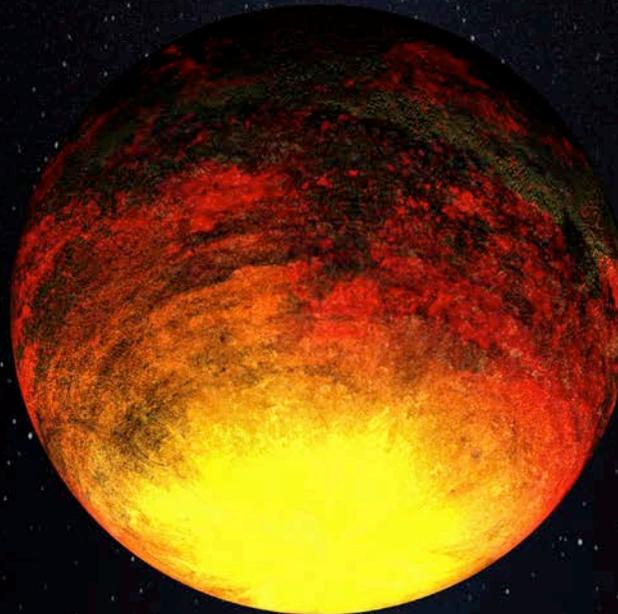
Rocky planet around a star a lot like the Sun  
560 light-years away

1.4 x Earth size

4.6 x Earth mass

Definitely rocky

(8.8 g/cm<sup>2</sup>, denser than iron)



(3)

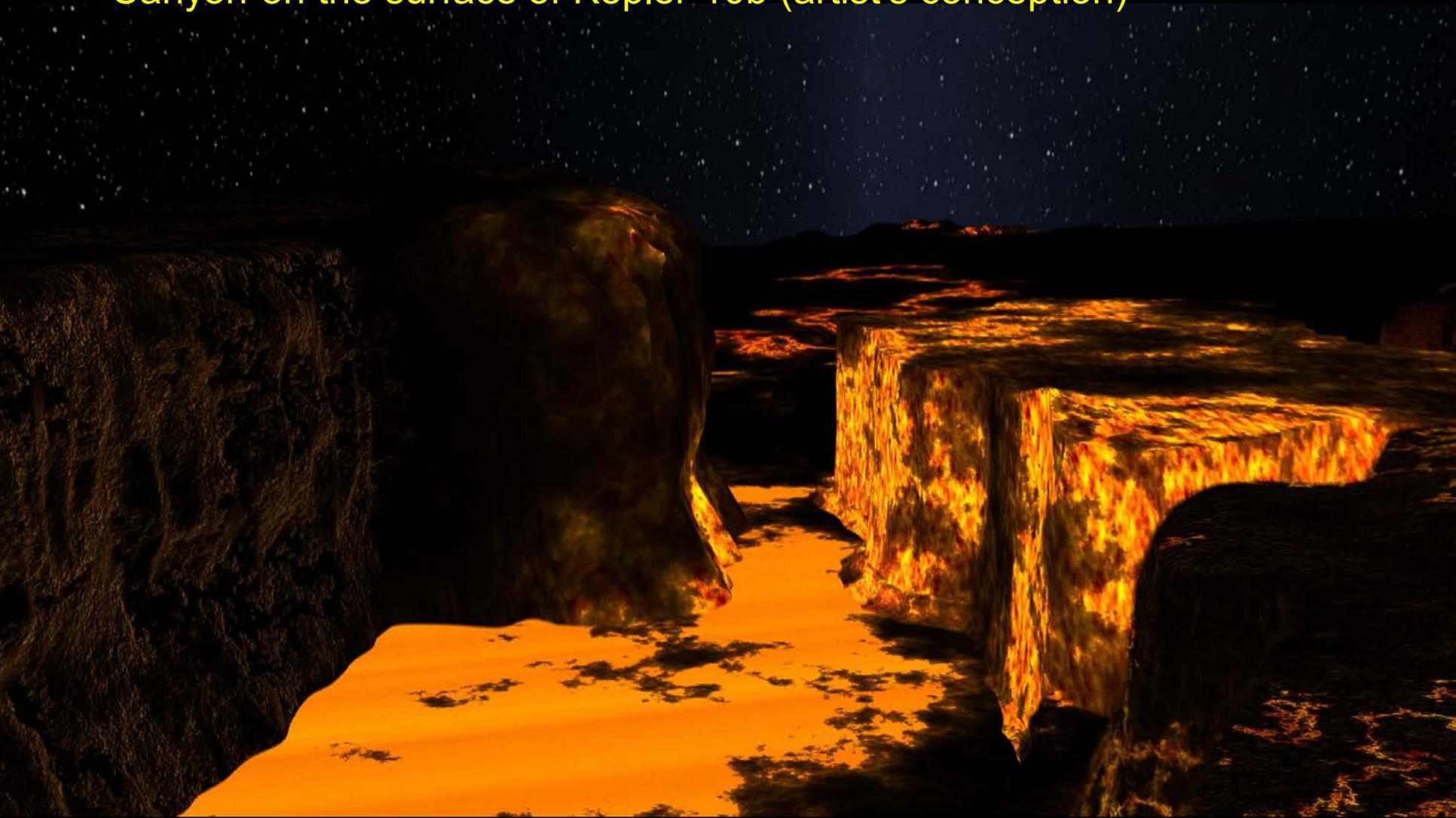
(3)

70 x closer to its star than Earth

19 hour orbit

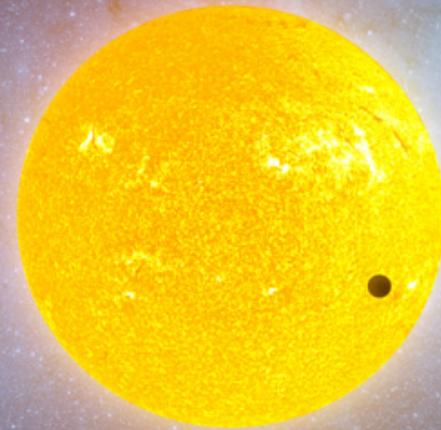
Scorching hot 2800 F (1833K) surface temperature

# Canyon on the surface of Kepler 10b (artist's conception)





# Kepler-10c



Companion to Kepler-10b  
About Neptune-size  
 $\leq$  Neptune mass  
42 day orbit  
410 F temperature

# Kepler-16



(2) Kepler-10

## Kepler-16b

Saturn-size planet

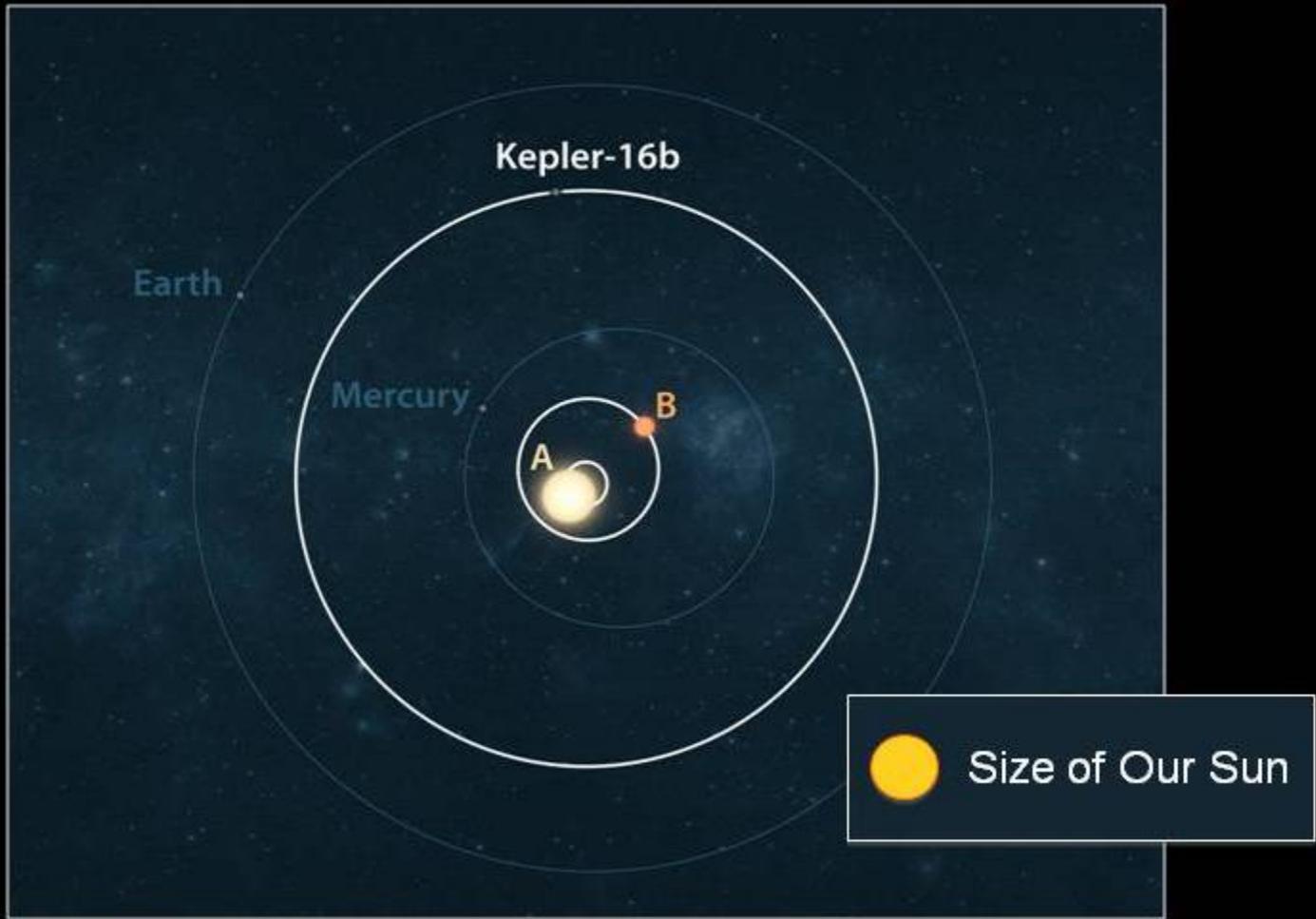
Circumbinary orbit

Like Tatooine in Star Wars

41 day orbit

Just outside of HZ: -126 F temperature

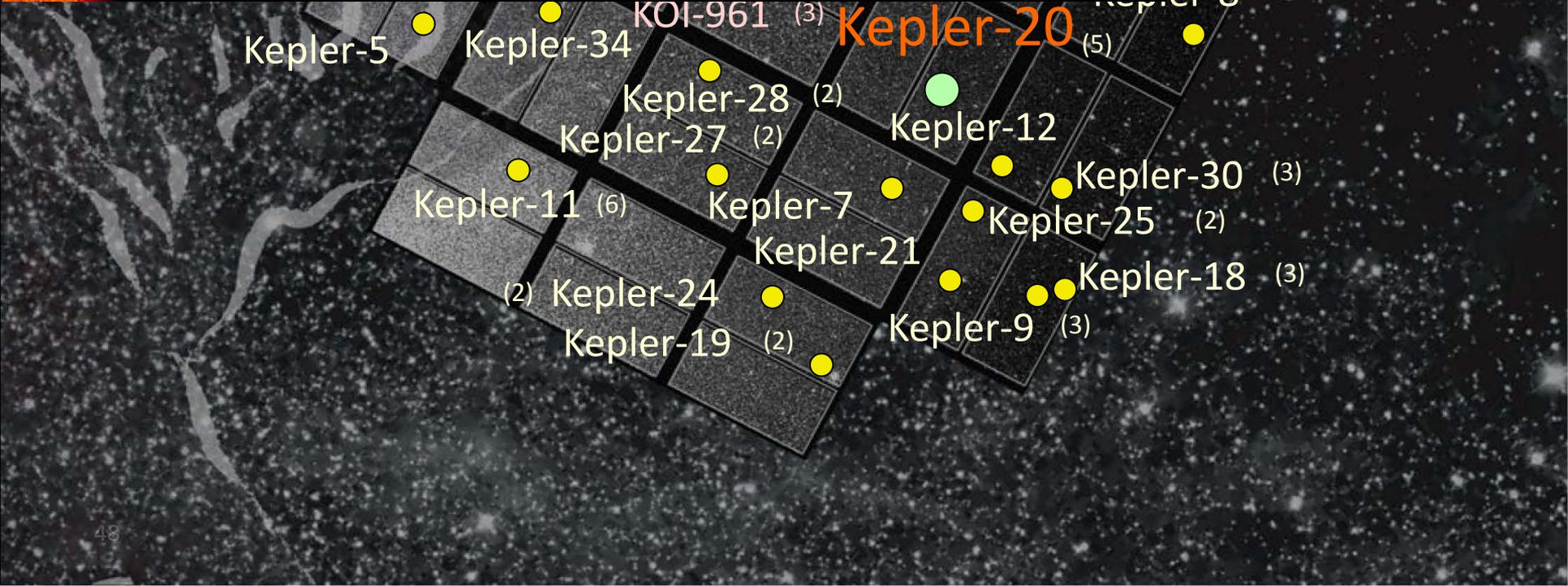
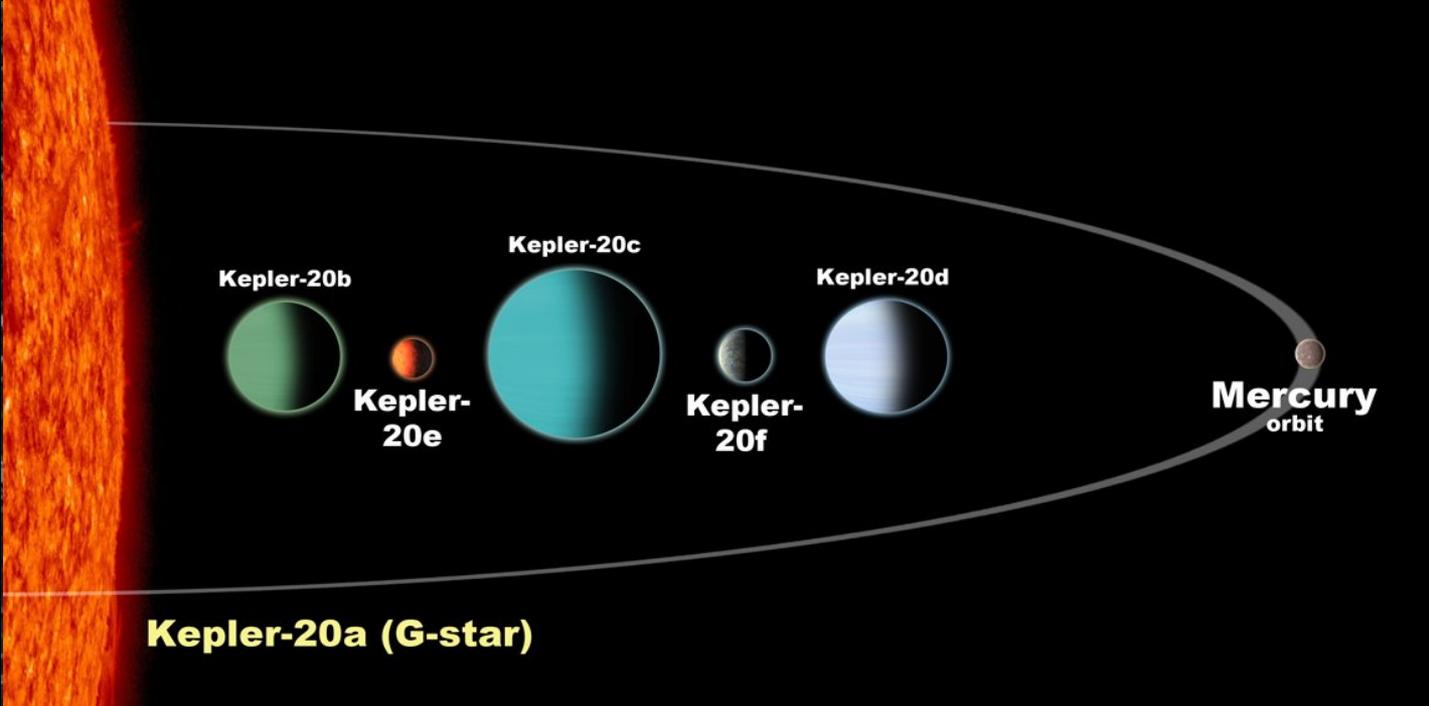




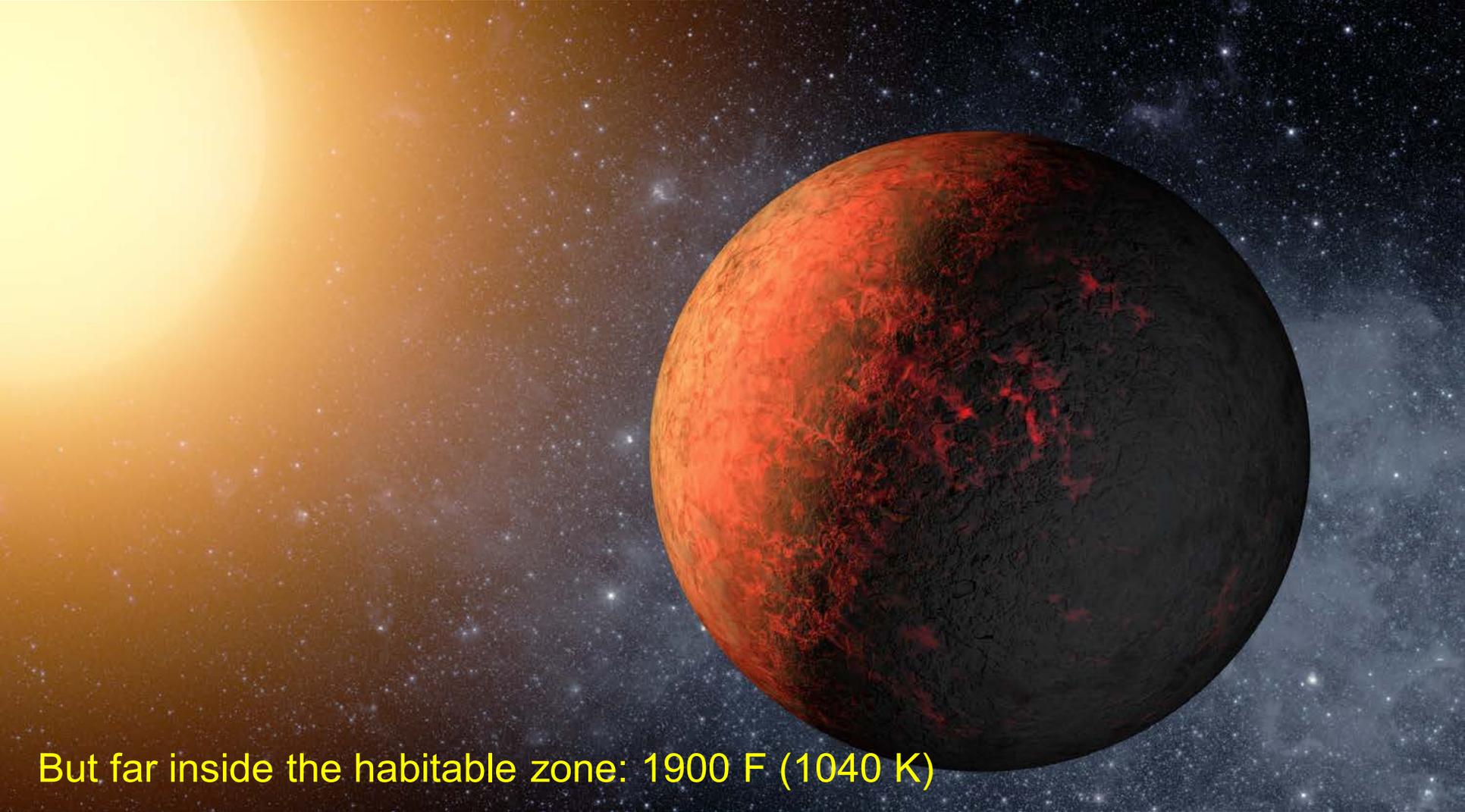




Sunset on Kepler 16b



# Kepler 20e: a world smaller than Earth around a Solar type star



But far inside the habitable zone: 1900 F (1040 K)

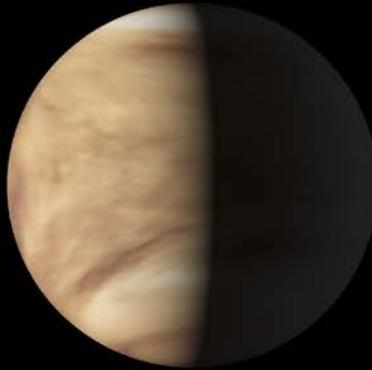
Kepler-20e

1040 K



Venus

299/753 K



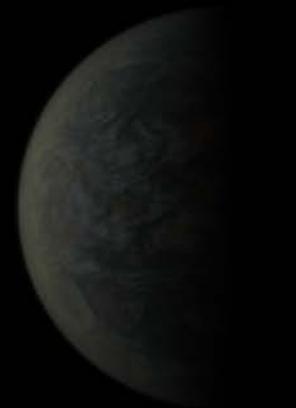
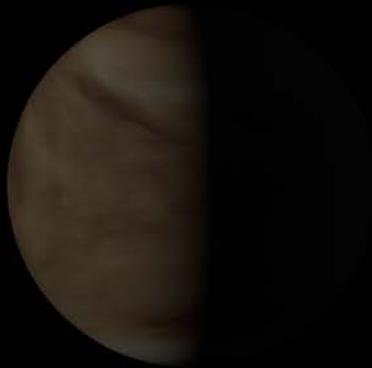
Earth

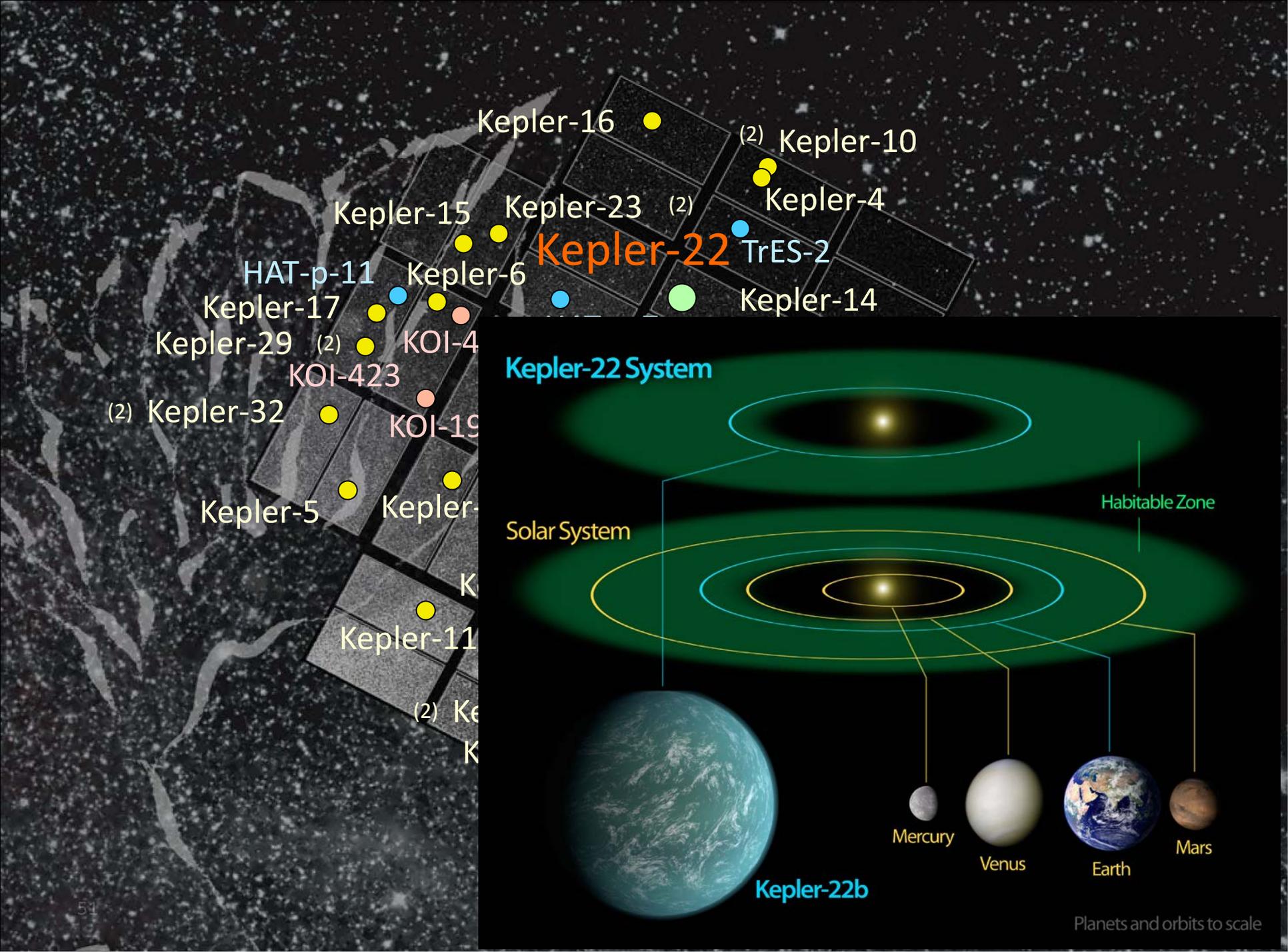
255/288 K



Kepler-20f

705 K





Kepler-16

(2) Kepler-10

Kepler-15

Kepler-23

(2)

Kepler-4

HAT-p-11

Kepler-6

Kepler-22

TrES-2

Kepler-17

Kepler-29

(2)

KOI-4

KOI-423

(2) Kepler-32

KOI-19

Kepler-5

Kepler-

Kepler-22 System

Solar System

Habitable Zone



Kepler-22b

Mercury

Venus

Earth

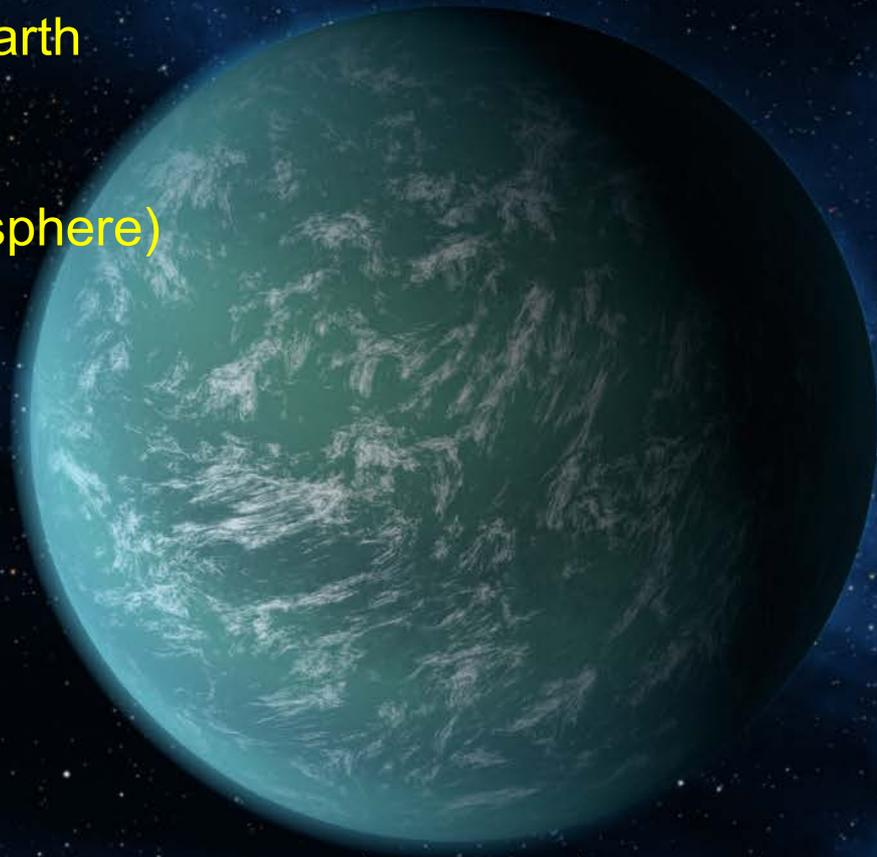
Mars

Planets and orbits to scale

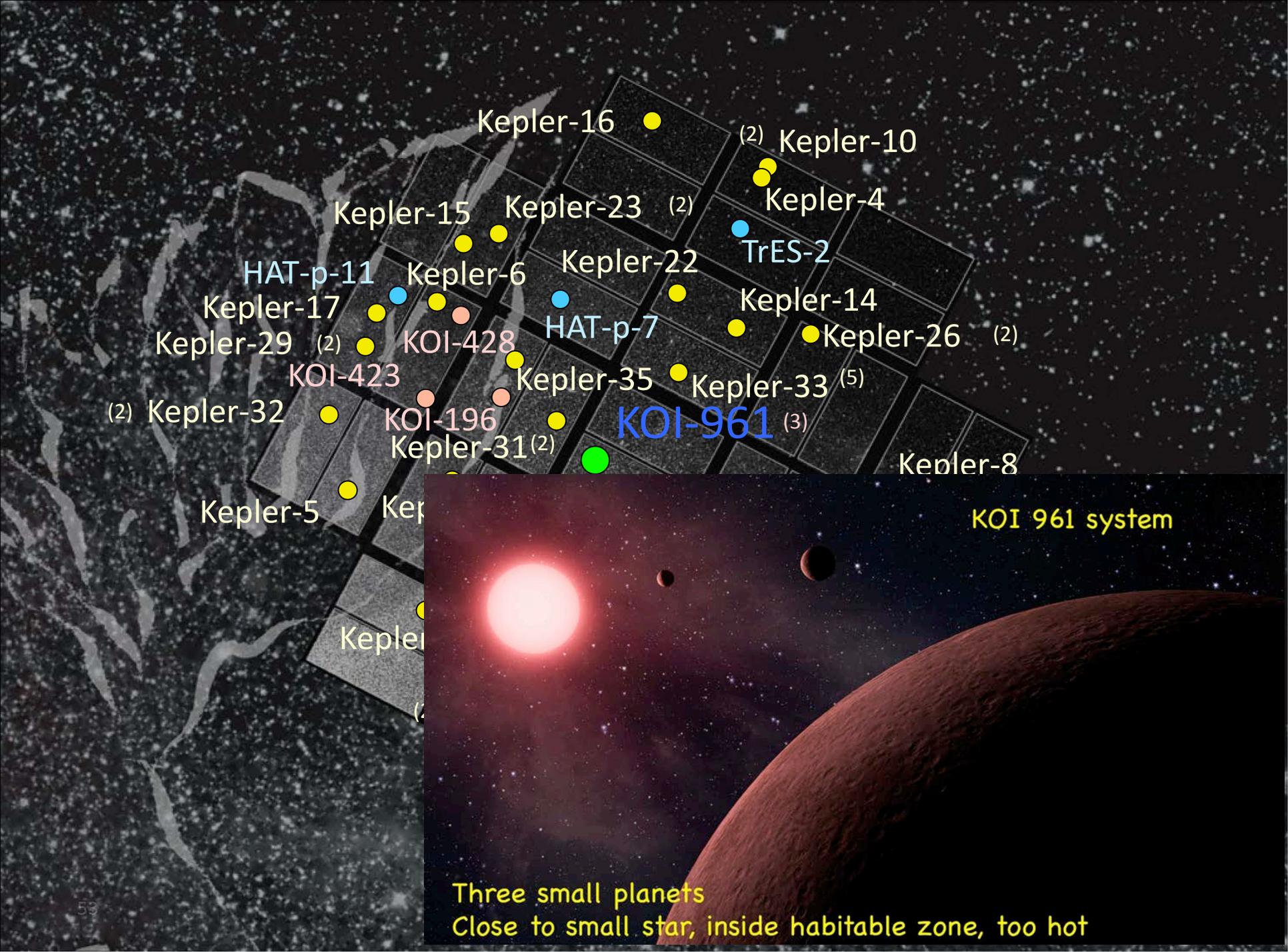
# Kepler 22b

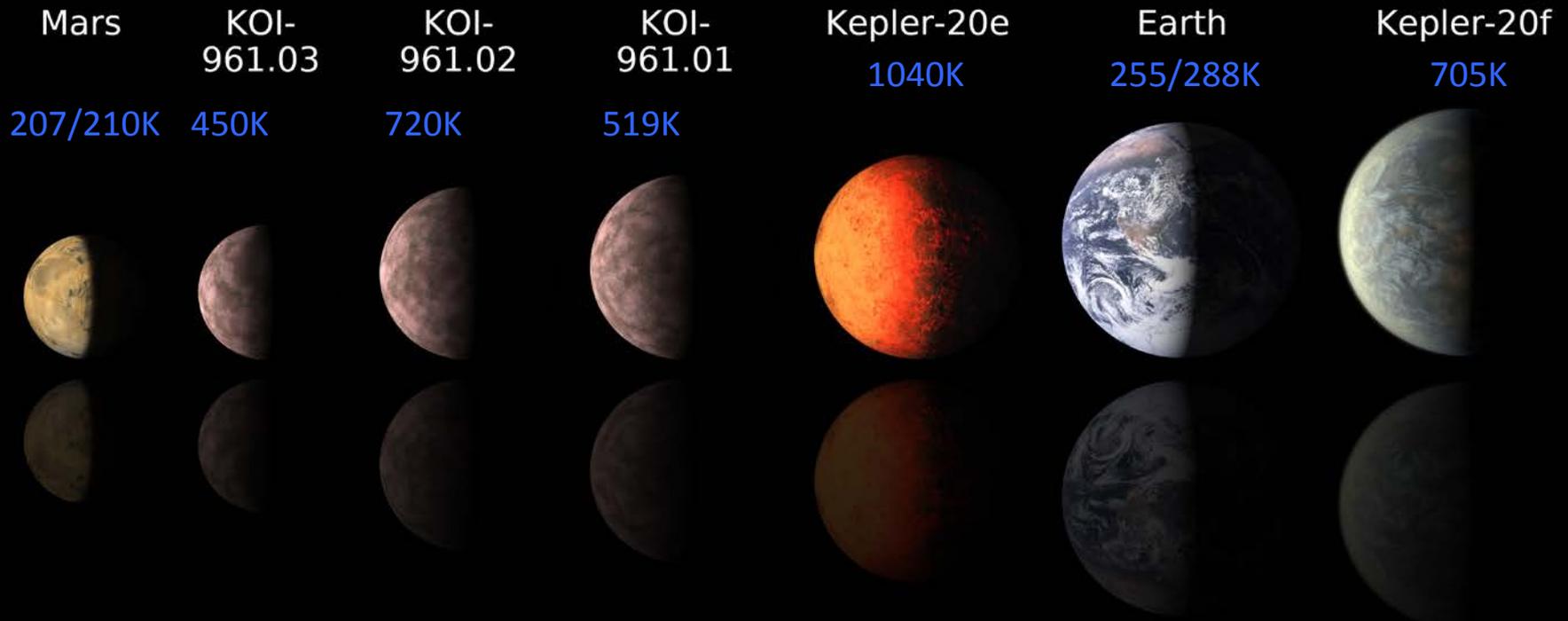
In the habitable zone of a Solar-like star  
but 2.4 times the size of Earth

Temperature 14 F  
(but warmer with an atmosphere)



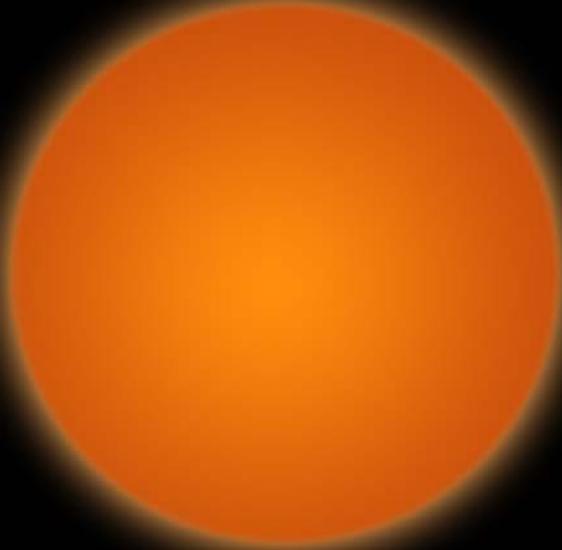
Water planet? Rocky surface? Deep atmosphere?





KOI 961 planet lineup

## KOI-961 and Its 3 Known Planets



02

01

03

## Jupiter and Its 4 Largest Moons



Io

Europa

Ganymede

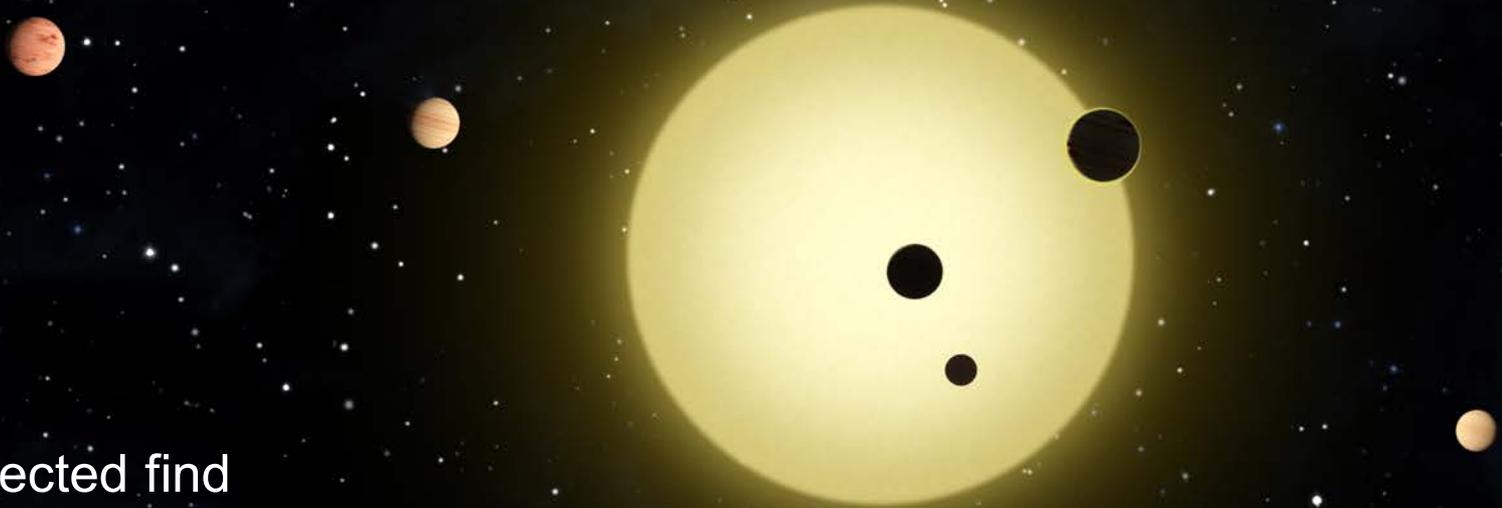
Callisto

# Multiple Planet Systems

Unexpected find

Tells us about planet system, not just planets

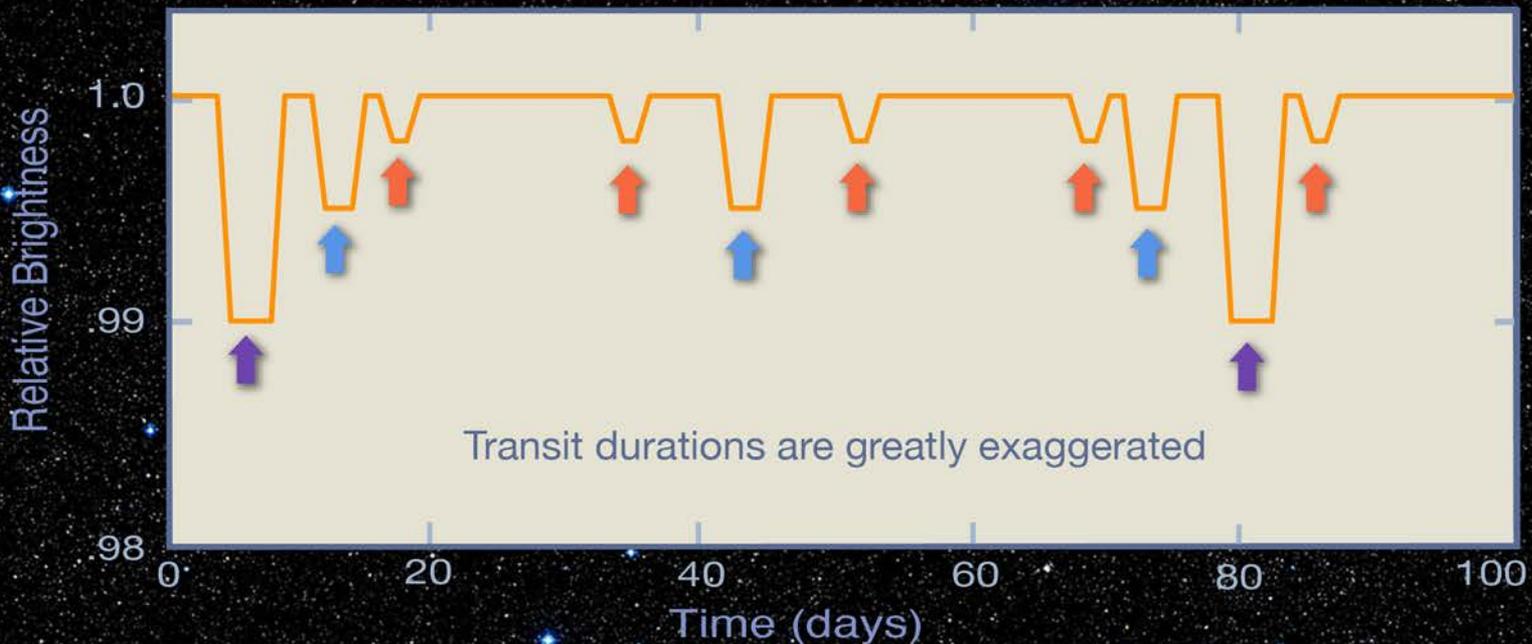
Self confirming because multiple occurrence of false positives is really unlikely



# Transit Signature of a Multiple-Planet System



- Planets can be distinguished by:
- Different periods
  - Different depths
  - Different durations



# Kepler's Confirmed Multiple Planet Systems

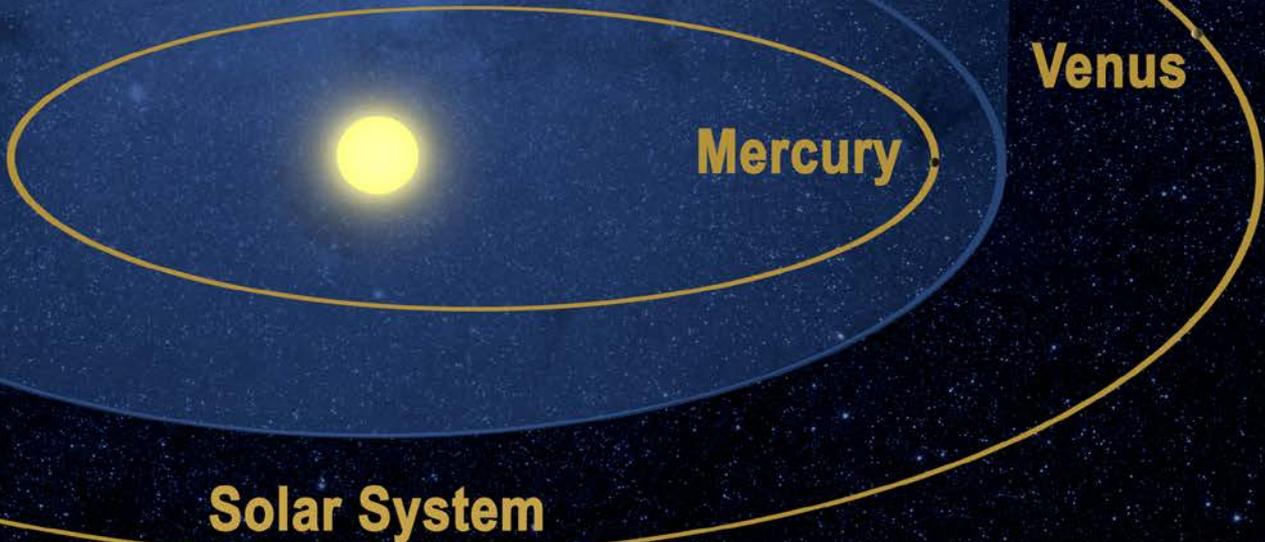
● Solar System    
 ● Planetary systems known prior to January 26, 2012    
 ● Planetary systems announced January 26, 2012    
 ● Unconfirmed planet candidates



>150 more multiple planet systems with >800 planet candidates

6 transiting planets

### Kepler-11 System

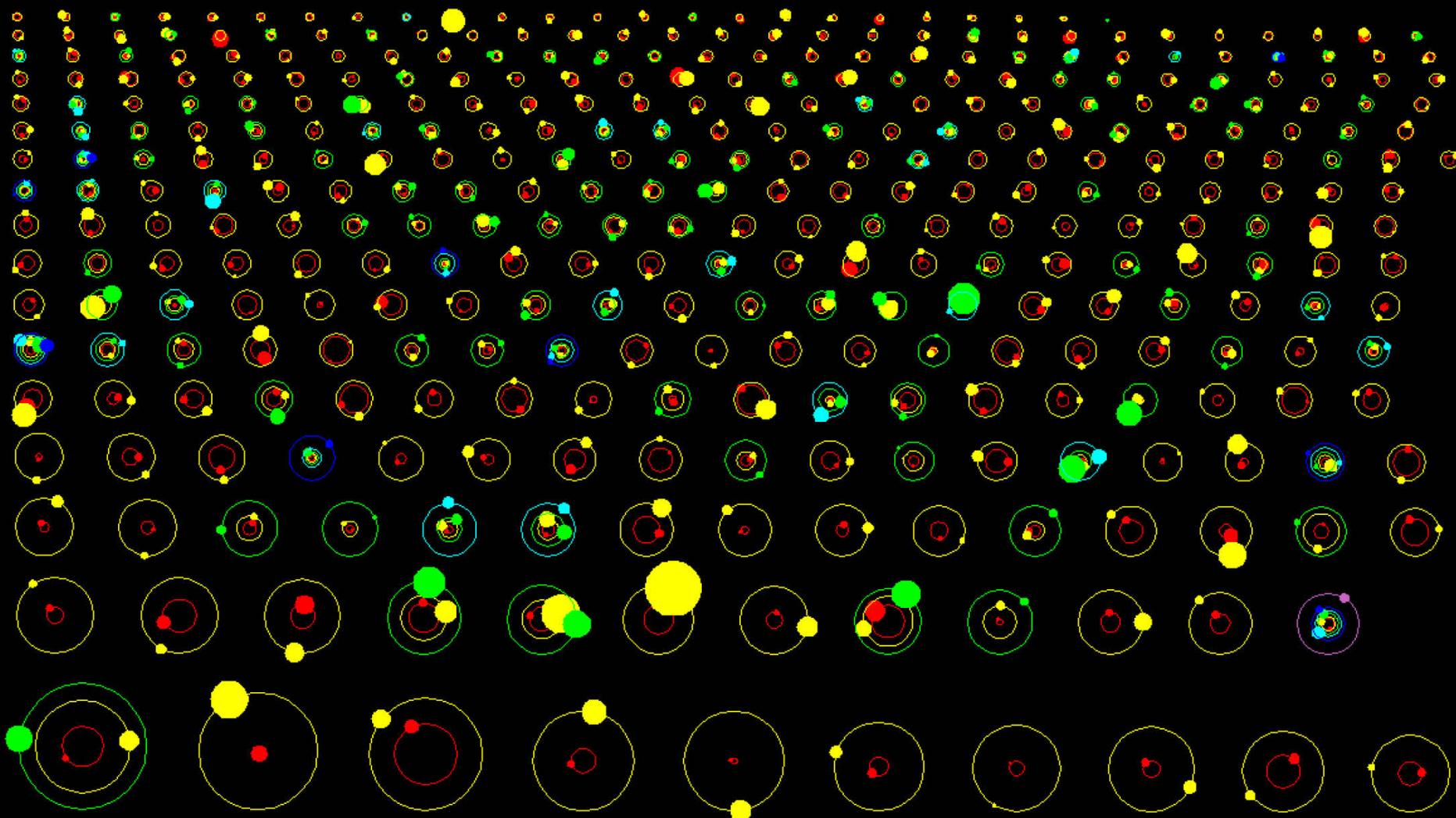


### Solar System

# The Kepler Orrery II

$t[\text{BJD}] = 2455879$

D. Fabrycky 2012



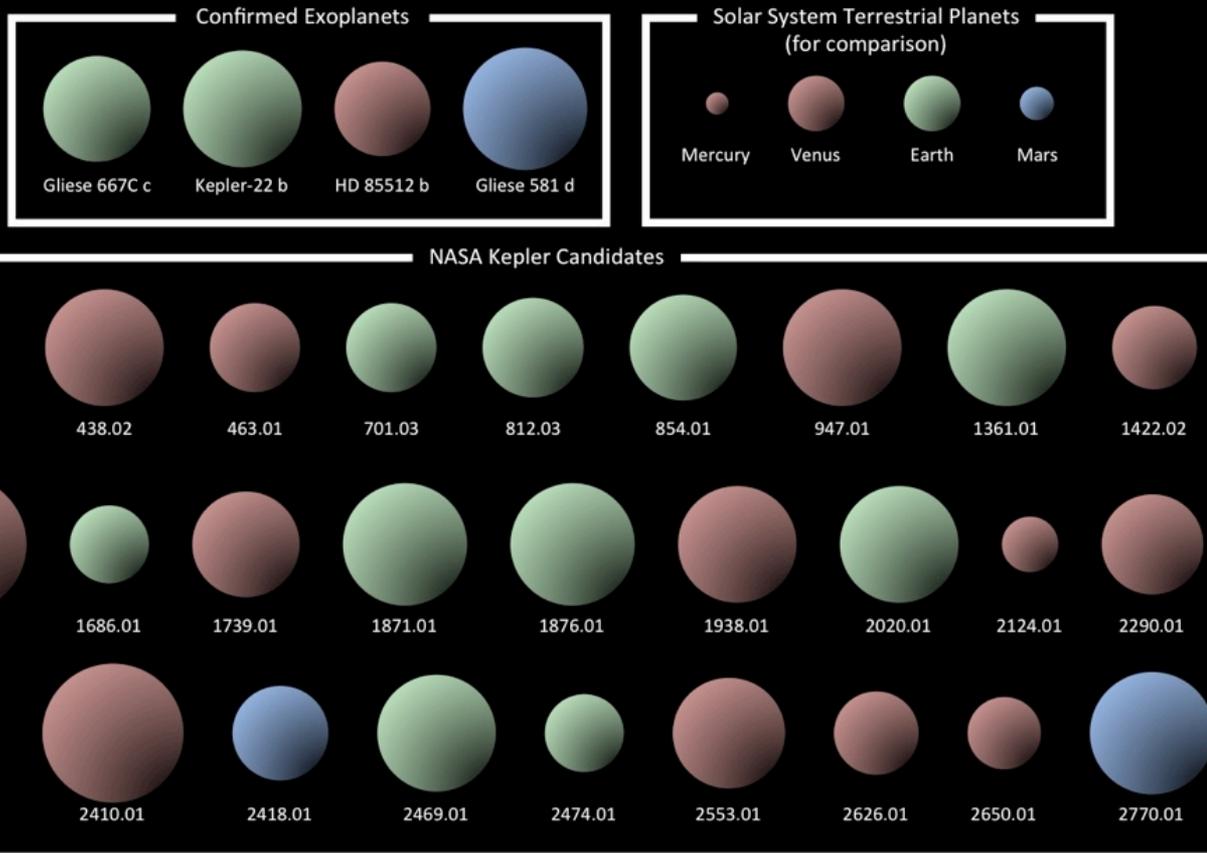
Many habitable zone planets.

Lots of Earth-size planets.

No confirmed Earth-size HZ planet yet. But soon!

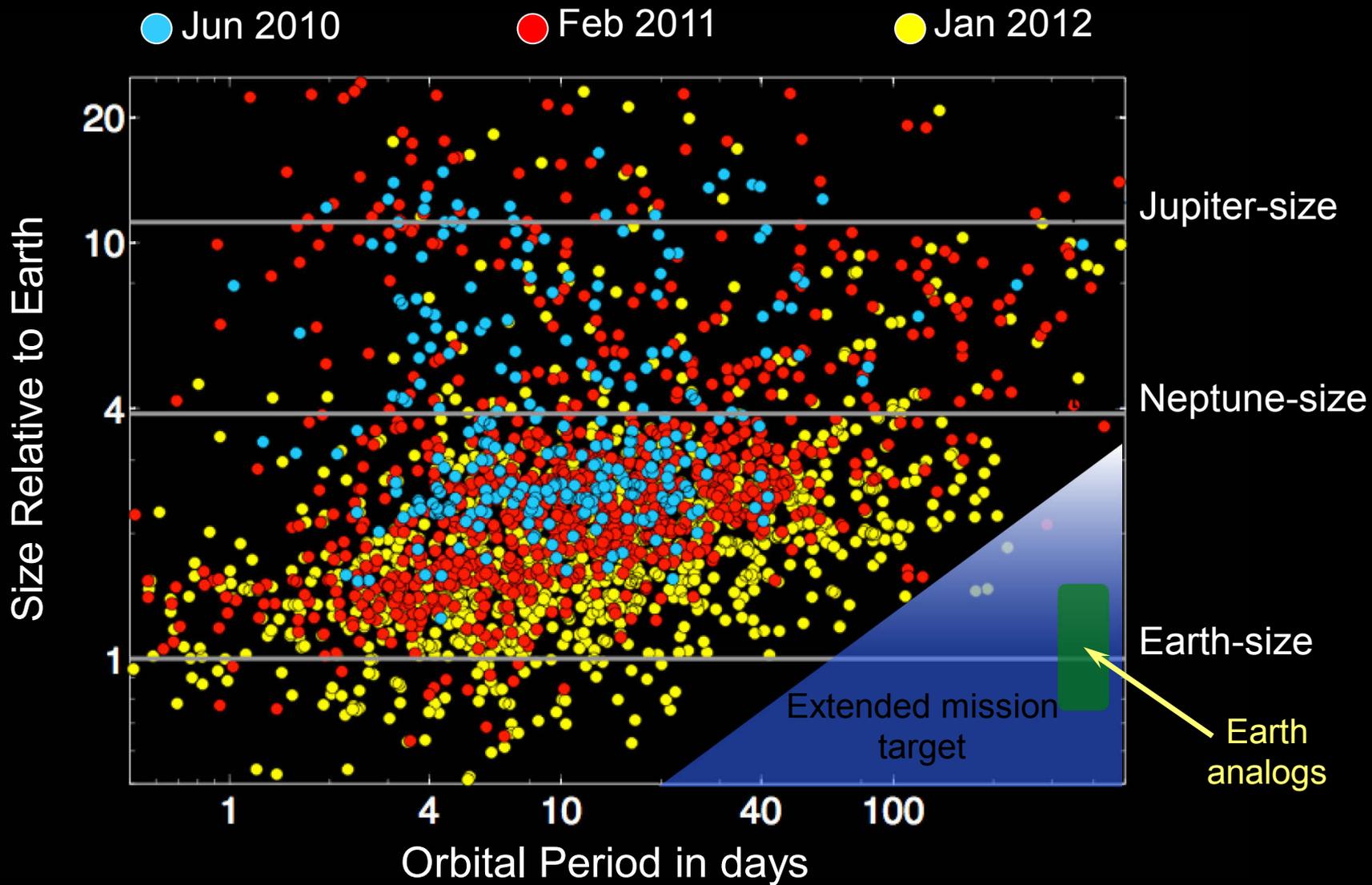
## Potential Habitable Exoplanets

(4 confirmed and 27 unconfirmed NASA Kepler Candidates)



Credit: Planetary Habitability Laboratory, UPR Arcibo (phl.upr.edu) April 2012

# Kepler Extended Mission Target: Habitable Zone Earths



# What about life?



Habitable zone planets (or their moons) COULD have life.

But DO they?

Compare Earth and Venus.

Same size

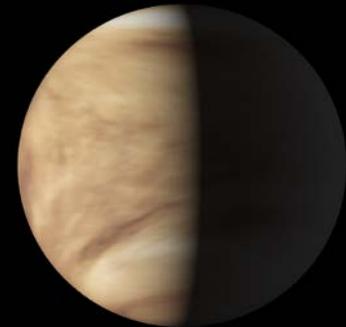
Both in the habitable zone

Earth: Surface temperature that melts water  
Life

Venus: Surface temperature that melts lead  
Probably no life

Can't tell the difference between Earth-like  
and Venus-like habitable zone exoplanets.

Need new telescopes to measure exoplanet  
atmospheres. (FINESS)



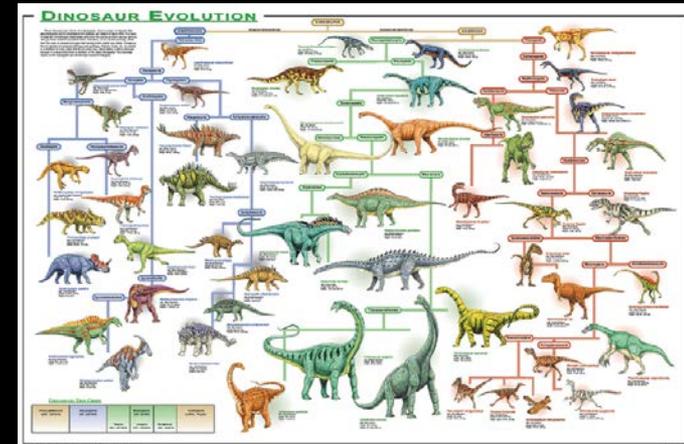
# Intelligent Life?

Another matter altogether

Life is probably frequent  
How often does it become intelligent?  
How long does intelligent life last?  
We are now performing this experiment.



Deep ocean worm feeds on methane eating bacteria



or

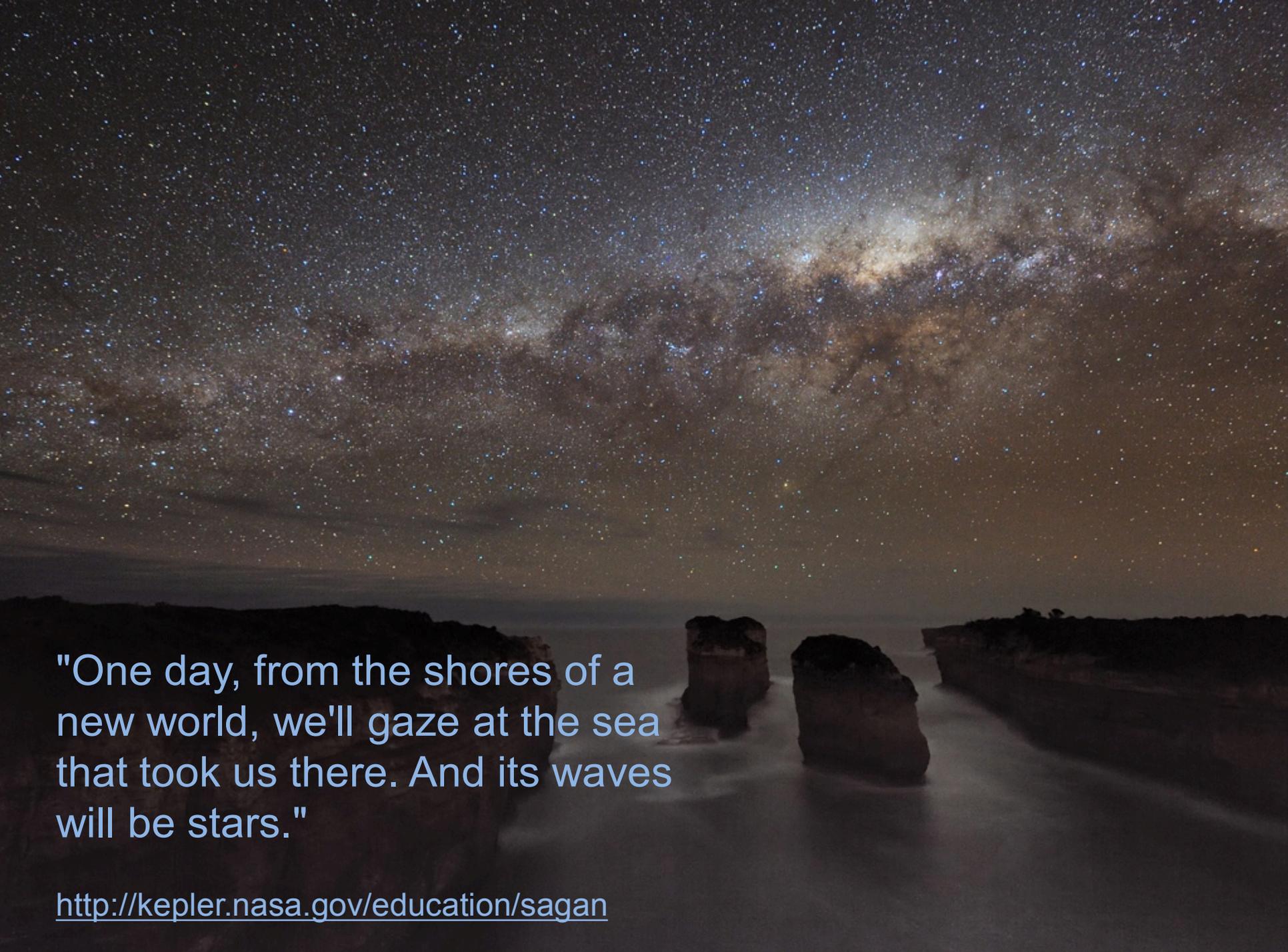


200 million years of evolution  
Walnut-size brains  
No tools

# Are we alone ?

A composite image of Jupiter and Saturn in space. Jupiter is on the left, showing its characteristic orange and white bands. Saturn is on the right, showing its rings. A small blue and white planet (Earth) is visible in the lower left. A bright star is in the upper right. The background is a dark field of stars.

**We don't know!**  
*(But we're working on it.)*



"One day, from the shores of a new world, we'll gaze at the sea that took us there. And its waves will be stars."

<http://kepler.nasa.gov/education/sagan>