



# Laser Technologies for Deep Space Optical Communications

*Malcolm Wright, Keith Wilson, Abi Biswas, Gerry Ortiz*

Optical Communications Group,  
Jet Propulsion Laboratory  
California Institute of Technology  
Pasadena, CA 91109



- Introduction to free space optical communications
  
- Requirements
  - Technology development for JPL/NASA missions
  
- Laser development
  - Diode and fiber based MOPA devices for high data rates
  - Solid state lasers with high peak powers
  - Status of solid state laser performance
  - Technology challenges
  - Laser characteristics - efficiency, pulse jitter
  
- Summary



## Advantages & Challenges of Lasercomm



- Higher data rates
  - For example, near Earth RF typically limited to 10 - 100s Mbps but  
⇒ Gbps with complex modulation schemes
- Lower beam divergence ⇒ smaller transmit aperture  
⇒ lower mass
- No bandwidth allocation restrictions - severe limitation for high BW RF modulation
- Challenges
  - clouds - availability of coverage  
⇒ *site diversity*
  - atmospheric induced scintillation effects - beam breakup and wander  
⇒ *adaptive optics, aperture averaging*
  - stringent acquisition, pointing and tracking requirements  
⇒ *spread beam, fast steering mirrors etc*

Descanso, Sept 1999

MWW

3/total



## JPL/NASA Mission Requirements



- Mission determines distance and data rate or volume  
⇒ drives size, aperture, transmit power required
- Near Earth:
  - LEO / GEO - ground
  - Example:
    - ISS demonstration: LEO - ground,  $\lambda = 1.5 \mu\text{m}$ , > 1 Gbps, 200 mW laser power, 10cm → 1m, unidirectional link with ground beacon acquisition
- Deep Space:
  - Planetary to Interstellar missions, distances AU - > 100 AU
  - Example:
    - Mars downlink, 20 kHz,  $P_{\text{peak}} \sim 30 \text{ kW}$ ,  $P_{\text{avg}} \sim 1 \text{ W}$ ,  $\lambda = 1\text{-}1.5 \mu\text{m}$ , 30cm → 1m apertures

Descanso, Sept 1999

MWW

4/total



- High rate optical channel using direct detection determined by application:
- Near Earth links:
  - high data rates desired, 100 Mbps - 10 Gbps
  - high average power sufficient, 1-5 W
  - ⇒ Semiconductor diode or fiber based lasers with OOK modulation
- Deep space links:
  - maximize photon efficiency ⇒ high average and peak powers, kW - MW/pulse
  - medium data rates, 10 kbps - 1 Mbps
  - ⇒ Diode pumped solid state lasers with PPM techniques

Both require compact, efficient, reliable, rad. hard lasers with good beam quality

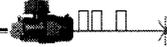


### Link Analysis

	<u>Mars downlink</u>	<u>ISS downlink</u>
• Link Range	1.5 AU (2.10 <sup>8</sup> km)	1000 km
• Data rate	1 Mbps	1.2 Gbps
• Modulation	PPM, coding	OOK
• Transmit power, avg.	3 W	200 mW
• Transmit aperture	30 cm	10 cm
• Losses - space, pointing, atmospheric	25% + dist (-368 dB)	25 % + dist (-258 dB)
• Receiver gain - telescope	10 m	1 m
• Required power - detector sensitivity	3 nW	0.2 uW
• Link margin	3 dB	5 dB



# Link Analysis



Mars(Down,X2000)

## Link Summary

Link Range	km	1.00 AU	
Data rate	kbps	PPM (M = 256)	
Coded BER		Reed-Solomon Coding	
Transmit power	1.00 W average	3.19 kW (peak)	65.04 dBm
Transmit losses	57.0 % transmission		-2.44 dB
Transmitter gain	6.3 urad beamwidth		116.03 dB
Pointing losses			-0.97 dB
Space loss			-364.97 dB
Atmospheric losses	67.2 % transmission		-1.72 dB
Receiver gain	3.50 m aperture diameter		140.24 dB
Receiver optics losses	25.3 % transmission		-5.96 dB
Received signal	1.98E-02 photons/pul se	3.36 nW (peak)	-54.74 dBm
Background signal level	7.60E+00 photons/slot	0.10 nW	
Required signal level	7.32E+01 photons/pul se	1.24 nW (peak)	-59.06 dBm
Allowances and Adjustments			-1.00 dB
<b>Link Margin</b>			<b>3.31 dB</b>

## Detailed Link Table

### Transmitter (Mars)

Laser (pulsed Q-switched Nd:YVO4)			
Peak power		3.19 kW	65.04 dBm
Average power	W		
Wavelength	1.06 um		
Pulse width	11.00 ns		
Pulse width to slot width ratio	0.80		
Slot width or integration time	13.75 ns		
Dead time	31.62 us		
Pulse repetition rate	28.46 kHz		

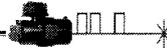
Descanso, Sept 1999

MWW

7/total



# Deep Space Laser Requirements



	Jupiter 6 AU	Interstellar Probe 200 AU
<b>Downlink Laser</b>		
Average Power, W	3	15
Peak Power, MW	0.022	11.3
Repetition Rate, Hz	6650	135
Energy per pulse, mJ	0.44	110
Pulse width, nsec	20	10
<b>Uplink Laser</b>		
Average Power, W	7.3	135
Peak Power, MW	1.46	12000
Repetition Rate, Hz	250	1.1
Energy per pulse, J	0.04	120
Pulse width, nsec	25	10

Assume: 30 cm aperture on spacecraft, 10 m on ground

Descanso, Sept 1999

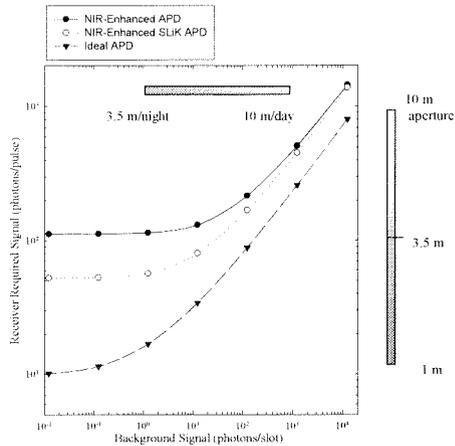
MWW

8/total

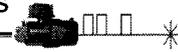


Laser requirement also driven by detector capabilities:

- Transmitter:
  - 1 Watt Q-switched Nd:YAG laser
  - 1064 nm
  - 25 ns pulse width
  - $10^{-5}$  modulation extinction ratio
  - 10 cm telescope
  - 27.8 % transmitter optics losses
  - 16 kbps
  - $10^{-5}$  BER
  - PPM (M=256)
  - Reed-Solomon coding
- Channel:
  - 2 AU link range
  - 32.8 % atmospheric loss
- Receiver:
  - Telescope diameter -various
  - 55.1 % receiver optics losses
  - Silicon APD Detector - various
  - High-Impedance Pre-amplifier
  - 31.25 ns slot width



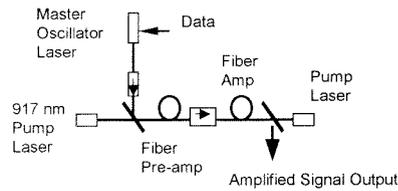
Depending on background level, link is viable with 3.5 or 10 m aperture



Lab Prototypes

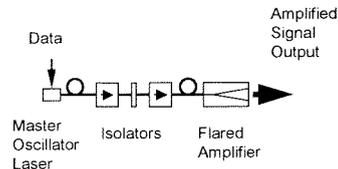
Fiber MOPA

- MO: InGaAs diode oscillator
- 5 W 2-stage YDFA
- 2.5 Gbps, OOK, 1070 nm
- Good beam quality



Diode MOPA

- MO: InGaAs FP oscillator capable of 10 Gbps
- InGaAs flared amplifier
- Packaged: 1 Watt InGaAs flared amplifier,
- 2.5 Gbps, OOK,  $\lambda = 960$  nm, efficiency  $\sim 24\%$
- Table-top: 3.6 W at 2.5 Gbps
- Poor modulation extinction ratio due to ASE



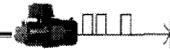
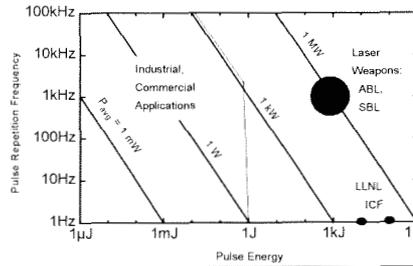


*Laser Transmitter:*

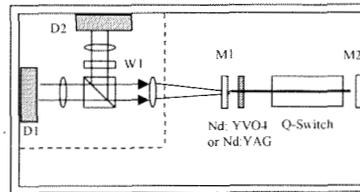
- Good beam quality, ~ diffraction limited,  $M^2 < 1.3$
- Reasonable overall efficiencies, ~ 6 - 10%
- Good modulation extinction ratio, 50 dB

*Operational modes of Candidate Solid State Lasers:*

- Q-Switched  $\Rightarrow$  direct detection, limited to  $< 1$  Mbps (100 kHz)
- Cavity dumping  $\Rightarrow$  direct detection, will support 1- 10 Mbps
- Continuous wave  $\Rightarrow$  coherent detection

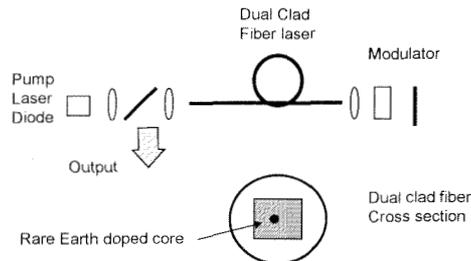


**Bulk**  
 Nd:YAG  
 Yb:YAG  
 Nd:YVO<sub>4</sub>  
 Nd:YLF



**Fiber based**

- Fiber Amplifier or Laser
- Er or Yb/Er co-doped core
- Increased peak power with large effective area core

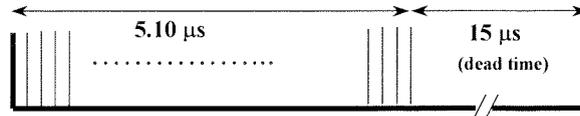




Modulation techniques:

Pulse position modulation (PPM) provides a suitable modulation and detection scheme for solid state Nd:YAG type lasers

25 ns per slot ( 256 slots => 8 bit/pulse )



50 KHz rep rate Nd:YAG laser

- PPM requires:
  - that pulse width be < 20 ns slot width
  - that pulse jitter be such that probability of pulse over-shooting intended slot be minimized
  - slot width is restricted in order to limit background photons
- Example: night time link from Mars using above modulation and 5 W of average laser power will result in a downlink of ~ 200 photons/pulse using a 10 m receiver on the ground



High Power Q-switched Lasers

		Nd:YAG 1.06 μm		Nd:YLF 1.05 μm	Yb:YAG 1.03 μm	Doubled .532 μm		YALO 1.08 μm
Pav/	W	10	690	29	280	120	175	140
Ppulse	mJ	100						
PRF	kHz	0.1	2.5	5	10	5	2.5	0.1
Pulsewidth	ns	20	~100	20	30	-	~100	-
Efficiency	%	6	?	6	6	2	?	1.8
BQ/ D.L.		1.05	1.1	1.1	5	5	1.5	1.1
Comments			P- kW with osc/amp					MOFA
Developer		ESA	TRW	SEO	LLNL	DOS	TRW	German



# Status of Laser Development



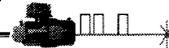
## High Power cw Solid State Lasers

Wavelength nm	Nd:YAG				Nd:YLF	Yb:YAG		Doubled	
	1060				1050	1030		532	
P <sub>cw</sub> W	1	10	45	107	50	430	40	5	.2
Efficiency %	6.3	<8	6	11	8	6.3	2.5	0.4	~0.5
BQ/D.L.	~1	1.05	<1.3	5	~1.1	5	1.02	-	1.1
Linewidth kHz	10	Q.L.*	?	-	?	?	?	5000	<10
Timescale	10ms							50 ms	1 ms
Comments		Freq. Stable	300 W poor BQ	137 W BQ=45 η=14			T=100K	commercial	commercial NPRO†
Developer	JPL	Japan	German	Japan	SEO	LL NL	DOS	Coh.	Light-wave

\* Injection locked to a single frequency NPRO laser.  
 Frequency noise  $2 \times 10^{-5}$  /√Hz at 300 Hz. Shot noise limited above 1 MHz.  
 † Two 40 mW Nd:YAG devices frequency locked to high finesse Fabry-Perot cavity;  $\Delta\nu = 330$  mHz.



# Laser Dev. for X2000 Laser Requirements

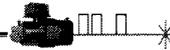


Specifications	Continuum EPO 5000	O-peak 1047Q	Spectra-Physics B10-106Q	Uniphase Stablalight
Material	YAG	YLF	YVO <sub>4</sub>	YLF:YAG
λ, nm	~ 1060	1064	1047	1064 1047
P <sub>average</sub> W	1	1	10	3 2
PRF kHz	1.5-20	0-5	3-100	50 (0-70)
P <sub>peak</sub> kW	30 at 3kHz	83	50	18 25
Pulse width ns	< 25	3	40	10 10 10
z/ freq	1, 5, 20		75	12 20 250
Pulse energy μJ	> 300	250	2000	180 250 110
Beam Quality	1.1	< 2	< 1.2	TEM <sub>00</sub> TEM <sub>00</sub>
Air cooled	Yes	yes	no	yes no/yes
Efficiency %	10	~ 1	~ 3	~ 75 (air cooled)
Q-switch	E.O.	F.O.	A.O.	A.O. A.O.
End-Pumped	yes	yes	no	yes
Comment		Single Diode	8W YVO <sub>4</sub> air cooled, end pumped, 20 kHz	Remote diode bar
Customize	yes	~ no	yes - YVO4	no Yes



### Issues in High Power Solid State Laser Development

- Nd:YAG - well established technology, robust, peak output at PRF ~ 5 kHz
- Nd:YLF - no thermal birefringence but lower thermal fracture strength which limits high power capability, gain peaks at PRF ~ 2 kHz
  - Peak absorption ~797 nm so less absorption from 808 nm pump diode energy
  - Longer upper state lifetime therefore higher output energy for given pump energy
- Nd:YVO<sub>4</sub> - high gain and high absorption
  - Suitable for high PRF as able to maintain high pulse energy
  - Short pulsewidths possible due to short upper state lifetimes
  - Not as robust thermally, start to lose BQ at high PRF
- Yb:YAG - broader absorption band at 940 nm relaxes diode  $\lambda$  constraint
  - quasi 3 level system
  - InGaAs pump diodes have less heating and higher efficiency from lower quantum defect with potentially increased reliability
- Fiber laser - Peak power limited but can increase fiber area
  - High efficiency
  - Good thermal design



Efficiency	Ideal	Expected at High Power	Demonstrated
$\eta_p$ - diode	.65	.35	.3-.5
$\eta_T$ - transfer	.95	.85	.7-.93
$\eta_a$ - absorption	.9	.9	
$\eta_s$ - q. defect	.76	.76, .89*	
$\eta_Q$ - q. conversion	.95	.95	$\eta_{opt-opt} = .4-.5$
$\eta_E$ - extraction	.9	.8	
$\eta_B$ - beam overlap	.9	.85	
$\eta_{PS}$ - power supply	.9	.9	.9
$\eta_D$ - doubler	(0.6)	(0.45)	(.3-.5)
Total $\eta$	.29 (.17)	.14 (.06)	.02-.14 (<.07)

\* Yb

### Strategies to increase efficiencies:

- To obtain maximum diode efficiency, provide better thermal management, improved material quality  $\Rightarrow > 40\%$  at high power
- Increase diode coupling by novel optical techniques such as microlens design
- $\eta_{opt-opt}$  limited by gain medium; can be increased by using Yb, end pumping, and improved resonator design for mode matching. Side pumping required for  $> 50\text{ W}$  due to thermal stress though.

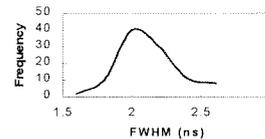
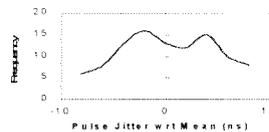


- A Nd:YVO<sub>4</sub> master oscillator (Lightwave Electronics) along with a diode pumped amplifier was modulated using a custom PPM modulator
- Three types of detectors used

Detector	BW	Area (μm)	QE	Pulse Jitter	FWHM ns	Sensitivity Photons
InGaAs PIN	1.0 GHz	100	78%	+/- 8 ns	2+/- .5	11000
Si APD	500 MHz	500	4-5 %	+/- 9 ns	1.2-3	100
Si SPCM APD	-	120	2-3 %	+/- 6 ns	8.7	600

Extremum of pulse jitter + pulse width ~ 21 ns

InGaAs PIN



- JPL lasercomm development
  - Near Earth ⇒ high data rate diode or fiber based laser transmitters
  - Deep space ⇒ DPSS laser transmitters
- Laser development:
  - *Diode and fiber based lasers:*
    - High efficiency, good beam quality, high data rates
    - High ASE degrades modulation extinction ratio
    - Commercial devices available that meet near Earth requirements
  - *Solid State lasers:*
    - High peak power, good beam quality, high modulation extinction ratio
    - Medium data rates, improvements in efficiency being explored
    - Commercial systems meet requirements for various missions
    - With the right detector, a PPM modulated Nd:YVO<sub>4</sub> can provide the power and timing characteristics needed for a Mars link