

On the splitting and outburst of 73P/Schwassmann-Wachmann 3

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The major outburst of 73P/Schwassmann-Wachmann 3, in 1990, occurred ~2 weeks before perihelion as a sudden increase in the production of hydroxyl, which is found to have been followed by a steady brightening in the visible light until ~20 days past perihelion when the comet was a factor of ~200 more luminous than in its quiescent phase. Within 80 days after perihelion the nucleus, whose effective diameter is estimated at ~3 km, appeared to be triple (with unconfirmed reports of up to four additional companions), the issue of primary interest was the nature of the relationship between the splitting and the initial stages of the outburst. Application of a standard model for the split comets indicates that, nominally, the first breakup (the separation of the component B from the principal nucleus C (the easternmost component)) occurred most probably in late October, some six weeks following the outburst's onset. A secondary breakup (the separation of A (the westernmost component) from C) followed some two weeks later. These fragmentation episodes correlate well with two less prominent secondary flare-ups on the light curve. Also examined are the circumstances of separation for the (111) additional, suspected companions. The observed delayed response of the 111C nucleus to the major disturbance (if its surface is a phenomenon not previously reported for any other split comet) was proposed in the early 1980s that some companion nuclei of non tidally split comets may represent large, 111C early inner mantle, the surface mantle of refractory material that were torn off, thus exposing the previously protected reservoir of ice beneath. If so, the case of 73P/Schwassmann-Wachmann 3 suggests that the disturbed area of the mantle first released considerable amounts of dust, 111C grains (micrometric to 111C, an outburst. The separation of a large fragment (or fragments) of 111C mantle apparently required continuing, more advanced manifestations of elevated activity (triggered possibly by a buildup of vapor pressure in localized subsurface pockets of confined volatiles), in order to overcome the resistance to fracture offered by the intrinsic cohesion of the nucleus and its mantle. In any case, the delayed response provides strong evidence against models of a strengthless, cometary nucleus.