

# Power Combining Considerations for Prometheus TWTAs

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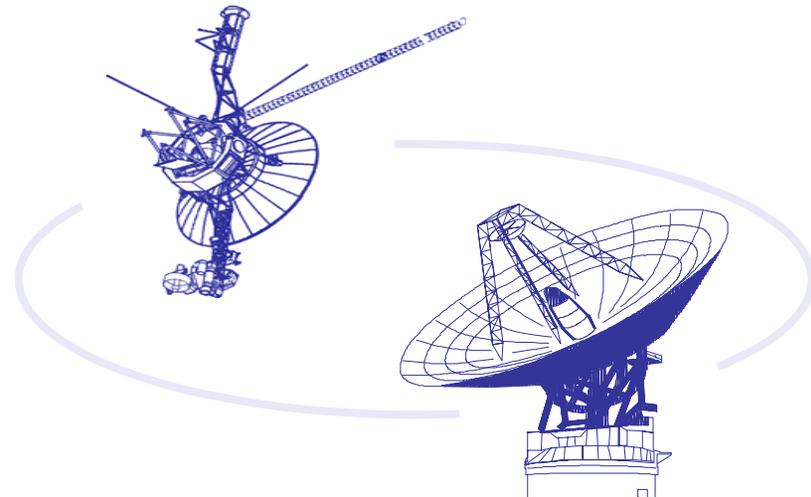
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# Introduction



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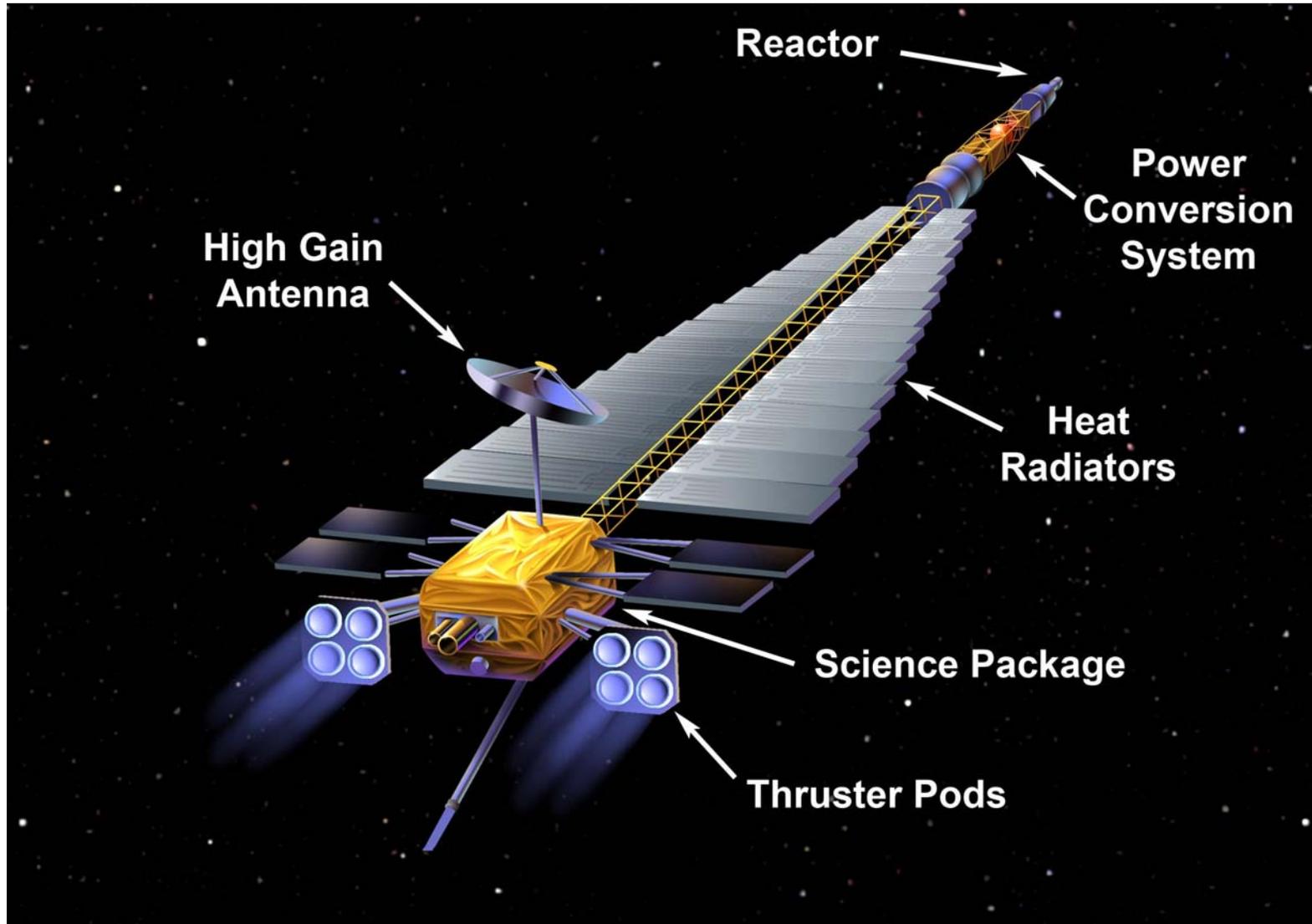
- **NASA planning a nuclear-electric spacecraft**
  - Higher exhaust velocity allows more ambitious missions than chemical propulsion
  - Spacecraft series known generically as “Prometheus”
- **One mission considered is to Jupiter’s icy moons**
  - Jupiter Icy Moons Orbiter (JIMO)
  - Callisto, Ganymede, Europa
- **Spacecraft will orbit each moon in turn**
  - Collect data for several months
- **Spacecraft will have large prime power capacity for scientific instruments and communications equipment**



# Prometheus Spacecraft



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- **Science package may generate up to 50 Mbps of data**
- **6.2 A.U. ( $9.3 \times 10^8$  km) path length**
- **Data rate requires use of 32 GHz deep space allocation**
- **Link calculations result in 1 kW RF CW power needed**
  - **Greatly exceeds current space-qualified 32 GHz RF sources (e.g., 35 W TWTA for MRO)**
- **Power and modest bandwidth suggest klystron or EIK**
  - **However no significant history of space use for these devices**
  - **(But see CloudSat presentation elsewhere in conference)**
- **Recent improvements in Ka-band TWTs also suggest power combining**
  - **1 kW obtainable from four 250 W devices in parallel**
  - **Combining inherently redundant**



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# Power Combining



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- **Combining has significant history of space use (e.g., DTH TV broadcasting, Sirius Satellite Radio, etc.)**
- **NASA actively pursuing higher power 32 GHz sources**
  - 100 W TWT developed, 100 W TWTA in work for MTO
  - JIMO developing 180 W TWT (See L-3 ETI TWT presentation elsewhere in conference)
- **Network combining usual method, but spatial combining offers several advantages**
  - Reduced susceptibility to multipaction & RF breakdown
  - Electronic beam steering as adjunct to mechanical
  - JPL has done considerable work investigating “cluster” antennas
- **Power combining susceptible to phase and amplitude errors**

999H (100 W, MTO)



999HA (180 W, JIMO)



# Four-Beam Cluster Antenna



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# Spatial Superposition of Four Carriers



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$$E(t) = A \sin(\omega t) + B \sin(\omega t + \phi_1) + C \sin(\omega t + \phi_2) + D \sin(\omega t + \phi_3)$$

$$\langle E^2 \rangle = \frac{\int_0^T E^2(t) dt}{\int_0^T dt} \quad (T = 1/f = 2\pi/\omega)$$

$$\langle E^2 \rangle_{\text{nom}} = 8A^2 \quad (\text{all amplitudes equal \& no phase errors})$$

$$P_{\text{rel}} = \langle E^2 \rangle / \langle E^2 \rangle_{\text{nom}}$$

$$P_{\text{rel}} = (1/8) [ \frac{1}{2} ( 1 + b_2 + c_2 + d_2 ) + b \cos \phi_1 + c \cos \phi_2 + d \cos \phi_3 + bc \cos (\phi_1 - \phi_2) + bd \cos (\phi_1 - \phi_3) + cd \cos (\phi_2 - \phi_3) ]$$



- Prediction of  $\Delta P_{rel}$  requires estimate of amplitude and phase variations
- Two major underlying sources:
  - $E_k$  and/or  $I_k$  variations due to EPC drift and aging
    - Effect of  $\Delta E_k$  completely dominates  $\Delta I_k$
  - TWT gain changes with age
- Magnitude of underlying sources estimated empirically
  - Values used in worst-case analyses
- Amplitude and phase variation then predictable via TWT modeling code such as NRL's CHRISTINE 3-D
- Modeled 999HA (180 W) as proxy for future 250 W TWT
  - Assumed underlying variations normally distributed & input drive levels adjusted to keep TWT at saturation



- **Two cases modeled: five years of TWTA operation and 20 years of operation**
- **Five year case represents expected usage for JIMO**
  - **Ka-band TWTAs only used to transmit high-rate science data to Earth**
  - **Only occasional “health checks” planned during Earth-Jupiter cruise phase**
- **20 year case is maximum reasonable usage for any other deep space mission (e.g., Neptune/Pluto probe, Kuiper Belt, etc.)**



# TWT Amplitude and Phase Predictions



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	Five Years	20 Years
Phase Variation	$\pm 8.6^\circ$	$\pm 20.6^\circ$
Amplitude Variation*	$+0.07 \pm 0.02$ dB	$+0.13 \pm 0.03$ dB

\*Assumes  $I_k$  constant (by servo loop in EPC) and ignores any de-focusing or other  $I_b$  reductions

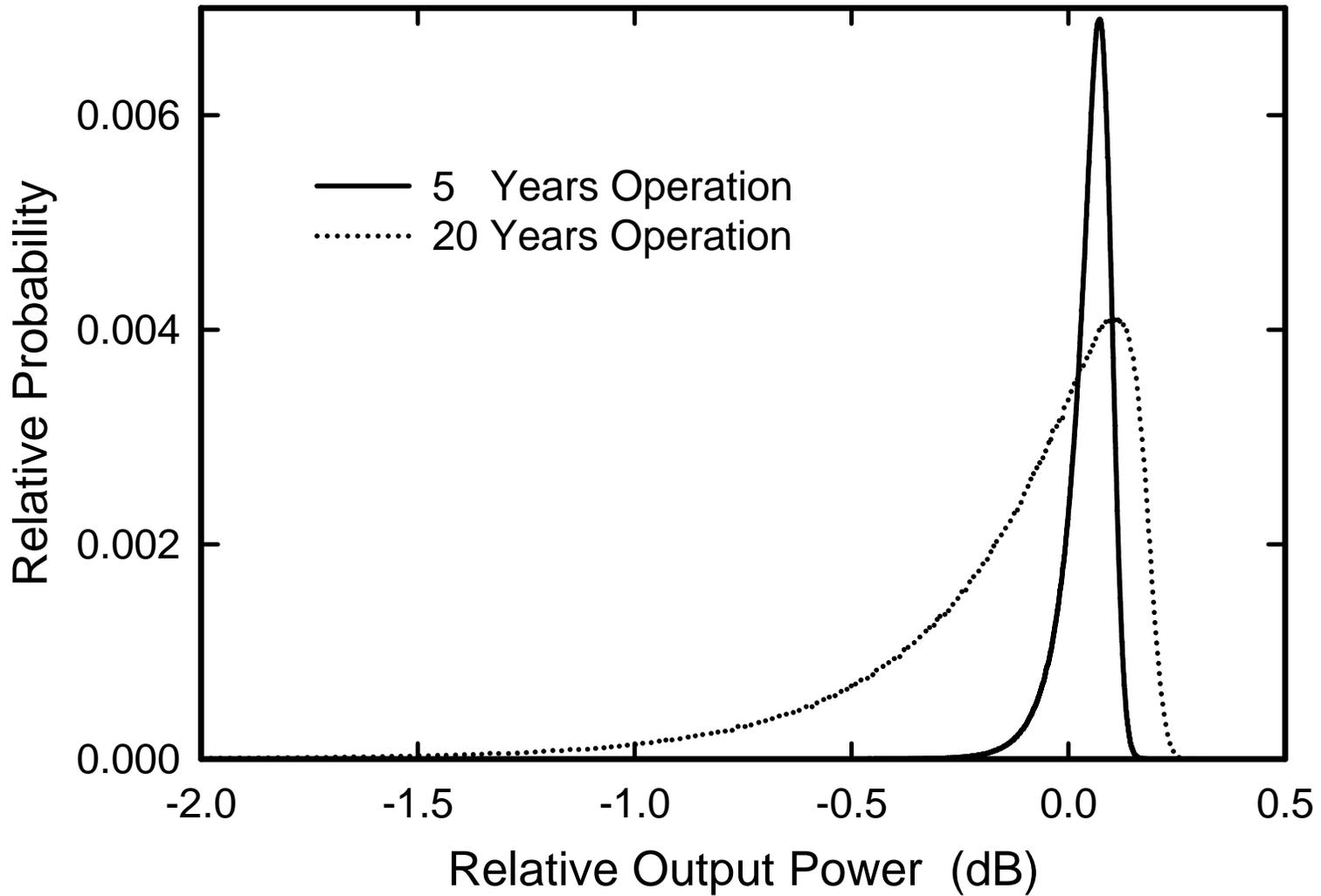
- These results input to 5 million trial Monte Carlo simulation to predict expected behavior of four power-combined TWTAs



# Distribution of Monte Carlo Results



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<b>(all in dB)</b>	<b>Five Years</b>	<b>20 Years</b>
<b>Best Case</b>	<b>+0.20</b>	<b>+0.29</b>
<b>Mean</b>	<b>+0.05</b>	<b>-0.11</b>
<b>Std. Dev.</b>	<b>0.01</b>	<b>0.10</b>
<b>Worst Case</b>	<b>-0.82</b>	<b>-4.5</b>



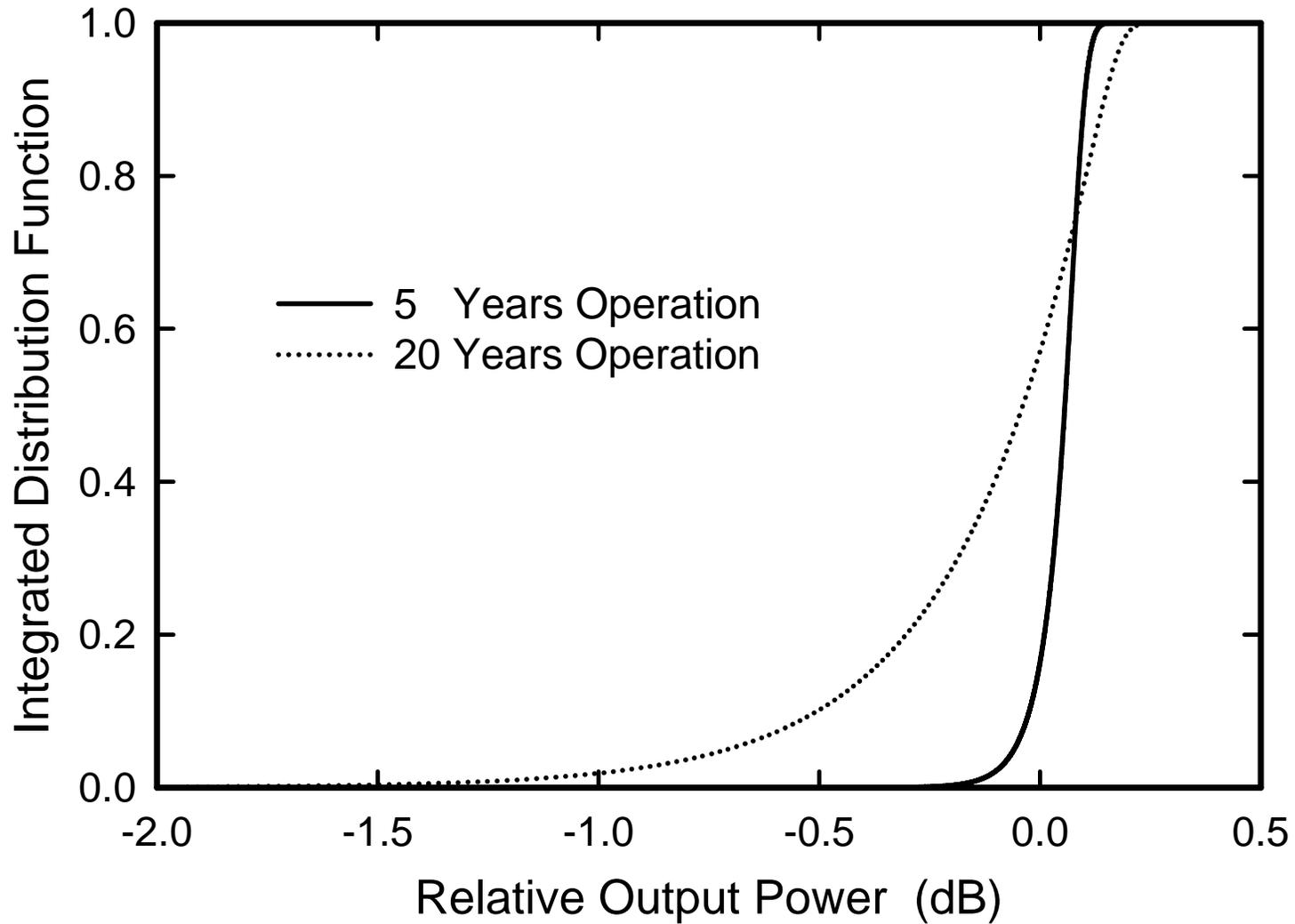
# Cumulative Distribution Function

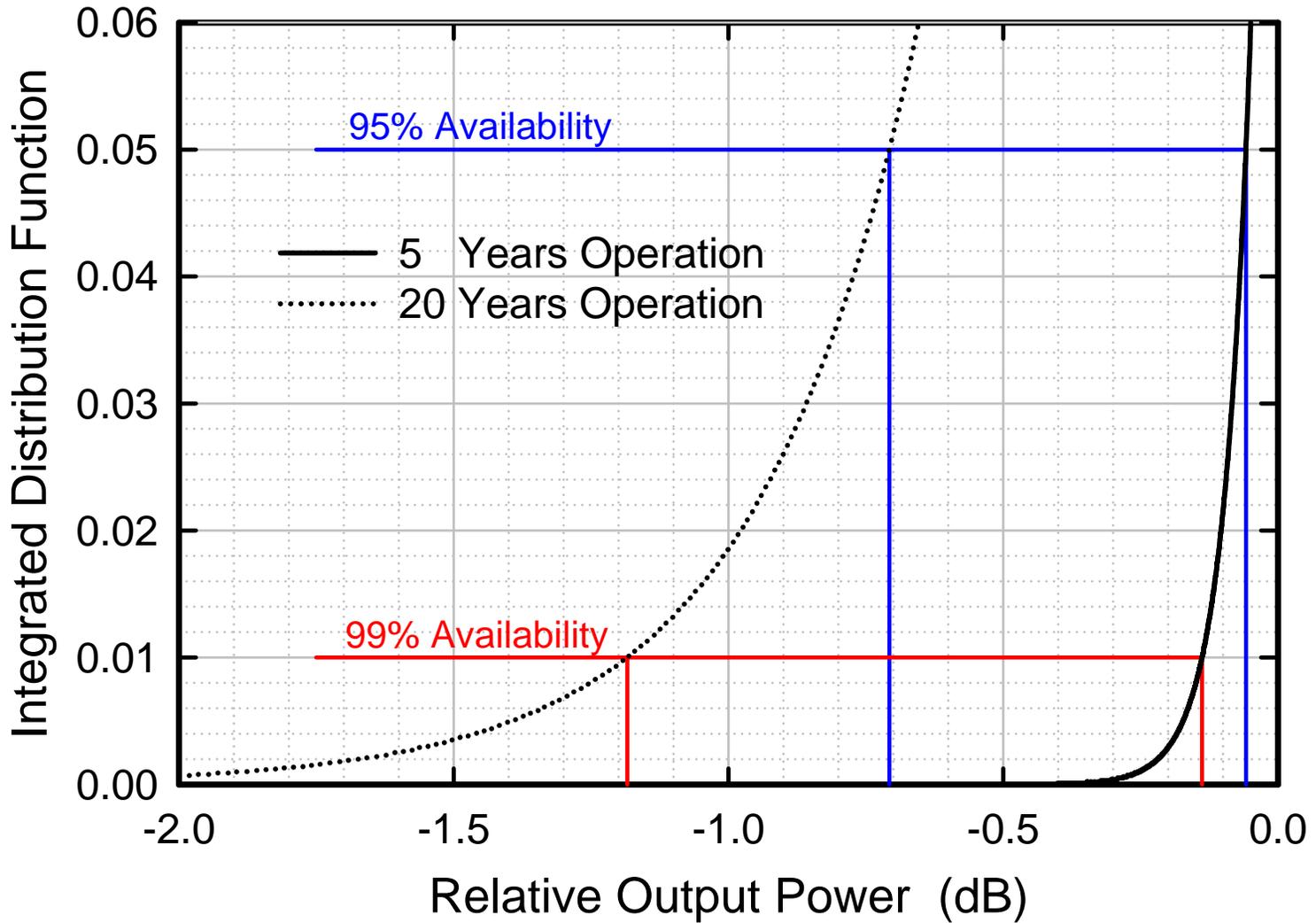


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- **Integral of distribution function**
  - Use histogram as proxy
  - More useful method of interpreting Monte Carlo results
- **Link budgets calculated to ensure a specified link availability**
- **What's important is percent of time RF power greater than some specified minimum**
  - 100% link availability usually not necessary
  - Data on deep space probes usually acquired, then stored on-board for possibility of multiple transmission to Earth







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# Conclusion



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- **The combining errors of four spatially-combined TWTAs have been modeled using characteristics of currently available devices**
  - **Errors occur due to inability to control phase and amplitude variations within the TWTAs**
- **During the JIMO Ka-band TWTAs five years of operation, 99% of the time the combining loss should be no worse than 0.14 dB**
  - **Expected output power 970 watts at 99% confidence level**
- **Even for 20 years of operation, loss should be less than 1.2 dB 99% of time**
  - **Expected output power 760 watts at 99% confidence level**
- **This analysis has not considered other elements in transmit chain**
- **Future TWTAs should have improved characteristics**