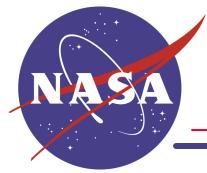


Characterization of the Visible Light Photon Counter for optical communications on a photon-starved channel

W. H. Farr, J. Kovalik, D. Jackson, L. Taylor, D. Q. Zhu

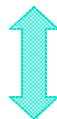
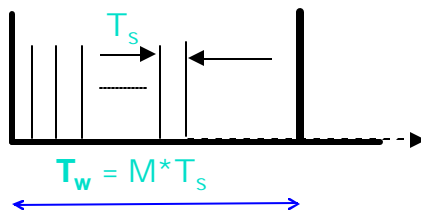
Jet Propulsion Laboratory
California Institute of Technology



- **Photon Counting: good for deep space optical communications**
- **The Visible Light Photon Counter (VLPC): what and why**
- **Test Configuration Overview: JPL's detector characterization facility**
- **VLPC Signal Processing: analog signal chain and digital analysis**
- **VLPC Characterization: pulse response, gain, noise, linearity**
- **Receiver Signal Processing: front end architecture**
- **Ongoing work and Summary**



T_s Slot Width
 T_w Frame (word) Time
 M Alphabet Size



$$C = B \left[\log_2 \left(1 + \frac{1}{M} \right) + \frac{1}{M} \log_2 (1 + M) \right] \xrightarrow{\text{PPM encoding}} C = \frac{B}{M} \log_2 M$$

from Brillouin's "negentropy principle"

- Photon counting can yield a higher channel capacity than phase-sensitive detectors
- Deep Space Optical Communications requires data encodings that maximize the (bits/sec) per (Joule/sec) metric
- PPM Encoding with Photon Counting detection is an attractive solution

| Detector Class | Examples | Photon Capacity Limit |
|-----------------------------|---|---|
| phase insensitive amplifier | parametric amplifier, Raman amplifier, laser amplifier | 1.44 |
| dual quadrature sensitive | coherent heterodyne | 1.44 |
| single quadrature sensitive | coherent homodyne, degenerate parametric amplifier | 2.88 |
| photon counting | photomultiplier tube, cooled avalanche photodiode, hot electron superconducting | $h\nu/2kT \ln 2$ (for instance, 69 bits /photon at 1 μm & 150K) |

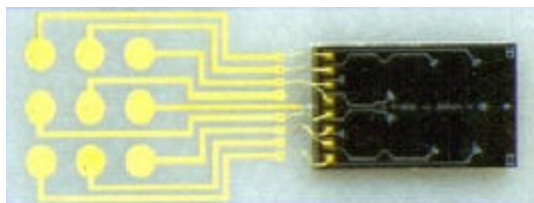
$$P_{\text{avg}} = \frac{B_{\text{bit/s}} N_{\text{photon/bit}} h\nu}{\eta_{\text{link}} \eta_{\text{detector}}}$$

Labels in diagram:
 - $B_{\text{bit/s}}$: bits/second
 - $N_{\text{photon/bit}}$: photon energy
 - $h\nu$: photon energy
 - η_{link} : link loss
 - η_{detector} : quantum efficiency
 - P_{avg} : average laser power

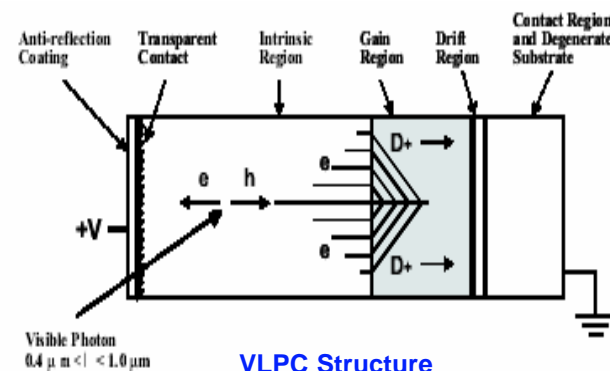
Higher efficiency means a choice of:

- Lower transmitter power
- Smaller receive aperture
- Higher data rate

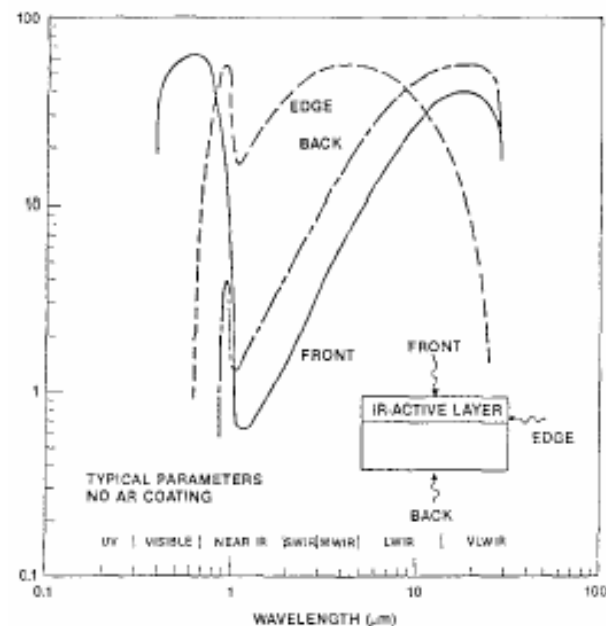
8-detector VLPC chip
1-mm diameter detector area

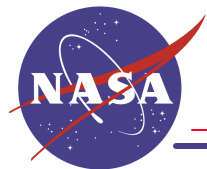


M. Petroff and M. Stapelbroek
IEEE Trans. Nuclear Sci., **36**, 158, (1989)

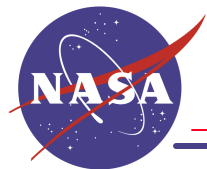


- **The VLPC is a variant of a Si:As detector with avalanche gain that was developed by DRS Technologies for use in Fermilab's D0 detector**
 - The combination of intrinsic Si and extrinsic As absorption gives a spectral response from 0.4 to 28 microns
 - Operating temperature range is typically 6 to 10K
- **A non-Markovian gain process allows the device to exhibit a gain variance near one**
 - With a typical gain of 20000 to 40000 this means photon counting with photon number resolution!

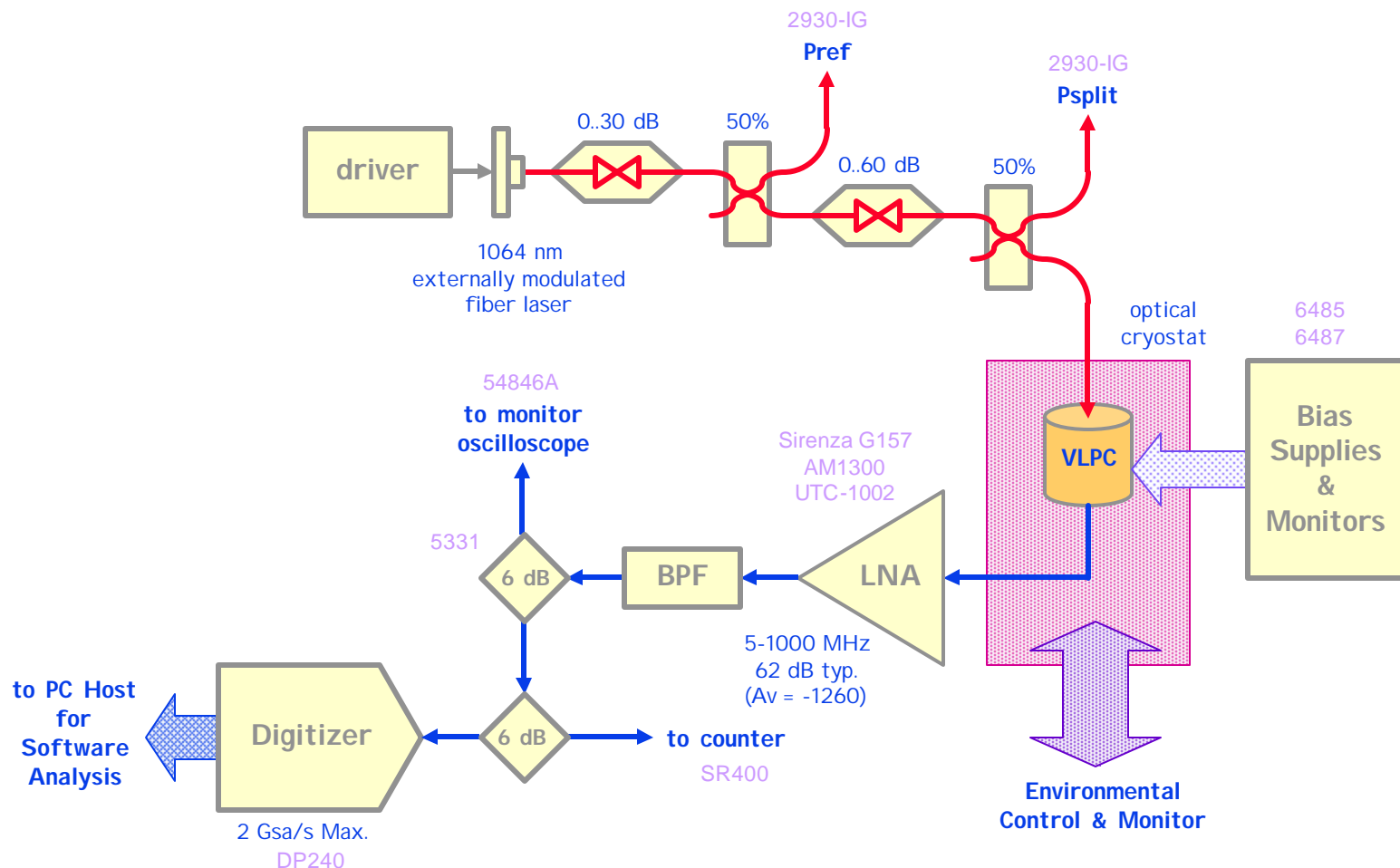




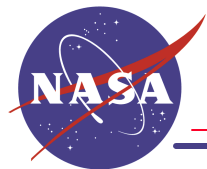
- **If you want to photon count in the visible, and can cool to $< 10\text{K}$, the VLPC is a great detector!**
- **Unfortunately, for 1 to 1.5 micron optical communications, where there are good laser sources, the detection efficiency is poor, as this falls between the intrinsic Si and extrinsic As absorption bands**
- **However, the VLPC is an excellent device to demonstrate high rate photon counting optical communications while other photon number devices with enhanced near-IR response are being developed**
 - Such as:
 - arrays of Geiger mode InGaAs/InGaAsP avalanche photodiodes
 - hybrid photomultiplier tubes with InGaAs/InGaAsP photocathodes
 - Si:As detectors with a separate absorber layer
 - InGaAs/InGaAsP MIS avalanche detectors



VLPC Test Configuration

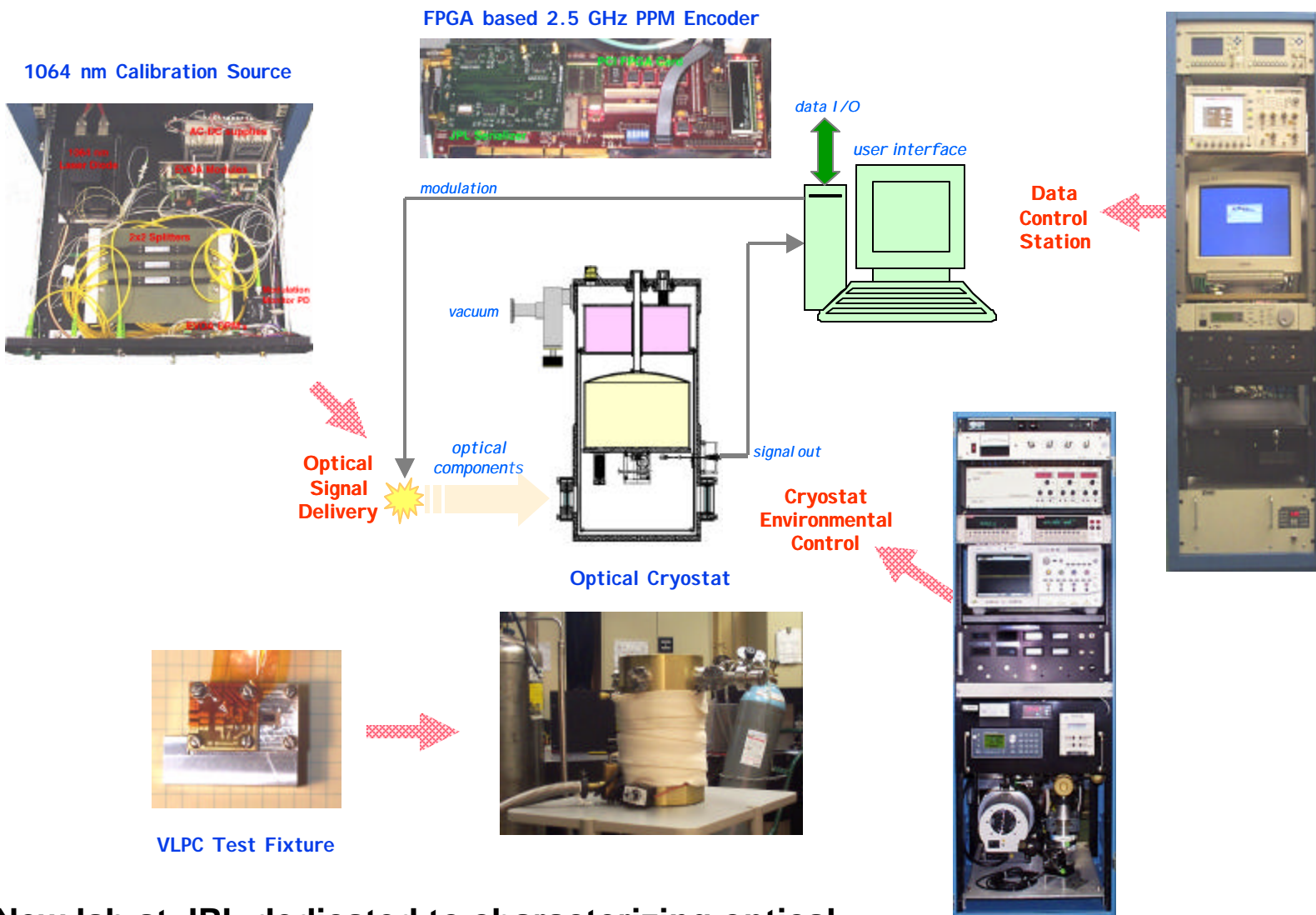


- The test configuration emulates deep space receive signal levels (pW) with CW background and a GHz rate intensity modulated optical data stream
- The detector output is typically digitized and post-processed in near-real time



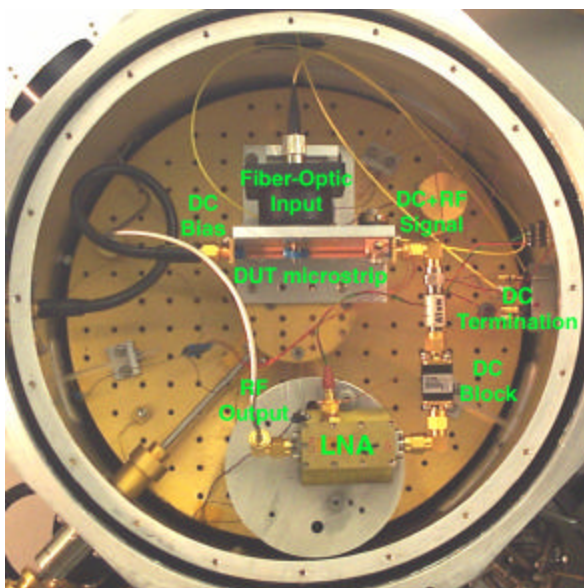
Detector Characterization Facility

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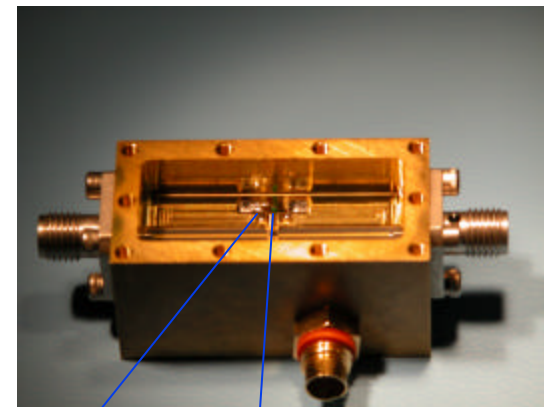
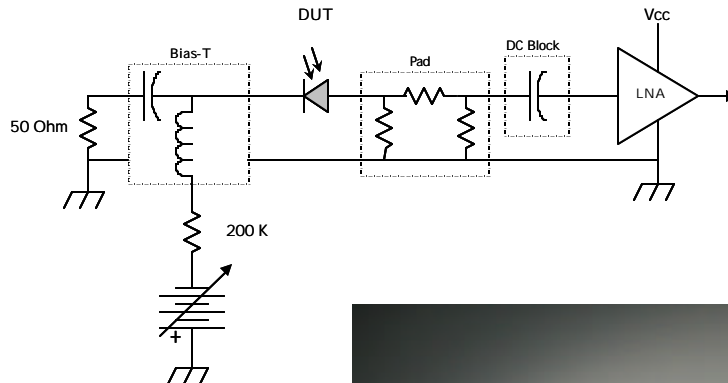


- **New lab at JPL dedicated to characterizing optical detectors for deep space optical communications**

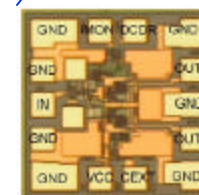
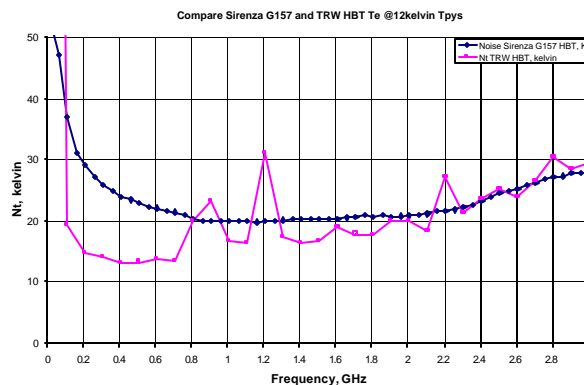
Optical Cryostat Layout with Cryogenic LNA



Microstrip Biasing Topology



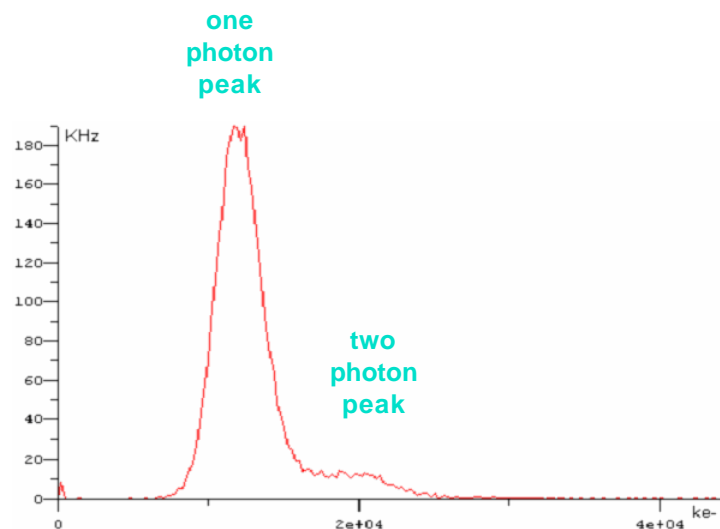
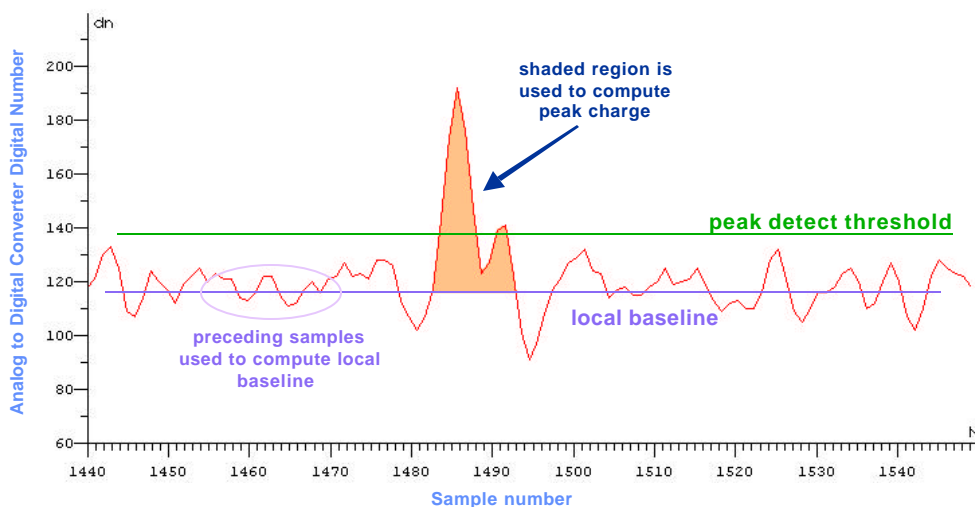
- The initial preamplifier is also cooled to $< 15\text{K}$ to minimize additive thermal noise for detector calibration
- A microstrip architecture allows the detector and amplifier to be physically separated



InGaP LNA in
12 GHz Chassis

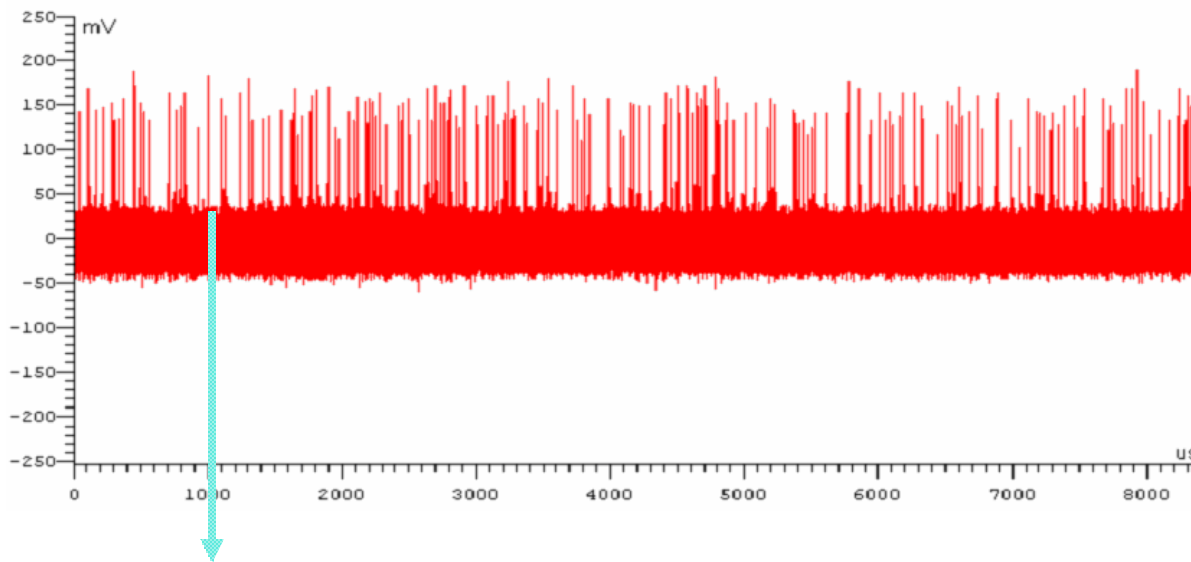
Commercial chip (Sirenza) with ultra-low noise and multi-GHz bandwidth that can operate at cryogenic temperatures

- In addition to time domain displays and Fourier analysis, pulse height and pulse area histograms are essential to characterize the detector performance
 - detector gain and excess noise are derived from pulse area histograms
 - dark rate and photon counting linearity are derived from thresholded peak height histograms
 - other histograms: pulse width distribution and interarrival time

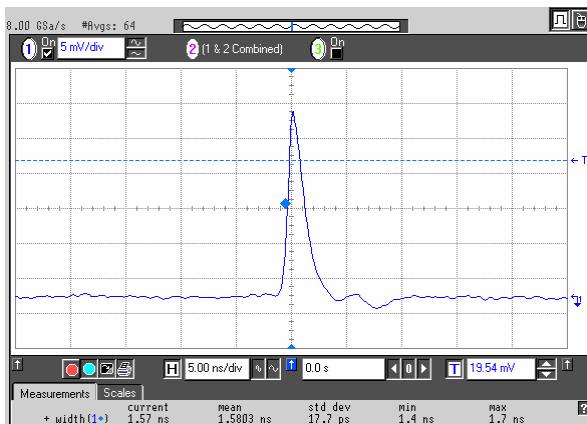




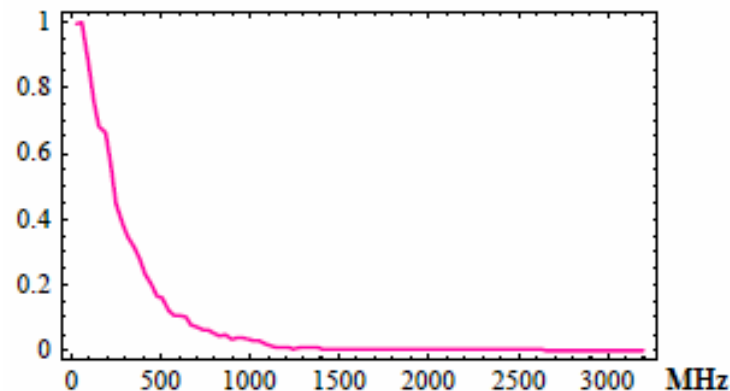
Temporal Pulse Response



1064 nm CW
Illumination



VLPC Single Photon Response

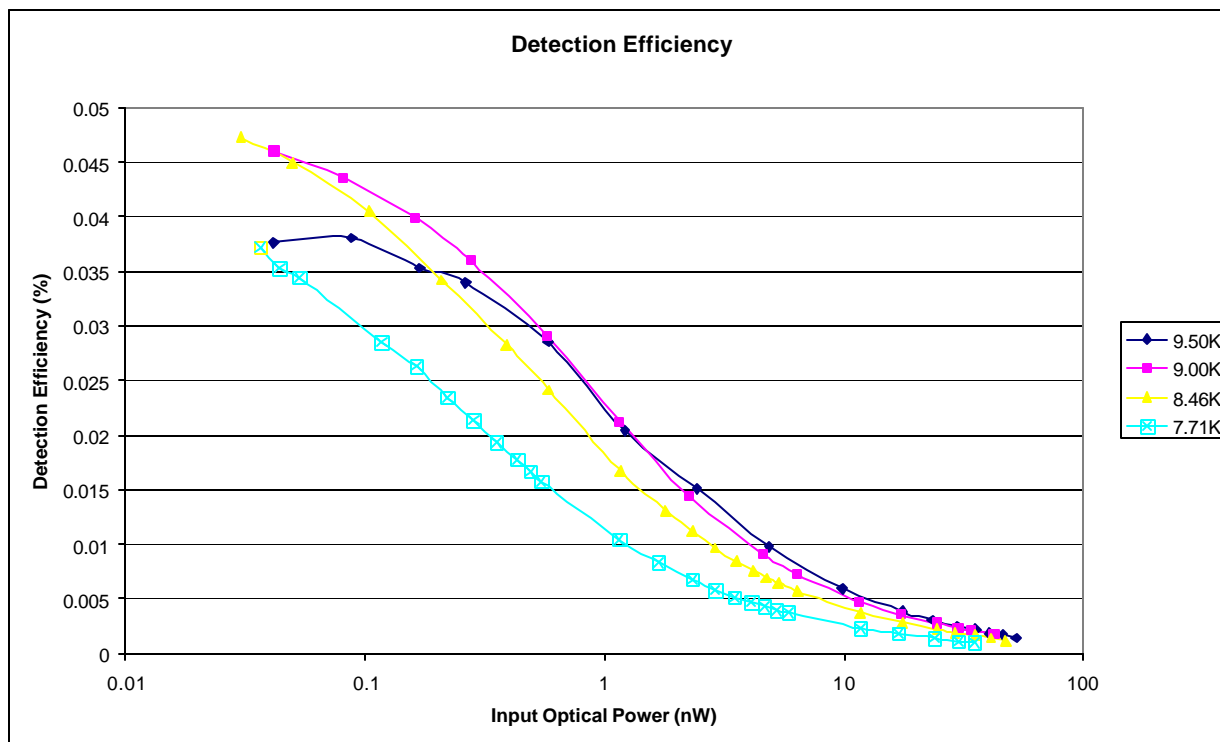


Normalized Power Spectrum

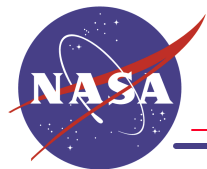
- Single photon output pulse is typically about 1 ns wide with 30000 electrons



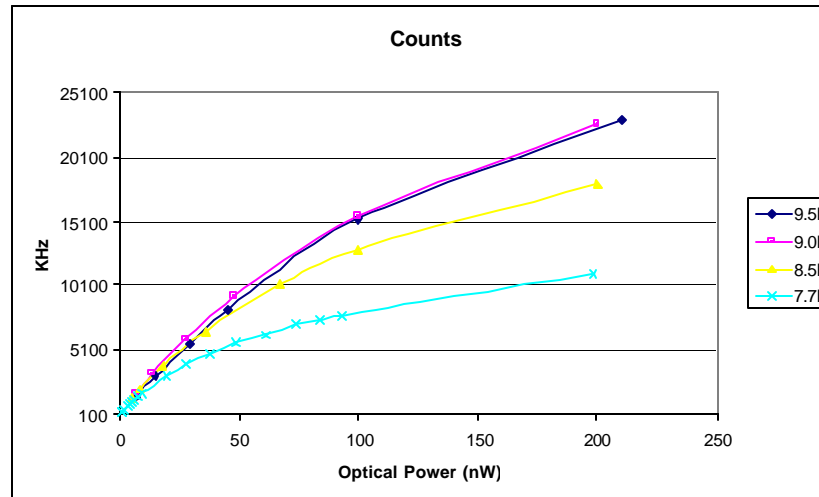
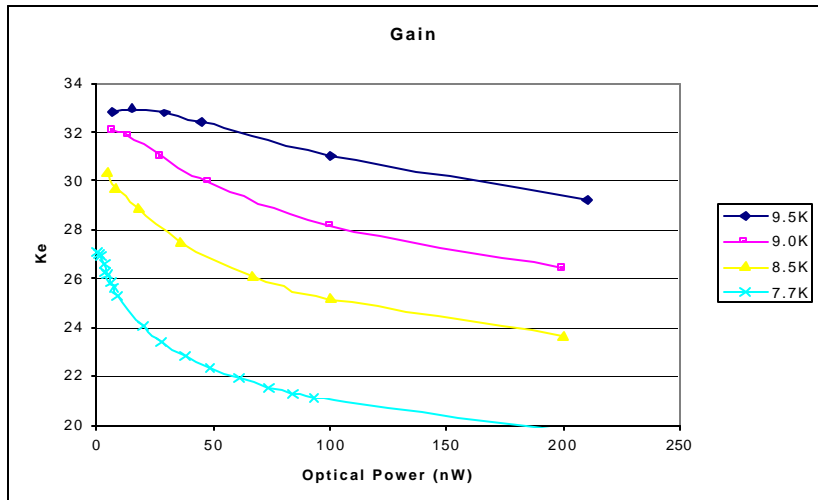
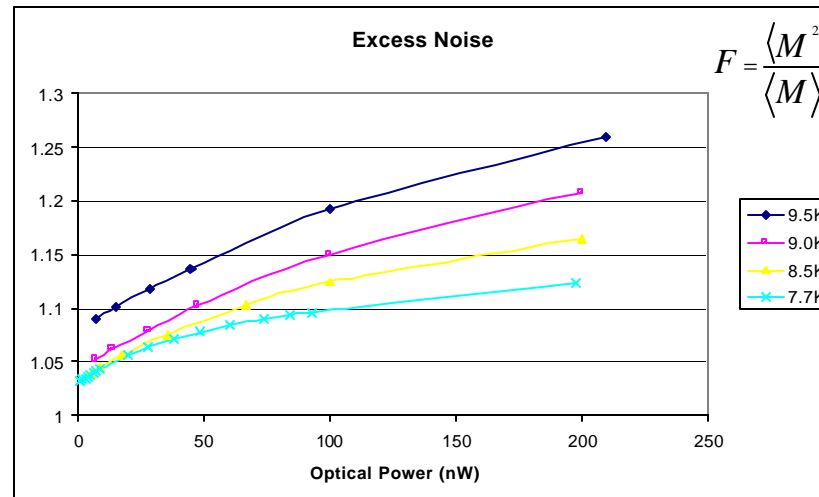
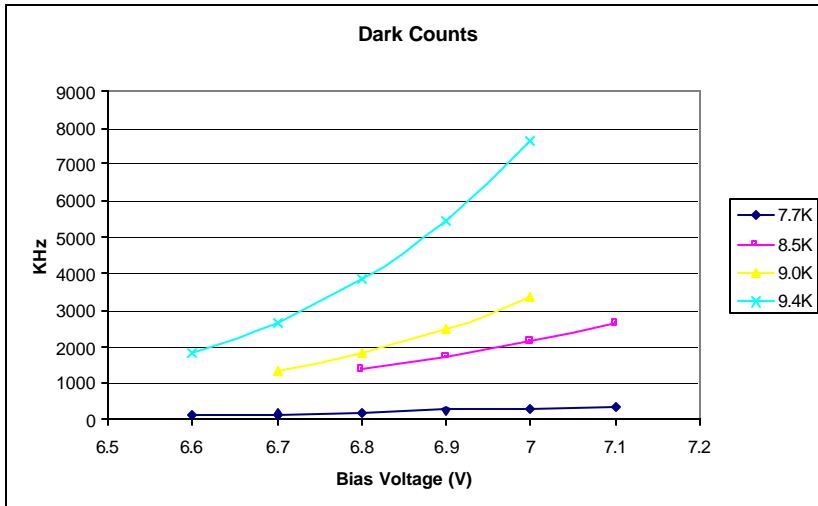
VLPC 1064 nm Detection Efficiency



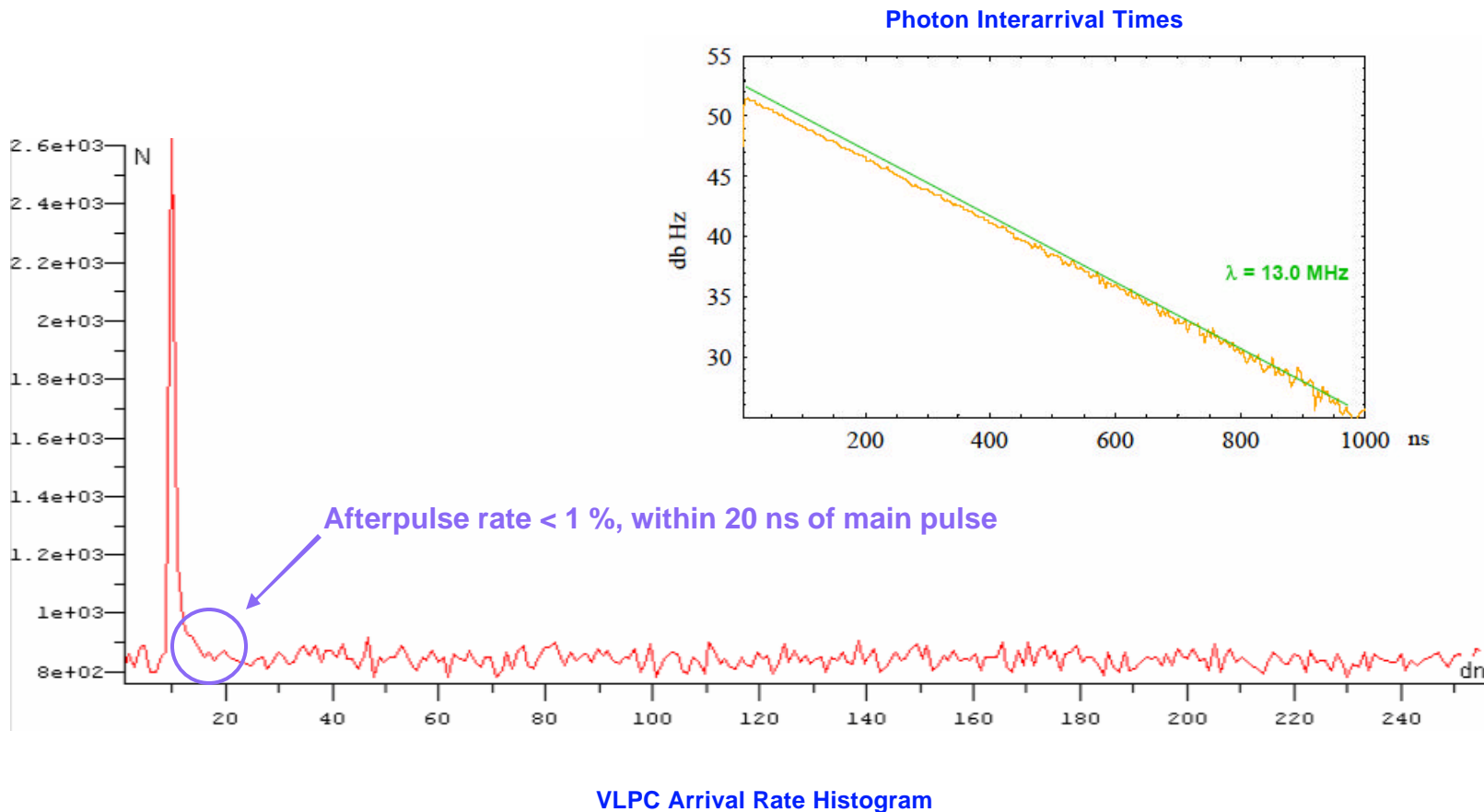
- **As expected, detection efficiency at 1064 nm is very poor**
 - However, electrical output pulse was verified to be independent of wavelength, so this wavelength can be used to verify optical communications performance
 - Observed loss in detection efficiency at high incident flux rates (many 10's of MHz detected photons) is due to loss of gain from space charge effects
- **We prefer to use 1064 nm as most of our calibrated sources and detectors are at that wavelength**



VLPC 1064 nm CW Performance

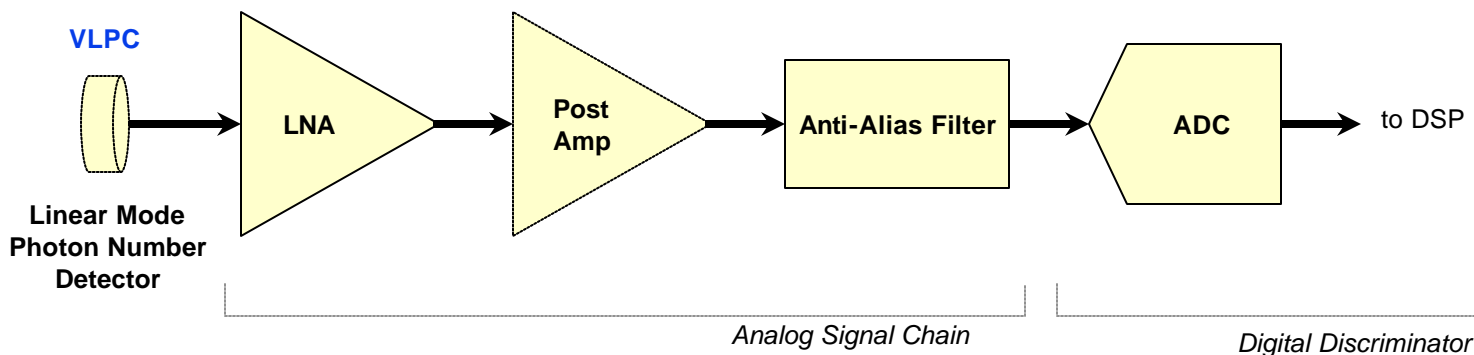


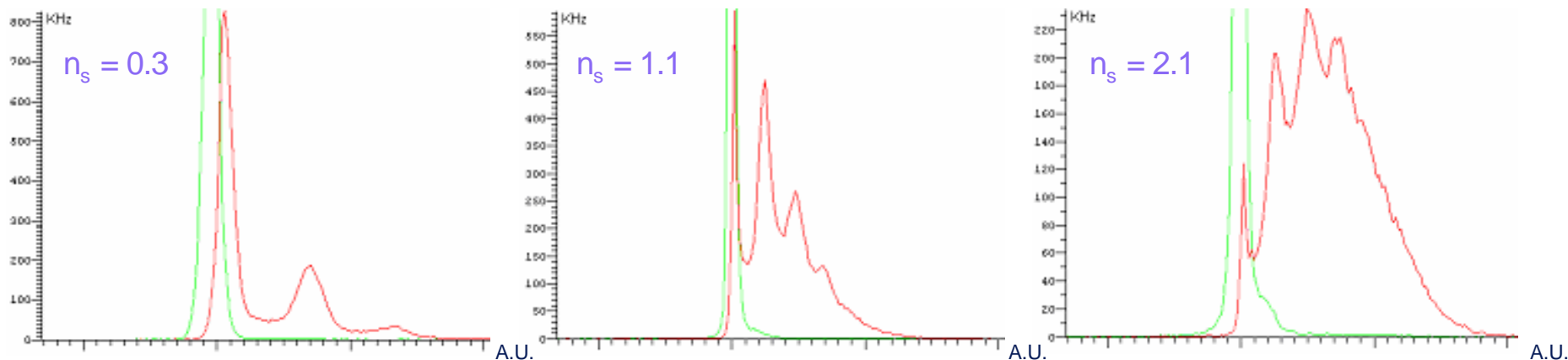
- **0.5 dB compression point at 17 MHz at 9.5K, for instance**
 - 5 dB compression at 50 MHz at 9.5K



- Exponential interarrival time distribution is expected for CW illumination

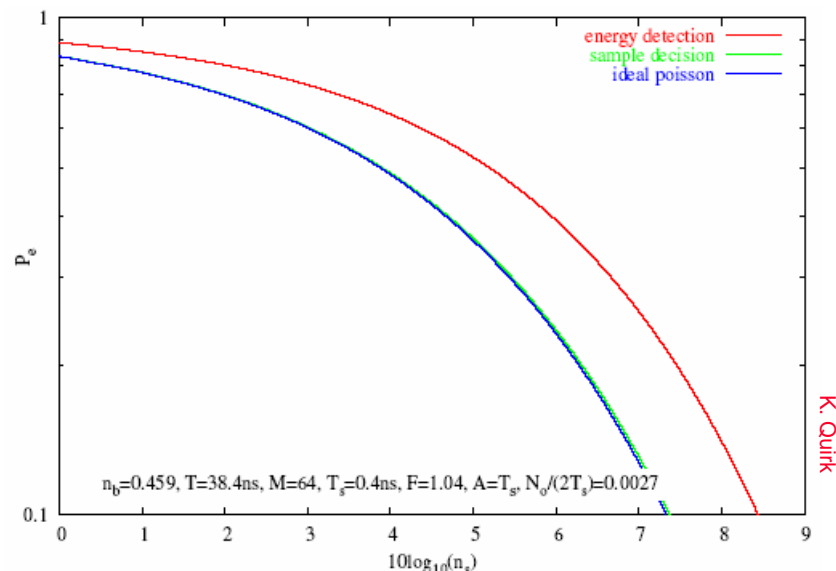
- **A flexible photon counting receiver front-end can be constructed by analog sampling of the detector output**
 - The pulse discriminator function is emulated by digital signal processing (DSP)
 - This is more flexible than a pure hardware solution, comprising a comparator and counter, for instance
 - DSP can implement FIR signal filtering, for instance
 - Note: a Geiger mode avalanche photodiode may be considered as having a hardwired “comparator” stage that implements a discriminator with only zero or greater than zero resolution





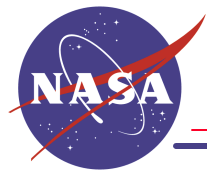
note: horizontal scale on slot energy histograms varies between plots

- Slot histograms represent the sum of signal and noise generated charge in a PPM slot
- Thresholding converts the slot integrated charge levels (“energy”) into a photon number
- Subsequent processing of the photon number signal is essentially “noise free”

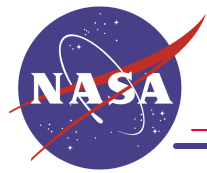


comparison of thresholded vs. non-thresholded signal processing for an example MLCD link

K. Quirk



- **Work is in progress to acquire symbol (no FEC coding) and bit (FEC decoded) error rate data using the VLPC as the detector for a software digital receiver for comparison with theoretical predictions**
- **Testing with the VLPC has established techniques that are now being used to characterized newer high bandwidth photon number resolving detectors with enhanced near-IR sensitivity**
 - VLPC with PtSi layer added to improve near-IR absorption
 - Intevac HPMT with InGaAs or InGaAsP photocathode
 - This has now demonstrated 34% single photon detection efficiency at JPL, for instance



- **A photon counting / photon number resolving detector can be used with a PPM encoded laser source to implement a high efficiency, high data rate optical communications link**
- **The VLPC is an excellent photon number resolving detector for the 0.4 to 28 micron spectral range (although poor in the 1 to 2 micron range)**
 - High bandwidth with large area
 - High gain with very low excess noise
 - Can handle high flux rates: many 10's of MHz per mm²
 - Although < 10K operation is problematic for many applications
- **Other detector technologies are rapidly maturing that will be able to provide similar performance in the 1 to 1.5 micron regime**