

## The Potential Benefits of Nuclear Power on the Surface of Mars: The Robotic Exploration Perspective

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### Overview

- **Mars Exploration Program**
  - Objectives and approach
- **Potential future robotic Mars missions (2011-2024):**
  - Approved and candidate Mars exploration missions
- **Power Systems**
  - Power Source Classification
  - Solar Availability and Solar Power Generation
  - Radioisotope Power Systems (SoP, development, future plans)
  - Batteries
  - Special considerations for using RPSs on Mars landed missions
- **Future Mission Concept Examples Utilizing RPSs**
- **Conclusions**



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## Mars Exploration Program



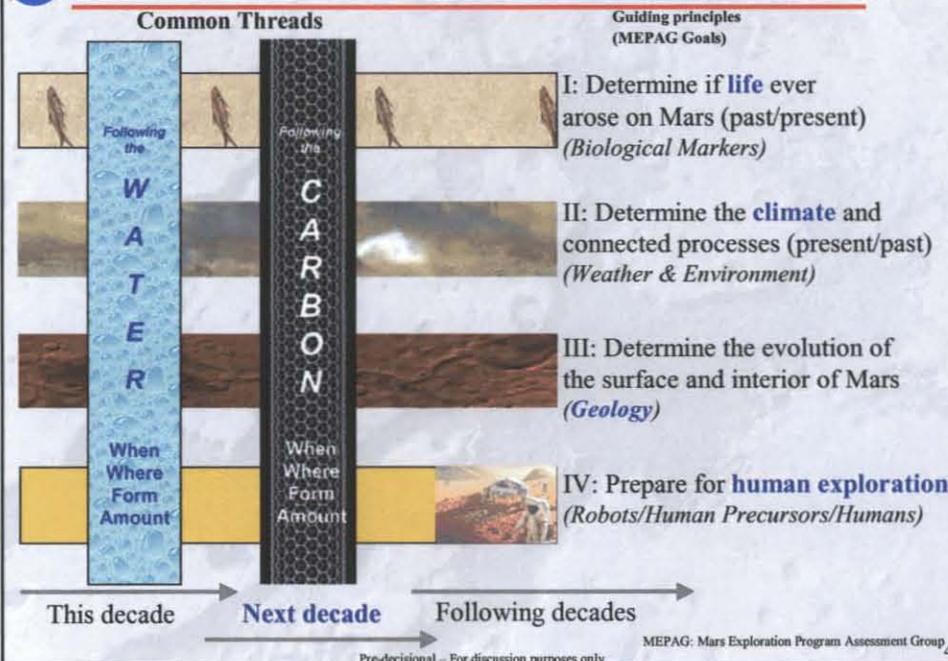
- Consists of a **science-driven** effort to characterize and understand Mars as a dynamic system. This includes its present and past environment; climate cycles; geology; and biological potential. A key question we are trying to answer is whether life ever arose on Mars.
- Exploration strategy: **“Follow the Water”** (present & early next decade)  
Search for sites on Mars with evidence of past or present water activity and with materials favorable for preserving either bio-signatures or life-hospitable environments.
- Exploration approach: **“Seek + In-Situ + Sample”**  
Orbiting and surface-based missions are interlinked to target the best sites for detailed analytical measurements, and for sample return.

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## Exploration Strategies



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## Recent Discoveries



### Based on Mars Exploration Rover (MER); Mars Global Surveyor (MGS); Mars Odyssey and Mars Express findings:

- Early Mars was wet and had an active hydrological cycle.
- Groundwater played an important role on early Mars.
- Periodic climate change influenced the hydrological cycle.
- Near subsurface is hydrogen rich at mid-latitudes and poleward.
- Detection (putative) of significant and spatially variable amounts of  $\text{CH}_4$  may indicate geologically or biologically active sites
- Evidence of an early dynamo suggests magnetic field could have provided shielding at the surface from solar and cosmic radiation.



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## Future Mission Concepts Evaluated



### Mars Science Orbiter and Telecom



### Midrovers



### Astrobiology Field Laboratory (AFL)



### Planetary Evolution and Meteorology Network

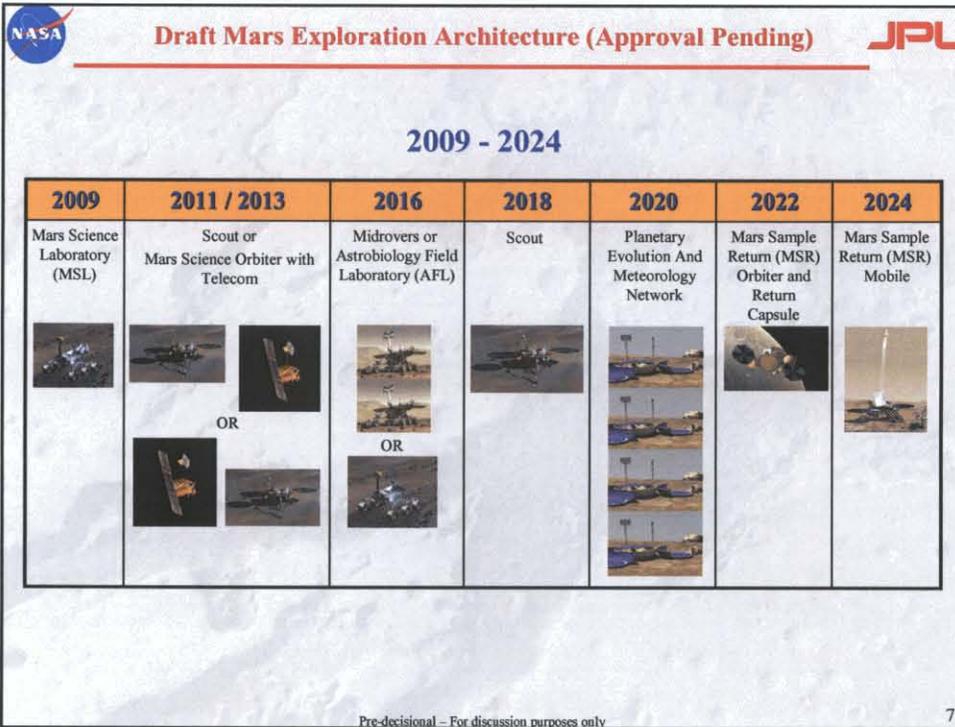


### Mars Sample Return (MSR)



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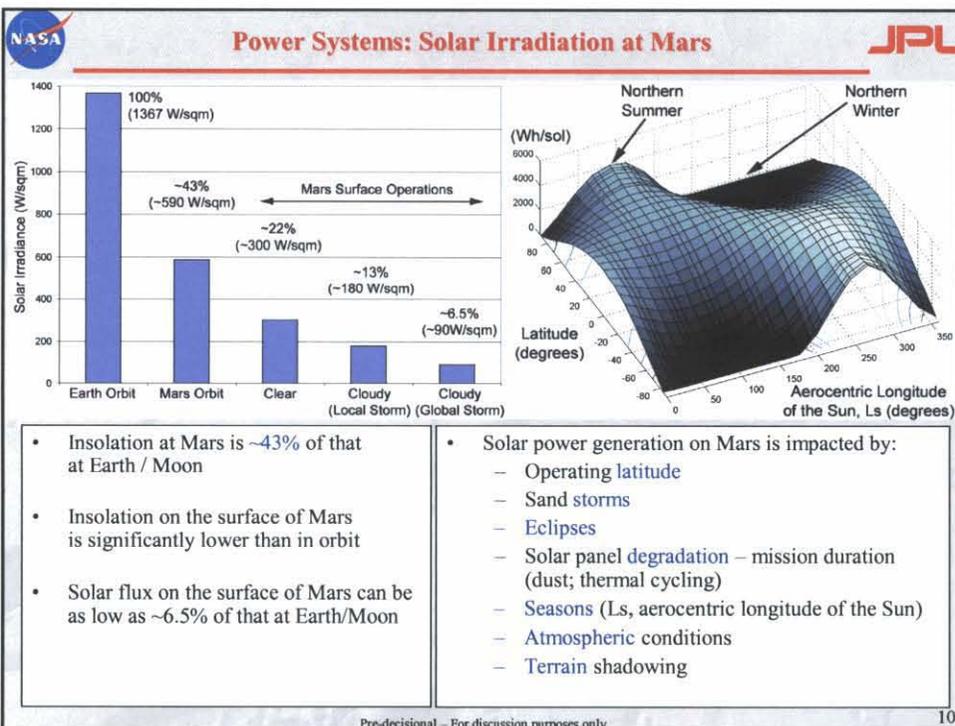
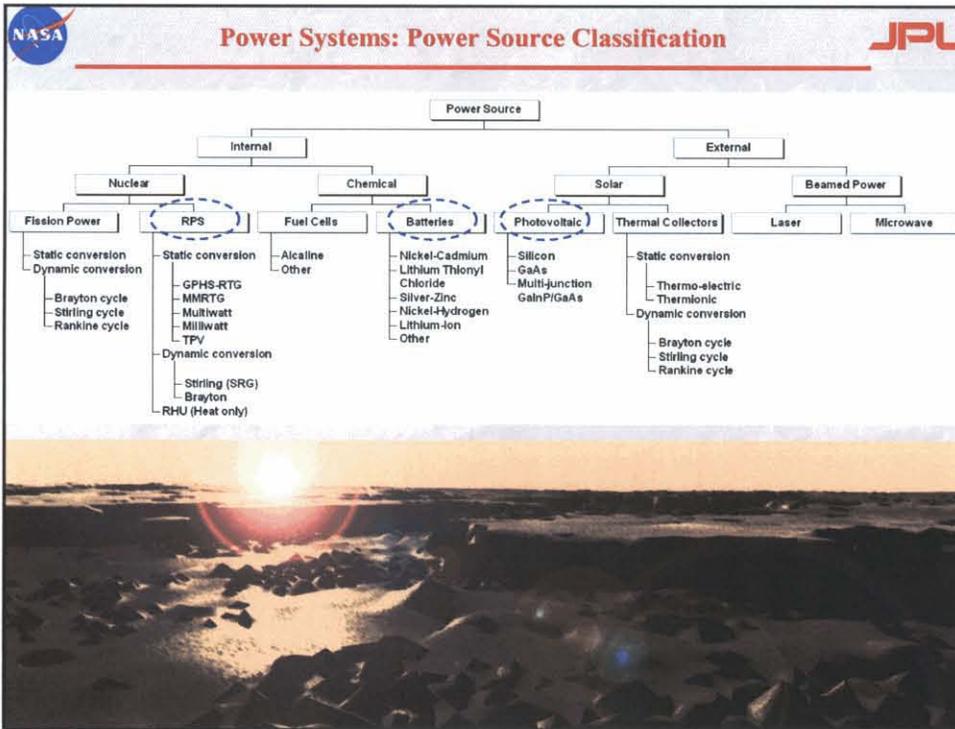
**Approved and Candidate Robotic Mars Missions**

Selected & Potential Missions	Mission Class	Power System Option(s)
Orbiters (e.g., MRO <sup>1</sup> )	Moderate/Large	Solar (typical for Mars orbiters)
<b>Phoenix<sup>1</sup></b>	Scout	Solar selected
<b>Mars Science Laboratory (MSL)<sup>1</sup></b>	Large	RPS (MMRTG) baselined (solar feasibility study in work)
Scouts (small missions)	Scout	Solar (RPS usage not allowed)
Multi-Lander Network	Moderate/Large	(Small-RPS or Solar) <sup>2</sup>
Astrobiology Field Lab (AFL) rover	Large	(To be determined) <sup>2</sup>
Mars Sample Return (MSR)	Flagship	(Solar or RPS) <sup>2</sup>
MSR Fetch rover (sub-MER class)	Scout-Large	(Small-RPS or Solar) <sup>2</sup>
Deep Drill	Large	(RPS or Solar) <sup>2</sup>

<sup>1</sup> **Funded** missions;

<sup>2</sup> **Everything below Scouts line should be referenced as based upon presumed mission requirements and preliminary power trade studies**

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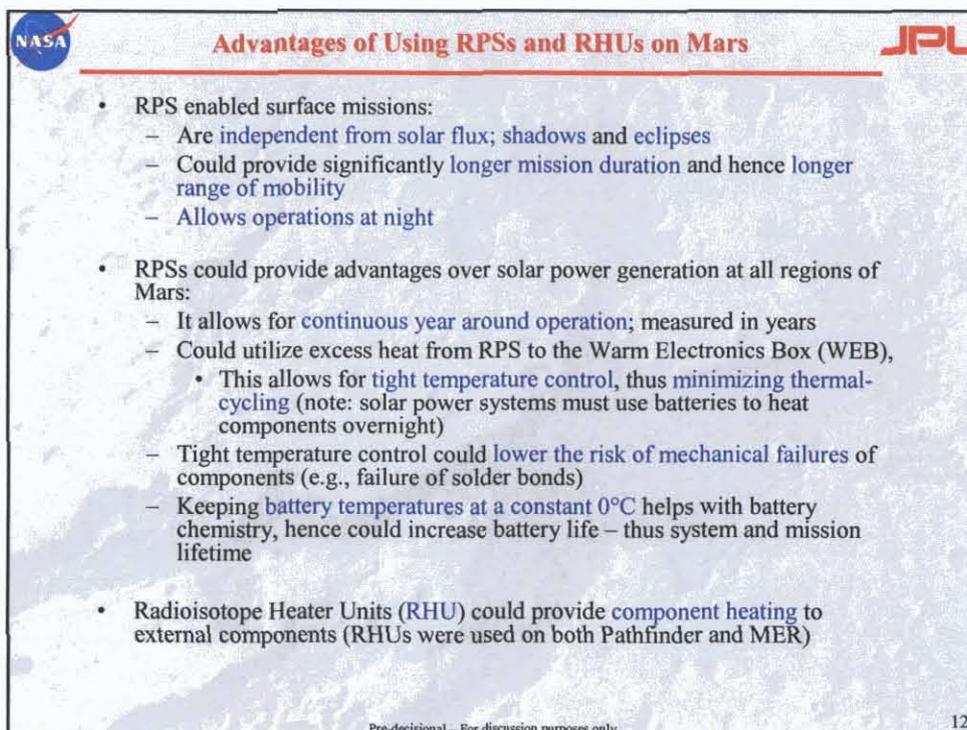
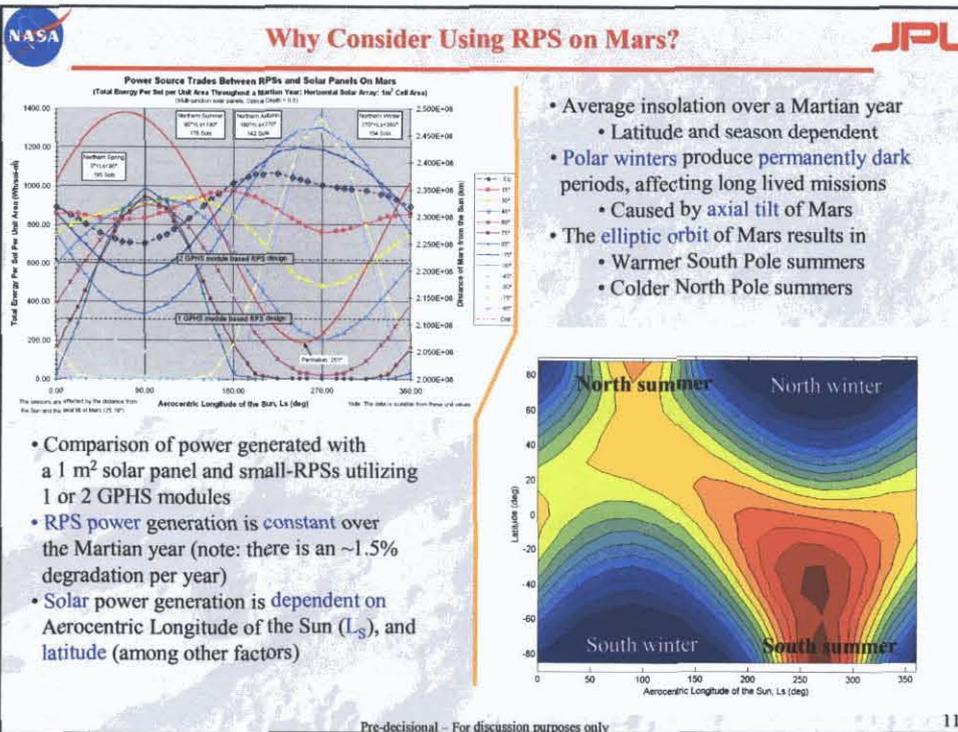


- Insolation at Mars is ~43% of that at Earth / Moon
- Insolation on the surface of Mars is significantly lower than in orbit
- Solar flux on the surface of Mars can be as low as ~6.5% of that at Earth/Moon

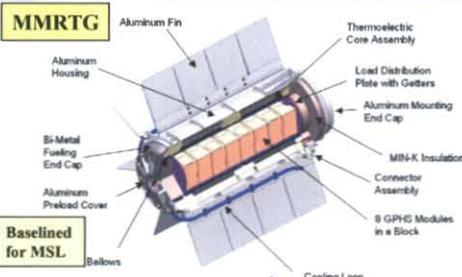
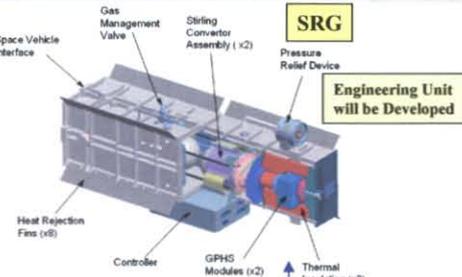
- Solar power generation on Mars is impacted by:
  - Operating latitude
  - Sand storms
  - Eclipses
  - Solar panel degradation – mission duration (dust; thermal cycling)
  - Seasons (Ls, aerocentric longitude of the Sun)
  - Atmospheric conditions
  - Terrain shadowing

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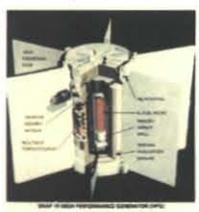
**RPSs for Mars: Past, Present and Possibly Future**

	MMRTG Future:	SNAP-19 (DISCONTINUED) Past:	SRG (Development)
Power per Unit (BOM), We	~125	~40.3	~116
Mass per Unit, kg	44	13.4	34
# of GPHS Modules per Unit	8	N/A	2
Thermal Power, Wt	2000	250	500
Specific Power, We/kg	2.9	3.0	3.4
Conversion type	Static	Static	Dynamic
Converter materials	PbTe/TAGS	PbTe/TAGS	Stirling
Technical Readiness level	TRL-5	TRL-9	TRL-3
Availability	MSL-2009	Discontinued (used on Viking 1 & 2)	2014+

**Baselined for MSL**

Present: Radioisotope Heater Unit 1Wt  
Used on MPF & MER (component heating)



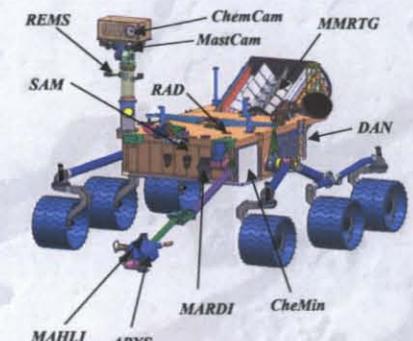
SNAP-19 was on Viking 1 operated for >6 years, spacecraft was shut down

Engineering Unit will be Developed

SRGs are only considered as trade options for next decade Mars missions

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**Mars Science Laboratory (MSL) Mission (2009)**



### Salient Features

- Mobile Science Laboratory
- One Mars Year nominal surface operational lifetime (669 sols / 687 days)
- Discovery responsive over wide range of latitudes and altitudes
- Controlled Propulsive Landing
- Precision Landing via Guided Entry
- Baselined with an MMRTG

### Remote Sensing (Mast)

ChemCam – Laser Induced Breakdown Spectrometer  
MastCam – Color Stereo Imager

### Contact Instruments (Arm)

MAHLI – Microscopic Imager  
APXS – Proton/X-ray Backscatter Spectrometer

### Analytical Laboratory (Front Chassis)

SAM – Gas Chromatograph/Mass Spectrometer/ Tunable Laser Spectrometer (Sample Composition / Organics Detection)  
CheMin – X-ray Diffraction / Florescence (Sample Mineralogy)

### Environmental Characterization (Body-mount)

MARDI – Descent Imager  
REMS – Meteorological monitoring  
RAD – Surface Radiation Flux Monitor (future human health & safety)  
DAN – Neutron Backscatter subsurface hydrogen (water/ice) detection

**In current studies, the main difference between MSL and AFL missions would be the science instrument suite and sample processing.**

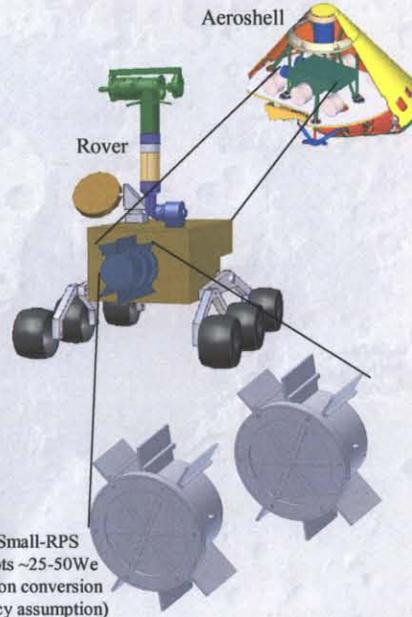
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### Rovers: MER-Class Rover and AFL Concepts



- Astrobiology driven goals to address MEPAG Goals 1 to 3
- Instruments could include: Microscopic imager; Raman spectrometer; APXS; Mini-TES; Pan Cam, RAT; TEGA; GC/MS etc.
- Could operate up to 2-3 years on Mars
- MER solar power generation:
  - ~1000 Wh/sol BOL;
  - ~600 Wh/sol EOL (primary mission);
- Small-RPS enabled mid-rovers with
  - 2 GPHS modules: ~620 Wh/sol BOL;
  - 4 GPHS modules: ~1240 Wh/sol BOL.
- Large MSL heritage rover:
  - MMRTG (8 modules): ~2700 Wh/sol
- Excess heat from RPS could be utilized to heat components, motors, actuators, etc.
- Typical high power operating modes:
  - mobility, telecom,
  - GC/MS, drilling



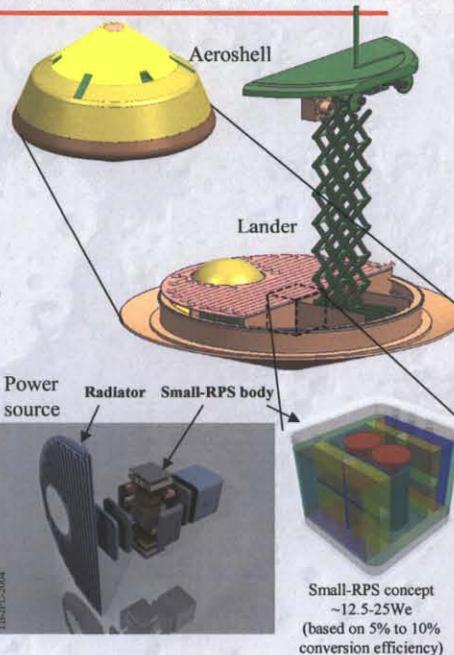
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### Mars Multi-Lander Network Concept



- Monitor weather & environment
- Address MEPAG Goals 2 to 3
- Instruments could include: seismometer; mini-MS; Pan Cam; meteorology station
- Baseline:
  - 4 landers (up to 10)
  - Solar powered
  - (RPS trades considered – shown here)
- High power requirement by telecom
- Small-RPS based configuration:
  - Long life (up to several years)
  - Independent from solar flux
  - Capable to operate at any location
  - Nominal hard landing requires high g-load tolerant RPS (~2000g, compared to 40g for the MMRTG). This requires significant technology development



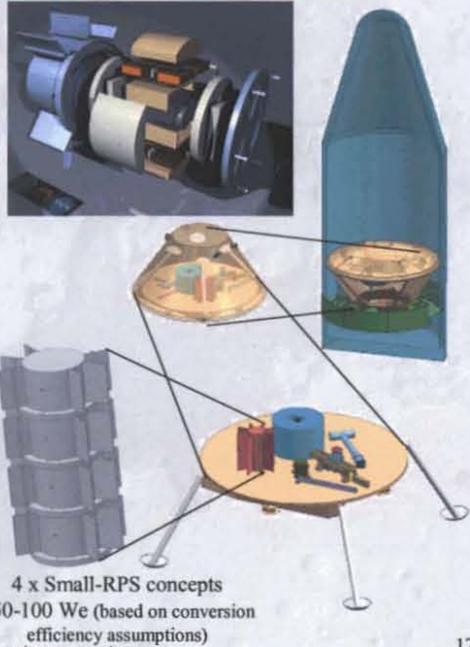
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## Mars Deep Drill Concept



- Deep Drill to 10 m or 50 m depth
- Key instruments: sample acquisition and processing; meteorology instrument; Pan Cam
- To address MEPAG Goals 1 to 3
- 10m drill could achieve mission goals with 4 small-RPSs, generating ~50We and ~1240 Wh/sol
- 50m drill would likely require an MMRTG (~110 We and ~2700 Wh/sol)
- High power requirements for drilling operations and telecom



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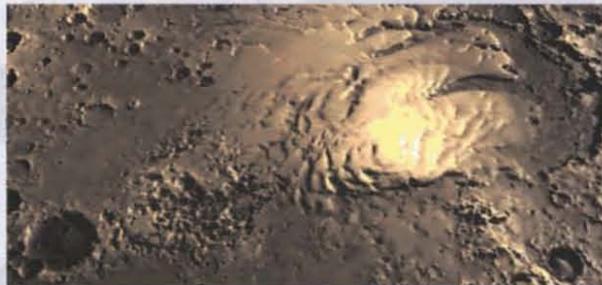
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## Conclusions



- NASA's post 2010 Mars program is in a pre-planning phase, therefore the listed missions are under study. Final program plan is expected in Summer 2006.
- Mars orbiter missions will likely continue employing solar power generation
- Some landers / rovers could use both RPSs or solar panels. The options depend on mission duration, landing location (latitude, longitude terrain), season, atmospheric conditions and planetary protection constraints
- Next decade landed missions will likely use MSL heritage, but if available then small-RPSs could also be considered
- Radioisotope Heater Units (RHU) will likely continue to provide local component heating on some surface missions



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